

### Fluid Dynamics for Heavy Ion Collisions

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# Hydrodynamics in heavy-ion collisions

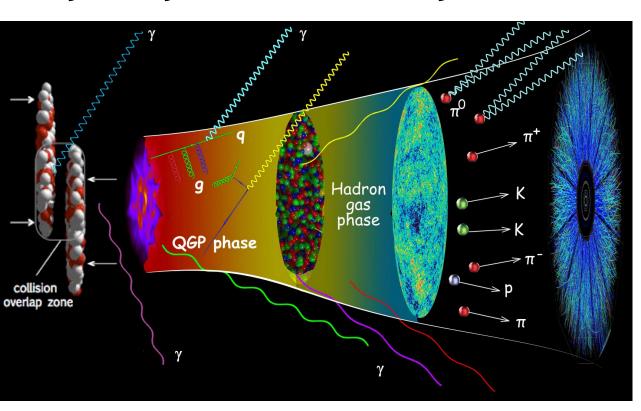
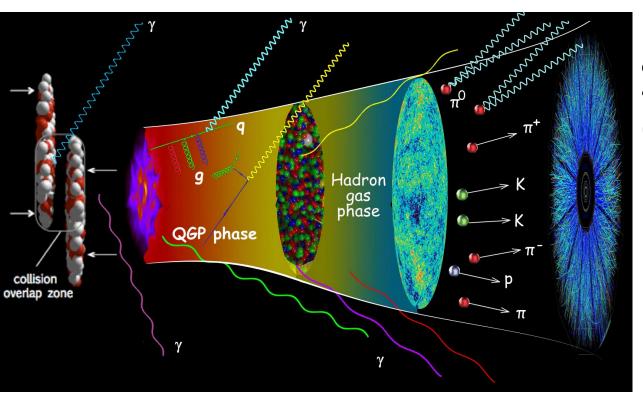


Figure ref.: Chun Shen

### Hydrodynamics in heavy-ion collisions



Hydrodynamics used as
effective description of the
"low energy modes"
(long wavelengths)

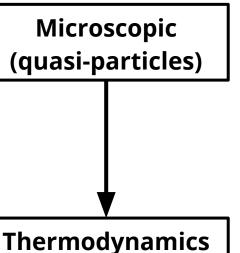
#### **Quasi-particles:**

(hard) partons, photons, heavy quarks, hadrons, ...

Figure ref.: Chun Shen

# Example of effective description: the ideal gas

#### (Non-relativistic) Ideal gas in equilibrium



Kinetic theory of gases (Maxwell, Boltzmann, ...)

$$f(\mathbf{p}) = (2\pi mT)^{-3/2} \exp\left[-\frac{\mathbf{p}^2}{2mT}\right]$$

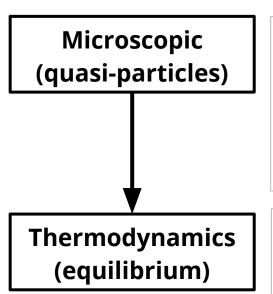
(Distribution of momentum of gas molecules)

**amics** Pressure "p", density "n/V", temperature "T"

Equation of state:  $p \propto \frac{n}{V}T$ 

# Effective description of the quark-gluon plasma

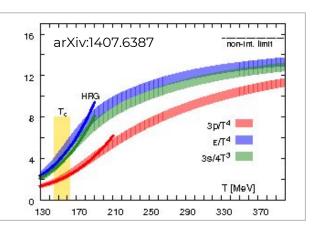
#### Quark-gluon plasma



- **Low energies (large a\_s):** Hadrons as quasi-particles
- ☐ Intermediate energies/ $\alpha_s$ : No general quasi-particle description
- $\blacksquare$  **High energies (small \alpha\_s):** Partons (quarks and gluons) as quasi-particles

Pressure "p", temperature "T", entropy density "s", energy density "ε"

