

*MEASUREMENT OF THE PROTON FORM
FACTOR IN e^+e^- ANNIHILATION AT
BABAR*



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SLAC

On Behalf of the *BABAR* Collaboration

2013 DPF
UC SANTA CRUZ

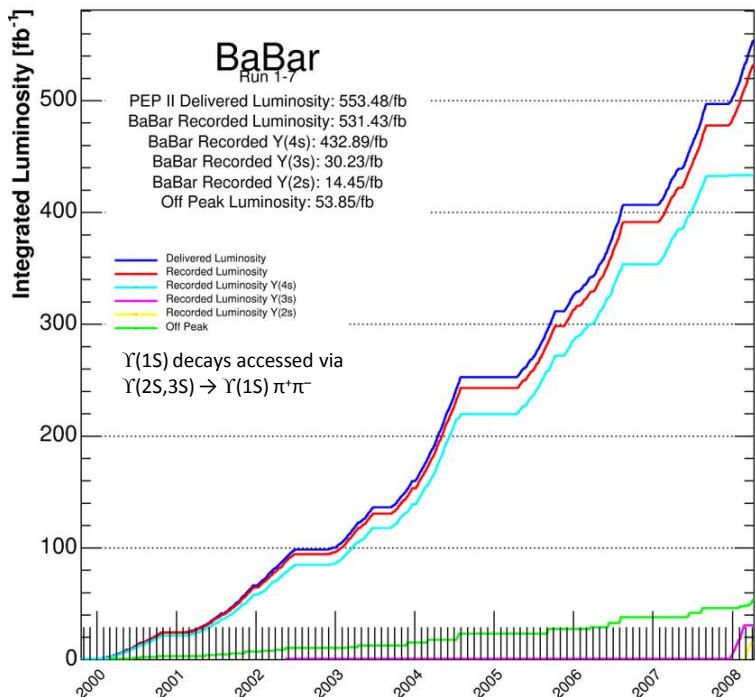
DPF 2013

Meeting of the American Physical Society Division of Particles and Fields

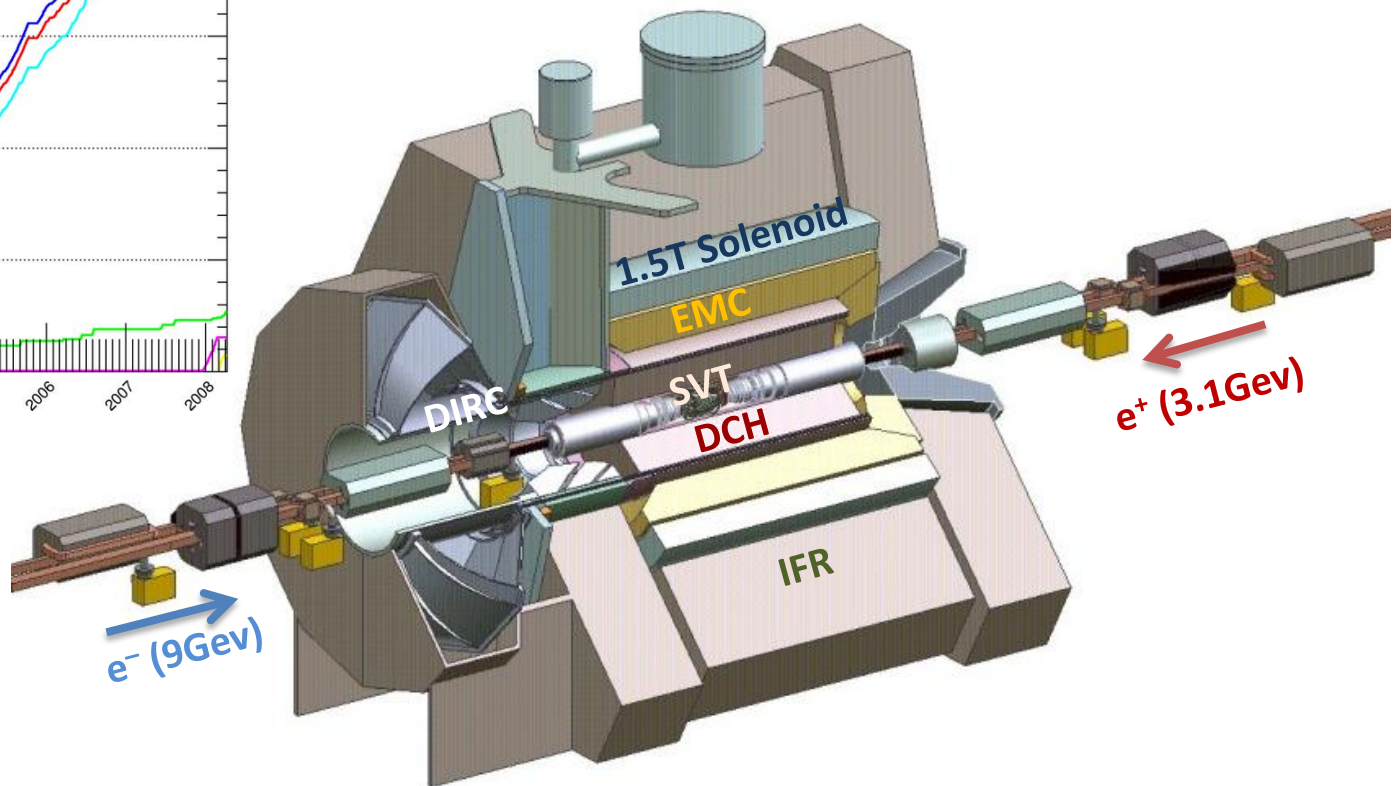
Santa Cruz, CA, August 13-17, 2013



BABAR DATASET



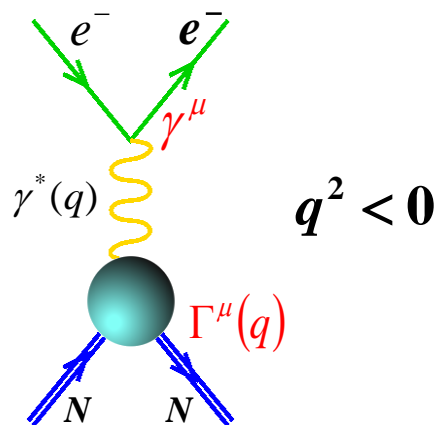
Data used for the analyses presented:
On $\Upsilon(4S)$ resonance (10.58Gev): 424 fb⁻¹
Off $\Upsilon(4S)$ resonance (10.54Gev): 45 fb⁻¹



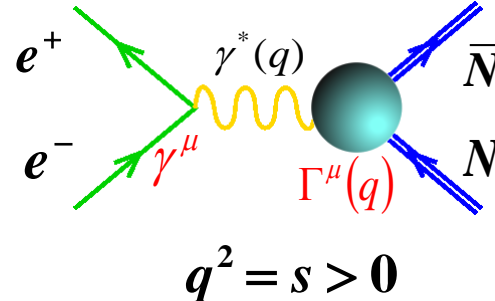


NUCLEONS FORM FACTORS

Space-like



Time-like



$$\Gamma^\mu(q) = \gamma^\mu F_1(q^2) + \frac{i\sigma^{\mu\nu}}{2M_N} q_\nu F_2(q^2)$$

- Dirac (F_1) and Pauli (F_2) FF

$$F_1^p(0) = 1 \quad F_2^p(0) = k_p$$

$$F_1^n(0) = 0 \quad F_2^n(0) = k_n$$

k_N =nucleon anomalous magnetic moment

- Sachs FF: Electric (G_E) and Magnetic (G_M)

$$G_E(q^2) \equiv F_1(q^2) + \frac{q^2}{4M^2} F_2(q^2)$$

$$G_M(q^2) \equiv F_1(q^2) + F_2(q^2)$$

- Notable points:

$$G_E(0) = 1 ; \quad G_M(0) = 2.79$$

$$|G_E(4m_p^2)| = |G_M(4m_p^2)| \quad (\text{p}\bar{\text{p}} \text{ production threshold})$$



