

# Multiple Probes of Lorentz Violation with Reactor Antineutrinos

Joshua Spitz  
DPF, 8/16/2013

Based on: [Phys. Rev. D 86 112009 \(2012\)](#) and [arXiv:1307.5789 \[hep-ex\] \(2013\)](#)

# Outline

- Introduction to Double Chooz
- Introduction to Lorentz violation
- A search for a time-dependent oscillation signal
- A search for neutrino-antineutrino mixing

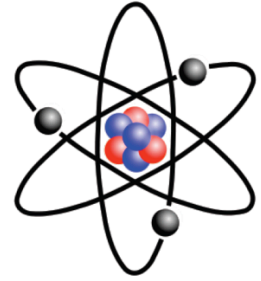
# Double Chooz

(The north of France, a few miles from Belgium)





Reactor ( $\bar{\nu}_e$  source)



$\bar{\nu}_e$

$\bar{\nu}_e$



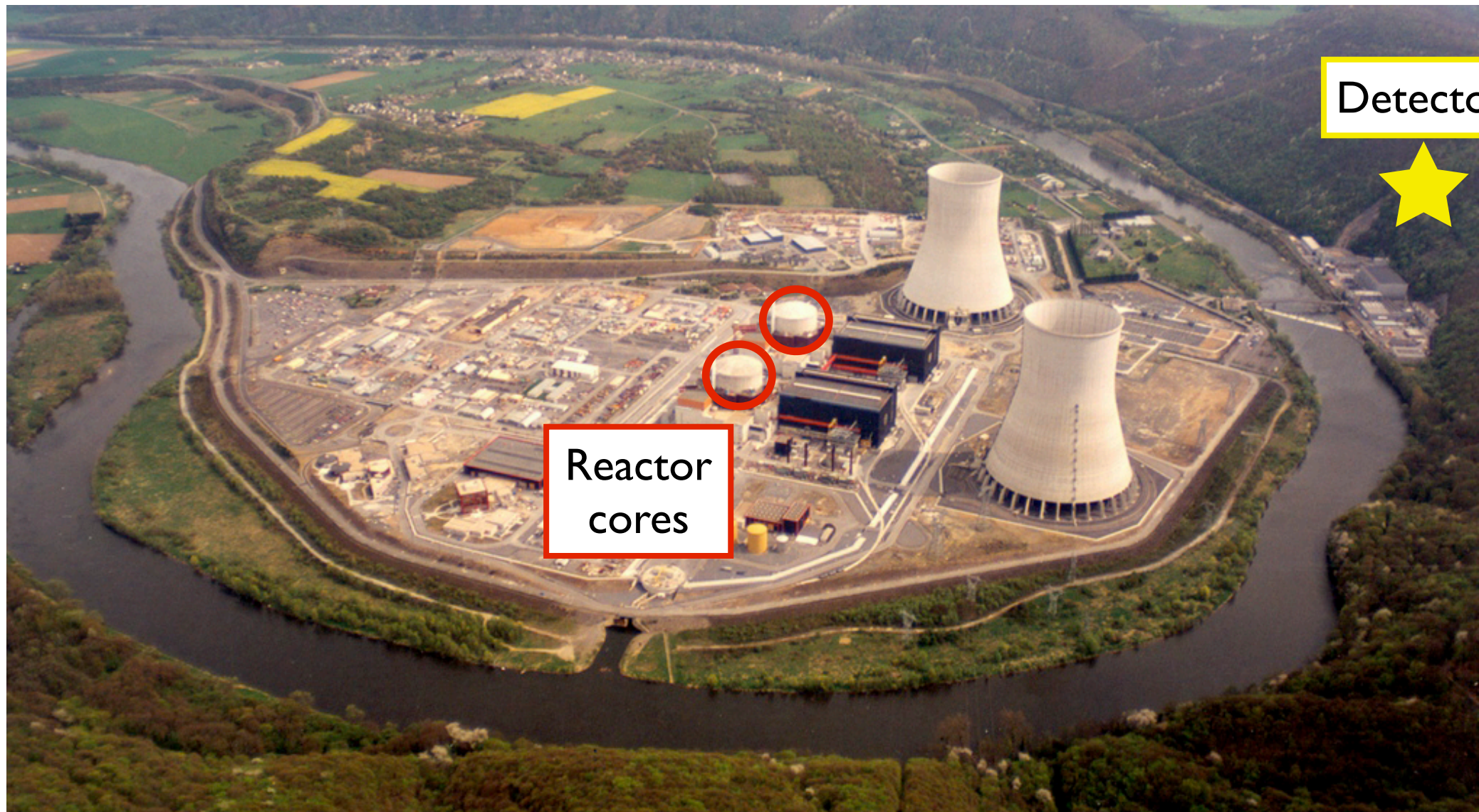
$\bar{\nu}_e$  detector

signal

$\bar{\nu}_\mu$

$\bar{\nu}_\tau$

In a *disappearance experiment*, we look for a deficit of antineutrinos



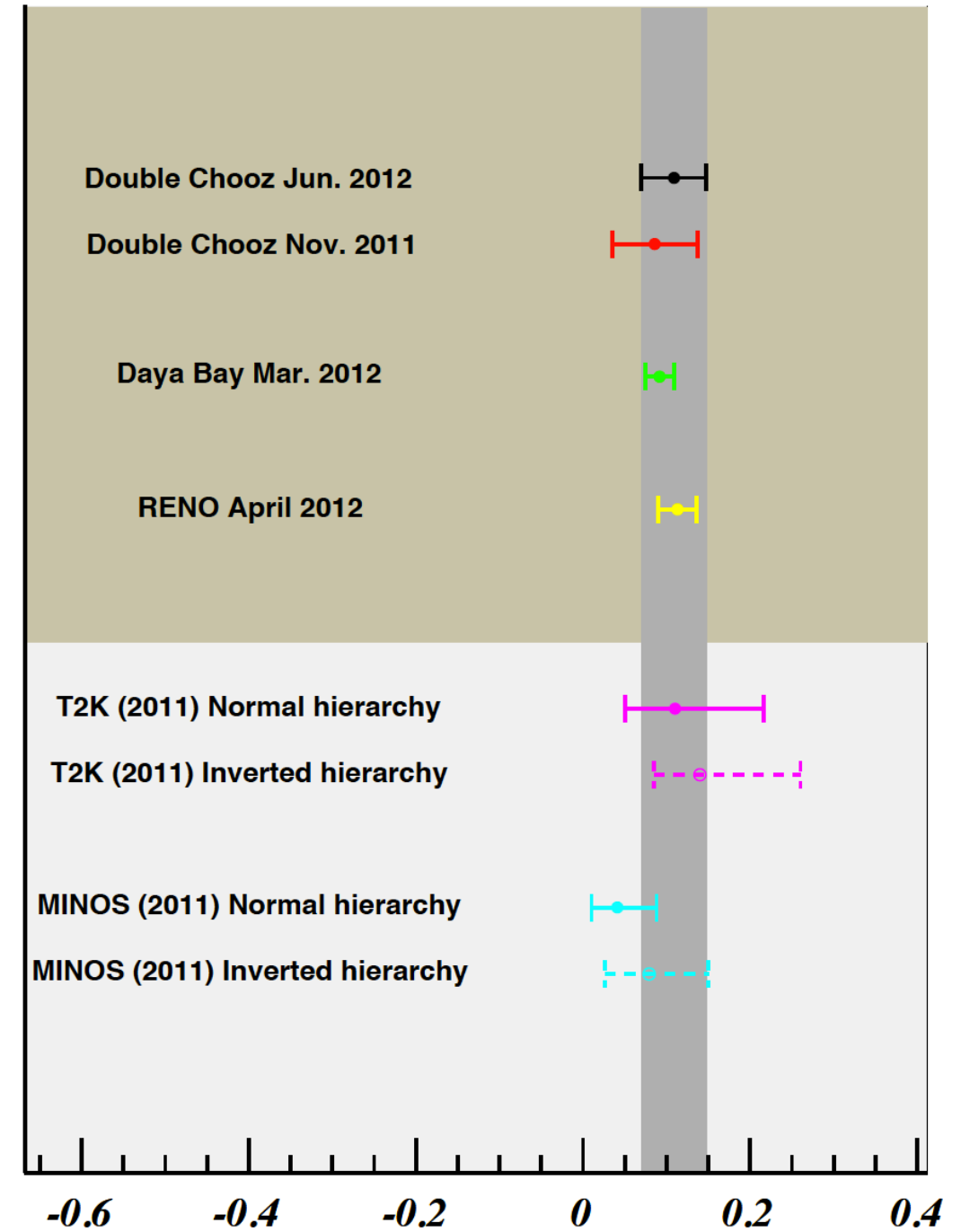
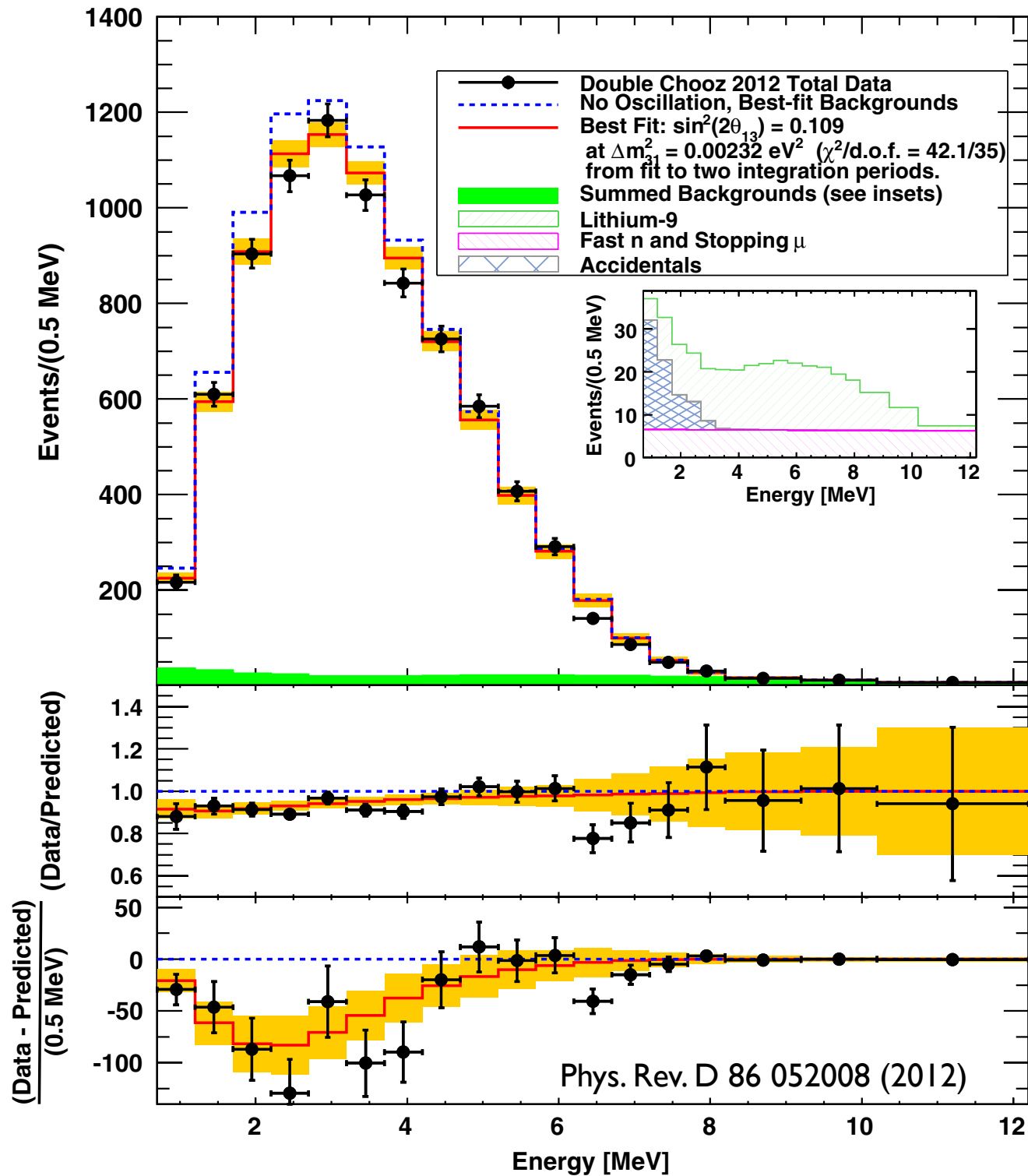
Detector



