

THE LBNE NEAR DETECTOR

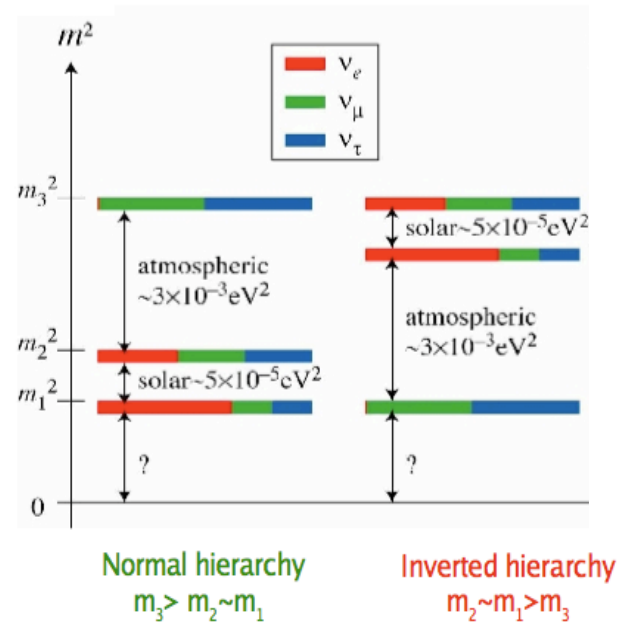
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Status of ν Physics and LBNE

Neutrino oscillation experiments proved that neutrinos are mixed and massive

Open questions

- What is the correct hierarchy pattern?
- Is there \mathcal{CP} and what is the value of the CP phases?
- Type of fermion: Dirac or a Majorana?
- What is the absolute mass scale?
- Are there sterile neutrinos?



- Principal focus of LBNE: simultaneous measurement of MH & \mathcal{CP}
- With a baseline of 1300 km and a beam energy peaking between 0.5 GeV and 10 GeV LBNE could study ν_μ disappearance and $\nu_\mu \rightarrow \nu_e$ oscillations with unprecedented sensitivity
- The resolution and redundancies of LBNE will allow unprecedented reach to search for physics beyond PMNS

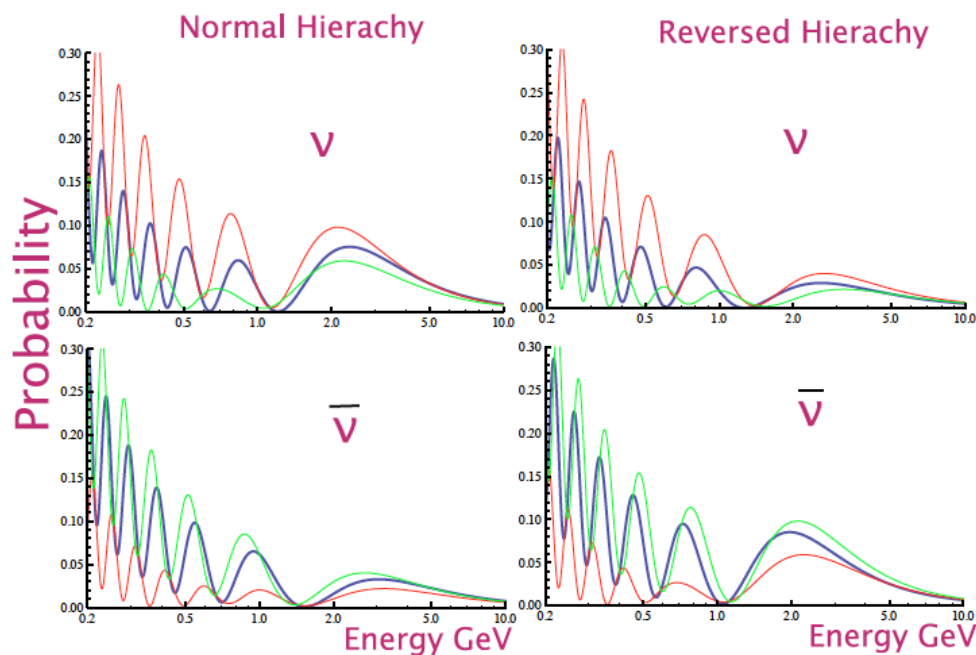
LBNE



LBNE: the signature

$$A(E_\nu) = \frac{P(\nu_\mu \rightarrow \nu_e) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)}{P(\nu_\mu \rightarrow \nu_e) + P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)}$$

- The asymmetry is due to
 - ~~CP~~: if $\delta_{CP} \neq 0, \pi$
 - Matter effect: matter interacts differently with neutrinos and antineutrinos
 - Neutrinos and antineutrinos experience a different index of refraction. $P_{\mu e}$ is different for neutrinos and antineutrinos.
 - The two effects can be searched for (and distinguished) by looking at the spectral features of $P(\nu_\mu \rightarrow \nu_e)$ and $P(\text{anti-}\nu_\mu \rightarrow \text{anti-}\nu_e)$



- $\delta_{CP} = 0$
- $\delta_{CP} = \pi/2$
- $\delta_{CP} = -\pi/2$
- $\sin^2 2\theta_{13} = 0.1$

Motivations for a near detector

- LBNE will not measure oscillation probabilities but event rates:

