Neutrino Physics

Deborah Harris Fermilab

August 17, 2013 DPF Meeting Santa Cruz, California

Preamble

- 40 minutes is a short time to cover...
 - 32 talks
 - (at least) 17 different detectors
- You heard two excellent talks in Monday's plenary session
 - Patrick Huber on What Neutrinos have Told Us
 - Bonnie Fleming on Possibilities for Neutrino Physics
- I have taken liberties with this talk, and not all parallel session talks will be covered at the same level
 - Priority given to talks with data, and topics not already covered by Patrick and Bonnie

Top ten reasons to study neutrinos

- 1. Neutrinos are among the most abundant particles in the universe
- 2. Neutrinos have been around since the universe was 1 second old
- 3. Neutrinos are signals from the highest energy accelerators in the universe
- 4. Neutrinos will teach us about how mass is generated
- 5. What other fermion could be its own antiparticle?
- 6. Neutrinos can see inside the nucleus like none else
- 7. Neutrinos will tell us if we really understand flavor
- 8. Neutrinos may be the reason that we enjoy such a healthy baryon asymmetry
- 9. Neutrinos can access physics up to the GUT energy scales
- 10. Neutrinos broke the standard model and will tell us what is beyond

Energy Scales in Neutrino Physics



from D.Williams

- Reason number 11: Neutrinos can be directly studied from keV (10³) to EeV (10¹⁸) energies
- Heard talks here from both ends of this spectrum, and much from the middle

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DPF 2013 Neutrino Parallel Sessions, I (Red means Data!)

Neutrinos are signals from the highest energy accelerators in the universe

- WHITEHORN, Nathan on high energy neutrinos in ICECUBE

- Neutrinos may teach us about how mass is generated
 - KUNDE, Gerd J on direct mass measurement with electron capture spectroscopy
 - OBLATH, Noah on Project 8
- What other fermion could be its own antiparticle?
 - CHAVES, Jason on EXO 200 Status and results
 - KRAVITZ, Scott on Barium Tagging in EXO-200
 - GIOVANETTI, Graham on the MAJORANA Demonstrator

DPF 2013 Neutrino Parallel Sessions, II

- Neutrinos can see inside a nucleus like none else
 - RAKOTONDRAVOHITRA, Laza on Quasi-elastic scattering at MINERvA
 - HANSEN, Damon on CCQE scattering in T2K P0D
 - ADAM, Jeanine on CC electron neutrino scattering in T2K P0D
 - BODEK, Arie on Transverse Enhancement Models and Meson Exchange Currents
 - Dr. SZELC, Andrzej on Argoneut Update

DPF 2013 Neutrino Parallel Sessions, III

- Neutrinos will tell us if we really understand flavor
 - WORCESTER, Elizabeth on Daya Bay results
 - CARR, Rachel on Double CHOOZ Results on Gd Capture
 - HIGNIGHT, Joshua on T2K Results on $\nu_{\rm e}$ appearance
 - IMBER, James on T2K /SK Joint fits
 - RADOVIC, Alexander on MINOS $\nu_{\mu} \, / \, \, \nu_{e}$ joint fits
 - WILLIAMS, Dawn on ICECUBE's Oscillation Results and PINGU
 - FRIEND, Megan on future T2K sensitivity
 - BIAN, Jianming on NOvA status and $\nu_{\rm e}\,$ sensitivities
 - SACHDEV, Kanika on predicting NOvA backgrounds
 - BAIRD, Michael on NOvA sensitivities on ν_{μ} disappearance

DPF 2013 Neutrino Parallel Sessions, IV

- Neutrinos may be the reason that we enjoy such a healthy baryon asymmetry
 - BASS, Matthew on LBNE Physics Sensitivities
 - Dr. SZELC, Andrzej on Lariat
 - GRANT, Christopher on CAPTAIN
 - GUARDINCERRI, Elena on LBNE Near Detector
 - ANGHEL, Ioana on Fast Photosensor development for Water Cerenkov neutrino detectors
 - SHAEVITZ, Mike on Daedalus Experiment
- Neutrinos broke the standard model and will tell us what is beyond
 - ZHURIDOV, Dmitry on neutrino magnetic moments from NSI
 - AHMED, Rashed on tau neutrinos as probe of NSI
 - SPITZ, Joshua on testing Lorentz Invariance with Double CHOOZ
 - AURISANO, Adam on MINOS sterile neutrino results
 - COLLIN, Gabriel on MicroBooNE
 - SHAEVITZ, Mike on IsoDAR Experiment

Seeing the highest energy accelerators

- ICECUBE:
 - 1km³ volume
 - 86 strings, 5160 PMTs
 - See atmospheric ν_{e} flux above 80GeV for 1st time
 - See evidence for astrophysical neutrinos: above prediction from atmospheric sources
 - No evidence (yet) for clustering