Reactor  
cores

# The last mixing angle

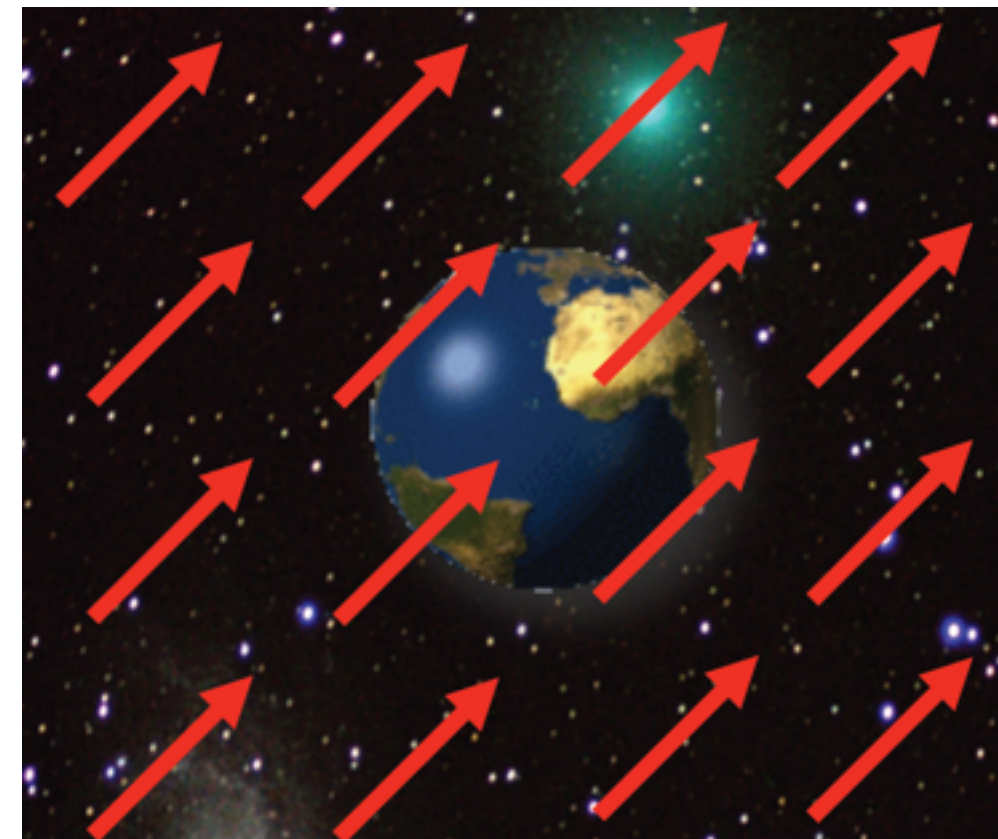
$\sin^2 2\theta_{13}$  Measurements



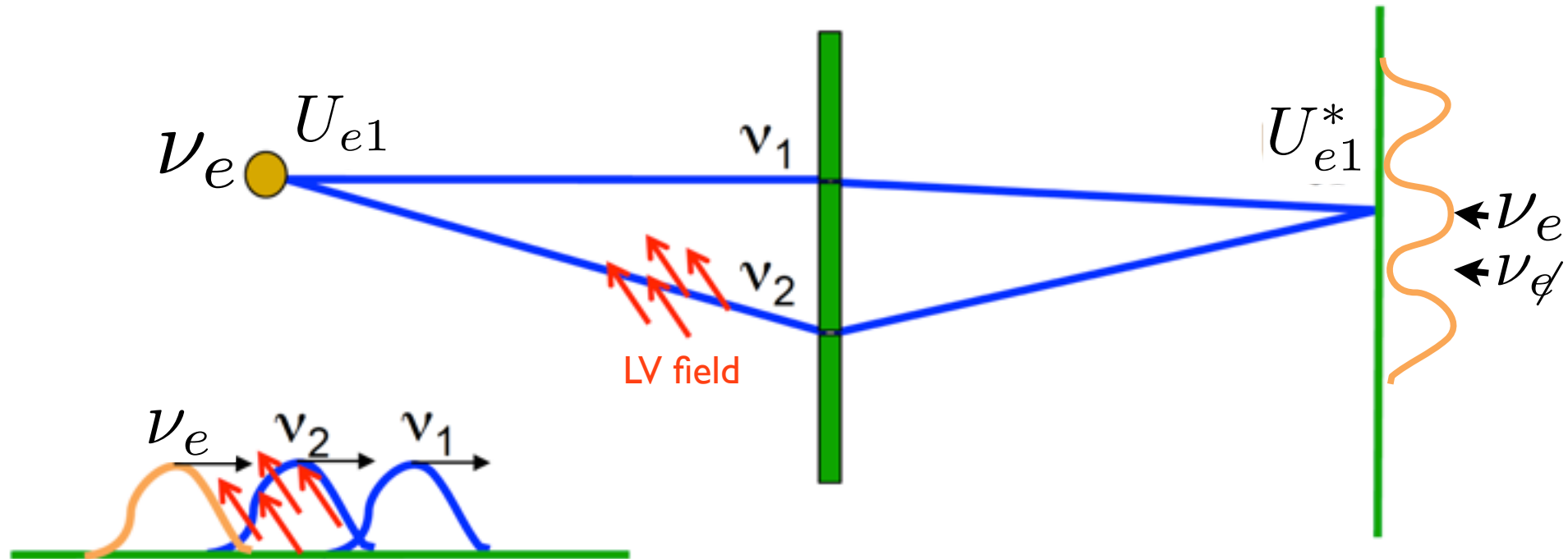


# What is Lorentz violation?

- Lorentz invariance requires that the behavior of a particle is independent of its direction or boost velocity.
- Basically, LV means that the universe has a preferred direction.
- Our SM particles can couple to this background field and create observable effects.
- LV has never been seen.



# Neutrino oscillation and LV



Neutrino oscillations are natural interferometers!

If the mass eigenstates have different couplings to a Lorentz violating field, the oscillation pattern will be affected.

Neutrino eigenvalue difference is comparable to the target scale of Lorentz violation (Planck scale)...  $< 10^{-19}$  GeV

# Is there any hope of ever seeing LV with neutrinos?

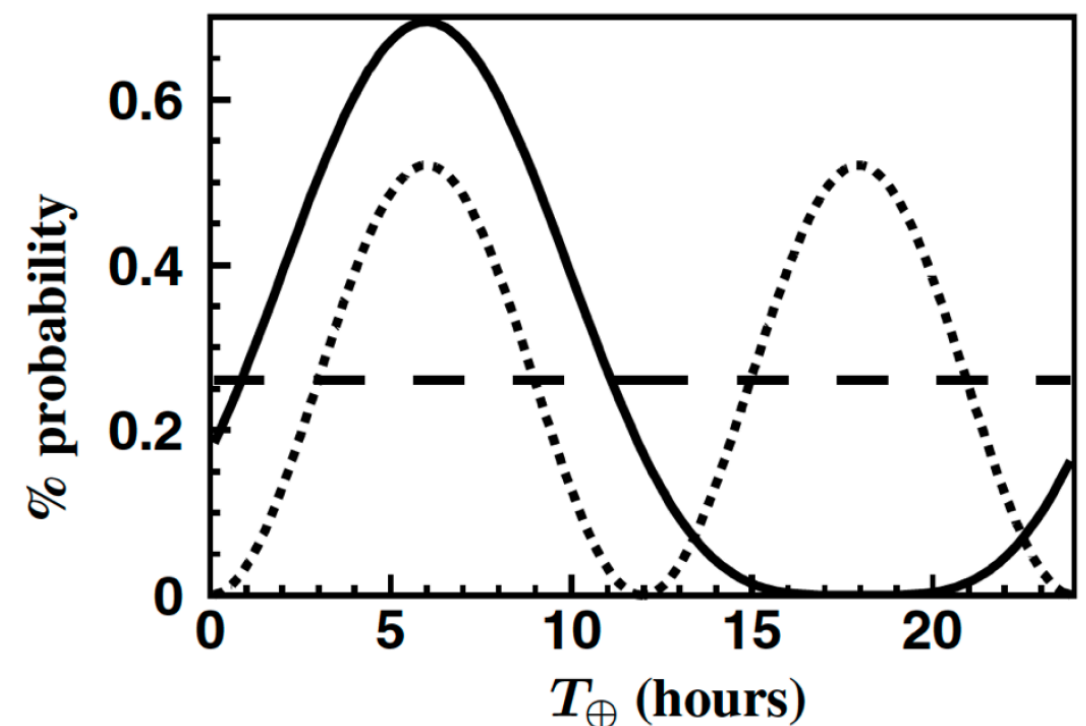