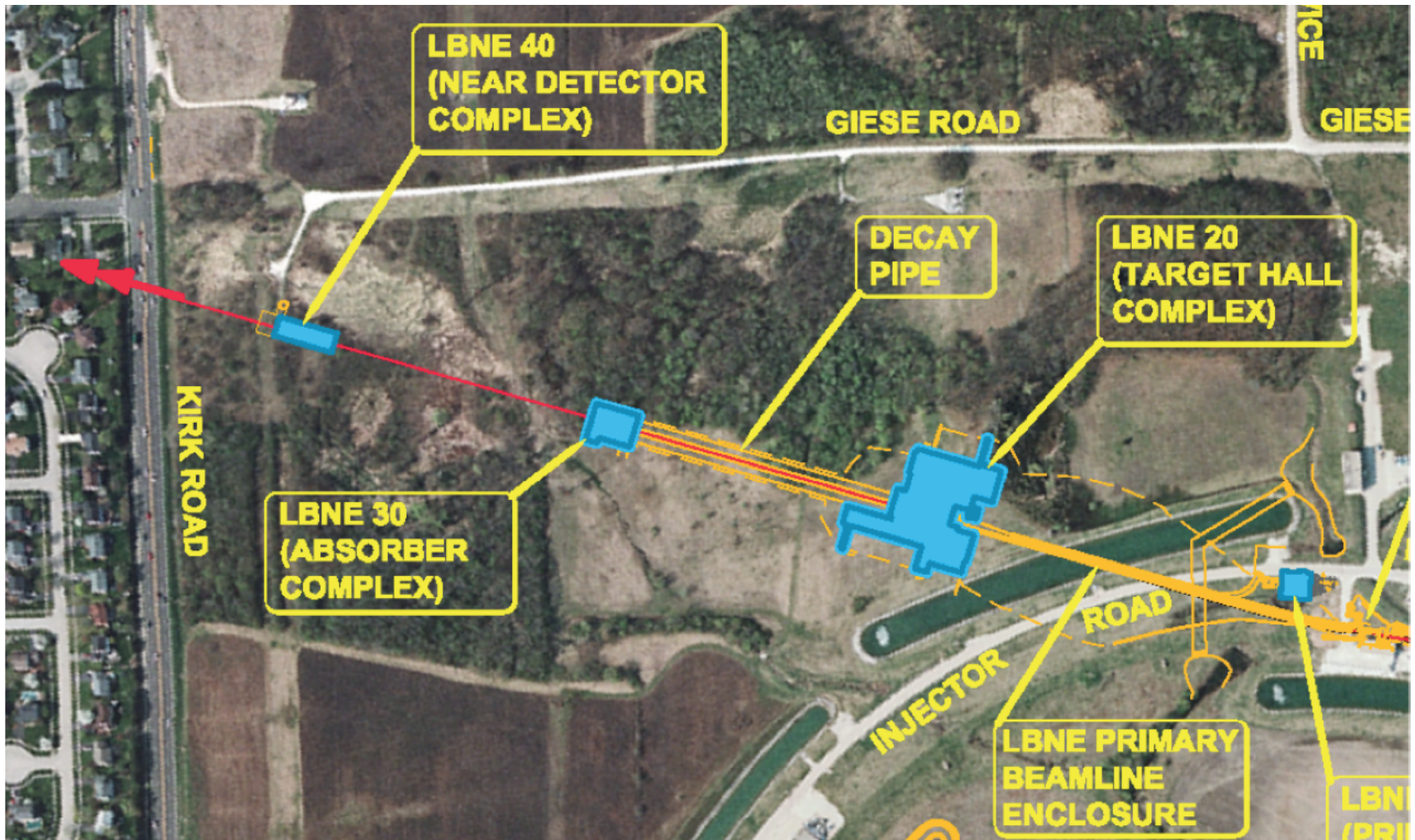
$$R(E)_{\text{FD}} = k \times \Phi(E) \times P(\nu_\alpha \rightarrow \nu_\beta) \times \sigma(E) \times \varepsilon(E) + B(E)$$

- At the far detector (FD), the first and the second oscillation maxima occur at about 2.4 GeV and 0.8 GeV, respectively
- In this energy range neutrino cross sections are poorly known (uncertainties ~20 %)
- crucial to measure the neutrino fluxes and interaction channels at a near site – before the fluxes have been affected significantly by neutrino oscillations – in order to be able to predict both signal and background at the far site
- It is also critical to identify and measure processes that can mimic oscillation signals at the liquid Argon far detector (FD)
 - Mainly charged and neutral pion production
 - backgrounds must be measured individually since different backgrounds have different extrapolations to the FD position

Status of the LBNE near detector

- Approved scope:
 - Beamline measurements (more on this next)
- Goal:
 - Beamline measurements
 - Near Neutrino Detector (more on this next)
 - Strong interest from India to build the neutrino detector with US scientists (STT)
 - In addition, explore the possibility of a LAr neutrino detector
 - Non-DOE (NSF or international) partners could provide a LAr TPC that could be placed upstream of the STT
 - We wish to build a hall that can accommodate more than one detector

