



Data Driven Triggers for the NOvA Experiment

On behalf of NOvA

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University of Minnesota

APS DPF Meeting – 08/15/2013

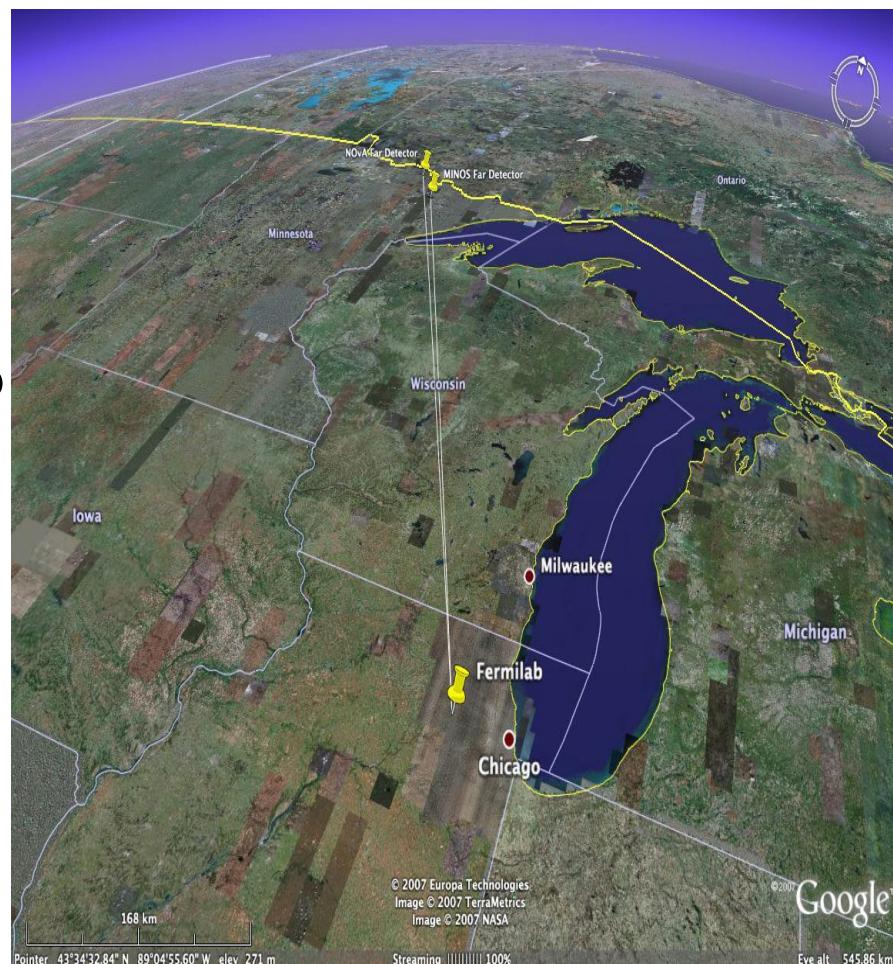


Introduction to NOvA



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- **Neutrino Oscillation Experiment**
 - $P(\nu_\mu \rightarrow \nu_e)$
 - $P(\nu_\mu \rightarrow \nu_\mu)$
 - As well as anti-neutrino channels
- **Physics Goals**
 - Precision measurement of θ_{23} and $|\Delta m^2_{32}|$
 - Measure θ_{13}
 - Resolve mass hierarchy
 - Constrain δ_{CP}

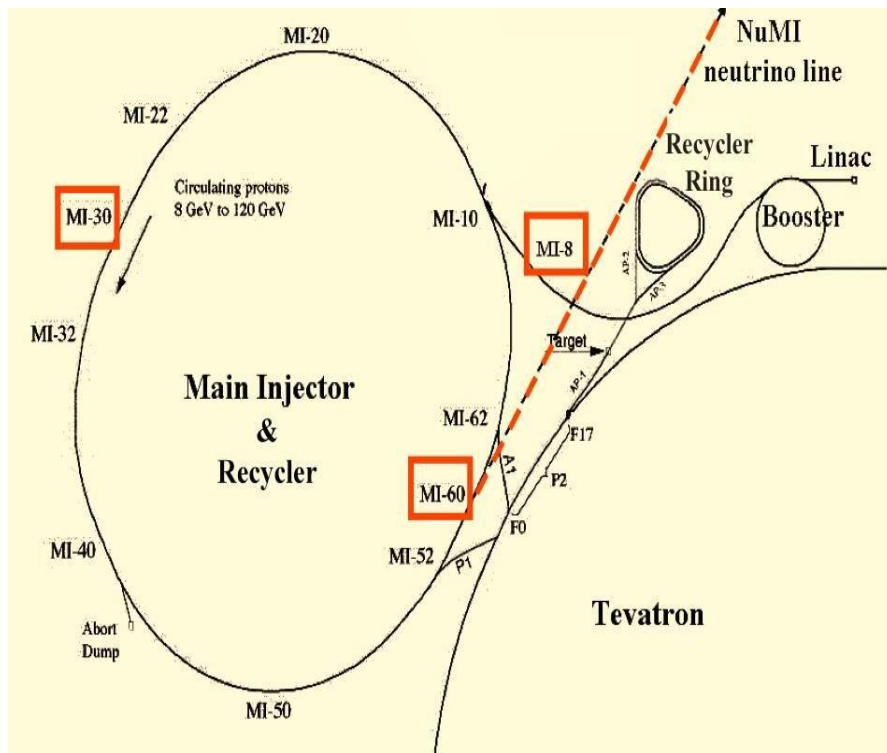
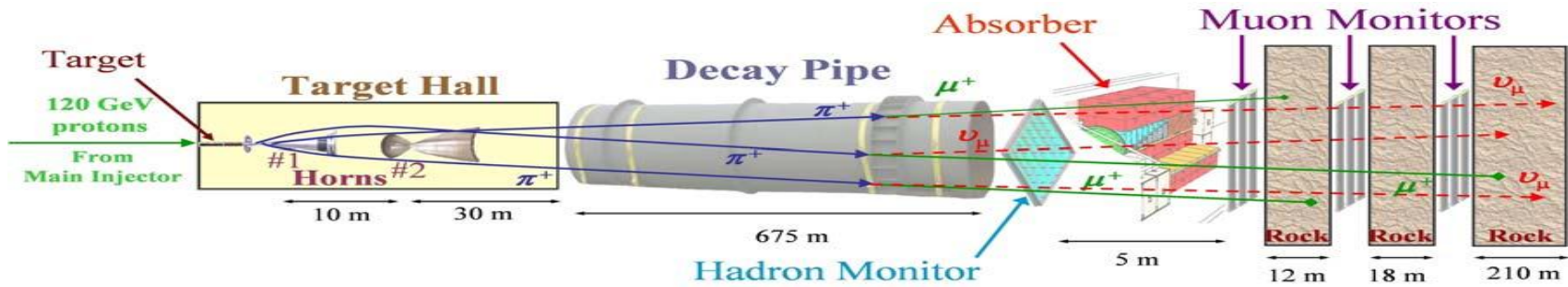




The Accelerator



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• NuMI Beamline

- Graphite target
- Magnetic focusing horns
- 10 μ s beam spill every 1.3 s
- $\pi^+ \rightarrow \mu^+ + \nu_{\mu}$
- 700 kW capable beam



Data Driven Triggers (DDTs)