Equation of state calculated with lattice QCD



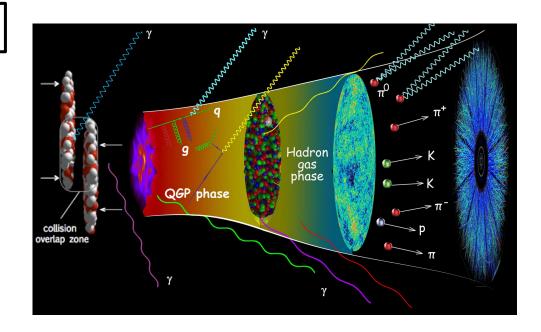
## Quasi-particles in heavy ion collisions

#### Quark-gluon plasma

#### **Microscopic (quasi-particles)**

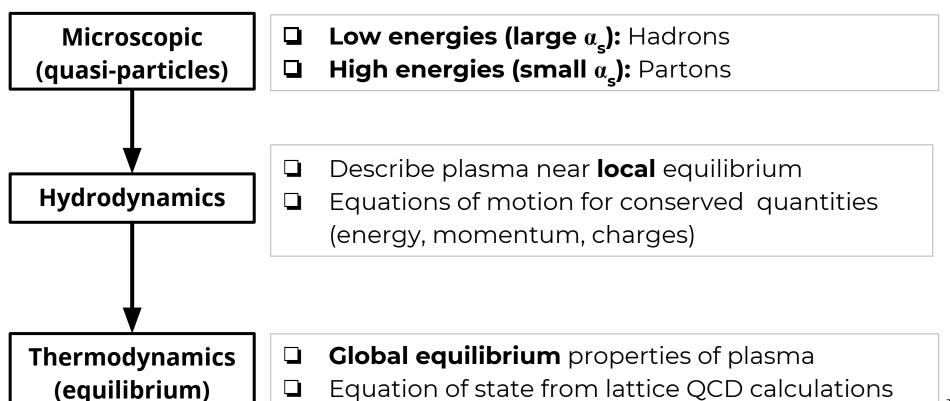
- Low energies (large  $\alpha_s$ ):

  Hadron as quasi-particles
- Intermediate energies/ $a_s$ :
  No general quasi-particle description
- $\Box$  High energies (small  $\alpha_s$ ): Quarks and gluons as quasi-particles



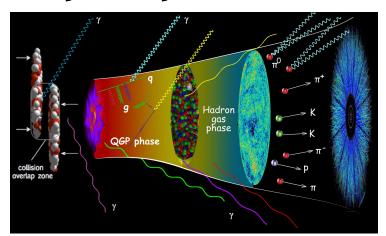
# Quasi-particles; thermodynamics; hydrodynamics

#### Quark-gluon plasma



7

### Hydrodynamics in heavy ion collisions



#### Effective description of the plasma?

- ☐ Global equilibrium/thermodynamics? No
- (Near) local-equilibrium:

Possible from a theory point of view, supported by phenomenological evidence

#### **Hydrodynamics**

Coarsegraining

- Describe plasma near local equilibrium
- Equations of motion for conserved quantities (energy, momentum, charges)
- Thermodynamics (equilibrium)
- ☐ Global equilibrium properties of plasma
- Equation of state from lattice QCD calculations

# Relativistic ideal hydrodynamics

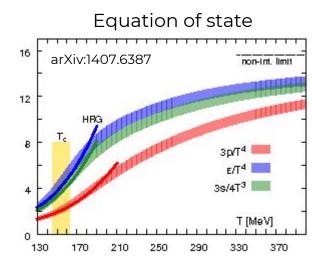
- Assume local equilibrium
- Energy-momentum tensor:

$$T^{\mu\nu} = (\epsilon + P(\epsilon)) u^{\mu} u^{\nu} - P(\epsilon) g^{\mu\nu}$$