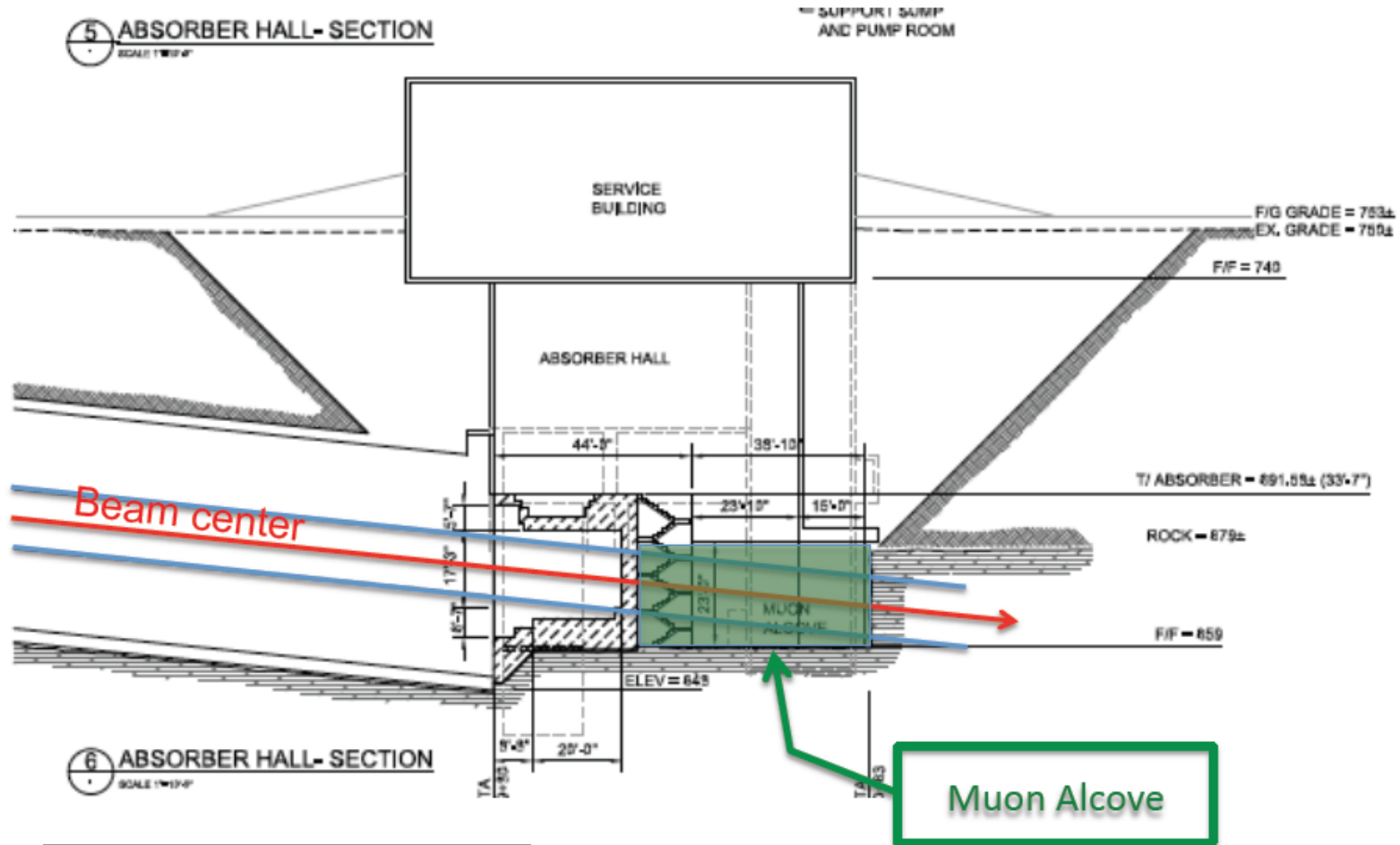
The near detector complex



Proposed near detector complex

- Consists of a beamline measurements system (BLM) and a neutrino-detection system (ND)
- The **ND** will be located at the Near Site (Fermilab), downstream of the beamline
 - Measure the ν_μ , anti- ν_μ , ν_e , anti- ν_e using CC
 - Measure the absolute flux and shape of neutrinos at ND, and predict FD/ND
 - Measure backgrounds to oscillation signal (π^0 , π^+ , π^- , ..)
 - Measure ν -Ar interactions
- The **BLM** will be located in the region of the Absorber Complex at the downstream end of the decay region to measure the muon flux from hadron decay
 - To determine the neutrino flux and spectrum and to monitor the beam profile on a spill-by-spill basis

Muon measurements facilities

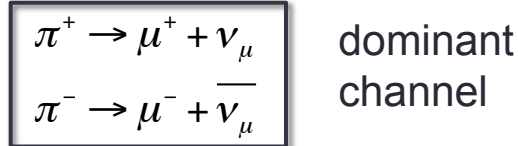


~ 275 m from the target

muon flux at the ND: $10^8/\text{cm}^2/\text{s}$

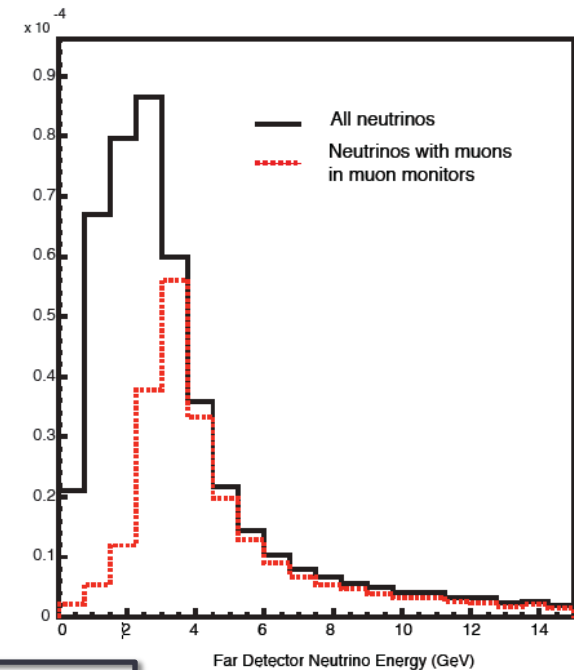
Muon measurements

- Muons and neutrinos come from the same parent pion and kaon decays



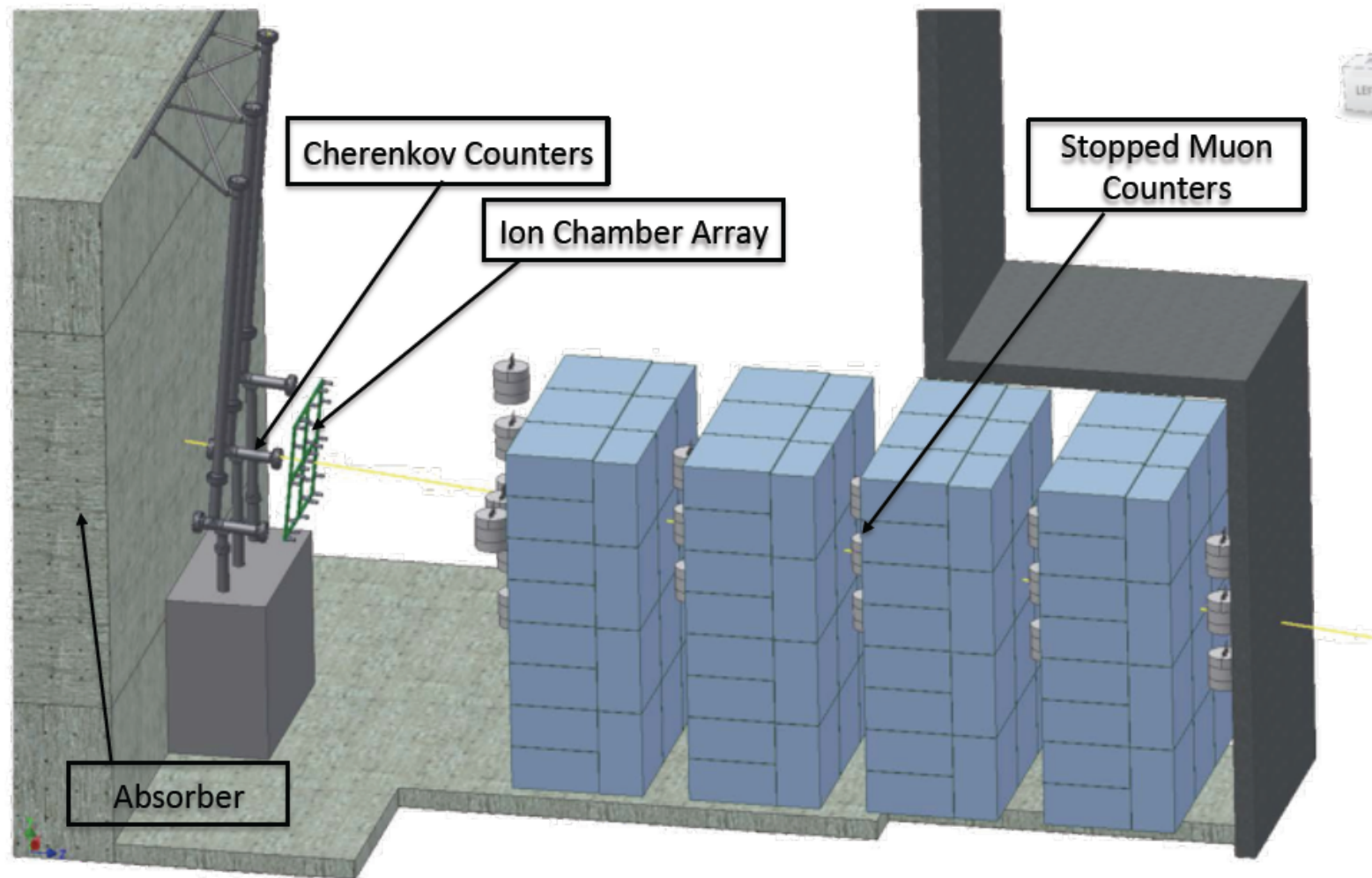
- a measurement of the absolute muon flux and energy spectrum can confirm the absolute neutrino flux
 - Goal: determine the absolute muon flux to an accuracy of 5% above $E_\mu \sim 6$ GeV (corresponding to $E_\nu \sim 1.6$ GeV)