$e^+e^- \rightarrow p\bar{p}$ CROSS SECTION

- Cross section of the annihilation process:

$$\sigma(s) = \frac{4\pi\alpha^2\beta C(s)}{3s} \left(|G_M(s)|^2 + \frac{2m_p^2}{s} |G_E(s)|^2 \right) \quad \sqrt{s} = 2E_{beam}^*$$

C is the Coulomb factor → The cross section is non-zero at the threshold

- The ratio $|G_E/G_M|$ can be determined from the analysis of the polar angle distribution of the proton:

$$\frac{d\sigma}{d\Omega}(s, \theta) = \frac{\alpha^2\beta C(s)}{4s} \left(|G_M(s)|^2 (1 + \cos^2 \theta) + \frac{4m_p^2}{s} |G_E(s)|^2 \sin^2 \theta \right)$$

- From the measured cross section we define an effective form factor:

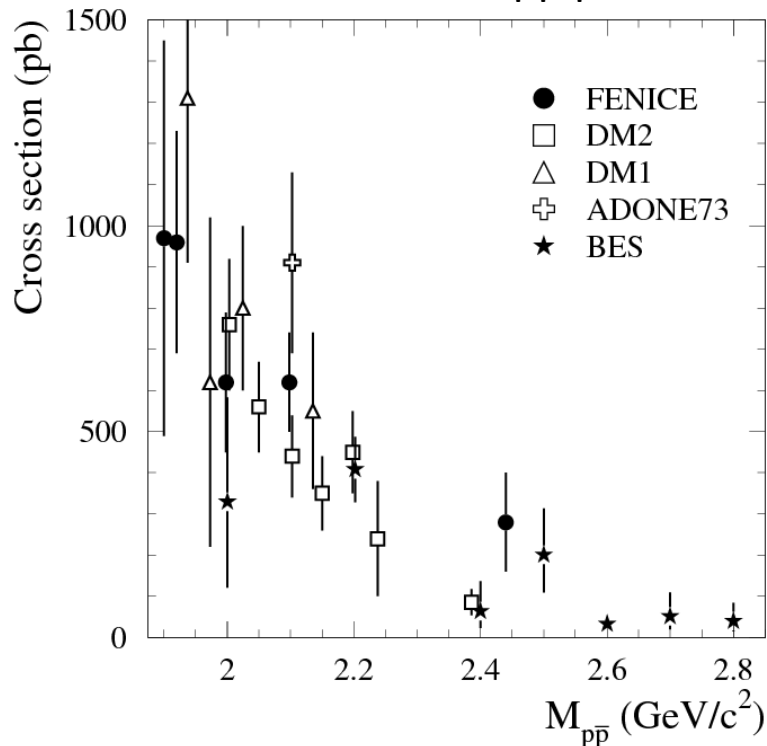
$$|F_p(s)| = \sqrt{\frac{|G_M(s)|^2 + (2m_p^2/s) |G_E(s)|^2}{1 + 2m_p^2/s}}$$

- At high energies G_E is suppressed as $1/E^2$ and $|F_p| \sim |G_M|$.



PREVIOUS MEASUREMENTS (1/2)

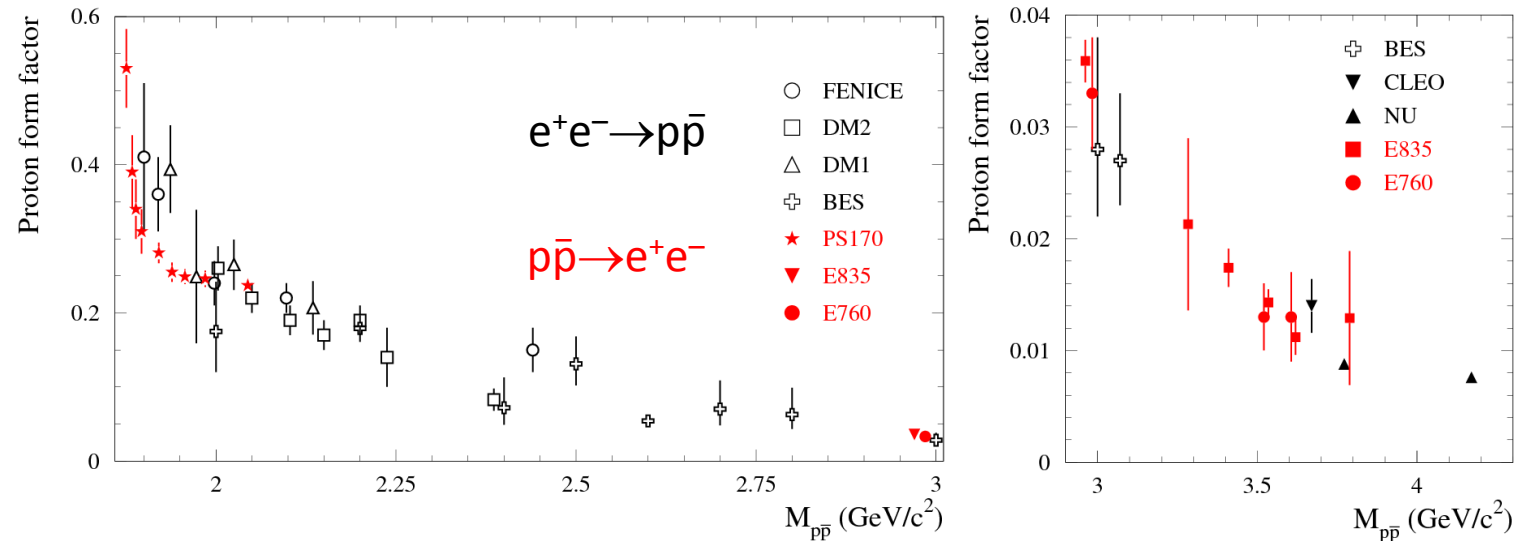
$e^+e^- \rightarrow p\bar{p}$ processes



- Below 3.2 GeV:

- ADONE73 (1973), DM1(1979), DM2(1983,1990), FENICE(1994), BES(2005).
- Statistical accuracy of these measurements is (20-30)%.
- Limited statistics does not allow to extract the G_E/G_M ratio from the analysis of the angular distribution.

PREVIOUS MEASUREMENTS (2/2)

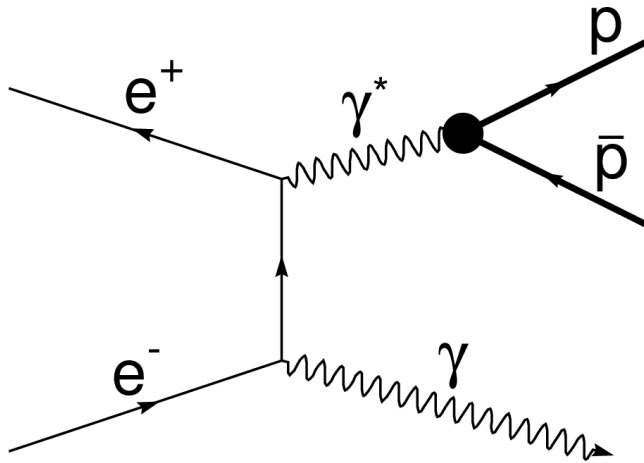


The statistics of these experiments are usually not enough to study angular distributions. All data were obtained under the assumption that $|G_E| = |G_M|$.