- First and foremost, “You can’t see if you don’t look”.
- Note that neutrinos generally do not provide the best sensitivity to LV. Measurements of gamma ray burst photons set the best limits.
- Gamma rays from GRBs all seem to arrive at the same time, despite having different energies/frequencies and traveling a very long way.
- Neutrinos are special because they only feel the weak force and thus can avoid QED constraints.
- Neutrinos are also special because we don’t understand them very well.



# How to look for it?

- Strange energy dependence (i.e. non-L/E behavior).
- CPT violation (differences between neutrinos and antineutrinos).
- Neutrino-antineutrino mixing.
- Periodicity (in time) of neutrino oscillation.

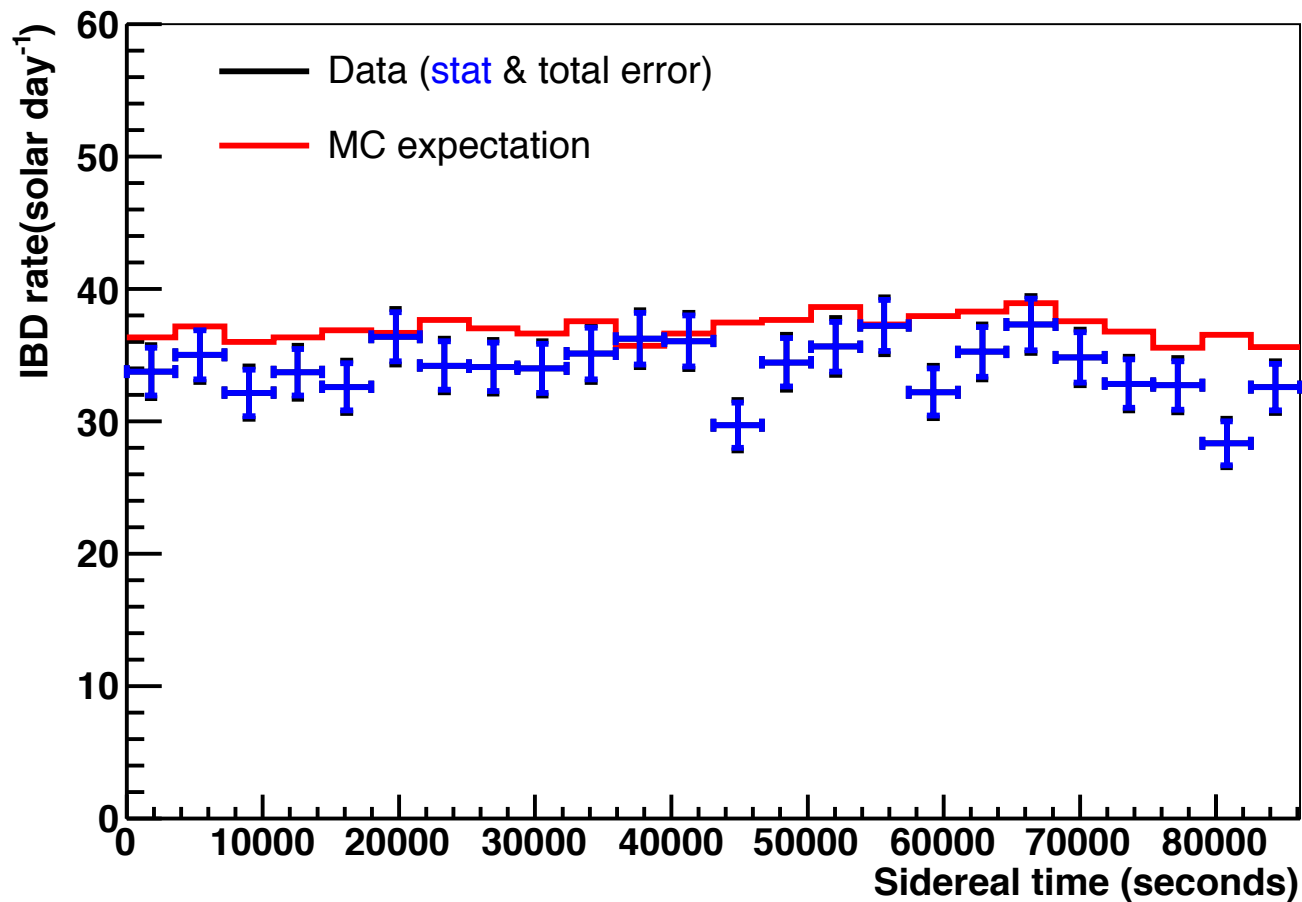
An example of a sidereal dependence of oscillation probability