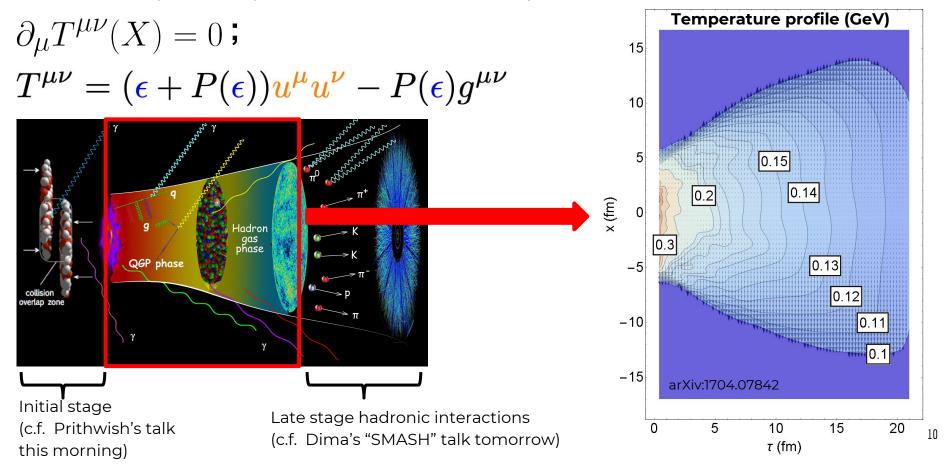
with

- ☐ "€" the energy density
- "u" the flow velocity ("rest frame")
- □ "P" the pressure which is related to "€" by the equation of state
- $\Box$  g<sup>µv</sup>=diag(1,-1,-1,-1)
- Local conservation of energy-momentum tensor:

$$\partial_{\mu}T^{\mu\nu} = \partial_{t}T^{t\nu} + \partial_{i}T^{i\nu} = 0$$



# Ideal hydrodynamics in heavy ion collisions



# Relativistic viscous hydrodynamics

- Near-local equilibrium
- Energy-momentum tensor:

$$T^{\mu\nu} = \epsilon u^{\mu}u^{\nu} - (P+\Pi)(g^{\mu\nu} - u^{\mu}u^{\nu}) + \pi^{\mu\nu}$$
 Bulk pressure Shear tensor

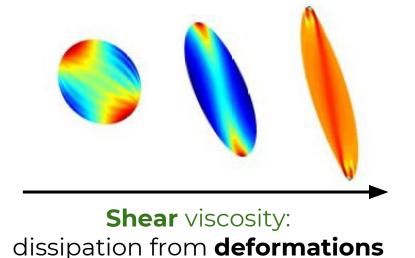
- $oldsymbol{\Box}$  Conservation of energy-momentum tensor:  $\partial_{\mu}T^{\mu
  u}(X)=0$
- Equations of motion for viscous part of energy-momentum tensor:

$$au_\Pi\dot{\Pi} + \Pi = -\zeta(\partial_\mu u^\mu) + ext{(second order terms)}$$

$$au_{\pi}\Delta^{\mu\nu}_{\alpha\beta}\dot{\pi}^{\alpha\beta} + \pi^{\mu\nu} = 2\eta\sigma^{\mu\nu} + \text{(second order terms)}$$

# Viscosity and transport coefficients

- ☐ Viscosities **convert kinetic energy into thermal energy** ("**dissipation**"; generates entropy)
- ☐ Shear and bulk viscosities are characteristic properties of the fluid that reflect this rate of dissipation



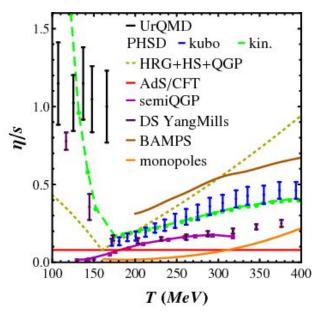


**Bulk** viscosity: dissipation from **expansion** 

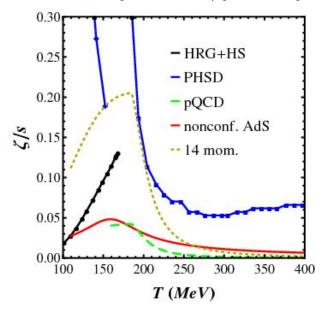
[ <u>Figures</u> <u>borrowed from</u> <u>Gabriel</u> <u>Denicol</u> ]