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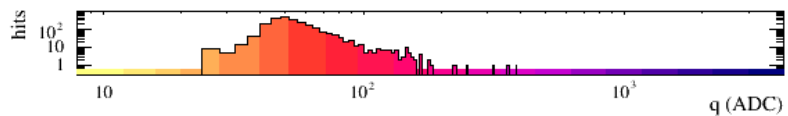
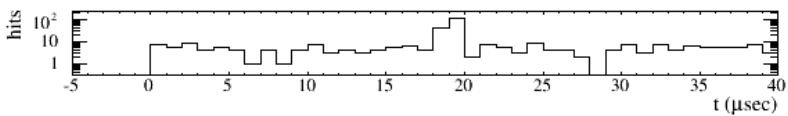
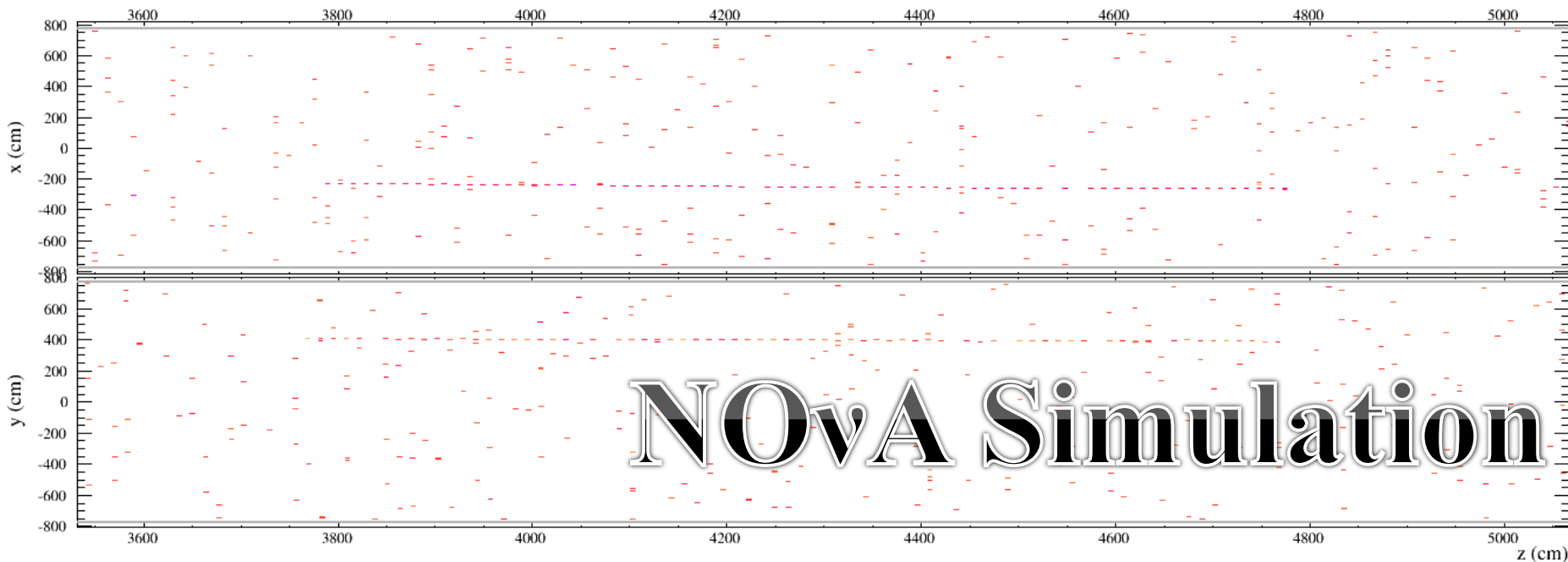
- **Aid beam related analyses**
 - ~600 beam neutrino events at **Far Detector** per year
 - **Every single one is precious**
- **Crucial for non beam related physics**
 - **Magnetic Monopoles**
 - **Supernovae**
- **Enhance calibration and alignment samples**



Rare Topologies



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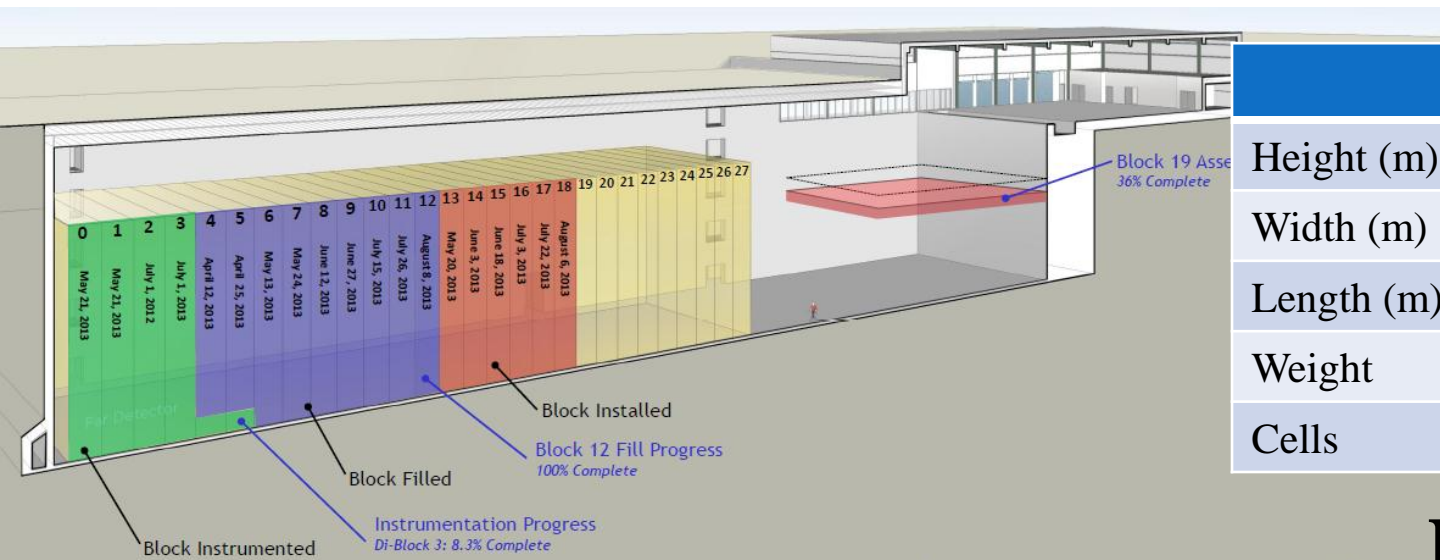


The Detector



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	FarDet	NearDet
Height (m)	15.6	4.2
Width (m)	15.6	4.2
Length (m)	63	14.3
Weight	14 kt	260 tons
Cells	344,064	20,256

14 kilotons = 28 NOvA Blocks
 19 blocks of PVC modules are assembled and installed in place
 13.0 blocks are filled with liquid scintillator
 4.17 blocks are outfitted with electronics