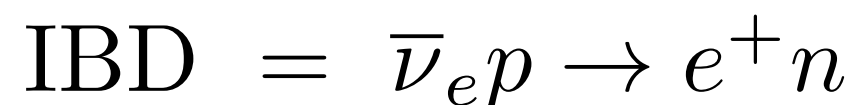
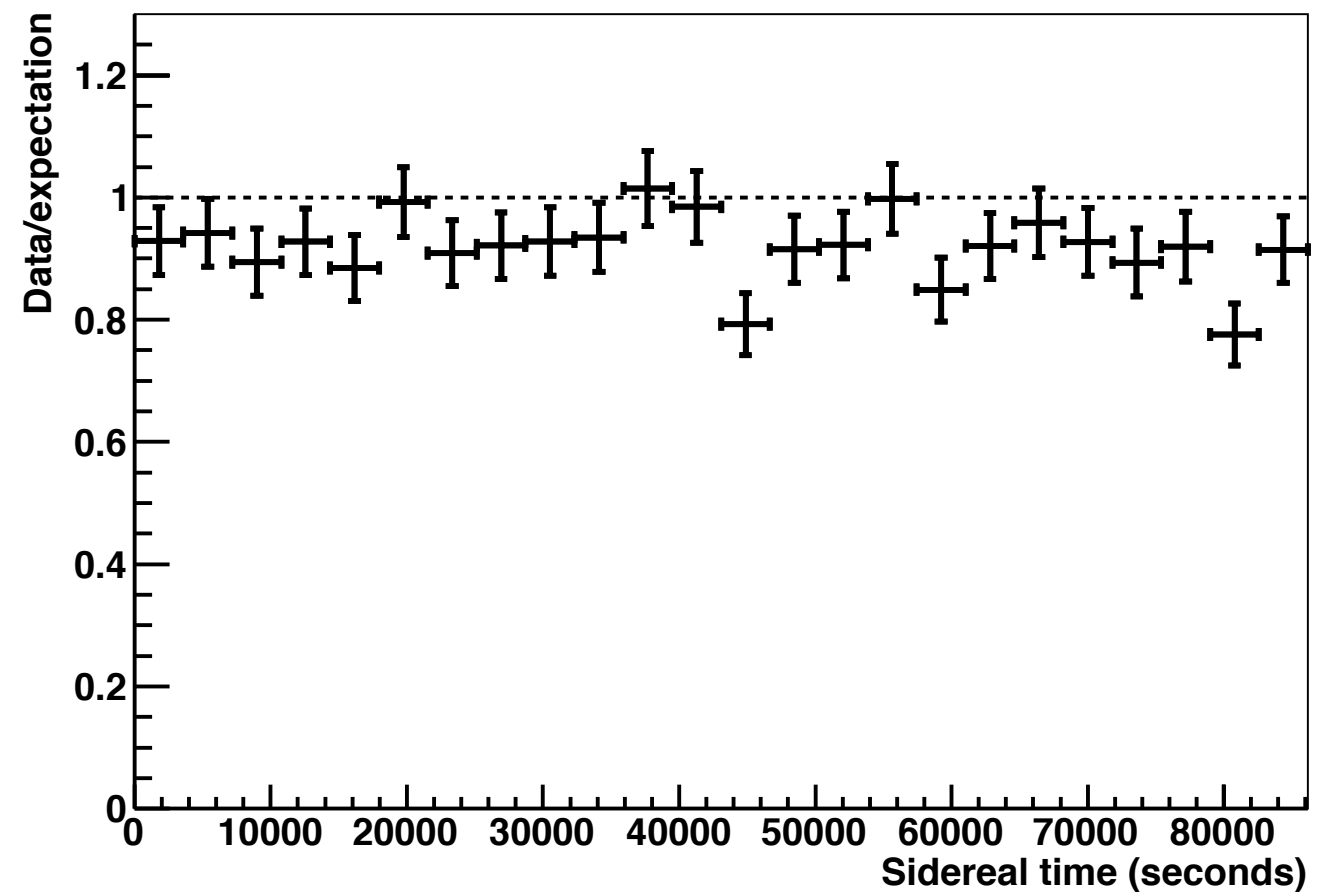
# Rate and time

We search for a sidereal modulation among the 8249 electron antineutrino candidates

IBD rate in sidereal time (227.9 live solar days)



IBD rate in sidereal time (227.9 live solar days)



# What to fit?

The Standard Model Extension (SME) is a framework for all possible types of Lorentz violation and separates the different experimental effects that may be seen.

The full SME equation for electron antineutrino disappearance:

Kostelecký & Mewes, PRD **69**, 016005 (2004)

Kostelecký & Mewes, PRD **85**, 096005 (2012)

$$\begin{aligned}
 P_{\bar{\nu}_e \rightarrow \bar{\nu}_e} &\simeq 1 - P_{\bar{\nu}_e \rightarrow \bar{\nu}_\mu} - P_{\bar{\nu}_e \rightarrow \bar{\nu}_\tau} \\
 &= 1 - \frac{|(h_{\text{eff}})_{\bar{e}\bar{\mu}}|^2 L^2}{(\hbar c)^2} - \frac{|(h_{\text{eff}})_{\bar{e}\bar{\tau}}|^2 L^2}{(\hbar c)^2} \\
 &= 1 - \frac{L^2}{(\hbar c)^2} [ |(\mathcal{C})_{\bar{e}\bar{\mu}} + (\mathcal{A}_s)_{\bar{e}\bar{\mu}} \sin \omega_\oplus T_\oplus + (\mathcal{A}_c)_{\bar{e}\bar{\mu}} \cos \omega_\oplus T_\oplus \\
 &\quad + (\mathcal{B}_s)_{\bar{e}\bar{\mu}} \sin 2\omega_\oplus T_\oplus + (\mathcal{B}_c)_{\bar{e}\bar{\mu}} \cos 2\omega_\oplus T_\oplus|^2 \\
 &\quad - |(\mathcal{C})_{\bar{e}\bar{\tau}} + (\mathcal{A}_s)_{\bar{e}\bar{\tau}} \sin \omega_\oplus T_\oplus + (\mathcal{A}_c)_{\bar{e}\bar{\tau}} \cos \omega_\oplus T_\oplus \\
 &\quad + (\mathcal{B}_s)_{\bar{e}\bar{\tau}} \sin 2\omega_\oplus T_\oplus + (\mathcal{B}_c)_{\bar{e}\bar{\tau}} \cos 2\omega_\oplus T_\oplus|^2 ]
 \end{aligned}$$

contains 28 SME  
coefficients

10 params.





# We perform the following fits:

Assume e- $\mu$  terms are zero:

$$P_{\bar{\nu}_e \rightarrow \bar{\nu}_e} \simeq 1 - \frac{L^2}{(\hbar c)^2} [ | (\mathcal{C})_{\bar{e}\bar{\tau}} + (\mathcal{A}_s)_{\bar{e}\bar{\tau}} \sin \omega_{\oplus} T_{\oplus} + (\mathcal{A}_c)_{\bar{e}\bar{\tau}} \cos \omega_{\oplus} T_{\oplus} + (\mathcal{B}_s)_{\bar{e}\bar{\tau}} \sin 2\omega_{\oplus} T_{\oplus} + (\mathcal{B}_c)_{\bar{e}\bar{\tau}} \cos 2\omega_{\oplus} T_{\oplus} |^2 ] \quad \mathbf{5 \text{ params.}}$$

Assume e- $\tau$  terms are zero:

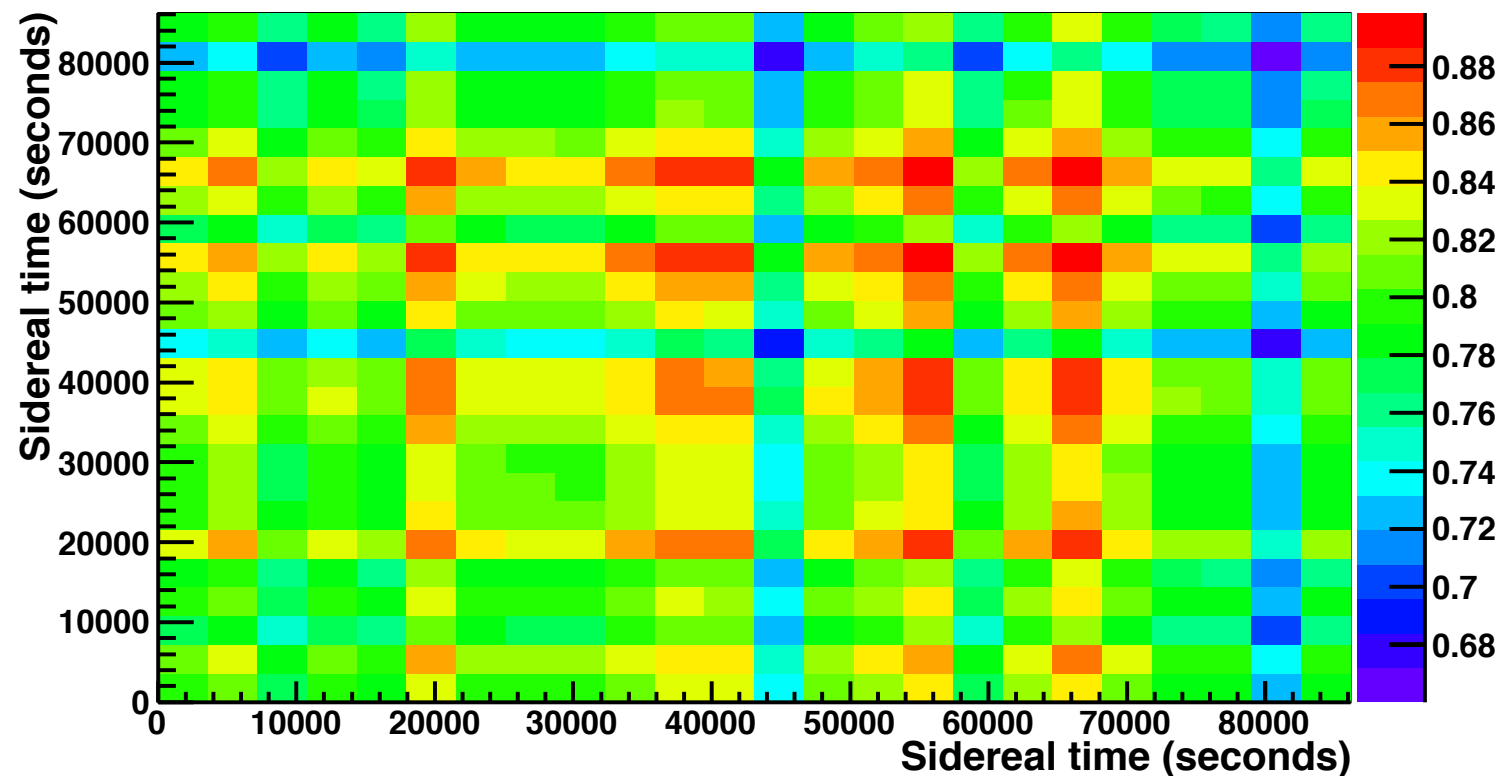
$$P_{\bar{\nu}_e \rightarrow \bar{\nu}_e} \simeq 1 - \frac{L^2}{(\hbar c)^2} [ | (\mathcal{C})_{\bar{e}\bar{\mu}} + (\mathcal{A}_s)_{\bar{e}\bar{\mu}} \sin \omega_{\oplus} T_{\oplus} + (\mathcal{A}_c)_{\bar{e}\bar{\mu}} \cos \omega_{\oplus} T_{\oplus} |^2 ] \quad \mathbf{3 \text{ params.}}$$

# Fit technique

- We use a least squares fitting technique to extract the best fit parameters.

$$X^2 = \sum_{ij}^N [r_{i,data} - P_{\bar{\nu}_e \rightarrow \bar{\nu}_e}(t, SME) \cdot r_{i,MC}] \cdot M_{ij}^{-1} [r_{j,data} - P_{\bar{\nu}_e \rightarrow \bar{\nu}_e}(t, SME) \cdot r_{j,MC}]$$

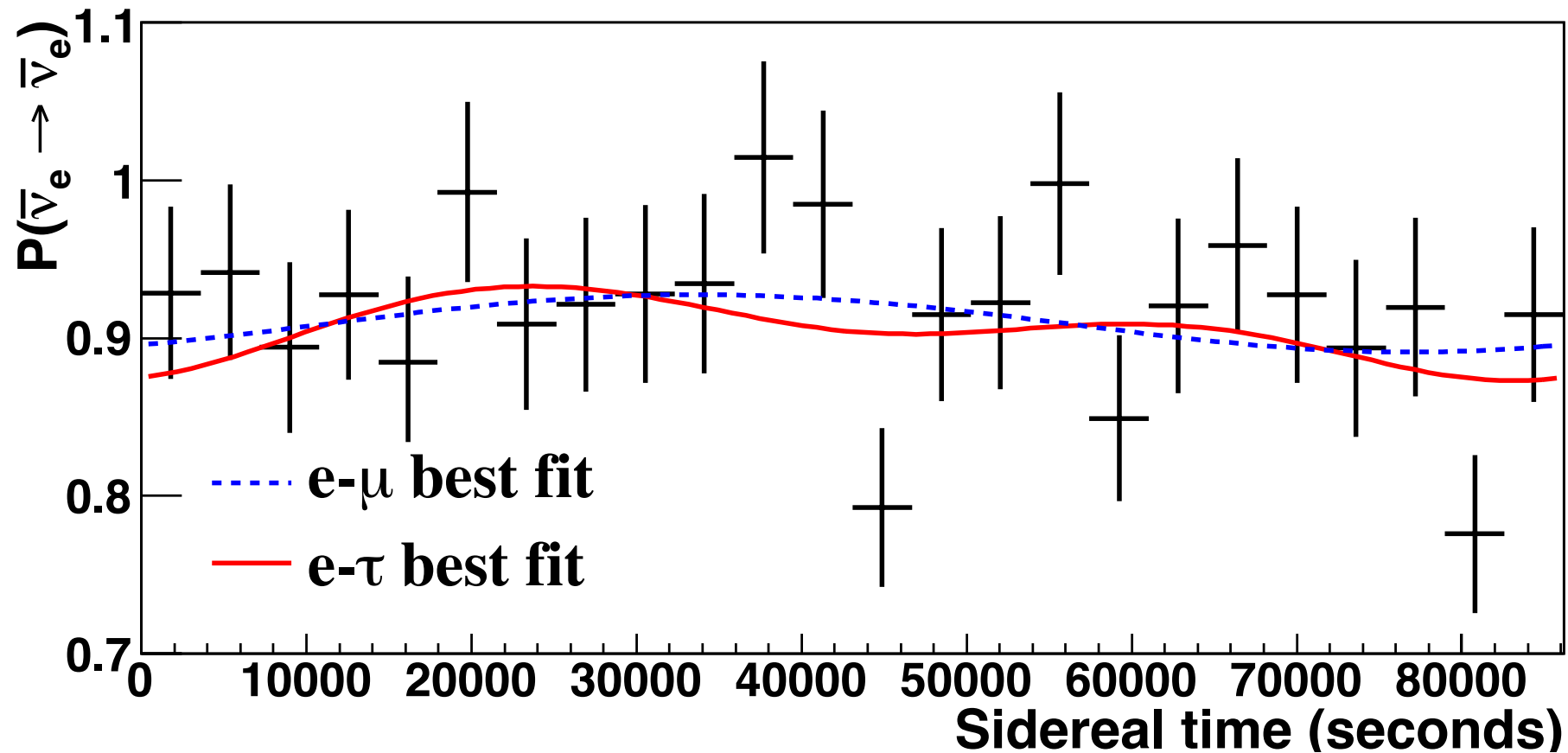
Systematics covariance matrix (sidereal)



full covariance matrix

Source	Variance wrt data
Stats.	1.10%
MC correction	1.01%
Background	1.69%
Reactor+detector	1.75%
Total	2.85%