- Near threshold data were obtained at the PS170 experiment at the antiproton storage ring LEAR (CERN) :
 - Steep growth of the form factor near threshold.
 - The ratio $|G_E/G_M|$, measured with 30% accuracy in five energy points, agrees with unity
- Measurements above 3GeV were performed at FNAL (E835 and E760)
- The very precise points marked “NU” (Phys. Rev. Lett. 110 (2013) 022002) were obtained using CLEO data($\sim 1/\text{fb}$) collected at 3.77 and 4.17 GeV



THE ISR METHOD (1/2)



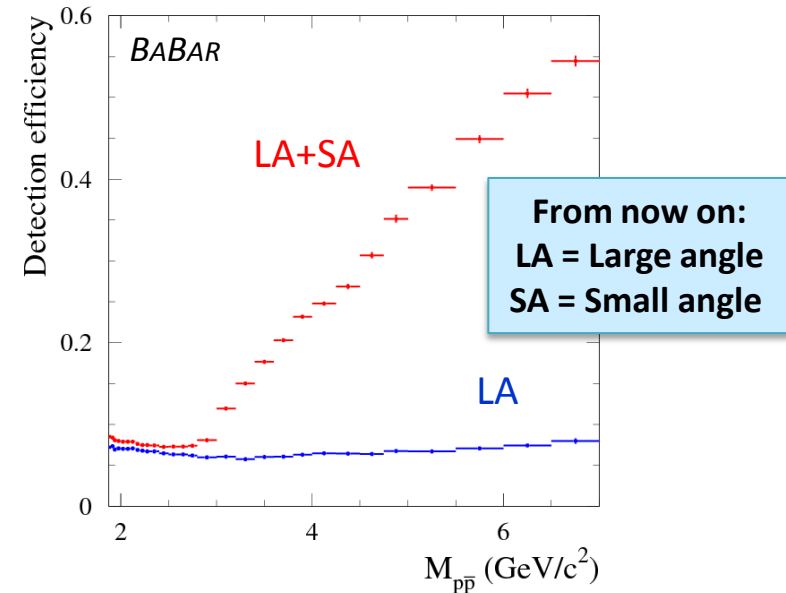
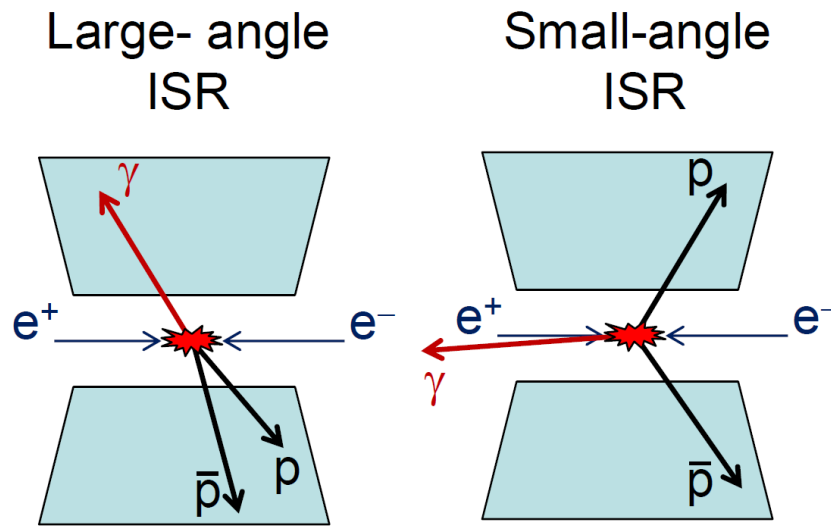
$$\frac{d\sigma_{e^+e^- \rightarrow p\bar{p}\gamma}}{dm d\cos\theta} = \frac{2m}{s} W(s, x, \theta) \sigma_{e^+e^- \rightarrow p\bar{p}}(m)$$

$$x = \frac{2E_\gamma}{\sqrt{s}} = 1 - \frac{m^2}{s}$$

- The $p\bar{p}$ mass spectrum in $e^+e^- \rightarrow \gamma p\bar{p}$ is related to the cross section for the non radiative process $e^+e^- \rightarrow p\bar{p}$
- $W(s, x, \theta)$ is calculated from QED and describes angular ($1/\sin^2\theta$ at $\theta \gg m_e/\sqrt{s}$) and energy ($1/x$) distributions of the ISR photon.



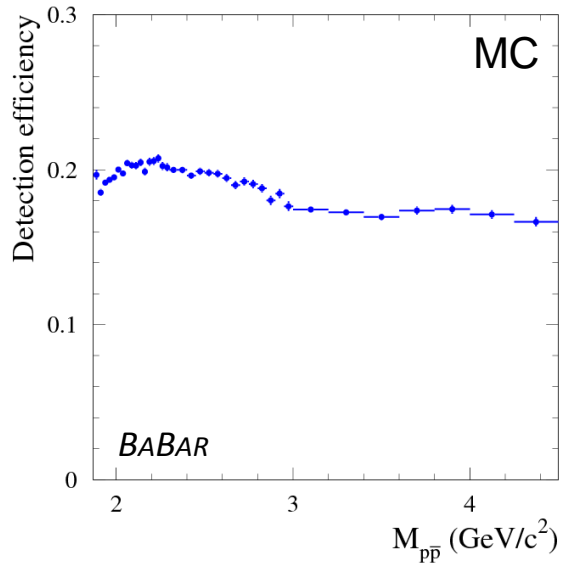
THE ISR METHOD (2/2)



- The ISR photon is emitted predominantly along the beam axis. The produced hadronic system is boosted against the ISR photon. Due to limited detector acceptance the mass region below 3 GeV can only be studied with detected photons (about 10% of ISR events).
- The statistics can be significantly increased above 3 GeV by using small-angle ISR.

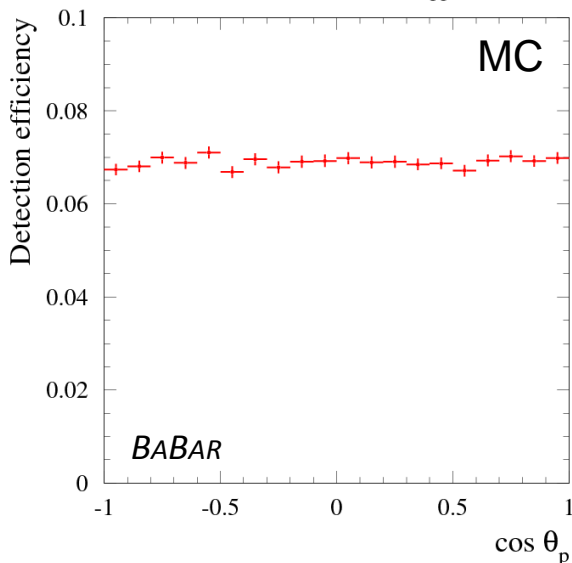


ADVANTAGES OF THE ISR METHOD



- A wide energy region is studied in a single experiment.
 - The effective ISR luminosity ($\text{pb}^{-1}/\text{GeV}$) increases with mass, partly compensating a decrease of the measured cross section.

- Low dependence of the detection efficiency on the hadronic invariant mass.
 - Measurement near and above threshold with the same selection criteria.