Large detector

– Supernovae

- DDT

- Mostly active
- Staggered planes
- Construction underway

Large surface area

– Magnetic Monopoles

- DDT



Far Detector



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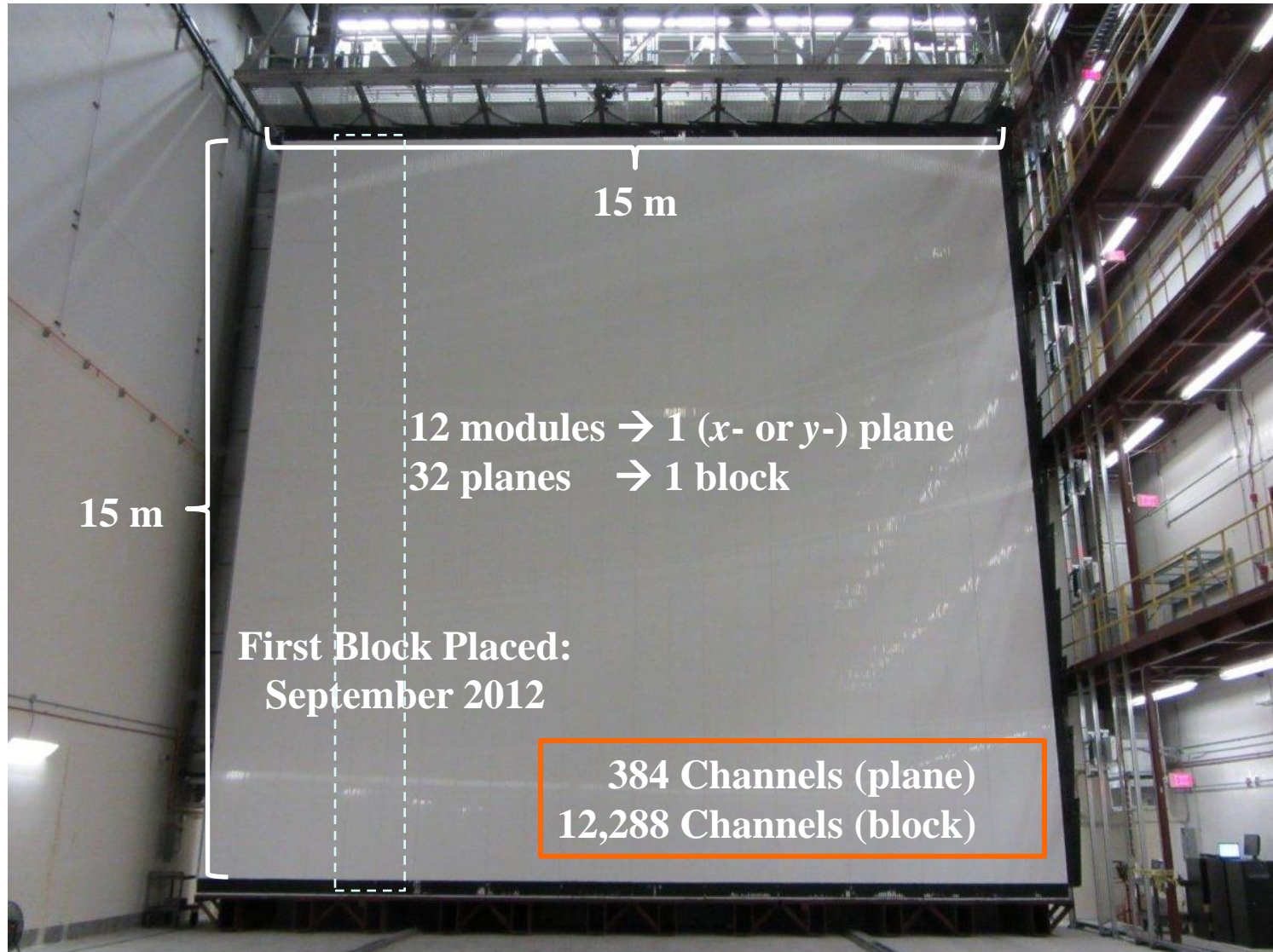




Far Detector Face

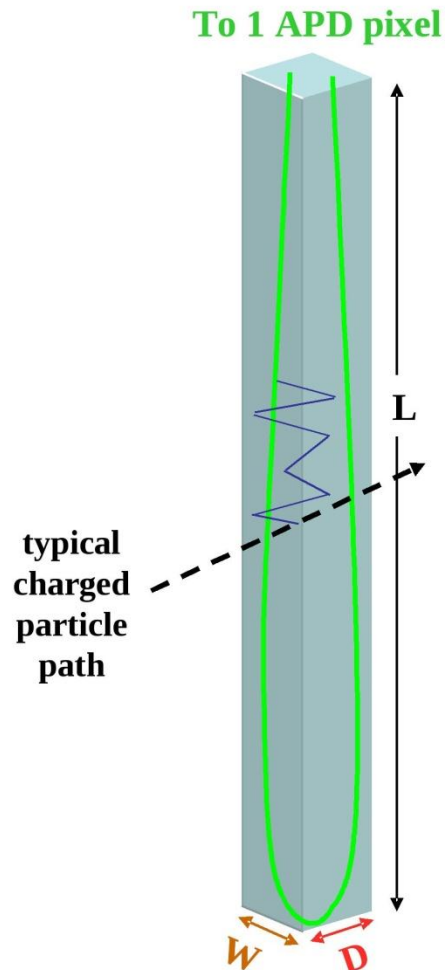


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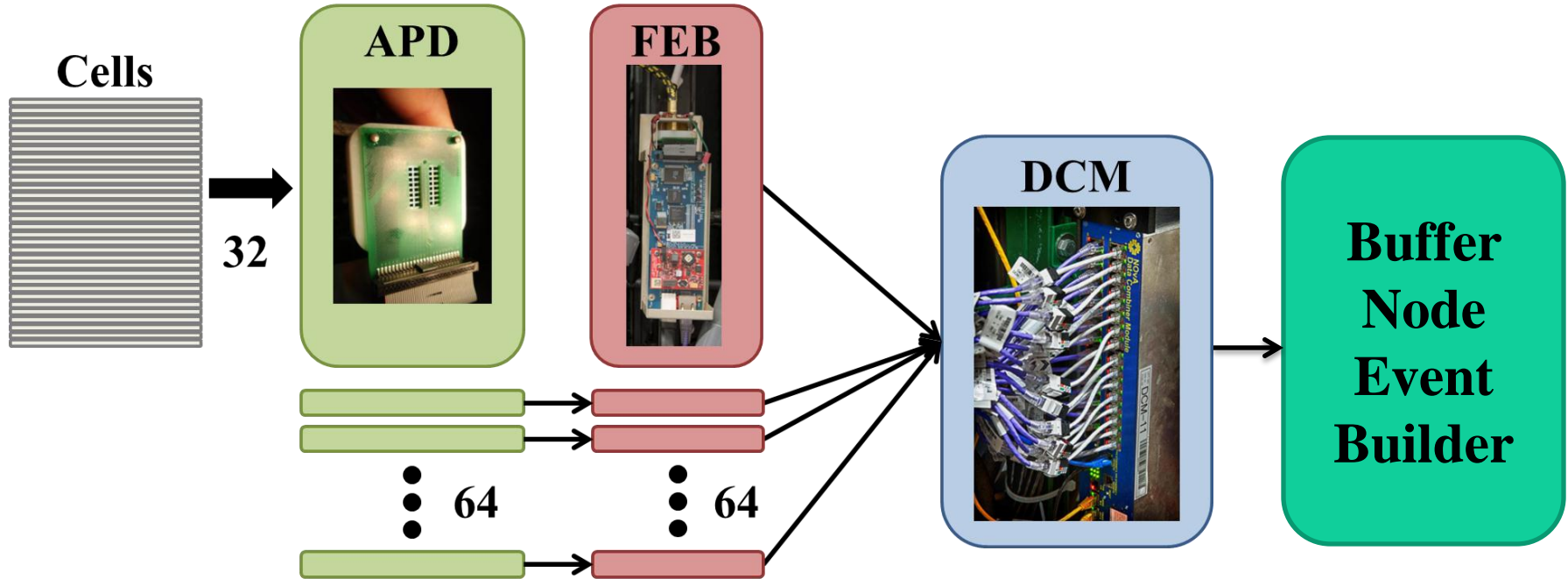
The Cells



- 32 cells per module
- Scintillation light travels to Avalanche Photo Diode (APD) via Wavelength Shifting Fiber
- Data at this stage
 - Pulse height
 - Analog to Digital Conversion (ADC) counts
 - Time stamp
 - Cell ID