### Shear and bulk viscosities of QCD

Shear viscosity to entropy density

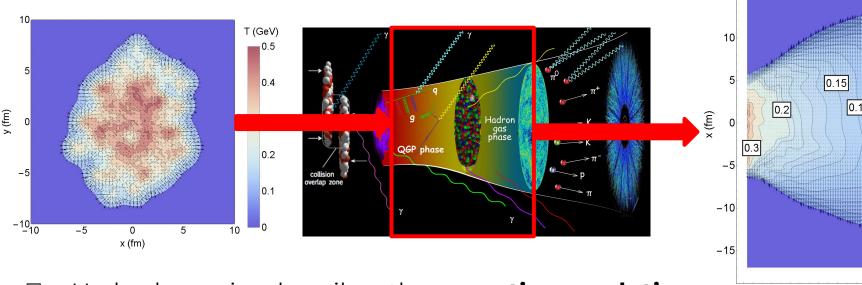


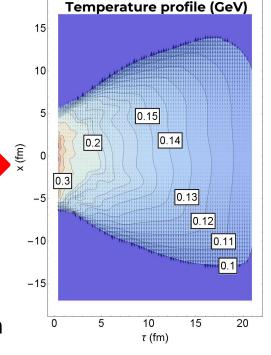
Bulk viscosity to entropy density



- Multiple theoretical attempt to compute shear and bulk viscosity
- Heavy ion collisions are used to constrain them experimentally

### Initial conditions



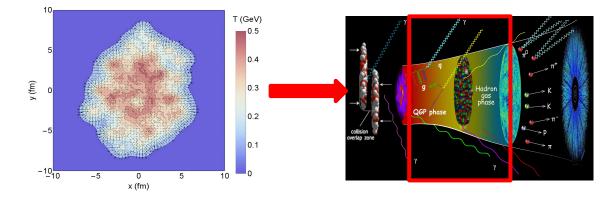


- Hydrodynamics describes the spacetime evolution of the plasma given initial ("boundary") conditions
- The full energy-momentum tensor must be initialized:

$$T^{\mu\nu} = \epsilon u^{\mu} u^{\nu} - (P + \Pi)(g^{\mu\nu} - u^{\mu} u^{\nu}) + \pi^{\mu\nu}$$

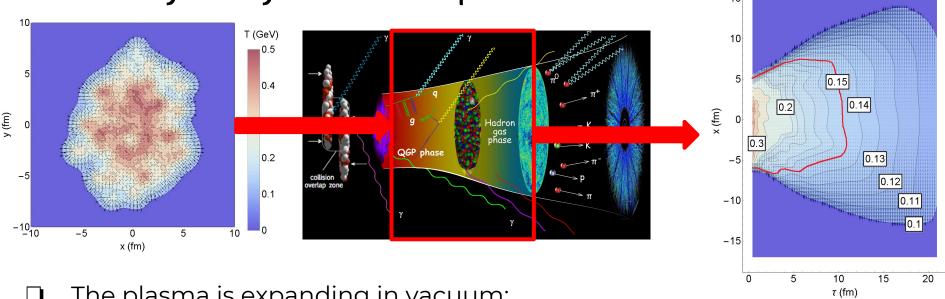
#### Initial conditions

(c.f. Prithwish Tribedy's talk this morning)



- General approach:
  - Describe microscopic degrees of freedom in the early stage of the plasma
  - lacktriangle Coarse-grain the degrees of freedom to get initial hydrodynamic  $T^{\mu\nu}$
- "Microscopic" degrees of freedom can be:
  - ☐ Gluons/quarks (e.g. Color Glass Condensate, IP-Glasma)
  - Nucleons (e.g. MC Glauber, "Trento" model)
- Initial conditions can be obtained from multistage model themselves
   e.g. (IP-Glasma or Trento) + (KøMPøST or free-streaming)