Deborah Harris, Fermilab: Neutrino Physics Summary

Neutrinos and Mass Generation

- Neutrino masses are so much lower than other fermion masses, they could have been generated in a different way
- Want to learn the absolute mass spectrum for the neutrinos, that will guide us to how the mass was generated
- Current best limit from KATRIN looking at Tritium beta decay in magnetic spectrometer (2.2eV)
- We heard at DPF about two other techniques
 - Tritium decay at Project 8
 - Electron Capture on Holmium



Measuring Mass with Tritium in new way

- Project 8 Concept: (Oblath)
 - trap a volume of tritium (atomic)
 - watch decay electrons process around constant magnetic field
 - measure the cyclotron frequency to get e- Energy
 - Highest Energy electrons give lowest frequency signals
- Working to prove the principle:
 - Start with ^{83m}Kr volume
 - Look for 18 and 30keV conversion electrons
 - Added bonus: see e-lose energy as it processes in field
 - 7TB of data taken in Jan. 2012, must improve data-taking technique



$$\frac{w_0}{\gamma} = \frac{eB}{K + m_e}$$



 $\omega(\gamma) = -\frac{\omega}{2}$



Mass Measurement using Electron Capture Spectroscopy

- Study electron capture in ¹⁶³Ho (2.3KeV-2.8KeV electron capture energy) (Kunde)
- See small changes in energy using transition edge sensor



- See huge dynamic range and get several lines of these decays
- Energy resolution achieved: 9.0±0.2 eV!
- Expect improvements with ¹⁶³Ho from atomic physics
- Will use ¹⁶³Ho later this year



Counts are per 1.00 eV bin

6420 6440 6460 6480 6500 6520 6540 6560 6580 6600

Can a Fermion be its own antiparticle? $0\nu\beta\beta$

Neutrino-less Double Beta Decay

- Also a sensitive probe of absolute v mass
- EXO-200: 200kg of liquid Xenon in WIPP
- MAJORANA Demonstrator: 30kg of ⁷⁸Ge at SURF
- EXO: Study 2v mode as calibration (and background)
 - Presented here: best measurement of the 2ν mode half-life
- EXO-200: Trying to improve sensitivity by x1 with barium tagging! (Daughter of decay)
- Expected EXO sensitivity: ~0.2eV
- Expected Majorana Sensitivity: 0.1-0.2eV

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\begin{array}{r} T_{1/2}^{2\nu\beta\beta} = \\ (2.172 \pm 0.017 \pm 0.060) \cdot 10^{21} \, \text{yr} \\ (\text{stat.}) \quad (\text{sys.}) \end{array}
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EXO: Jason Chaves, Kravitz, Majorana: Giovanetti



Energy (ke)

Using Neutrinos to see inside nucleus

- Why is this in a DPF session? Isn't it nuclear physics? Two answers
 - Yes, nuclear physicists are interested in using neutrinos as probe of the nucleus and that's why they've joined neutrino experiments
 - Yes, but we need to understand it in order to measure oscillation probabilities
 - Signal is affected: visible energy in detector must be used to reconstruct neutrino energy, but this could be affected by nuclear environment
 - Background is also affected: bare nucleon models can't predict the whole story here either, and Near Detectors can't tell you everything

"Simplest" v interaction on nuclei: Quasi-elastic

- Important because they are dominant channel for T2K and significant fraction of NOvA events
- Clean identification of outgoing lepton possible
- "Theoretically" clean kinematic reconstruction
 - But have to assume something about initial state of proton/neutron inside the nucleus to get Energy and Momentum transferred (Q²) $\nu + n \rightarrow \mu^{-} +$



$$\nu_{\mu} + n \to \mu^{-} + p$$
$$\bar{\nu}_{\mu} + p \to \mu^{+} + n$$

$$E_{\nu}^{QE} = \frac{2(M_n - E_B) E_{\ell} - \left[(M_n - E_B)^2 + m_{\ell}^2 - M_p^2 \right]}{2[M_n - E_B - E_{\ell} + p_{\ell} \cos(\theta_{\ell})]}$$
$$Q_{QE}^2 = -m_{\ell}^2 + 2E_{\nu}^{QE} \left(E_{\ell} - \sqrt{E_{\ell}^2 - m_{\ell}^2} \cos(\theta_{\ell}) \right)$$

Seeing inside the nucleus at Argoneut

• "A picture is worth a thousand words"



- Argoneut: Liquid Argon TPC that was in the NuMI beamline (~3GeV)
- New results on antineutrinos (Szelc)
- This data will constrain standard neutrino event generators because final state is so clear
- Challenge will be to correctly simulate acceptance for these extra protons
- Two test beam detector plans described at DPF: LARIAT (Szelc) and CAPTAIN (Grant)