- Muon ionization detectors
 - Beam direction stability over time
- Stopped-muon (Michel-electron) detectors
 - Measure muon energy spectra in $\sim (6, 14)$ GeV range
- Muon Cherenkov detectors
 - Map the muon momentum distribution



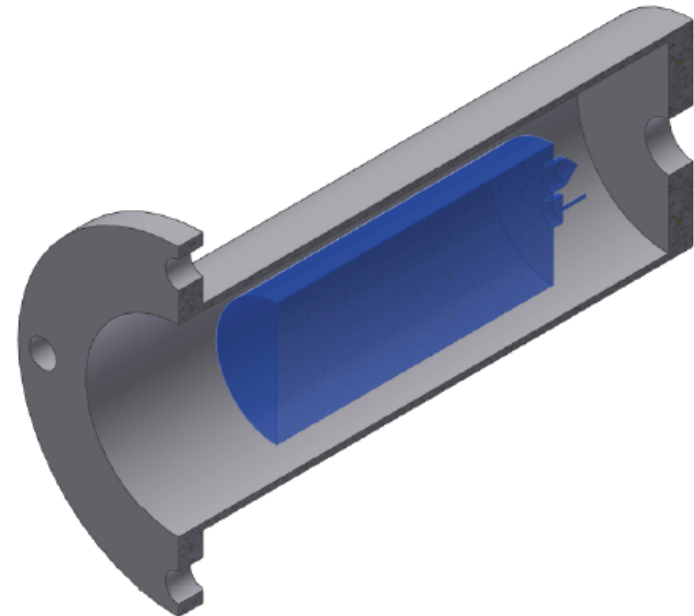
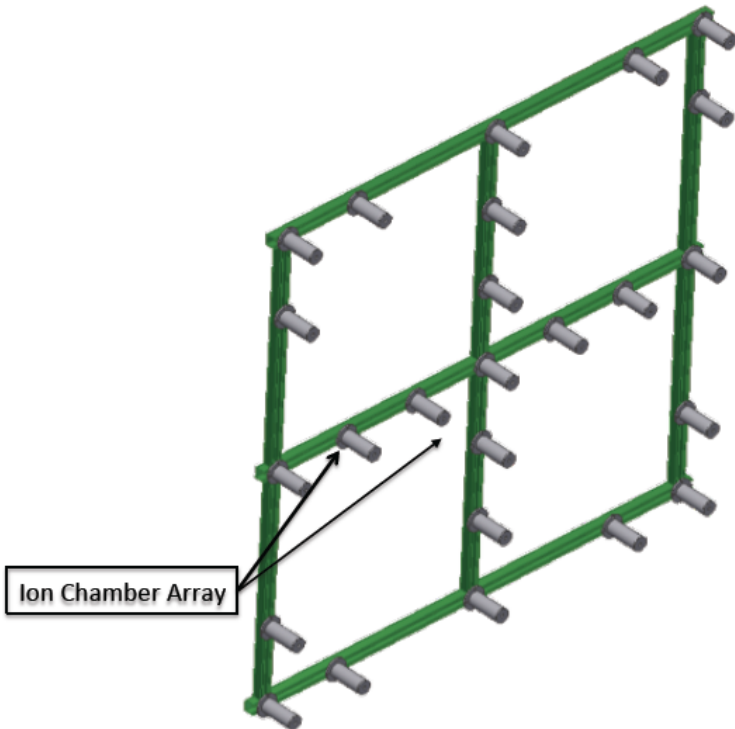
muon flux at the ND: $10^8/\text{cm}^2/\text{s}$

The beamline measurement



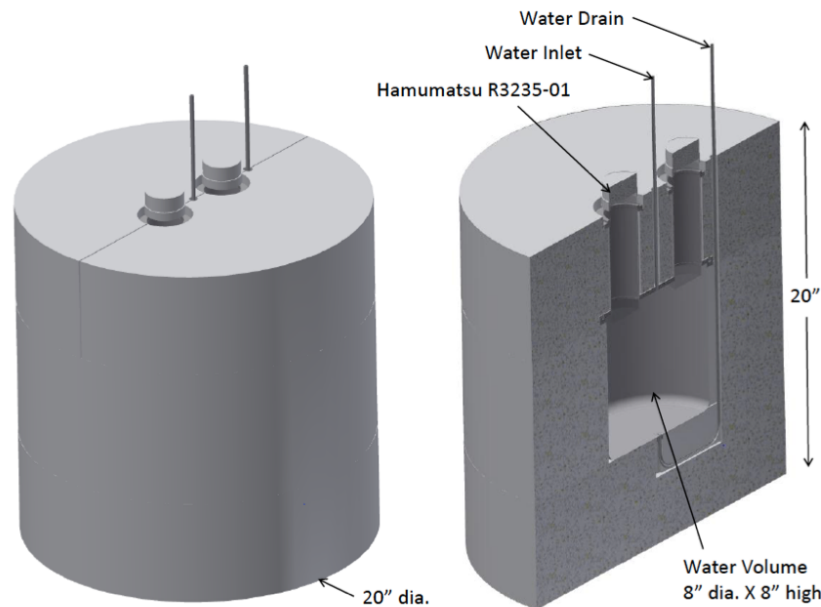
Muon ionization detectors

- A misalignment of the beam leads to a change in the ratio $\text{flux}_{\text{ND}}/\text{flux}_{\text{FD}}$: to keep the change lesser than 1% in all energy bins the beam direction must be known to ~ 0.2 mrad \rightarrow 5 cm
- Sealed ionization counters (rad hard)