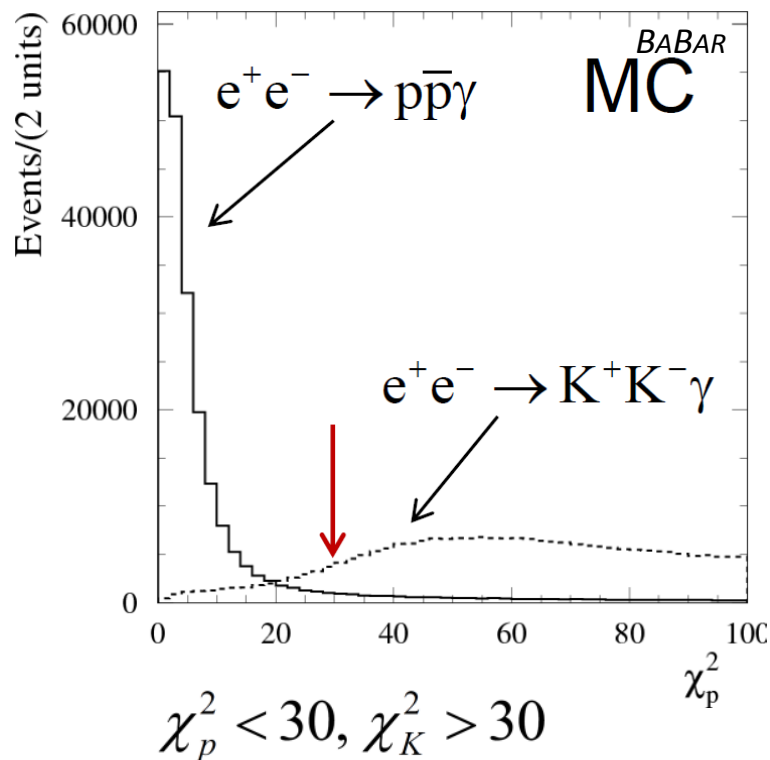


- Low dependence of the detection efficiency on hadron angular distributions (in the hadron rest frame).
 - For protons this significantly increases sensitivity for measurements of the G_E/G_M ratio.



LA EVENT SELECTION

- The LA analysis is published in:
 - **Phys. Rev. D 87, (2013) 092005 (469 fb⁻¹).**
 - **Phys. Rev. D 73, (2006) 012005 (232 fb⁻¹).**

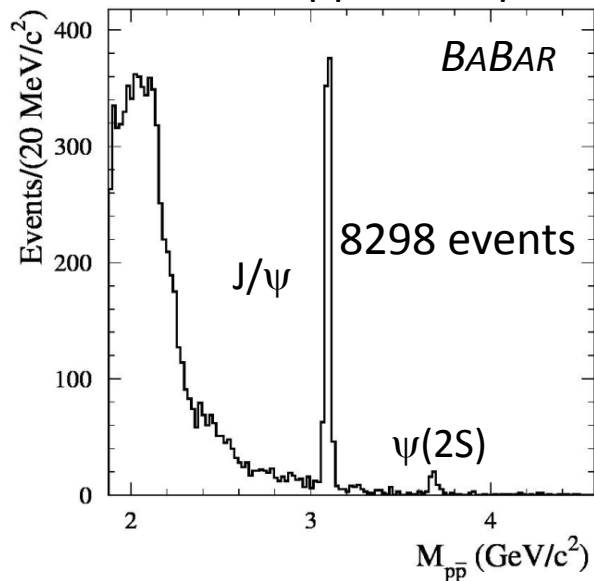


- 2 tracks of opposite charge originating from the IP and identified as protons ($25.8^\circ < \theta < 137.5^\circ$).
- A photon with $E_{\text{CM}} > 3 \text{ GeV}$ ($20.0^\circ < \theta < 137.5^\circ$).
- Kinematic fit to $e^+e^- \rightarrow \gamma h^+ h^-$ ($h=p, K$) requiring p and E conservation.
- $e^+e^- \rightarrow \gamma\pi^+\pi^-, \gamma\mu^+\mu^-, \gamma K^+K^-$ backgrounds suppressed by PID requirements and χ^2 cuts.



BACKGROUNDS FOR LA SELECTION

Reconstructed $p\bar{p}$ mass spectrum



- Main background from $e^+e^- \rightarrow p\bar{p}\pi^0$
 - Undetected low-energy photon or merged photons from the π^0 decay.
 - Estimated from data using a control sample of $e^+e^- \rightarrow p\bar{p}\pi^0$ events.
 - Increases from 5% at $M_{pp} < 2.5$ GeV to 50% at 4 GeV.
 - All data events observed above 4.5 GeV are explained by this background.

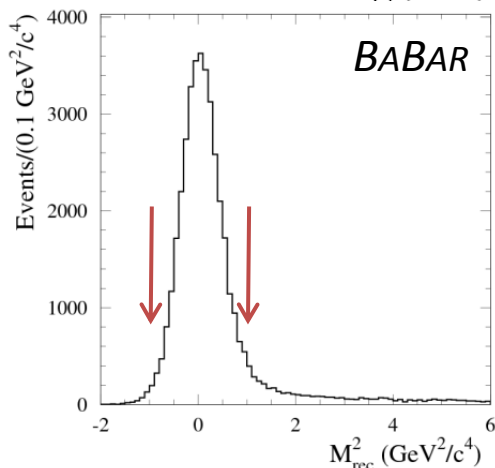
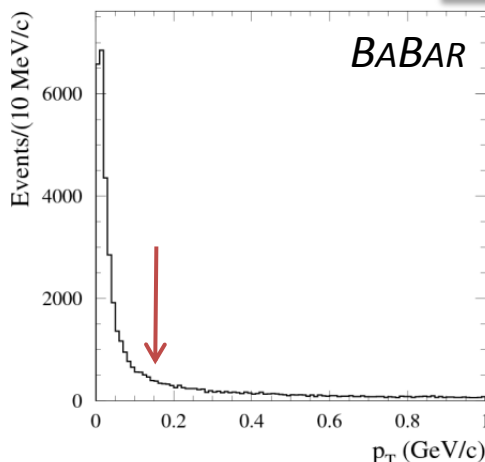
Data	$pp\gamma$	$pp\pi^0$	ISR and e^+e^-
8298	7741 ± 113	448 ± 42	109 ± 30



SA EVENT SELECTION

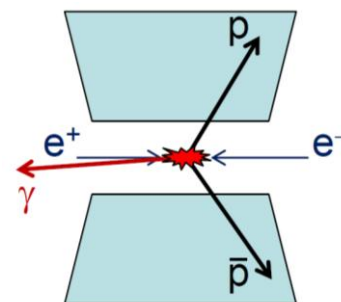
- SA analysis is a preliminary result:
 - <http://arxiv.org/abs/1308.1795>

$e^+e^- \rightarrow \gamma p \bar{p}$ (MC)



- 2 tracks of opposite charge originating from the IP and identified as protons ($25.8^\circ < \theta < 137.5^\circ$)
- $P_T < 0.15 \text{ GeV}/c$
- $|M_{\text{rec}}^2| < 1 \text{ GeV}^2/c^4$

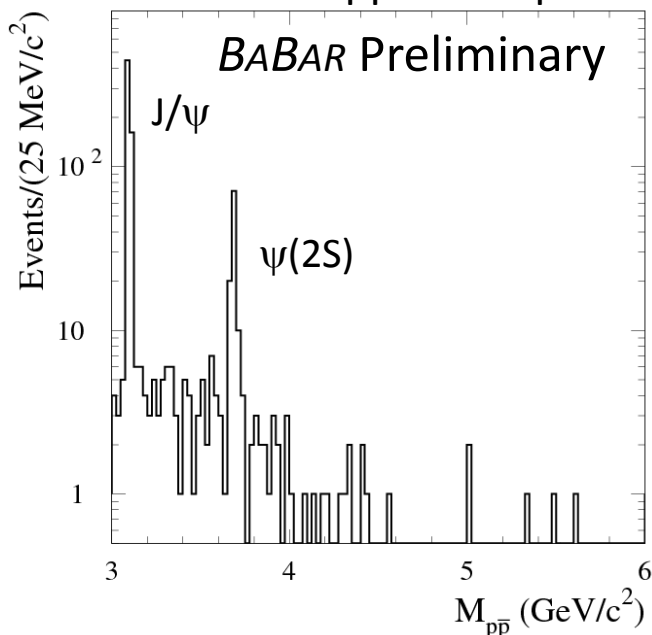
No requirement on the ISR photon!