The Readout

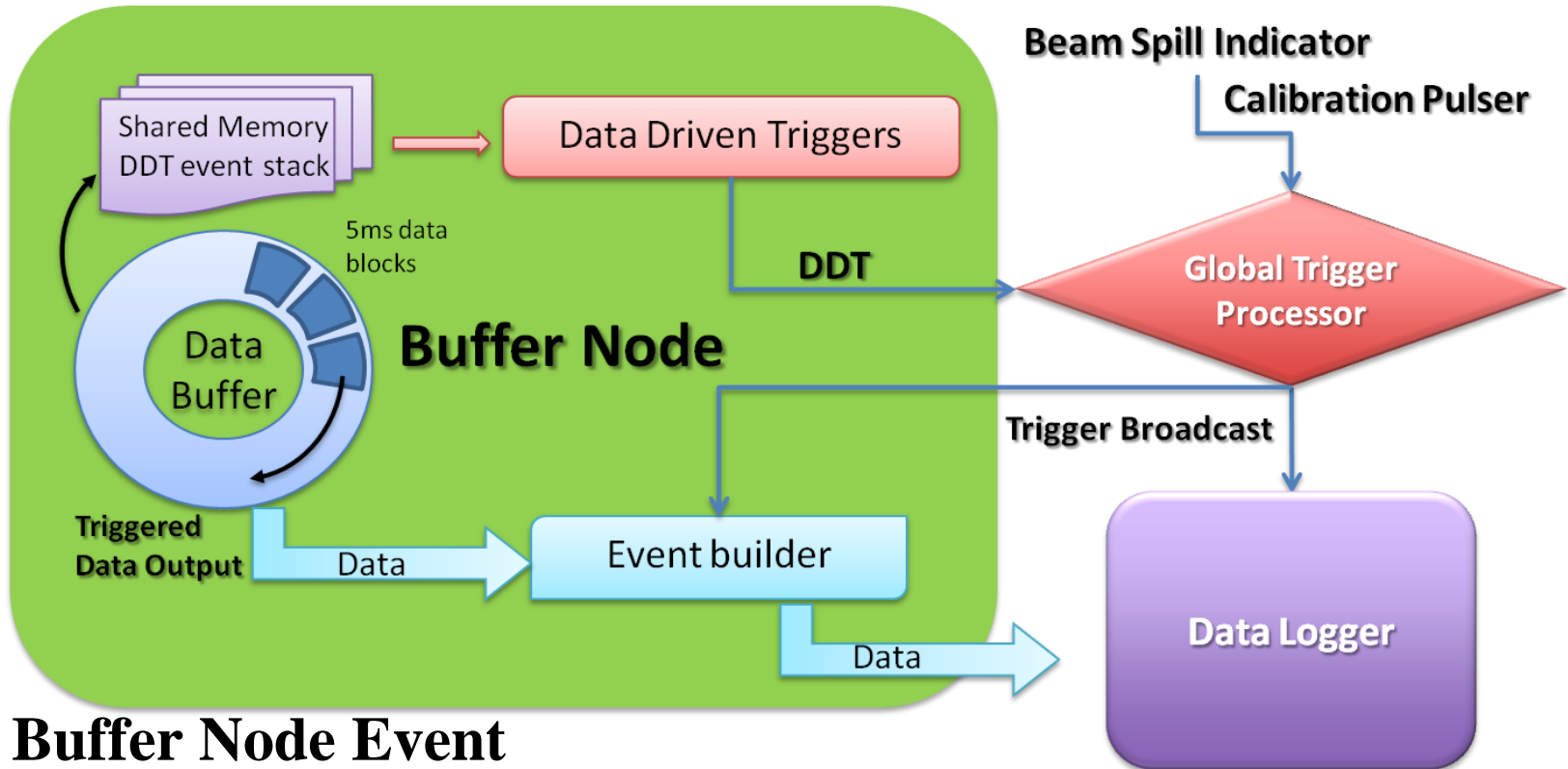


- **APD reports photo electrons**
- **Front End Board decides what's a hit**

- **Data Concentrator Module**
 - **Collects discretized data**
 - **Micro- and Millislices**



Data Buffering



- **Buffer Node Event Builders**

- Assembles millislices into milliblock

- **Snapshot of the detector**

- **Global Trigger**

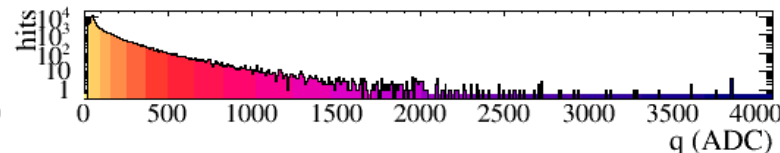
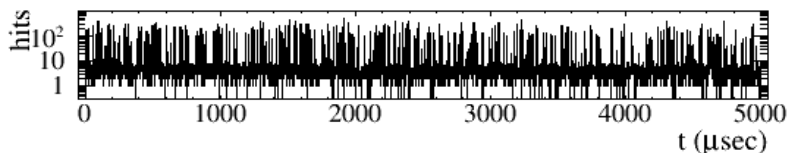
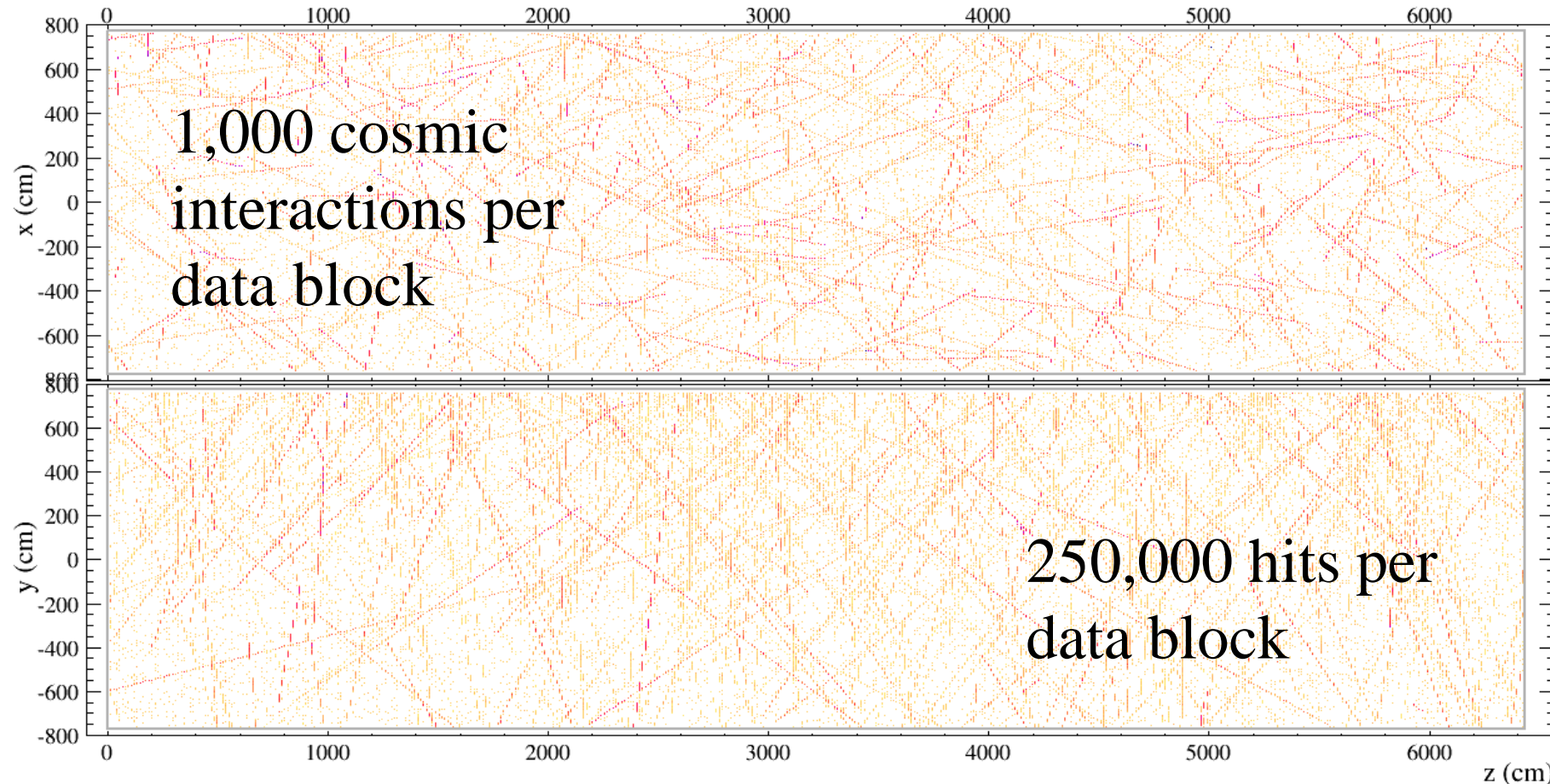
- **Manages trigger messaging**