# Best fit results



---  $P_{\bar{\nu}_e \rightarrow \bar{\nu}_e} \simeq 1 - \frac{L^2}{(\hbar c)^2} [ |(\mathcal{C})_{\bar{e}\bar{\mu}} + (\mathcal{A}_s)_{\bar{e}\bar{\mu}} \sin \omega_{\oplus} T_{\oplus} + (\mathcal{A}_c)_{\bar{e}\bar{\mu}} \cos \omega_{\oplus} T_{\oplus}|^2 ]$

—  $P_{\bar{\nu}_e \rightarrow \bar{\nu}_e} \simeq 1 - \frac{L^2}{(\hbar c)^2} [ |(\mathcal{C})_{\bar{e}\bar{\tau}} + (\mathcal{A}_s)_{\bar{e}\bar{\tau}} \sin \omega_{\oplus} T_{\oplus} + (\mathcal{A}_c)_{\bar{e}\bar{\tau}} \cos \omega_{\oplus} T_{\oplus} + (\mathcal{B}_s)_{\bar{e}\bar{\tau}} \sin 2\omega_{\oplus} T_{\oplus} + (\mathcal{B}_c)_{\bar{e}\bar{\tau}} \cos 2\omega_{\oplus} T_{\oplus}|^2 ]$



# Is oscillation probability independent of time?

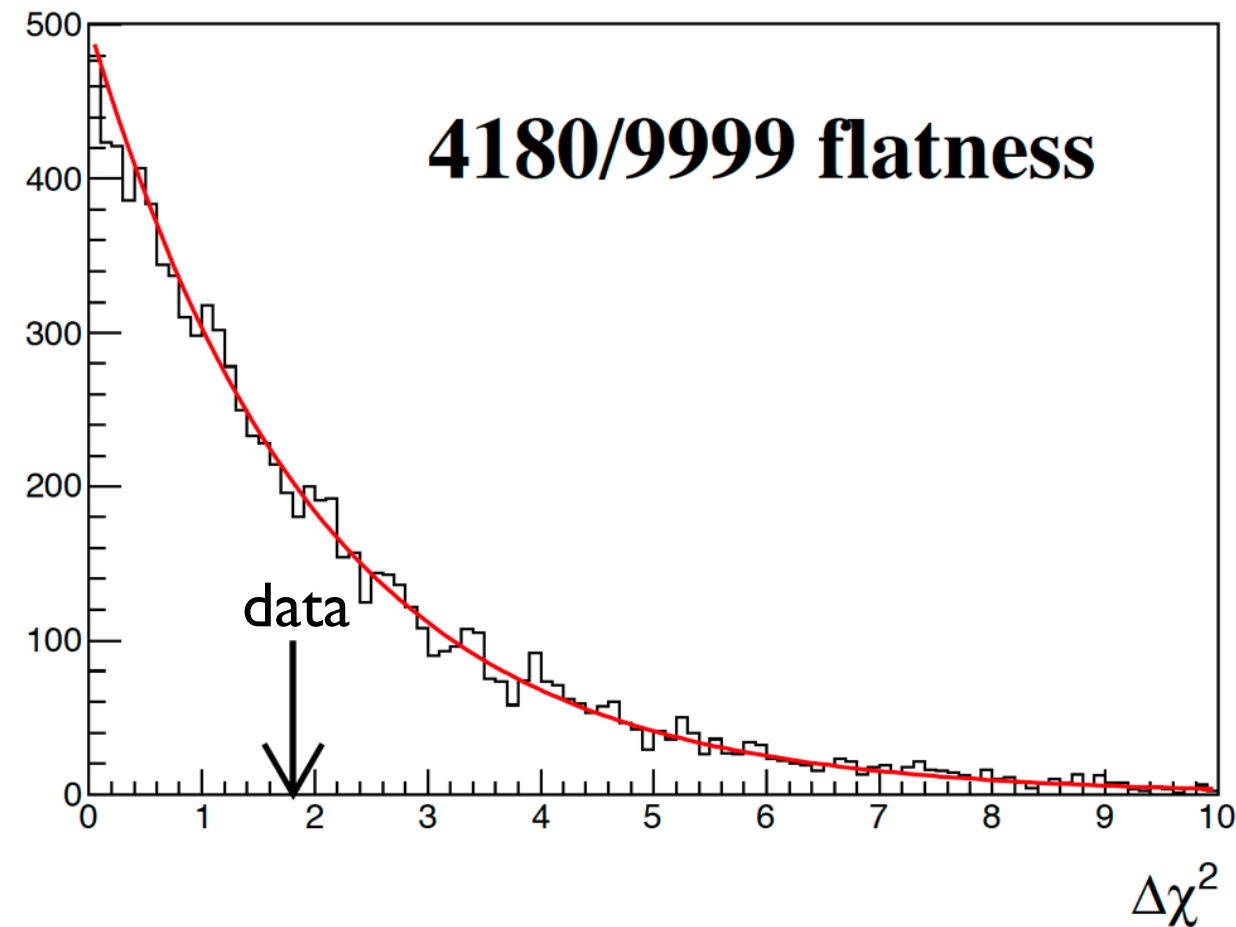
A frequentist study is used to determine the significance of the results

$$\Delta\chi^2 = \chi^2 \text{ (1 parameter fit)} - \chi^2 \text{ (3 parameter fit)}$$

Sidereal time independent

Sidereal time dependent

Fake data  $\Delta\chi^2$  distribution



Data  $\Delta\chi^2$  is in the middle of the fake data  $\Delta\chi^2$  distribution.  
In other words, the data is consistent with a flat hypothesis.

**No Lorentz violation is seen!**

# Limits on SME coefficients

With no sidereal time dependence apparent, we proceed to set limits on the relevant SME coefficients

$2\sigma$  limits

SME coefficients	$e - \tau$ fit	$e - \mu$ fit
$\text{Re}(a_L)^T$ or $\text{Im}(a_L)^T$	$7.8 \times 10^{-20}$ GeV	—
$\text{Re}(a_L)^X$ or $\text{Im}(a_L)^X$	$4.4 \times 10^{-20}$ GeV	$1.6 \times 10^{-21}$ GeV
$\text{Re}(a_L)^Y$ or $\text{Im}(a_L)^Y$	$9.0 \times 10^{-20}$ GeV	$6.1 \times 10^{-20}$ GeV
$\text{Re}(a_L)^Z$ or $\text{Im}(a_L)^Z$	$2.7 \times 10^{-19}$ GeV	—
$\text{Re}(c_L)^{XY}$ or $\text{Im}(c_L)^{XY}$	$3.4 \times 10^{-18}$	—
$\text{Re}(c_L)^{XZ}$ or $\text{Im}(c_L)^{XZ}$	$1.8 \times 10^{-17}$	—
$\text{Re}(c_L)^{YZ}$ or $\text{Im}(c_L)^{YZ}$	$3.8 \times 10^{-17}$	—
$\text{Re}(c_L)^{XX}$ or $\text{Im}(c_L)^{XX}$	$3.9 \times 10^{-17}$	—
$\text{Re}(c_L)^{YY}$ or $\text{Im}(c_L)^{YY}$	$3.9 \times 10^{-17}$	—
$\text{Re}(c_L)^{ZZ}$ or $\text{Im}(c_L)^{ZZ}$	$4.9 \times 10^{-17}$	—
$\text{Re}(c_L)^{TT}$ or $\text{Im}(c_L)^{TT}$	$1.3 \times 10^{-17}$	—
$\text{Re}(c_L)^{TX}$ or $\text{Im}(c_L)^{TX}$	$5.2 \times 10^{-18}$	—
$\text{Re}(c_L)^{TY}$ or $\text{Im}(c_L)^{TY}$	$1.1 \times 10^{-17}$	—
$\text{Re}(c_L)^{TZ}$ or $\text{Im}(c_L)^{TZ}$	$3.2 \times 10^{-17}$	—

# Neutrino coverage of LV

Current coverage of renormalizable SME coefficients ( $d = 3, 4$ )

(CPT-odd)

$$(a_{\text{eff}}^{(3)})_{jm}^{ab}$$

	$ee$	$e\mu$	$e\tau$	$\mu\mu$	$\mu\tau$	$\tau\tau$
00	Yellow	Orange	Green	Yellow	Blue	Yellow
10	Yellow	Orange	Green	Yellow	Blue	Yellow
11	Yellow	Orange	Green	Yellow	Blue	Yellow
11	Yellow	Orange	Green	Yellow	Blue	Yellow
20	Yellow	Orange	Green	Yellow	Blue	Yellow
21	Yellow	Orange	Green	Yellow	Blue	Yellow
21	Yellow	Orange	Green	Yellow	Blue	Yellow
22	Yellow	Orange	Green	Yellow	Blue	Yellow
22	Yellow	Orange	Green	Yellow	Blue	Yellow

$$(c_{\text{eff}}^{(4)})_{jm}^{ab}$$

(CPT-even)