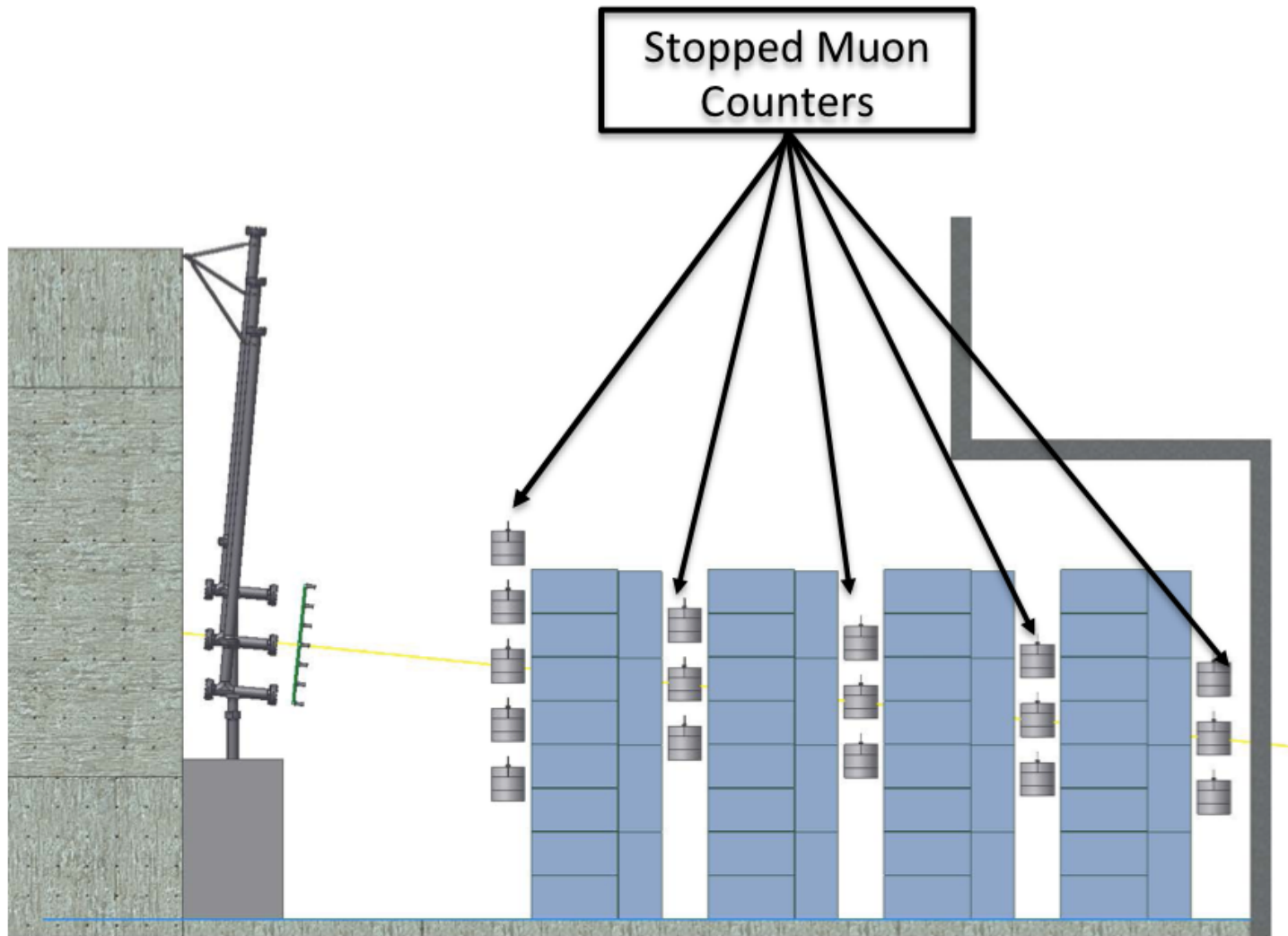


Stopped-muon detectors

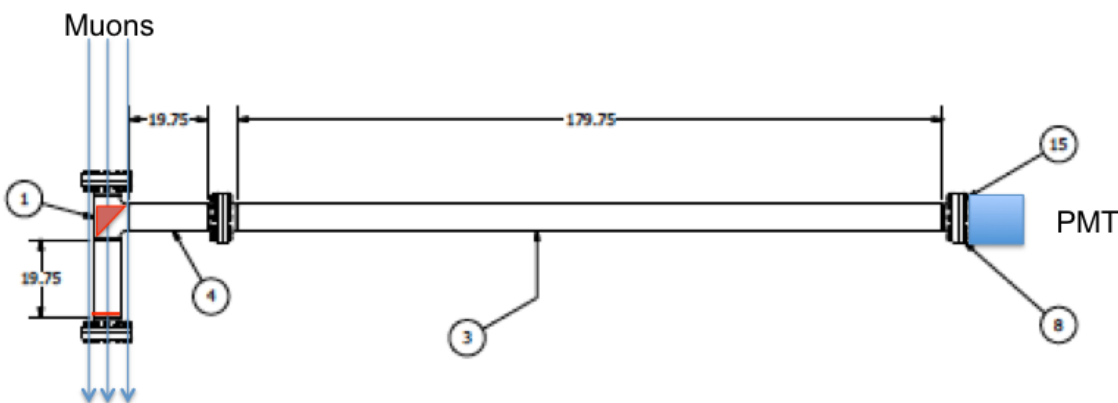
- Detect stopped muons through their decay (μ^+ and μ^-) or capture (μ^-)
 - capable of distinguishing μ^+ and μ^-
 - modules at multiple depths in the shielding/rock behind the absorber to sample the muon flux from different energies