# From hydrodynamics to particles



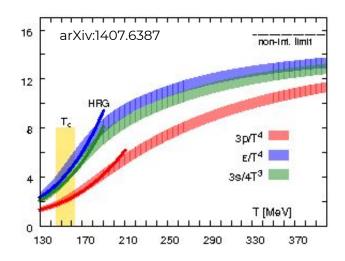
- The plasma is expanding in vacuum: its energy density/temperature is decreasing rapidly
- ☐ Low enough energy density: can't use a hydrodynamic description

Hydrodynamics drives the plasma out of equilibrium

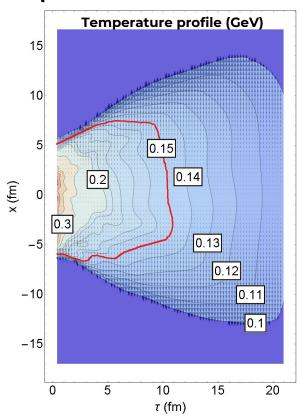
Temperature profile (GeV)

### Particlization: from hydrodynamics to particles

■ At temperatures T~160 MeV, the degrees of freedom of the plasma are hadrons



At T< ~160 MeV, dual description of the plasma: (i) hydrodynamics and (ii) gas of interacting hadrons [recall the ideal gas example]



# From hydrodynamics to hadron resonance gas

■ Energy-momentum tensor of the hadron gas:

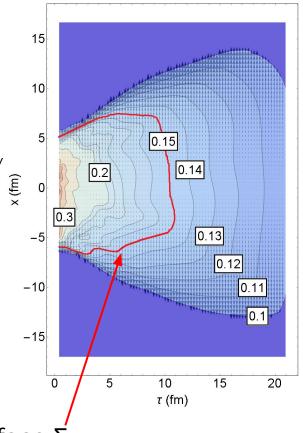
$$T^{\mu\nu}(X) = \sum_h^{\text{Sum over all hadron species and resonances}} \frac{d^3p}{P^0(2\pi)^3} P^\mu P^\nu g_h f_h(P,X)$$
 Hadron degeneracy

- In ideal hydrodynamics, f<sub>h</sub>(P,X) is the Fermi-Dirac or Bose-Einstein distribution
- Converting the hydrodynamic fluid to particles:

#### "Cooper-Frye"

$$E\frac{d^3N}{d\mathbf{p}} = \frac{g_h}{(2\pi)^3} \int_{\Sigma} d\Sigma_{\mu} P^{\mu} f_h(P \cdot u, X)$$

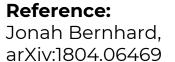
Particle flux flows out a surface  $\Sigma$ 

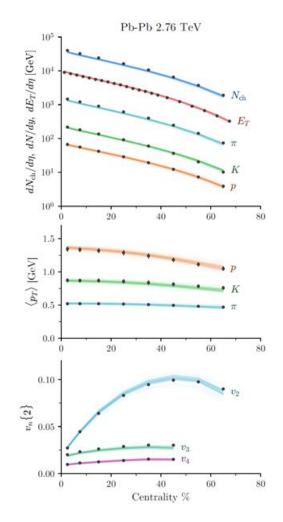


# Applications in heavy ion collisions

- By combining relativistic viscous hydrodynamics with
  - (i) a model of initial conditions
  - (ii) a late-stage hadronic afterburner

we can **describe** a wide range of **hadronic measurements** in heavy ion collisions