BACKGROUND FOR SA ISR

Reconstructed $p\bar{p}$ mass spectrum

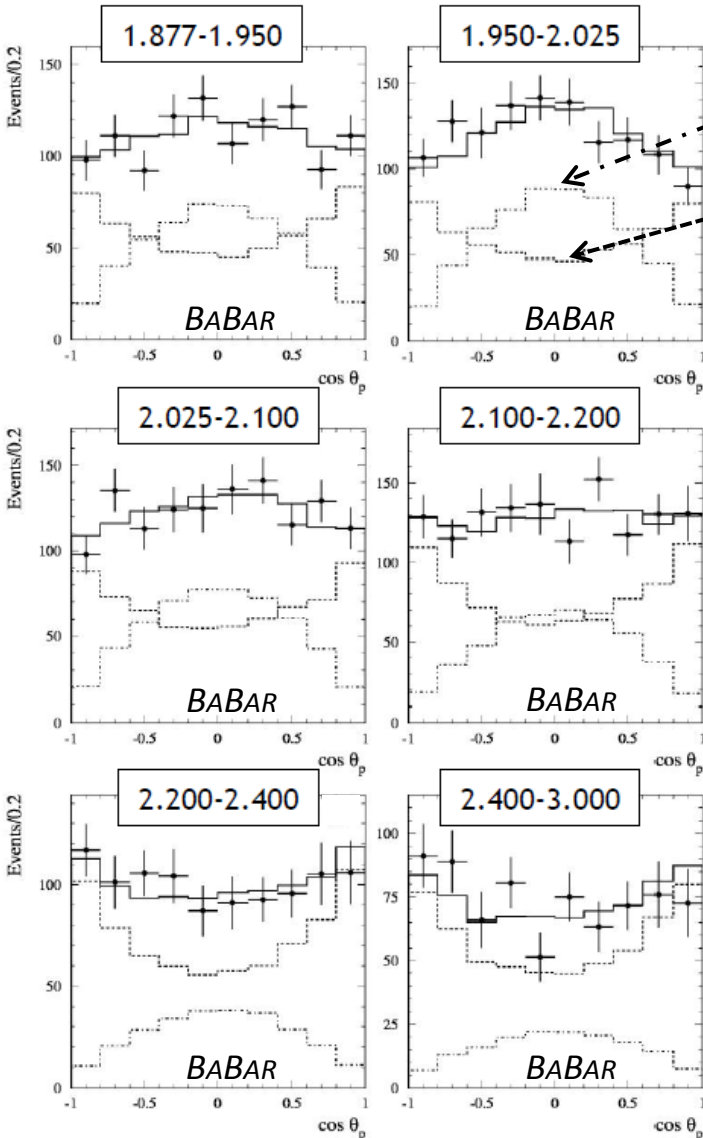


- The dominant background process is $e^+e^- \rightarrow \gamma p\bar{p}\pi^0$
- The main background in our large-angle ISR analysis from $e^+e^- \rightarrow p\bar{p}\pi^0$ is found to be negligible.
- Selected events: 845.
- Non resonant events: 149.
- No events are observed with mass above 6 GeV/c^2 .

Data	$\mu\mu\gamma$	2-photons	ISR
149 ± 12	< 1.2	1.6 ± 0.9	6.3 ± 3.5

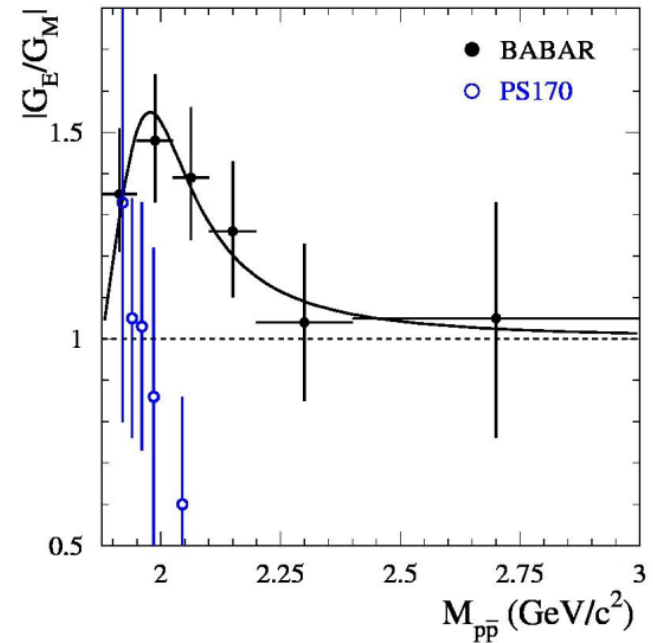


LA ANGULAR DISTRIBUTION



$$\frac{d\sigma(G_E)}{d\cos\theta} \sim \sin^2\theta_p$$

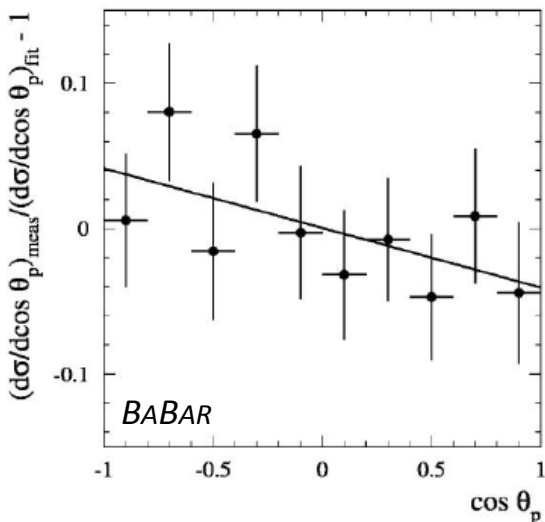
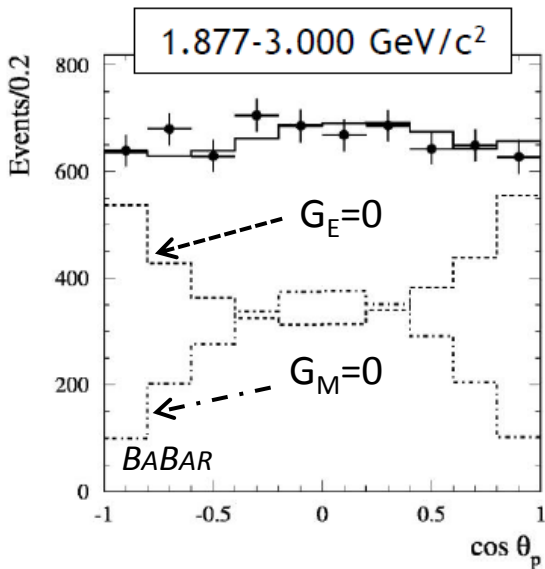
$$\frac{d\sigma(G_M)}{d\cos\theta} \sim 1 + \cos^2\theta_p$$



- θ_p is the angle between the proton momentum in the $p\bar{p}$ rest frame and the momentum of $p\bar{p}$ system in the e^+e^- c.m. frame.
- The distribution is fitted by a sum of histograms from two simulated samples, one with $G_E=0$ (dashed) and other with $G_M=0$ (dot-dashed).