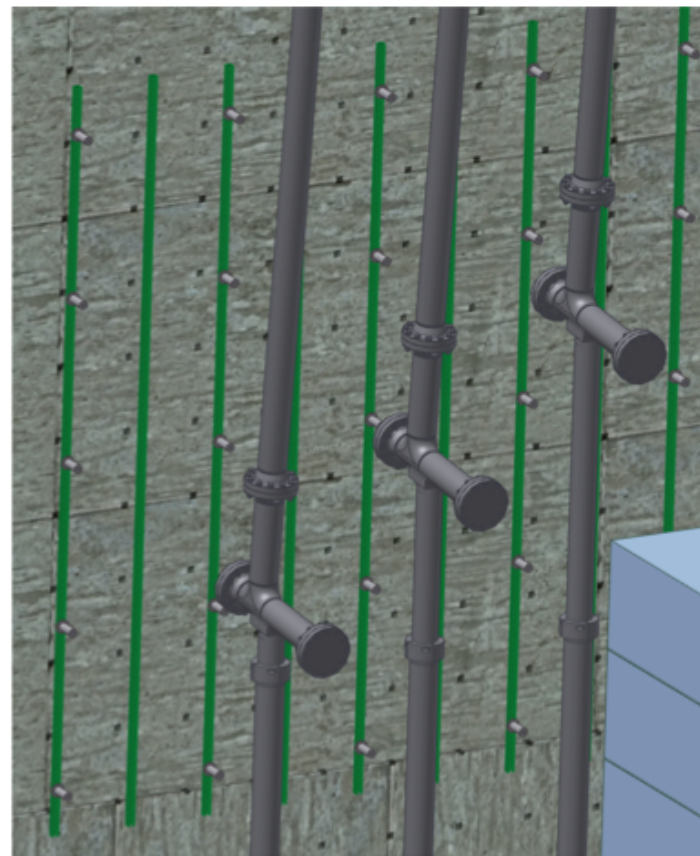
Stopped-muon detectors



Muon Cherenkov detectors



- Counters where a gas (noble gas) radiator is contained in a pressurized tube
- Cherenkov light collected at the end of the tube by a mirror that reflects it 90 degrees towards a PMT
- Gas pressure can be varied from vacuum to several atmospheres
 - this changes the index of refraction, and hence the muon momentum threshold
 - the pressure scan would give the momentum distribution of the muons

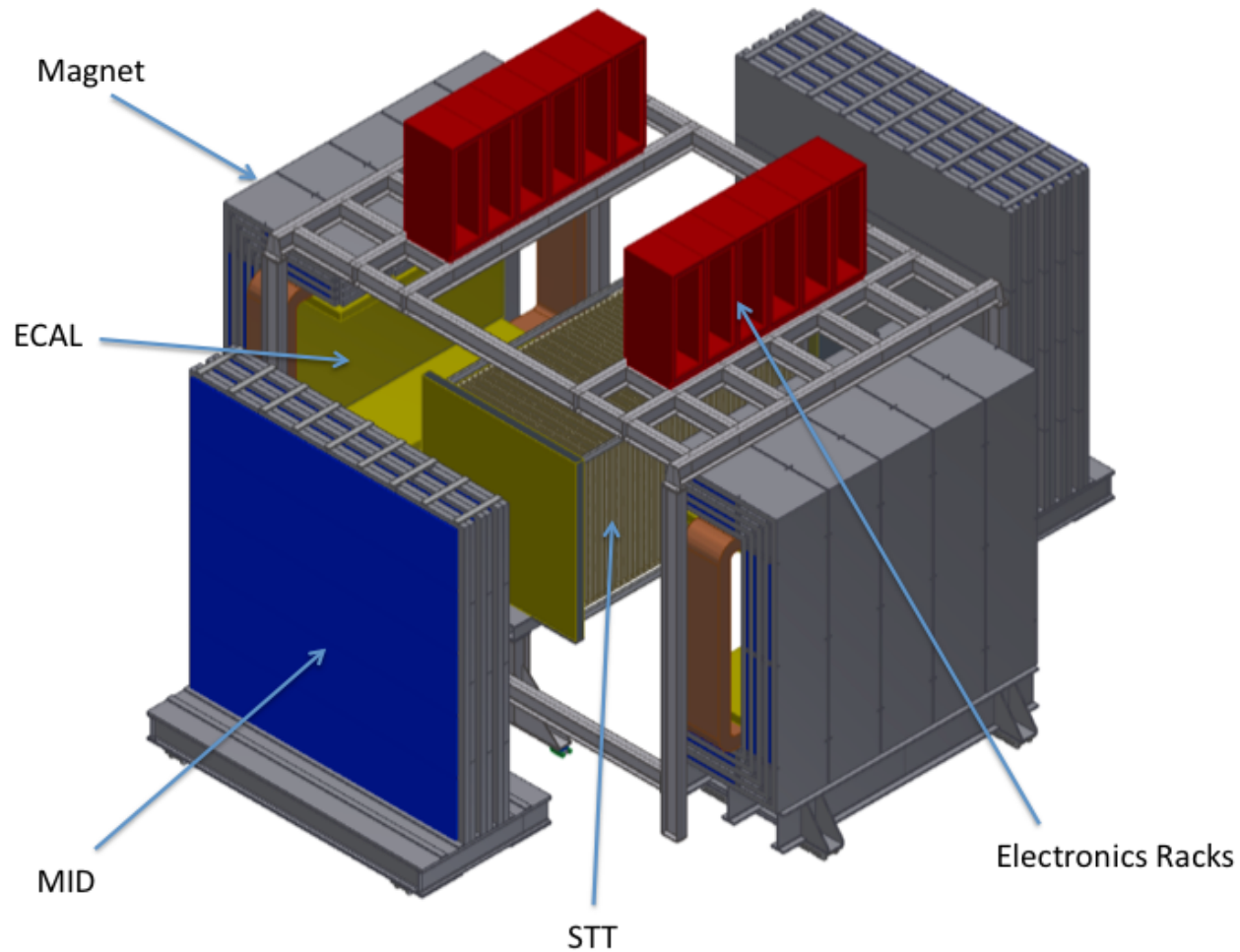


Prototyping effort

- Beam-line detectors being prototyped now
- Assembly of the Cherenkov detector on the surface at Fermilab in progress, transfer underground (alcove 2 of the NuMI beamline) scheduled in August.
 - A stopped-muon counter prototype is being prepared at Drexel also
- Tests scheduled from August 2013 through 2014 using the NuMI beamline parasitically