Seeing inside the nucleus at MINERvA

- MINERvA is a scintillator-based detector in the NuMI beamline (~3GeV) designed to look at interactions on plastic as well as a range of nuclear targets
- MINERvA has measured Q² distributions for v and anti-v, and also looks at the energy near the interaction vertex
- Sees evidence for np correlations in the nucleus: would give pp final state in v scattering, nn final state in vbar scattering (*Rakotondravohitra*)



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Seeing inside the nucleus with Electron Scattering

- Nuclear physicists have made much progress on understanding the nucleus with charged electron scattering
- They also see differences in cross sections from what you would see if there were no correlations between nucleons in the nucleus (Bodek)
- Applying this to neutrino Quasi-elastic measurements agrees better with data than current "out of the box" models







Seeing inside the nucleus at T2K

- T2K is a long baseline experiment using Super-Kamiokande and a beam with a peak energy of ~700MeV
- T2K has a near detector to measure cross sections specifically on water
- First glimpse at v_e events at high energies (Adam)
- Also trying to make CCQE measurement on water (Hansen)







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Do we really understand flavor?

Everything I needed to know about flavor I learned from the quarks...
 Lesson Learned from CKM: 3 mixing angles and a phase Call them θ₁₂, θ₂₃, θ₁₃, δ if s_{ij} = sin θ_{ij}, c_{ij} = cos θ_{ij}, then

Additional Complication: Matter Effects

 The oscillation probability changes differently for electron neutrinos vs antineutrinos when they propagate through matter in a straightforward way



- Can't treat neutrinos propagating through earth simply as mass eigenstates, have to take into account electron flavor
- This would give an apparent CP violation just because the earth is not CP-symmetric

Additional Bonus: Matter Effects

- Bonus: the way the oscillation probabilities change depend on whether the "electron-like" neutrino mass eigenstates are on the top or the bottom of the spectrum.
- We know from numerous experiments that there are two mass differences: large and small, but we don't know how they are ordered

• One of these looks much more like the quark masses, all the more reason to measure it!

v Oscillation Probabilities

- v_{μ} Disappearance: 1-sin²2 Θ_{23} sin²(Δm_{32}^2 L/4E)
- v_e Disappearance:

 $P_{\bar{\nu}_{e} \to \bar{\nu}_{e}} \approx 1 - \sin^{2} 2\theta_{13} \sin^{2} \left(\Delta m_{31}^{2} L / 4E \right) - \cos^{4} \theta_{13} \sin^{2} 2\theta_{12} \sin^{2} \left(\Delta m_{21}^{2} L / 4E \right)$

- v_e appearance in a v_μ beam: even more complicated...

•
$$P(\nu_{\mu} \rightarrow \nu_{e}) = P_{1} + P_{2} + P_{3} + P_{4}$$

$$P_{1} = \sin^{2} \theta_{23} \sin^{2} 2\theta_{13} \left(\frac{\Delta_{13}}{B_{\pm}}\right)^{2} \sin^{2} \frac{B_{\pm}L}{2}$$

$$P_{2} = \cos^{2} \theta_{23} \sin^{2} 2\theta_{12} \left(\frac{\Delta_{12}}{A}\right)^{2} \sin^{2} \frac{AL}{2}$$

$$P_{3} = J \cos \delta \left(\frac{\Delta_{12}}{A}\right) \left(\frac{\Delta_{13}}{B_{\pm}}\right) \cos \frac{\Delta_{13}L}{2} \sin \frac{AL}{2} \sin \frac{B_{\pm}L}{2}$$

$$P_{4} = \mp J \sin \delta \left(\frac{\Delta_{12}}{A}\right) \left(\frac{\Delta_{13}}{B_{\pm}}\right) \sin \frac{\Delta_{13}L}{2} \sin \frac{AL}{2} \sin \frac{B_{\pm}L}{2}$$



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Stages of ν Flavor Physics

- Stage 1: Denial (1960's through 1998)
- Stage 2: Are all the mixing angles non-zero? (1998-2012)
- Stage 3: Can a 3-generation mixing matrix explain every example we have of flavor change? (2012 and beyond)
- Stage 4: Are neutrino masses arranged like the quark masses?
- Stage 5: Do neutrinos actually violate CP?