	$ee$	$e\mu$	$e\tau$	$\mu\mu$	$\mu\tau$	$\tau\tau$
00	Yellow	Orange	Green	Yellow	Blue	Yellow
10	Yellow	Orange	Green	Yellow	Blue	Yellow
11	Yellow	Orange	Green	Yellow	Blue	Yellow
11	Yellow	Orange	Green	Yellow	Blue	Yellow
20	Yellow	Orange	Green	Yellow	Blue	Yellow
21	Yellow	Orange	Green	Yellow	Blue	Yellow
21	Yellow	Orange	Green	Yellow	Blue	Yellow
22	Yellow	Orange	Green	Yellow	Blue	Yellow
22	Yellow	Orange	Green	Yellow	Blue	Yellow



This result

From J. Diaz

- Double Chooz/LSND/MiniBooNE/MINOS ND
- Double Chooz
- MINOS FD/IceCube
- Unexplored



# Physics statements

Phys. Rev. D 86 112009 (2012)

- Double Chooz's second publication disappearance result by itself rules out a number of alternative oscillation models motivated by Lorentz violation.
- Double Chooz finds no evidence of Lorentz violation. We set the first limits on fourteen of the SME coefficients associated with e-tau transitions and set two competitive limits on coefficients associated with e-mu transitions.

NEW!

# A search for neutrino-antineutrino oscillations

(work with J. Diaz, T. Katori, and J. Conrad)

arXiv:1307.5789 [hep-ex] (2013),  
Submitted to Phys. Lett. B

- Lorentz violation can lead to the coupling of neutrinos and antineutrinos.
- Neutrino-antineutrino mixing would lead to enhanced disappearance as well as an unconventional energy dependence of the events.

$$P_{\bar{\nu}_e \rightarrow \bar{\nu}_e} = P_{\bar{\nu}_e \rightarrow \bar{\nu}_e}^{(0)} - P_{\bar{\nu}_e \rightarrow \nu_x}^{(2)}$$

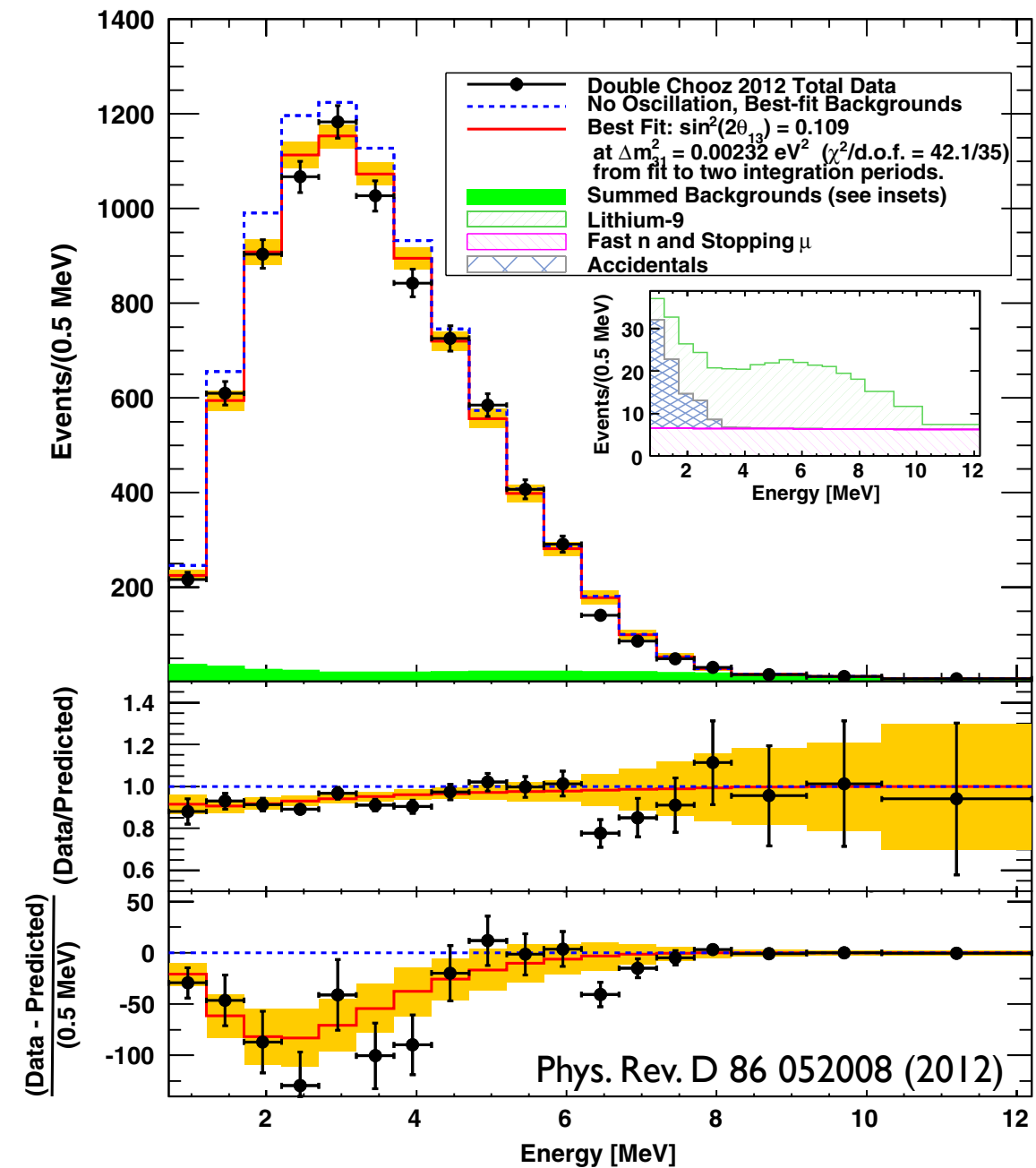
$1 - \sin^2 2\theta_{13} \sin^2(1.267 \Delta m_{\text{atm}}^2 L/E)$   
(conventional mass-based oscillations)

has terms containing E and E<sup>2</sup>

# Double Chooz spectral info

- We explore the possibility that the observed reactor antineutrino disappearance in Double Chooz may have two components: mass-based oscillations and neutrino-antineutrino oscillations due to Lorentz violation.
- Double Chooz has provided a public data release which includes energy spectra data, prediction, and background(s). Covariance matrices and software to analyze the information are also given.

[http://doublechooz.in2p3.fr/Scientific/Data\\_release](http://doublechooz.in2p3.fr/Scientific/Data_release)





# What to fit?

$$P_{\bar{\nu}_e \rightarrow \nu_x}^{(2)} = L^2 \left| \sum_{c=e,\mu\tau} \sum_{\bar{d}=\bar{e},\bar{\mu}\bar{\tau}} (\mathcal{M}_{x\bar{e}}^{(1)})_{c\bar{d}} \delta h_{c\bar{d}} \right|^2$$

factors depend on conventional oscillation parameters as well as experimental parameters including location, orientation, baseline, and energy