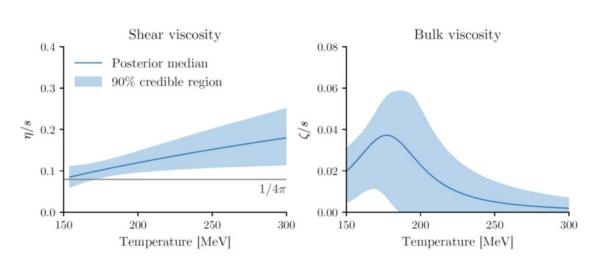


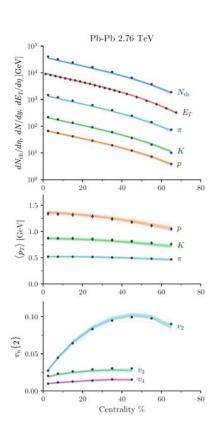


### Constraints on shear and bulk viscosities

**Reference:**Jonah Bernhard, arXiv:1804.06469

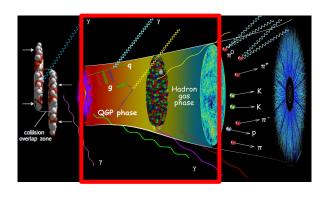
- By combining relativistic viscous hydrodynamics, initial conditions and a late-stage hadronic afterburner, we can
  - describe hadronic measurements
  - put constraints on shear and bulk viscosities





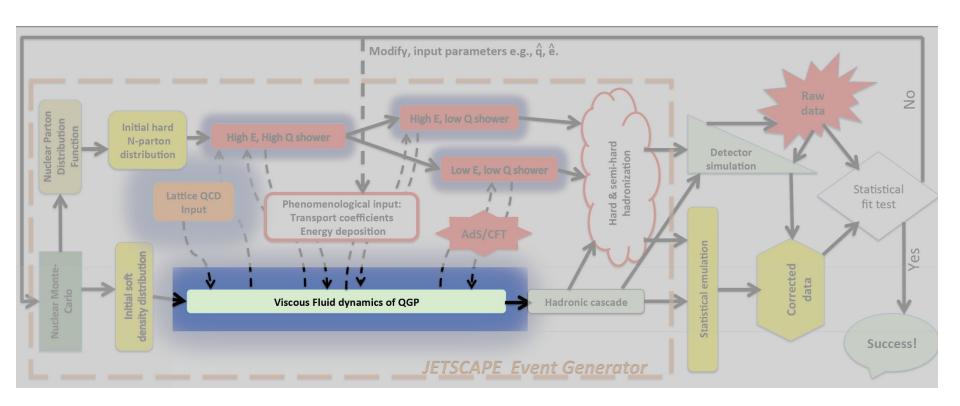
# Summary

Hydrodynamics is an effective description of matter created in heavy ion collisions



- Hydrodynamics is one key part of a multistage model of heavy ion collisions (initial stage and late stage described separately)
- Hybrid (multistage) model successful in describing (soft) hadronic observables --- allowing us to study transport properties of QGP
- Hydrodynamics provides realistic medium evolution event-by-event for parton energy loss

## Hydrodynamics with JETSCAPE



### Hands on Session

#### Goals

- Get familiar with running the JETSCAPE framework coupled with a hydrodynamic module (MUSIC)
- Understand the input options and output files from MUSIC\_evo
  - Change collision system (type of nucleus, collision centrality, collision energy)
  - Change the transport coefficient in hydrodynamic simulations (shear viscosity)
  - Make plots and animations for a hydrodynamic evolution (temperature, flow velocity)
- (Bonus) Further integrate a particlization module to generate a full event with particles
  - Trento + MUSIC + iSS



#### Instructions

- Download and compile JETSCAPE with MUSIC
  - Go to external\_packages/ folder and run ./get\_music.sh
  - Go to build/ folder and run cmake .. -Dmusic=ON
  - o After the compilation, run MUSIC\_evo to generate events
  - Change parameters in the jetscape\_init.xml
- Make plots and animations
  - Download the scripts
    - git clone <a href="https://github.com/JETSCAPE/WinterSchool2019">https://github.com/JETSCAPE/WinterSchool2019</a>
  - Open Jupyter notebook
    - Update your running directory
    - Have fun~