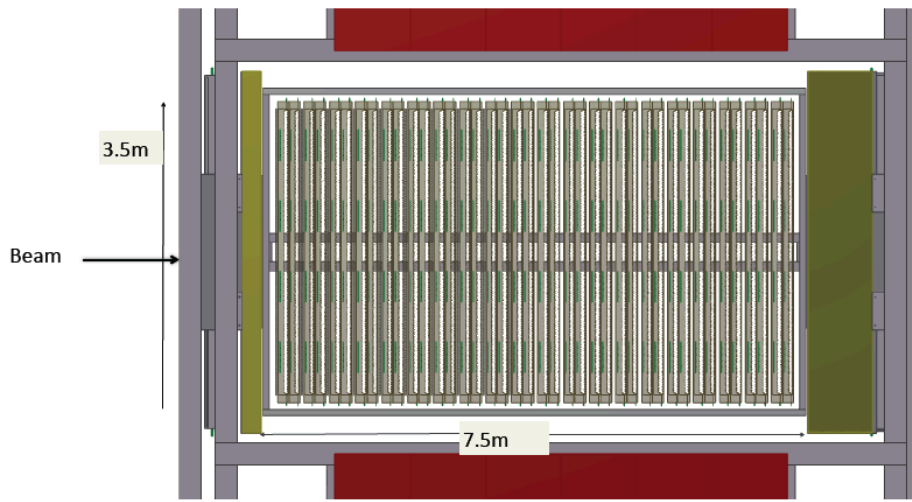


The neutrino detector



Neutrino detector: the Straw Tube Tracker

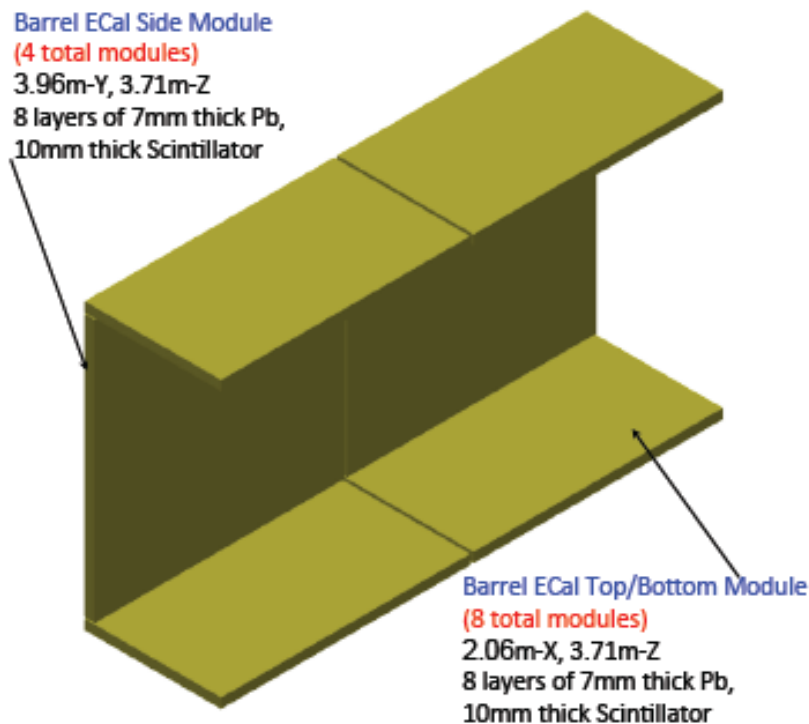
Straw Tube Tracker



- Nuclear targets at the upstream end to study ν -nuclear interactions (provide $\sim x5$ the FD statistics)
- TR \rightarrow e^-/e^+ ID (γ)
- $dE/dx \rightarrow \rho, \pi^{+/-}, K^{+/-}$

- $V = 350 \times 350 \times 750 \text{ cm}^3$
 - $\rho \sim 0.1 \text{ g/cm}^3, m = 7t$
- $\sim 1 \text{ cm}$ diameter ST
- Alternated Y and X planes of straws arranged in modules
- TR Radiator foils between the modules
 - Polypropylene (C_3H_6) $_n$ films alternating with spacers
 - $X_0 \sim 5m$ in the STT

Neutrino detector: the Electromagnetic CALorimeter



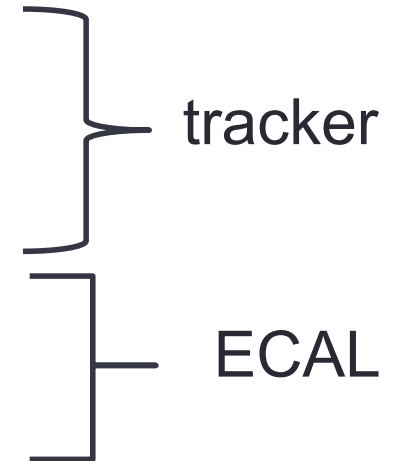
- Tracking volume completely surrounded by an ECAL
 - Layers of 10 mm thick scintillator bars with alternating directions separated by 1.75 mm Pb foils
 - DS: $18 X_0$
 - Side, US: $10 X_0$
-
- 4π coverage enables detection of neutral particles produced in ν interactions (γ , n , K^0) through their shower (longitudinal/transverse profile)

The magnet and the muon identifier detector

- Magnetic field (0.4 T) required to measure the momentum and charge-sign of the products from ν interactions. Purpose:
 - determine the beam flux at the near site
 - requires μ^+ and μ^- separation
 - measure $e^{+/-}$, necessary for the characterization of ν_e -CC and anti- ν_e -CC interactions and the determination of the neutrino beam content (intrinsic ν_e and anti- ν_e irreducible background for ν_e appearance search at the FD)
- Muon ID: Resistive Plate Chambers (RPC) operated in streamer regime
 - Muon Range Detector (MRD): RPCs in the gaps of the magnet yoke
 - External Muon Identifier (EMI): downstream.
 - Detects forward energetic muons