LA ANGULAR DISTRIBUTION ASYMMETRY



- The asymmetry is absent in lowest order ($\gamma^* \rightarrow p\bar{p}$). It arises from higher-order contributions (soft extra ISR and FSR interference, two-photon exchange). By measuring the asymmetry we control the higher-order contributions.
- We analyze the difference between the measured and fitted distributions.

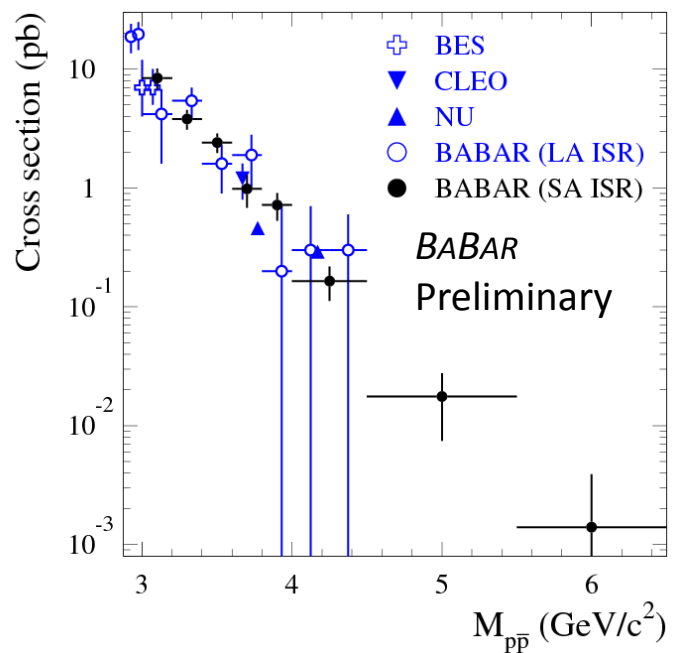
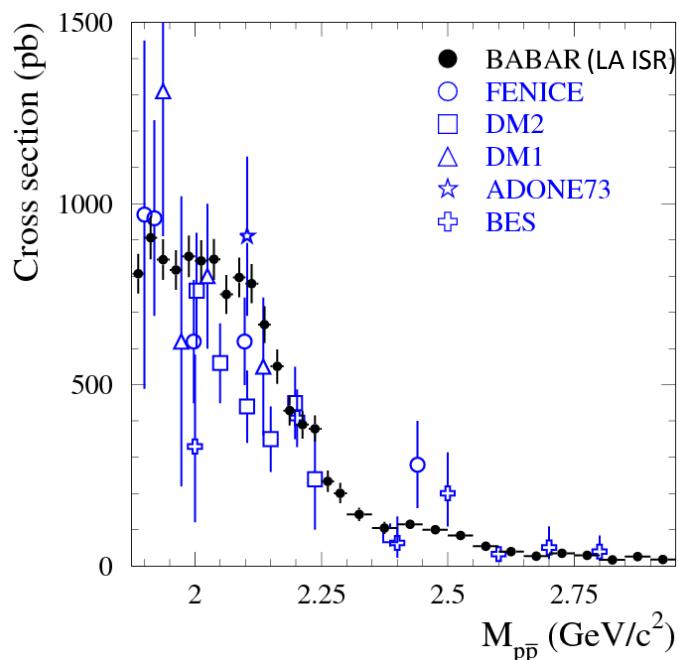
- The slope is $-0.041 \pm 0.026 \pm 0.005$

- The integral asymmetry :

$$A = \frac{\sigma(\cos \theta_p > 0) - \sigma(\cos \theta_p < 0)}{\sigma(\cos \theta_p > 0) + \sigma(\cos \theta_p < 0)} = -0.025 \pm 0.014 \pm 0.003$$



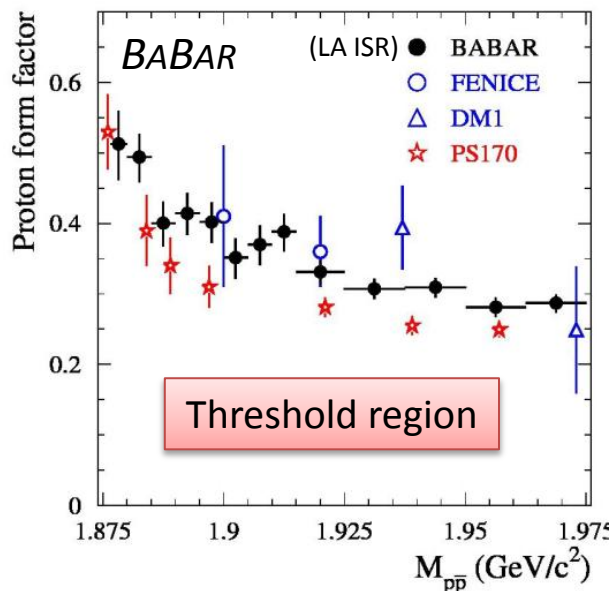
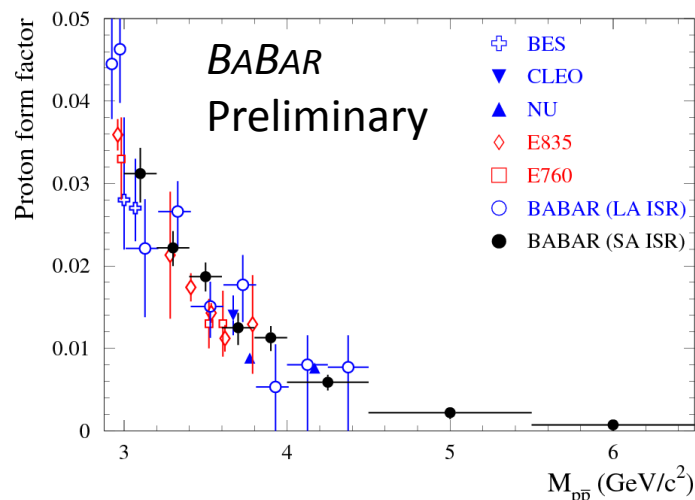
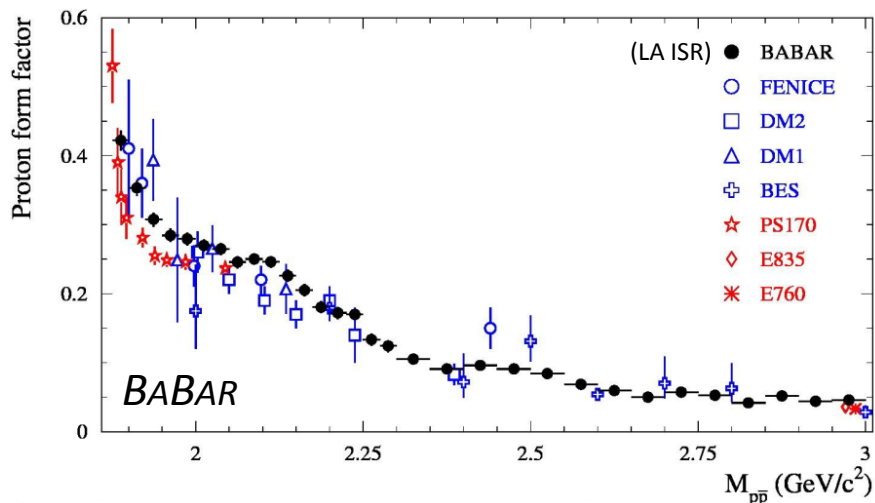
MEASURED CROSS SECTION $\sigma_{ee \rightarrow p\bar{p}}(m) = \frac{dN/dm}{\epsilon R dL/dm}$



- Mass-independent systematic uncertainty is 4-5% for LA ISR. For SA ISR it decreases from 16% at 3 GeV to 6% at 4.5 GeV.
- In the mass region under study the cross section changes by about six orders of magnitude.
- Our data are in reasonable agreement with previous measurements.
- We improve accuracy and extend the mass region of measurements.