Far Detector Simulation



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Data Rates

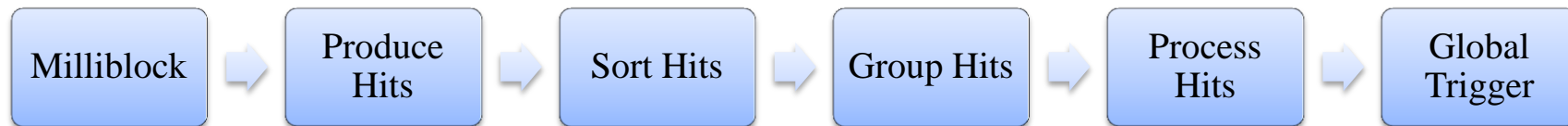
- **No dead time**
 - ~600 MB/s to write all to disc
- **200 kHz Cosmic rate**
- **0.8 Hz Beam neutrino rate**
 - 600 interactions per year
 - 20 μ Hz
- **DDT sieve needed**
- **Can save**
 - ~10% of live time to disc
 - ~2 PB/yr
 - Bandwidth limited
- **Planned upgrades to increase bandwidth**



DDT in Action



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- **Produce list of hits**
- **Sort the list**
 - Time
 - Detector region
- **Group hits within list**
- **Process grouped hits**
 - Tracking
 - Clustering
- **Trigger**
 - Send message to Global Trigger
 - Forwarded to Buffer Nodes



Magnetic Monopoles

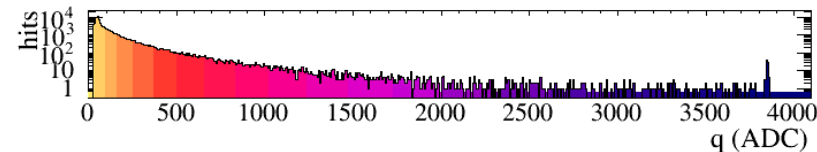
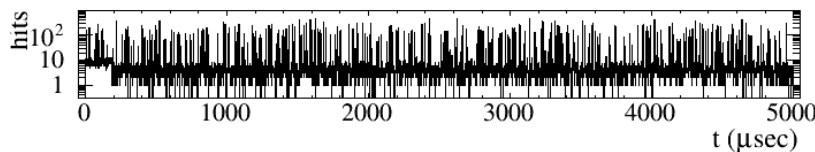
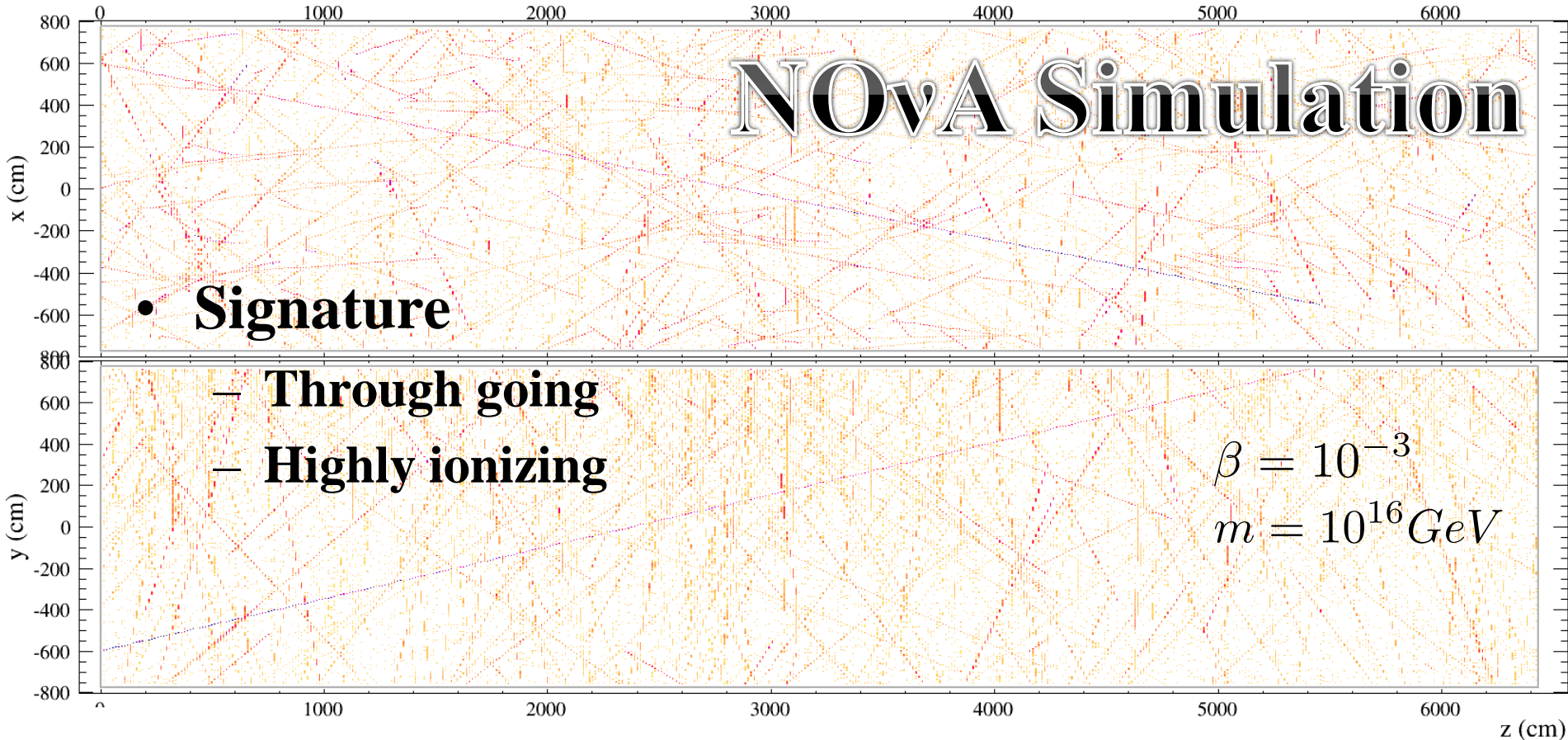