Comparison with Quark Sector

- CKM Matrix Elements were understood to be part of a unitary matrix, which implied the following triangle would close itself
- Measurements needed in kaon and B meson sectors to confirm or refute this





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Oscillation Results at DPF

- Daya Bay
 - Reactor experiment in China (3MeV vbars), near and far liquid scintillator detectors
- Double Chooz
 - Reactor experiment in France (3MeV vbars), currently only far liquid scintillator detector
- MINOS
 - 3GeV v_{μ} accelerator experiment in US, 735km baseline, magnetized steel scintillator detectors near and far
- ICE CUBE
 - Ice Cerenkov detector at South Pole looking at atmospheric vs
- T2K
 - 700MeV v_{μ} accelerator experiment in Japan, 295km, Water Cerenkov far detector, water/scintillator/TPC near detectors











(Other 5159 not shown)

Stage 2: Evidence last mixing angle (Θ_{13}) is >0

- Broad range of energies:
 - Reactors (Daya Bay, Double Chooz) at few MeV
 - Double Chooz reactor off, Gd capture: $sin2^2\Theta_{13} = 0.091\pm0.035$ (Carr)
 - Daya Bay has near and far detectors (Worcester)
 - Accelerators (T2K, MINOS) at 1-3 GeV
 - T2K: see 28 events, predict 4.6 background events, new improved π^0 background rejection technique (*Hignight*)
 - MINOS: see 152 events, 129 background in v mode, 20 events on 18.5 background events in vbar mode (Radovic)
 - Atmospheric Neutrinos (PINGU) to look, but at 10-100GeV, using mass hierarchy (Williams)



Stage 3: Are 3 angles and a phase enough?

- Seeing Largest Mass Splitting across many energies
 - MINOS, T2K, Deep Core (ICECUBE)
- Electron Neutrino Appearance:
 - T2K, MINOS
 - Immediate future
 - NOvA: 2GeV accelerator ν_μ's in US, 810km baseline, liquid scintillator detectors near and far, starting to take data in 2013 (*Bian, Baird, Sachdev*)
 - MicroBooNE: 1GeV v_{μ} accelerator experiment in US, 1km baseline, Liquid Argon TPC to start taking data in 2014 (Collin)
 - Farther ahead
 - LBNE (Bass, Guardincerri)
 - Other ideas: ISODAR, Deadalus (Shaevitz)





Stage 3: seeing the largest mass splitting (Δm_{23}^2)

- T2K: 700MeV at 295km
- MINOS: 3.5GeV at 735km, v and vbar, plus atmospheric v's (Radovic)
- ICECUBE (Deep Core): disappearance 20-100GeV, no disappearance 100GeV-1TeV (Williams)
- Stay tuned:
 - Daya Bay analyzing shape of reactor v's to measure mass splitting! (Δm_{13}^2 , instead of Δm_{23}^2), eventual sensitivity comparable to MINOS (Worcester)
 - T2K with full POT statistics would get to 4x10⁻⁵ eV² error (Friend)
 - NOvA will have strong constraints on this (Baird)



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Stage 4: seeing mass hierarchy through v_e appearance

• Recall that appearance depends on mass hierarchy and CP-violating phase δ , and even whether Θ_{23} is above or below 45°



Current Generation, Ultimate Sensitivities

- NOvA has statistical precision to making precise Δm^2_{23} measurement, regardless of CP phase or mass hierarchy (*Baird*)
- T2K sill also have good statistics and systematics to get to Δm^2_{23} (Friend)
- T2K also showed "full statistics" sensitivites, considering partial running in anti-neutrino beam (*Friend*)





Stage 5: Can we see if there really is CP violation in the lepton sector?

- Big range of ideas going forward
 - "preferences" for mass hierarchy or delta may come soon, but we can't forget what statistical tests really mean
 - We want to see this in as many ways as possible
 - Should not stop thinking about ways to improve on current plans
 - Should not think that 1 experiment is enough
- LBNE is gathering momentum
 - Clever ideas on beamline monitoring (Guardincerri)
 - New attention to accurate simulations (Bass)



- with knowledge of interactions and their current uncertainties
- With better description of detector capabilities

Are there any (more) surprises out there?

- Double Chooz: (Spitz)
 - Search for Lorentz violation

$$\overline{\nu}_e p \to e^+ n$$

- Search for n-nbar oscillations
- MINOS: (Aurisano)
 - Search for sterile neutrinos by comparing NC events in near and far detectors
 - Stay tuned: MINOS+ will have improved sensitivity to sterile neutrino sector through v_{μ} Observed: 1221 events disappearance



Summary

- These are exciting days in Neutrino Physics
- Have running experiments from keV to EeV to see very broad range of physics
- Progress is fast when you think about how recently we were in the denial phase of oscillation physics
- New ideas for absolute mass measurements, we should keep pressure on to make those work as well
- Current generation of accelerator-based experiments have many options ahead of us (v vs anti- v running), have to strive for best way to get the best measurements as a field
 - Very likely that the sum will be greater than each of the parts, especially if we coordinate!