EFFECTIVE PROTON FORM FACTOR



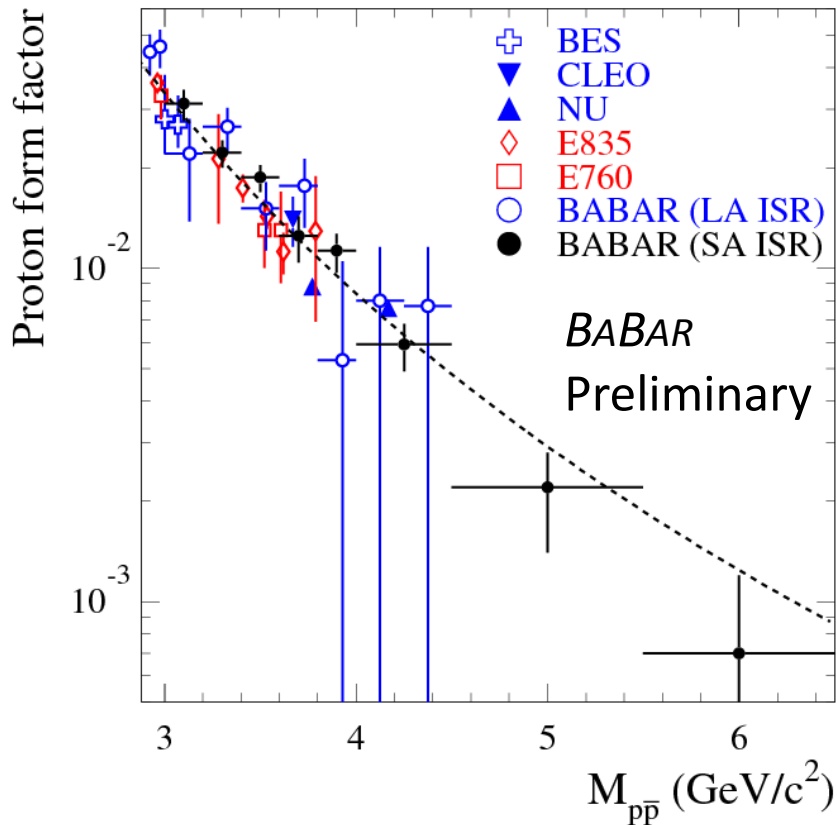
$$|F_p(s)| = \sqrt{\frac{|G_M(s)|^2 + (2m_p^2/s) |G_E(s)|^2}{1 + 2m_p^2/s}}$$

Effective proton form factor

- A steep fall off and $G_E/G_M \neq 1$ near threshold are explained by nucleons' final state interaction.
- A complex step-like behavior far off threshold (steps at 2.2, 2.55, 3 GeV/c²) is not described by existing models.



QCD MOTIVATED FIT

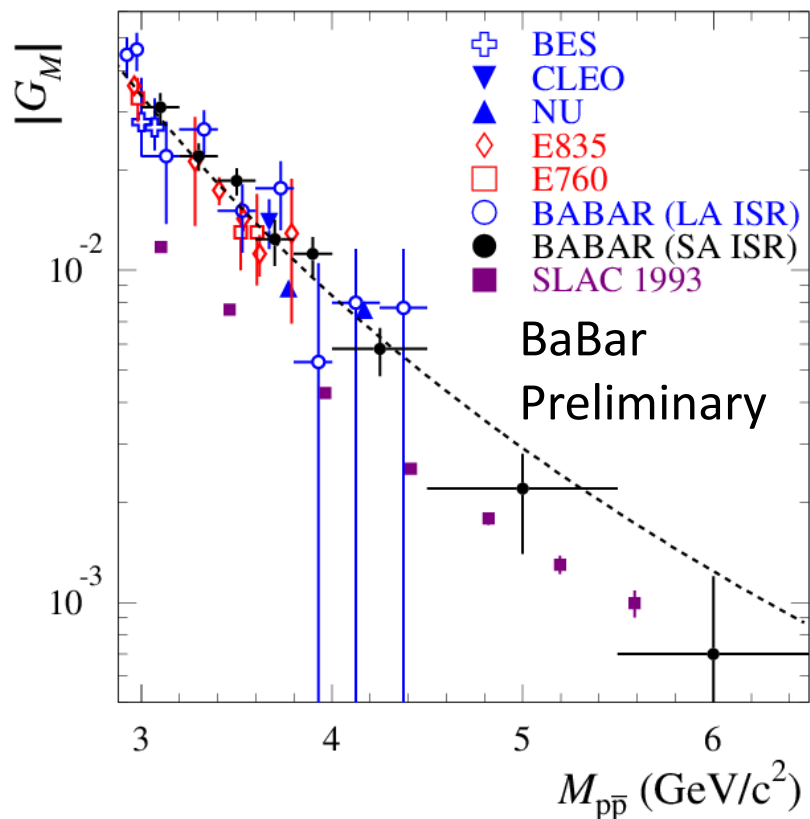


$$G_M \sim \frac{\alpha_s^2(m^2)}{m^4} \sim \frac{1}{m^4 \ln^2(m^2 / \Lambda^2)}$$

- All the data above 3 GeV except the two “NU” points are well described by this function.
- Adding the “NU” points change the fit χ^2/ν from 17/24 to 54/26.
- Our data shows that the form factor decreases in agreement with the asymptotic QCD prediction or even faster above 4.5 GeV.



COMPARISON WITH SPACE-LIKE G_M

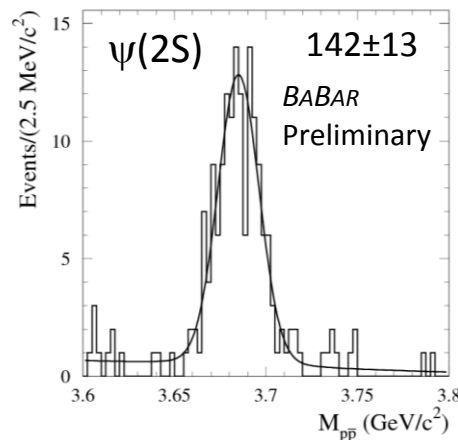
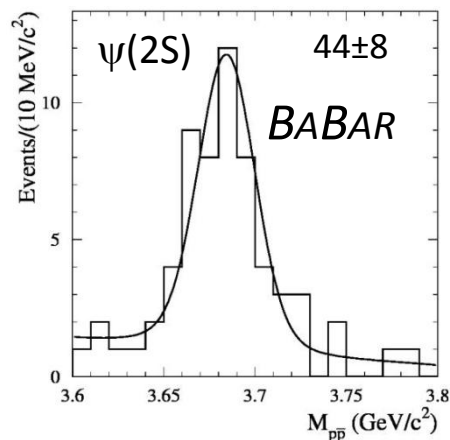
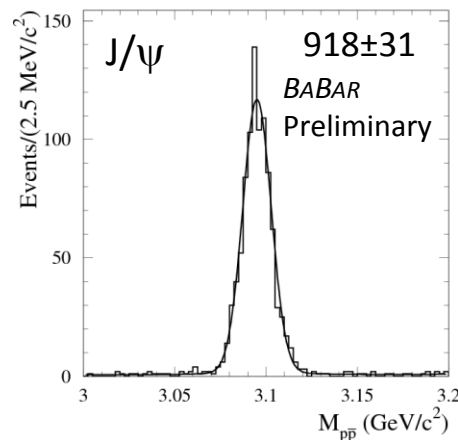
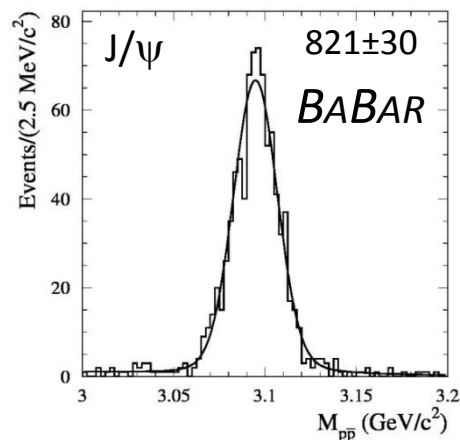


Our two points above 4.5 GeV may give an indication that the difference between the time- and space-like form factors begins to decrease.