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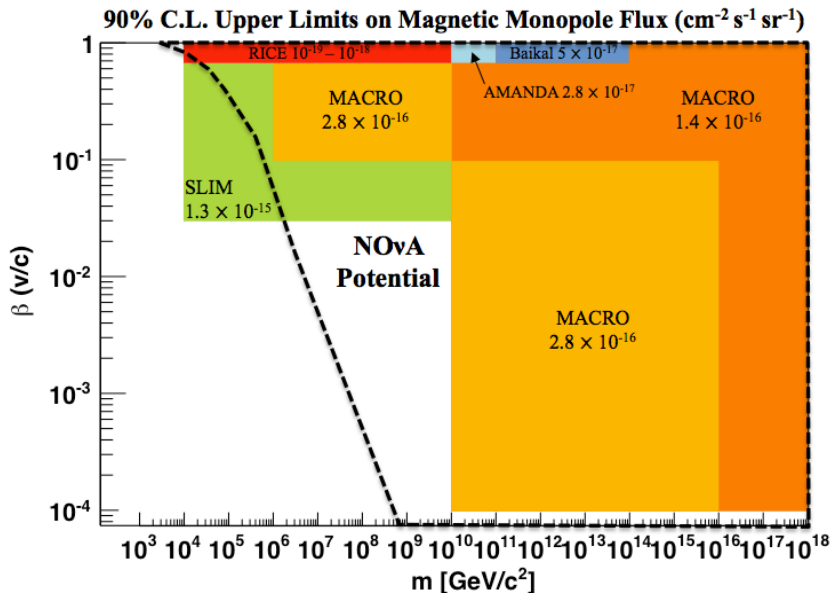
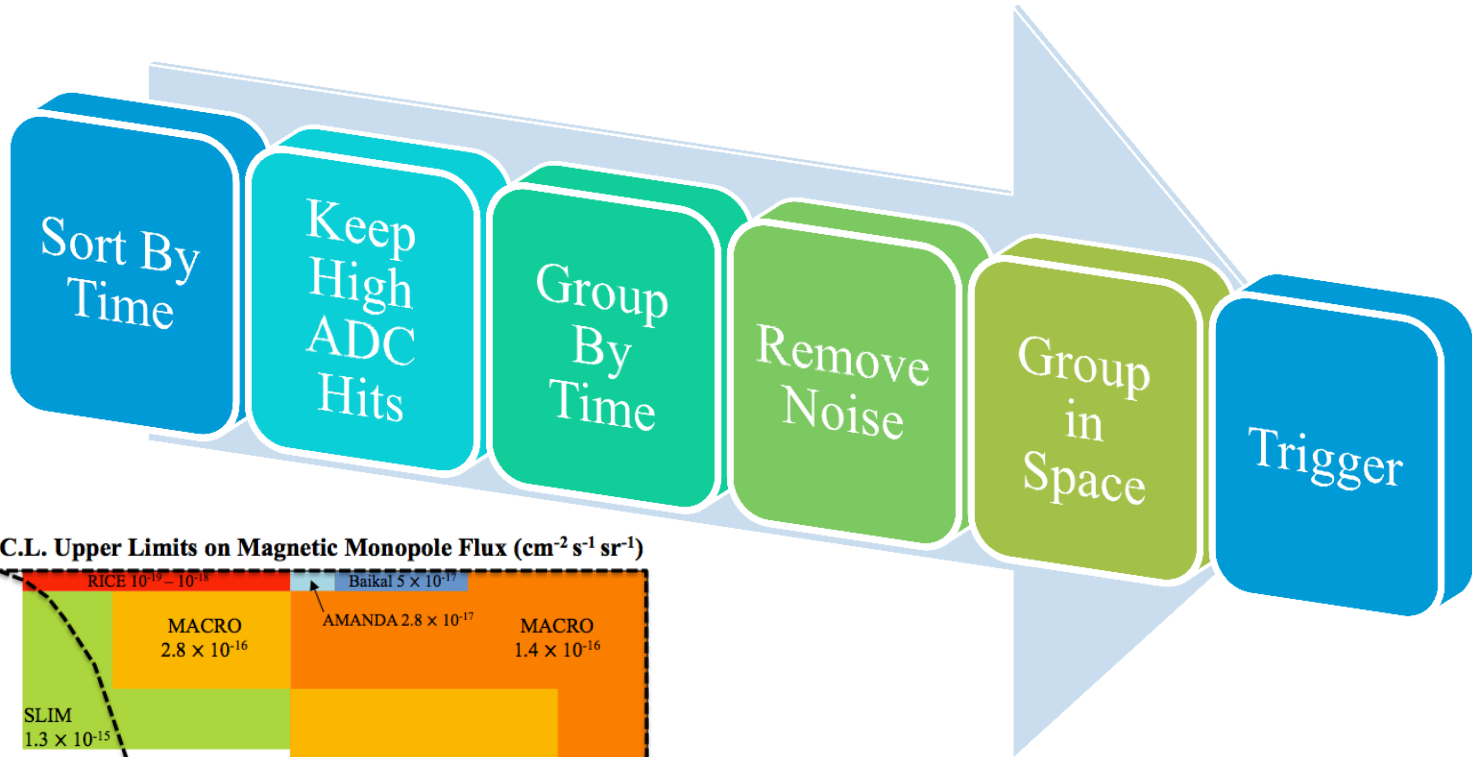


NOvA Simulation





Fast Monopoles in NOvA



Trigger on

- Min/Max position of grouped hits
- ADC Threshold



Supernova Trigger Challenges



- **Energy Spectrum**
 - ~5k neutrinos > 10 MeV
- **Provide time profile**
- **Framework**
 - 5 ms windows aren't long enough
 - Global noise level
- **Under active development**

Class	Channel	GKVM	Livermore
Inverse Beta Decay	IBD	3185	5355
ν_e CC on Carbon	ν_e (^{12}C)	134	50
	$\bar{\nu}_e$ (^{12}C)	131	195
Elastic Scattering	ν_e	72	125
	$\bar{\nu}_e$	31	55
	ν_μ	12	22
	$\bar{\nu}_\mu$	10	18
	ν_τ	12	22
	$\bar{\nu}_\tau$	10	18
	Total		3464

Class	Channel	GKVM	Livermore
Neutral Current on Carbon	ν_e (^{12}C)	61	30
	$\bar{\nu}_e$ (^{12}C)	64	96
	ν_μ (^{12}C)	61	223
	$\bar{\nu}_\mu$ (^{12}C)	61	223
	ν_τ (^{12}C)	61	223
	$\bar{\nu}_\tau$ (^{12}C)	61	223
Total		371	1020



Conclusions



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- **Successes**

- **Data framework developed**
- **Trigger messages sent and received**
- **Fast magnetic monopole trigger implemented**

- **Future Work**

- **Need studies of trigger rate and efficiency**
- **Need monitoring tools**
- **Implementing more triggers**
 - **Contained muons**
 - **Supernova trigger**
 - **Upward going tracks**



Thanks!

The DDT Group:

Gavin Davies, Martin Frank, Craig Group, Alec Habig, William Henderson, Adam Moren, Andrew Norman, Yuri Oksuzian, Brian Rebel, Louise Suter, Matthew Tamsett, Zukai Wang

Special Thanks to:

**Martin Frank, Andrew Norman, Zukai Wang, and
The ARTists**



Data Concentration



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- **Front End Board**
 - **Digitizes signal from one APD**
 - **2 MHz sampling**
 - **Assigns timing information**
 - **32 bit word gives unique stamp within 67 second window with 64 MHz clock**
 - **Produces one nanoslice per hit**
- **Data Concentrator Module**
 - **Reads out 64 FEBs**
 - **Assigns another 24 bits of timing information**
 - **Rolls over every ~36 years**
 - **Produces 50 μ s (microslice) activity window of detector region from individual nanoslices**
 - **Collects microslices into 5ms microslice for transfer**



Global Trigger



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- **Process**

- **Listening for trigger messages**
 - Calibration Pulsar
 - Beam Spill Indicator
 - DDT
- **Broadcast to Buffer Farm to send appropriate microslices to Data Logger**
- **Data Logger appends data to correct trigger stream**

