## Proton-oxygen collisions at the LHC

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## Air showers and cosmic ray mass



 Direction from particle arrival times
 Energy from size of eγ component
 Mass from size of muonic component and depth of shower maximum

> Nυ J<sub>μ</sub>Iroi at s

Number of muons and Mass Iron = 1.4 x proton yield at same CR energy



Shower depth and Mass Iron = proton - 100 g cm-2 at same CR energy



Based on Kampert & Unger, Astropart. Phys. 35 (2012) 660

### Motivation

Astrophysical origins of cosmic rays?

- Mass composition (<InA>) of cosmic rays carries imprint of sources and propagation
- Uncertainties of <InA> limited by uncertainty in description of hadronic interactions
- **Muon Puzzle:** Muon predictions in air showers are inconsistent with X<sub>max</sub>

Combined approach to get precise unambiguous <InA> data

- Cosmic ray community probes air showers and quantifies inconsistencies
- Collider community provides relevant reference measurements for model tuning

Indirect search for physics beyond the standard model at 100 TeV scale

## **Compilation of muon measurements**

- WHISP report at UHECR 2018 conference, Oct 8-12 2018
- Comprehensive compilation of muon measurements from air shower experiments



Systematic discrepancies reported by majority of experiments starting around 5x10<sup>16</sup> eV equiv. to s<sup>1/2</sup> = 10 TeV

Apart from shower energy, possible dependence on shower age, lateral distance, muon energy threshold

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## How to fix the issue?



- No simple key measurement
- Need to accurately know and extrapolate several features
- Focus on measuring these features accurately in references systems, use models to predict interactions in unavailable target systems

## Modeling air shower interactions

extrapolation to higher energy & different collision systems



- Light hadron production important
- Full rapidity range, mid-rapidity alone not enough



PbPb

## Oxygen beam at LHC

- Low-luminosity oxygen beam comparably easy to do
  - Oxygen in lead source as support gas
  - Rapid set-up following 2012 p-Pb and 2017 Xe-Xe runs
  - Almost scheduled at end of 2018, but didn't happen
- To be done during LHC Run 3 (2021-2023)
- Upcoming Yellow Report about physics at HL-LHC with subsection about proton+oxygen science case
- Questions asked
  - Project quantitative impact of measurements?
  - Is p+p and p+Pb data sufficient?

#### Important features in hadron production

Modify features at LHC energy scale with factor  $f_{LHC-pO}$  and extrapolate up to  $10^{19}$  eV proton initiating air shower R. Ulrich et al PRD 83 (2011) 054026





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#### Inelastic cross-section



- pp inelastic cross-section now known to 3 %, see e.g. ATLAS arXiv:1606.02625
- Similar for p+Pb, about 4 %, CMS arxiv:1509.03893
- Glauber interpolation to p+O could have similar precision
- p+O collisions to cross-check Glauber interpolation



- X<sub>max</sub> sensitive to: **inelastic cross-section**, hadron multiplicity
- $N_{\mu}$  sensitive to: energy fraction lost to  $\pi^{0}$ , hadron multiplicity
- Nuclear modification in forward-produced hadrons expected and important

## Nuclear effects poorly understood

- Hadronic interaction models used in air shower simulation must predict
  p-air (nitrogen & oxygen), but can only be tuned to p-p and p-Pb with current data
- Non-trivial nuclear effects severely affect forward production of particles (most important in air showers, because dominant for energy transport)

J/Psi production measured by LHCb, Physics Letters B 774 (2017) 159-178



Nuclear modification factor  $R_{pA} = \frac{\text{cross-section for pA}}{\text{A x cross-section for pp}}$ 

- Strong deviation from RpA = 1 for forward production
- 50 % uncertainty in PDF-based predictions
- Same effect observed in pion production at ALICE

Cannot translate this from p-Pb to p-O

### Hadron spectra

- Simulations done with CRMC by R. Ulrich et al. <u>https://web.ikp.kit.edu/rulrich/crmc.html</u>
- Model spread: EPOS-LHC, QGSJet-II.04, SIBYLL-2.3



Models mostly tuned to p+p data at  $|\eta| < 2$ : p+p 10 % model spread, p+O 50 % model spread



- Saturation visible in EPOS, not in QGSJet-II.04
- 7 % deviation in pO even if models could be fixed to same values in pp and pPb (50 % otherwise)
  - 4 % shift in  $N_{\mu}$ , 7 g cm-2 shift in  $X_{max}$  (comparable to exp. uncertainties)
- p+p and p+Pb may be able to constrain p+O, need measurement to confirm

# Energy flow ratio

- Hadronic energy "lost" to  $\pi^0$ s cannot produce muons in late shower
- "Energy loss" described by observable  $E_{e\gamma}/E_{hadrons}$



- Model predictions differ by **15 %** and in **shape**: only EPOS has forward peaks
- Translates to about **20 % shift in N\_{\mu} -> high impact on Muon Puzzle**

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## Energy flow ratio vs. system



- p+p and p+Pb together may be able to constrain p+O, but shape evolution not clear
- Need to measure it in p+O

## Summary

- Wanted: p-O collisions to accurately simulate hadronic showers in air
  - Current uncertainties 50 % in pion multiplicity, need better than 10 %
  - Needed by community of 900+ scientists (Auger, TA, IceCube, ...)
  - Moderate luminosity sufficient (100 M events)
  - Interest expressed by LHCf and members of LHCb, CMS, ATLAS
- Nuclear effects in proton-ion collisions cannot be accurately predicted (yet)
  - Cannot simply interpolate p-O from p-p and p-Pb
  - Effects largest in forward production which dominates air showers
- Measurements in p-O
  - Inelastic cross-section
  - Energy flow separated by hadrons and eγ
  - Spectra of light hadrons  $\pi$ , K, p
  - $\pi^0$ , n with LHCf in very forward range

## BACKUP

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#### LHCb



#### JINST 3 (2008) S08005 IJMP A 30 (2015) 1530022



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