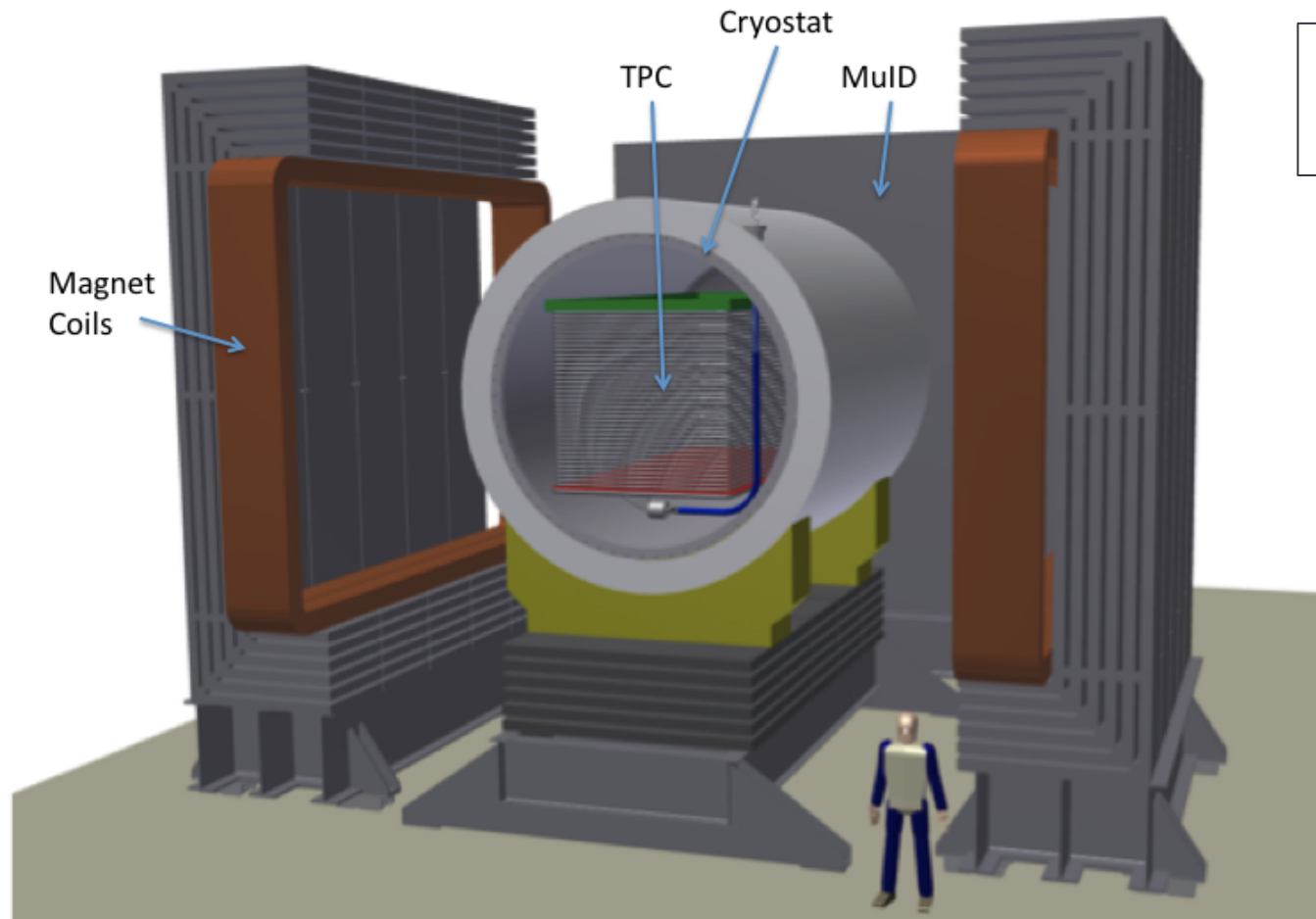
Expected performances

- 200 μm on individual hits, $\sim 100\mu\text{m}$ on the vertex of a ν_μ -CC event
- $\Delta t \sim 1\text{ns}$
- $\sigma_p/p = 0.05/\sqrt{L} \oplus 0.008p/\sqrt{L^5}$ (p in GeV/c², L in m)
- 1/1000 pion rejection from TR
- $\Delta E/E \sim 6\%/\sqrt{E}$
- $\Delta t \sim 1\text{ns}$
- Powerful discrimination of EM/hadronic showers



- Flux determination:
 - Absolute ν_μ flux $\approx 2.5\%$ for $E_\nu < 10$ GeV, absolute ν_μ flux $\sim 3\%$ at high energy
 - $\nu_\mu(E)/\text{anti-}\nu_\mu(E) < 2\%$ for $1.5 < E_\nu < 30$ GeV, $\nu_\mu(E)/\text{anti-}\nu_\mu(E) \sim 3\%$ for $0.5 < E_\nu < 1.5$ GeV
 - $\nu_e(E)/\nu_\mu(E) < 0.1\%$, $\text{anti-}\nu_e(E)/\text{anti-}\nu_\mu(E) < 0.1\% \rightarrow$ Absolute $\nu_e(\text{anti-}\nu_e) \sim 2\%$
 - Precision on shape of the $\nu_\mu(\text{anti-}\nu_\mu)$ dominant component $\leq 2\%$ for $1 < E_\nu < 30$ GeV
- Excellent event reconstruction and identification capabilities

The neutrino detector: alternative design



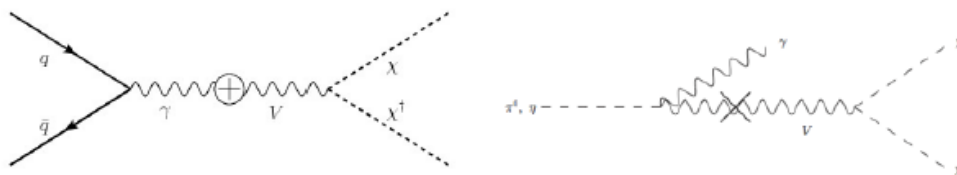
design developed
for the March
2012 CD-1

Physics opportunities

- The intense neutrino flux (expected $\sim 10^8$ CC interactions for 10^{22} POT) combined with the fine segmentation of the ND offer the opportunity for many physics measurements besides the long-baseline oscillation program
 - Weak mixing angle θ_W
 - Strangeness content of the nucleus
 - Nucleon Structure, Parton Distribution Functions and QCD studies
 - Nuclear effect impact on neutrino-nuclear interactions
 - Search for heavy neutrinos
 - Search for high Δm^2 neutrino oscillations
 - Light Dark Matter searches in the neutrino beam

Sub GeV dark matter searches

- Due to the lack of evidence for WIMPS at direct detection experiments and LHC (search for $m_W > 10$ GeV) new theories with light DM particles formulated
- Light mediator (dark photon V) needed to allow efficient DM annihilation in the early universe
 - In the simplest model an $U(1)$ gauge field mixes with the SM $U(1)$ gauge field
 - Detectable at high flux ν experiments (LBNE)
- V produced by protons striking the target either directly or via π^0 or η production and decay
 - If $m_V > 2m_{DM}$ $V \rightarrow 2\chi$



- χ detectable through NC-like interactions on nucleons/electrons. Same signature as neutrinos but:
 - different timing ($v_\chi \ll v_\nu$)
 - Electron recoil much more forward
 - A special run with no focusing would reduce the flux of neutrinos

Thank you!