- **DDS Messaging**

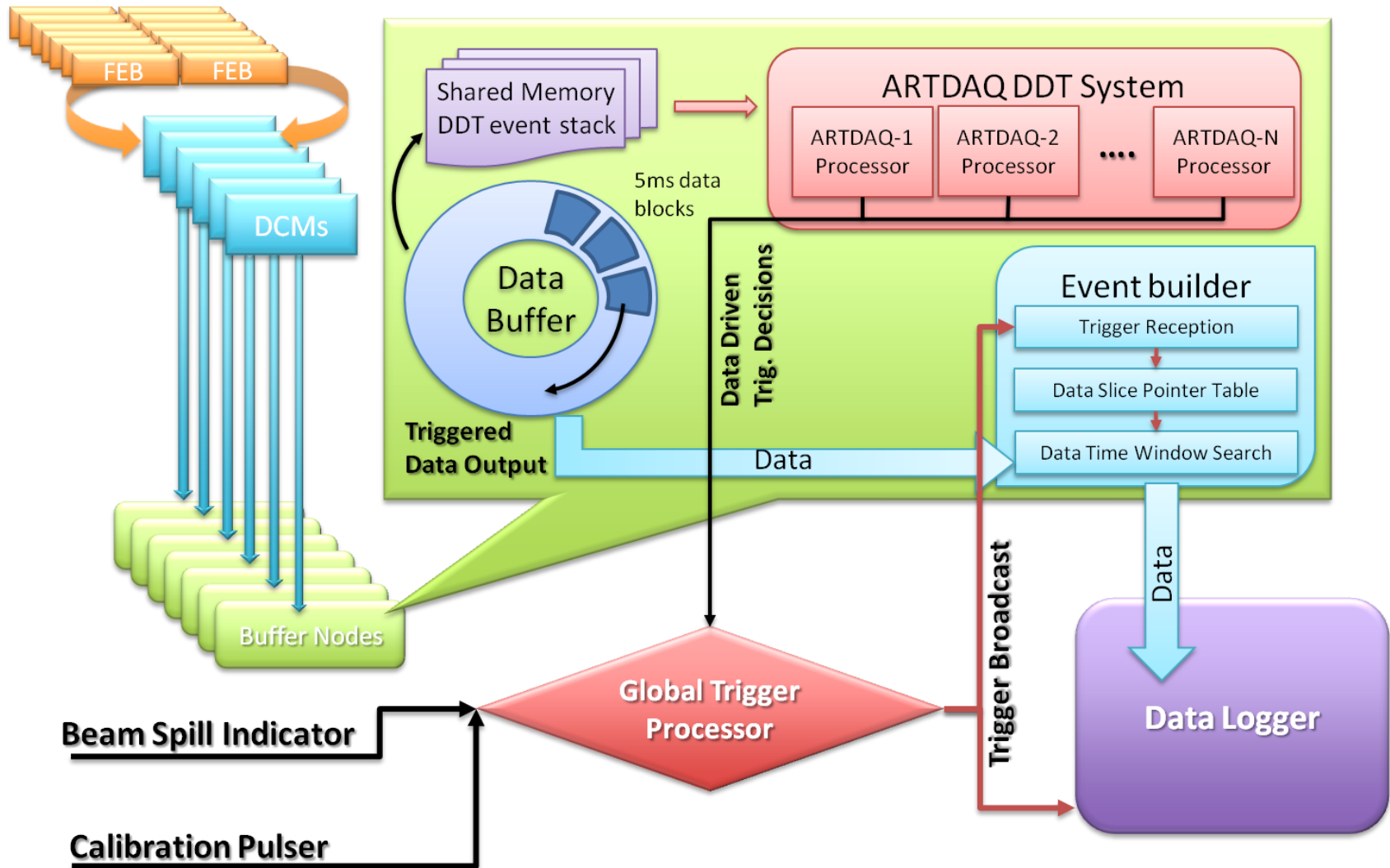
- **Data Distribution Service**
- **Publishers and Subscribers**
 - Mailboxes
- **How NOvA's DAQ functions**



DAQ Summary



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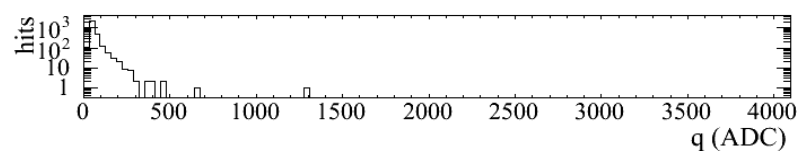
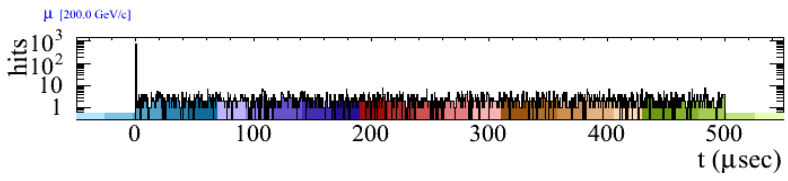
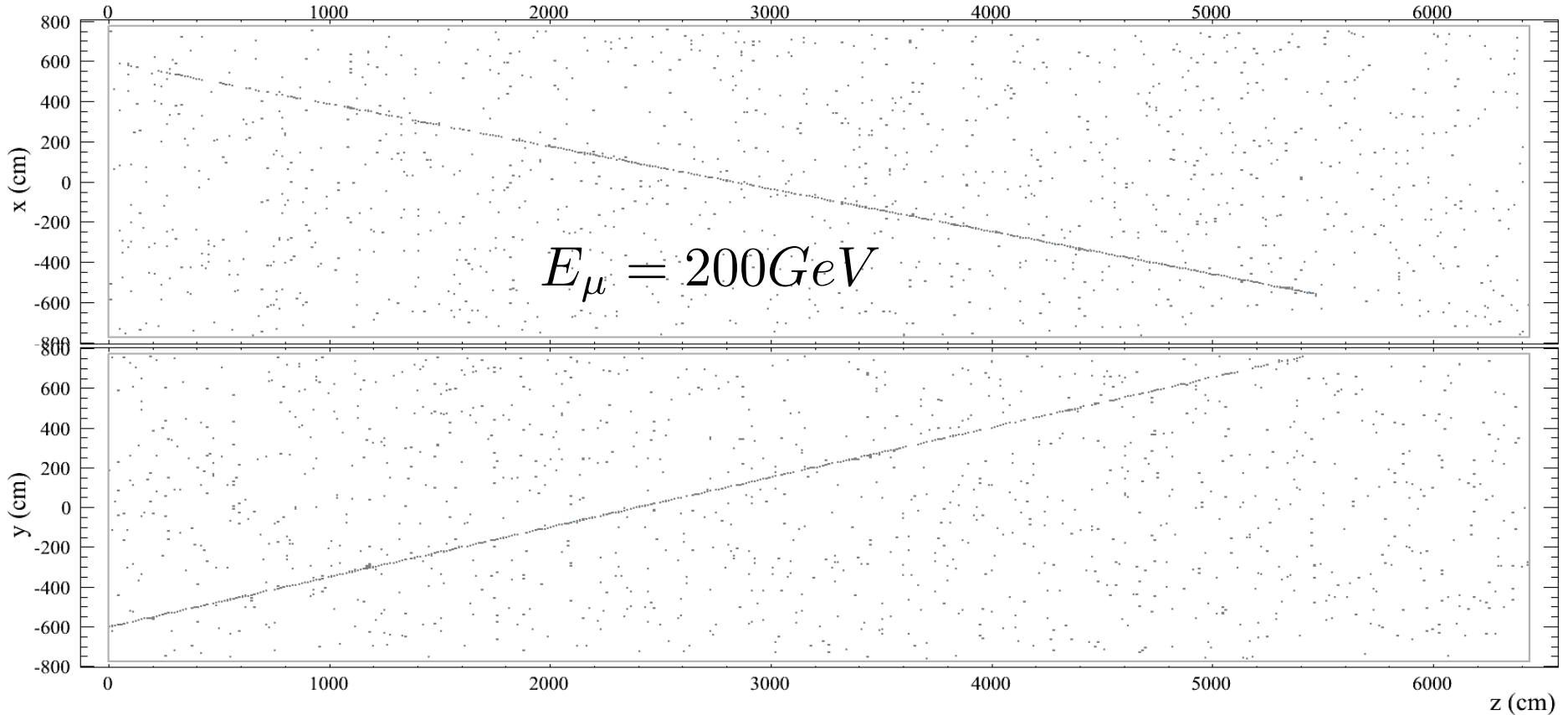