- The points “SLAC 1993” represent data on the space-like magnetic form factor measured in ep scattering as a function of $\sqrt{-q^2}$.
- The asymptotic values of the space- and time-like form factors are expected to be the same.
- In the mass region from 3 to 4.5 GeV the time-like form factor is about two times larger than the space-like one.



J/ψ AND $\psi(2S)$ DECAYS



- $B(J/\psi \rightarrow p\bar{p}) \times 10^{-3}$
 - SA: $2.33 \pm 0.08 \pm 0.09$
 - LA: $2.04 \pm 0.07 \pm 0.07$
 - PDG: 2.17 ± 0.07
- $B(\psi(2S) \rightarrow p\bar{p}) \times 10^{-4}$
 - SA: $3.14 \pm 0.28 \pm 0.18$
 - LA: $2.86 \pm 0.51 \pm 0.09$
 - PDG: 2.76 ± 0.12

• LA ISR

• SA ISR



SUMMARY

- The $e^+e^- \rightarrow p\bar{p}$ cross section and the proton effective form factor have been measured from threshold up to 6.5 GeV using the full BaBar data sample.
- The form factor shows a complex mass dependence. There are a near-threshold steep fall off and a step-like behavior at higher masses.
- The $|G_E/G_M|$ ratio has been measured from threshold up to 3 GeV/c². A large deviation of this ratio from unity is observed below 2.2 GeV/c².
- Asymmetry in the proton angular distribution has been measured.
- At masses above 3 GeV the observed decrease of the form factor agrees with the asymptotic dependence $\alpha_s^2(m)/m^4$ predicted by QCD.
- New measurements expected : VEPP2000, BES, FAIR,...



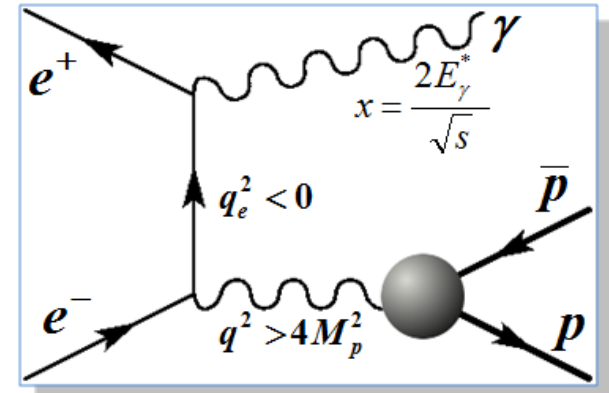


ISR CROSS SECTION

Born approximation

$$\frac{d\sigma_{e^+e^- \rightarrow p\bar{p}\gamma}(s, m_{p\bar{p}})}{dm d\cos\theta_\gamma^*} = \frac{2m}{s} W(s, x, \theta_\gamma^*) \cdot \sigma_{e^+e^- \rightarrow p\bar{p}}(m_{p\bar{p}})$$

$$x = \frac{2E_\gamma^*}{\sqrt{s}} \quad m_{p\bar{p}}^2 = q^2 = s(1-x) \quad \theta_\gamma^*: \text{ISR angle in } e^+e^- \text{ c.m.}$$



Radiator function:

$$W(s, x, \theta_\gamma^*) = \frac{\alpha}{\pi \cdot x} \cdot \left(\frac{2 - 2x + x^2}{\sin^2\theta_\gamma^*} - \frac{x^2}{2} \right)$$

$e^+e^- \rightarrow p\bar{p}$ cross section

$$\sigma_{p\bar{p}}(m) = \frac{(dN/dm)}{\mathcal{E}(1 + \delta_{rad})(d\mathcal{L}/dm)}$$

reconstruction efficiency
radiative corrections

ISR differential luminosity

$$\frac{d\mathcal{L}}{dM} = \frac{2m}{s} \frac{\alpha}{\pi \cdot x} \cdot \left((2 - 2x + x^2) \log \frac{1+C}{1-C} - x^2 C \right) L_{ee}$$

Machine integrated luminosity

$\cos\theta_{\min}^*$

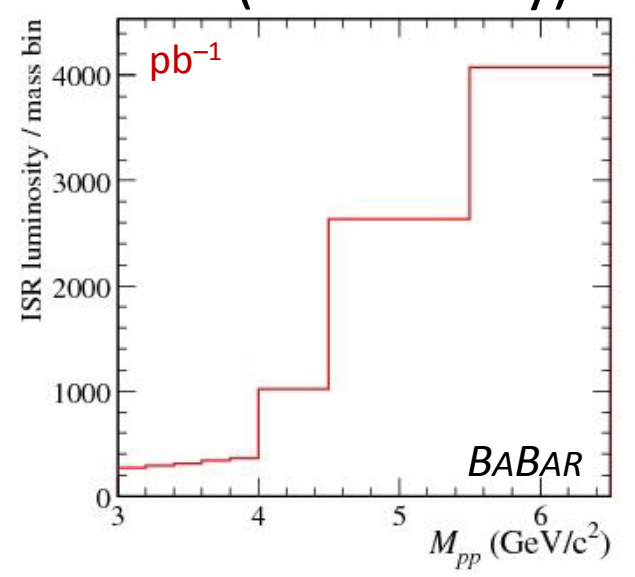
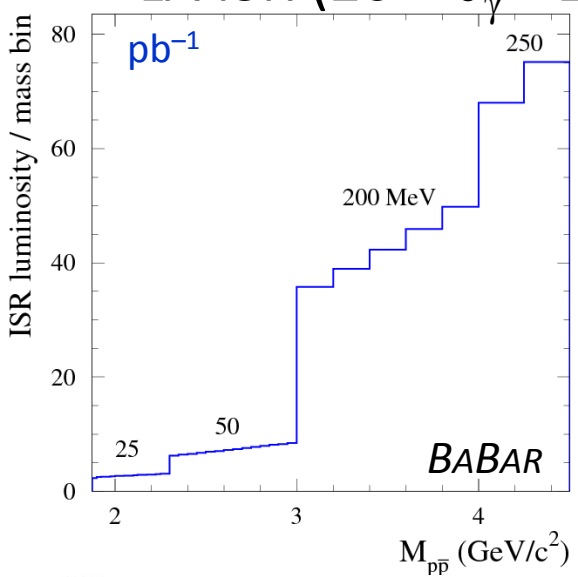
- Obtained from integration of the radiator function over θ_γ^*
- Known at 1% level (from MC simulation)
- $20^\circ < \theta_\gamma^* < 160^\circ \rightarrow$ acceptance for ISR photon $\sim 10\text{-}15\%$ in BABAR



CROSS SECTION

- LA ISR ($20^\circ < \theta_\gamma^* < 160^\circ$)

- SA ISR (Preliminary)



$$\sigma_{ee \rightarrow p\bar{p}}(m) = \frac{dN/dm}{\epsilon R dL/dm}$$