High Energy Muon



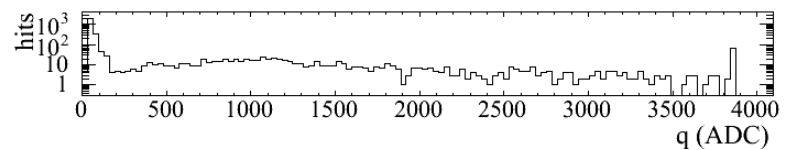
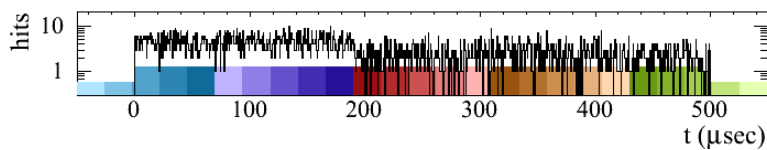
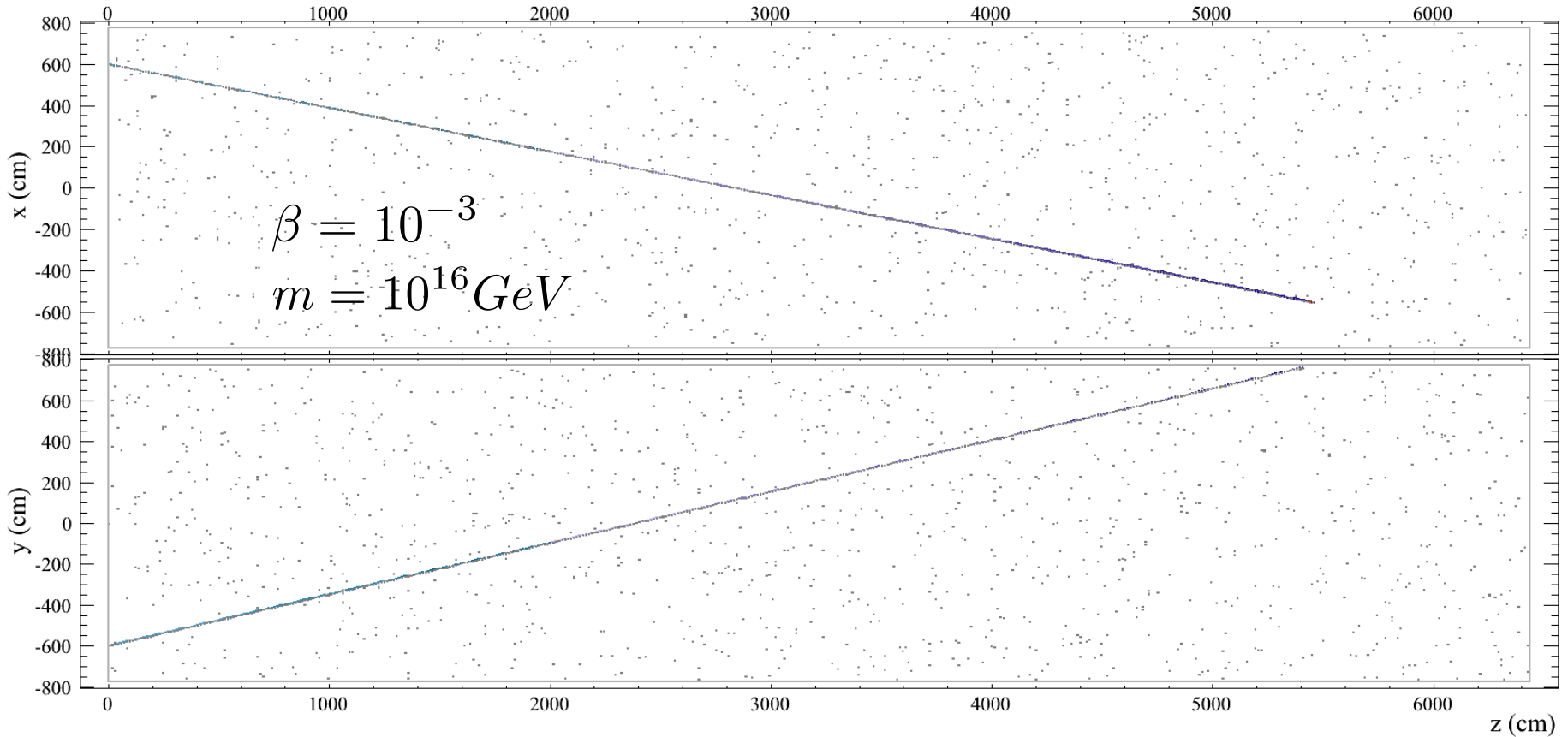
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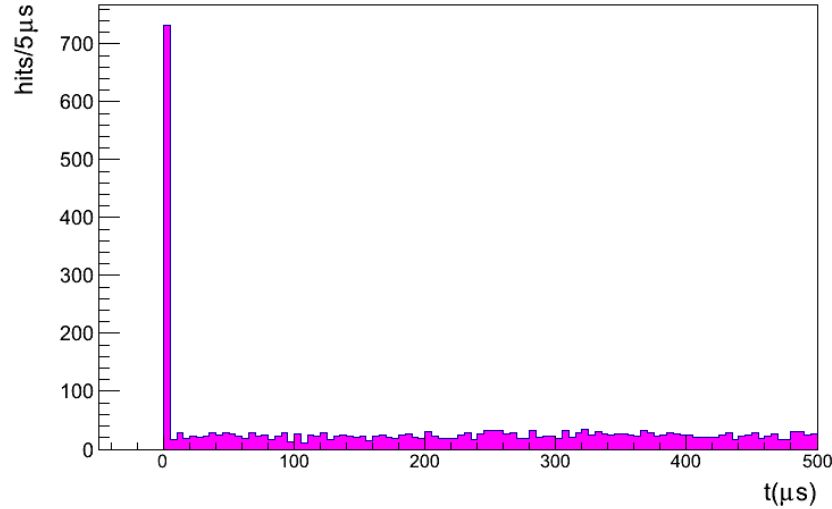
Magnetic Monopole



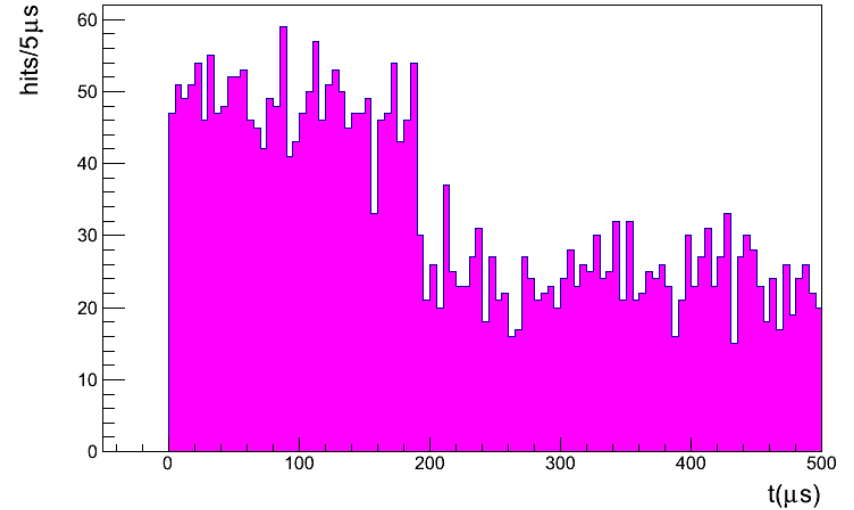


High Energy Muon vs Monopole

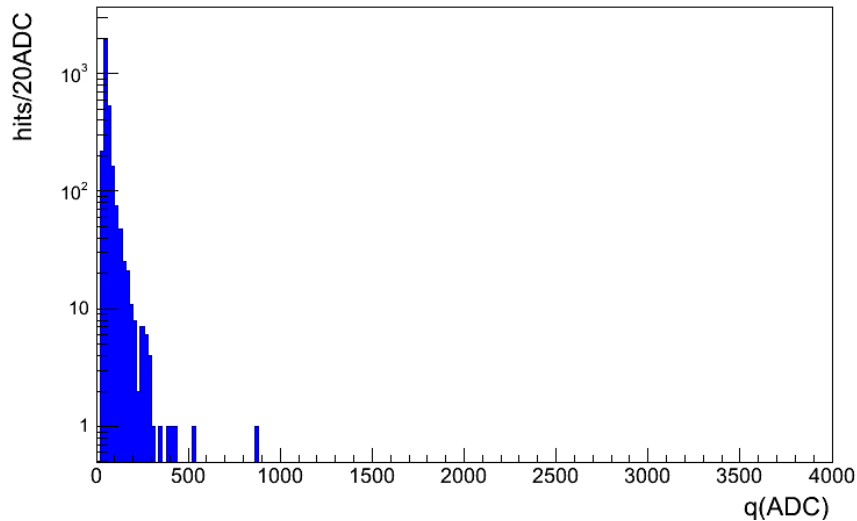
hit register time



hit register time



amplitude of signal from cell



amplitude of signal from cell