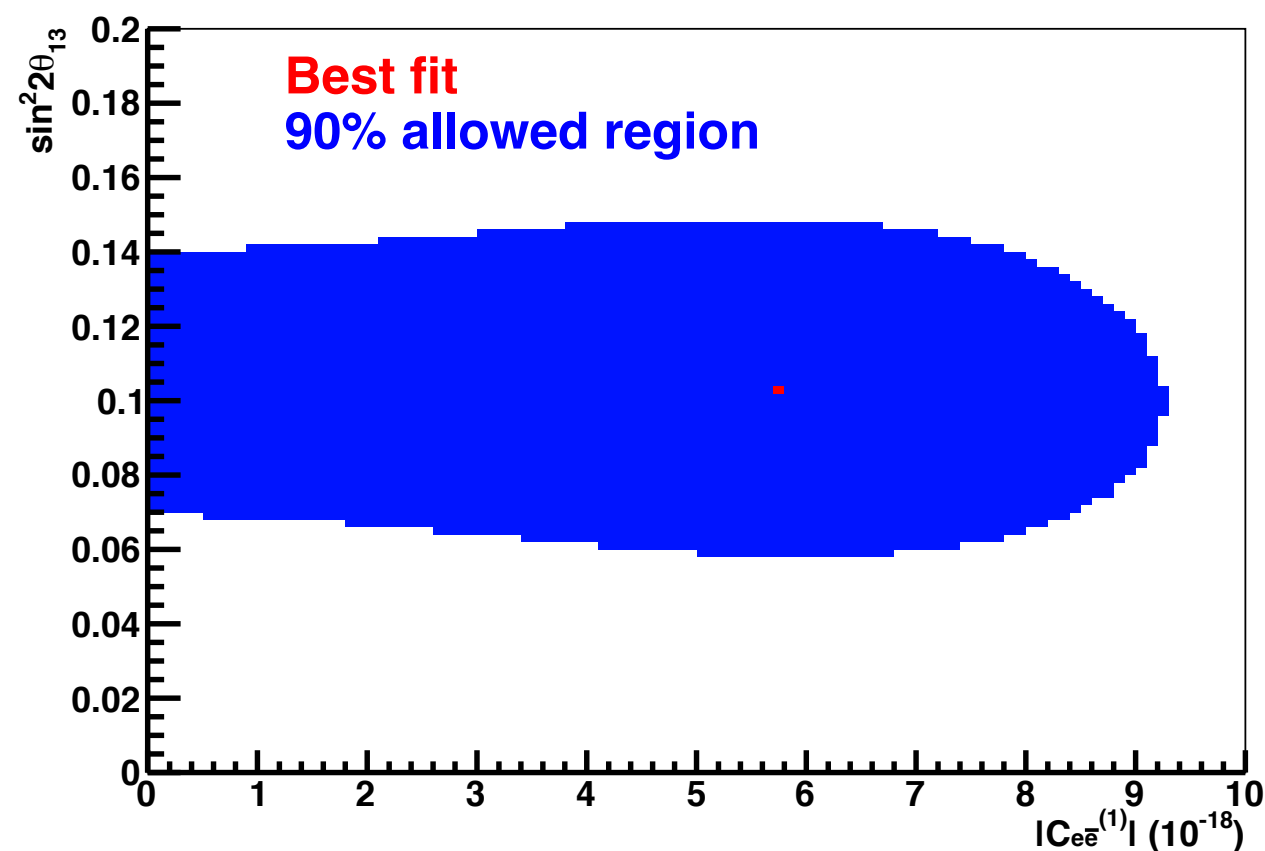
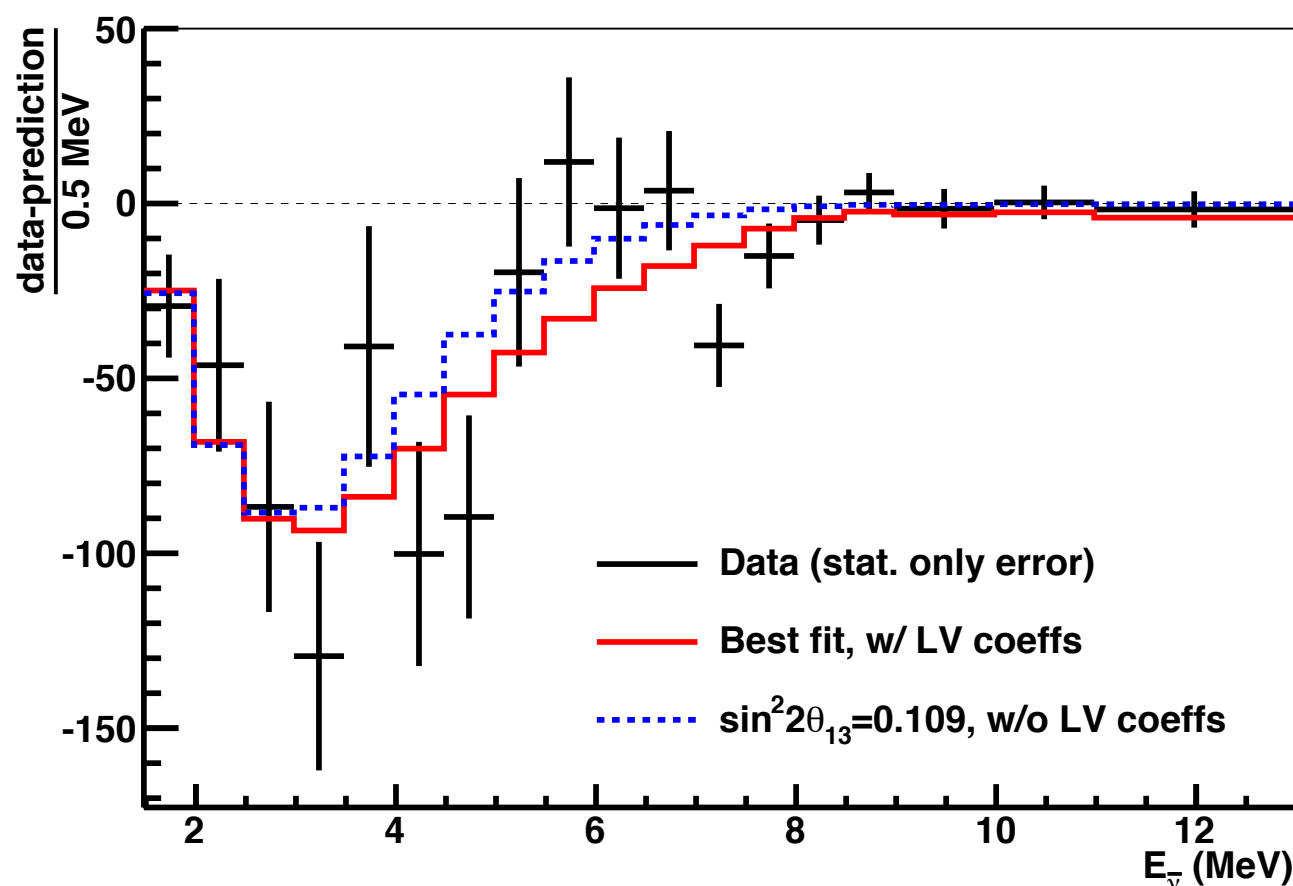
encodes the coefficients controlling Lorentz Violation; parameterized as:

$$\delta h_{c\bar{d}} = C_{c\bar{d}}^{(0)} + C_{c\bar{d}}^{(1)} E$$

A fit to the energy spectrum can potentially distinguish between L/E (mass-based) oscillations and (LV) oscillations that grow with  $E^2$ .

# The fit

- The technique is basically identical to the sidereal-time one. We employ a least squares fitting technique for comparing the Monte Carlo prediction plus background expectation and the data and extracting the best fit parameters.
- After minimization, a confidence region map is formed and then checked with a frequentist study.
- Six different fits are performed, for each of the  $c\bar{d}$  terms.



# Results

- We find no evidence for Lorentz violating neutrino-antineutrino oscillations.
- We set limits on 15 previously unexplored SME coefficients.
- This work completes the coverage of operators in the minimal SME producing neutrino-antineutrino mixing.

—	$ \tilde{g}_{e\bar{e}}^{ZT}  < 9.7 \times 10^{-18}$	$ \tilde{g}_{e\bar{e}}^{ZZ}  < 3.3 \times 10^{-17}$
—	$ \tilde{g}_{\mu\bar{\mu}}^{ZT}  < 2.3 \times 10^{-16}$	$ \tilde{g}_{\mu\bar{\mu}}^{ZZ}  < 8.1 \times 10^{-16}$
—	$ \tilde{g}_{\tau\bar{\tau}}^{ZT}  < 2.3 \times 10^{-16}$	$ \tilde{g}_{\tau\bar{\tau}}^{ZZ}  < 8.1 \times 10^{-16}$
$ \tilde{H}_{e\bar{\mu}}^Z  < 1.4 \times 10^{-19}$	$ \tilde{g}_{e\bar{\mu}}^{ZT}  < 2.7 \times 10^{-17}$	$ \tilde{g}_{e\bar{\mu}}^{ZZ}  < 9.3 \times 10^{-17}$
$ \tilde{H}_{e\bar{\tau}}^Z  < 1.4 \times 10^{-19}$	$ \tilde{g}_{e\bar{\tau}}^{ZT}  < 2.7 \times 10^{-17}$	$ \tilde{g}_{e\bar{\tau}}^{ZZ}  < 9.3 \times 10^{-17}$
$ \tilde{H}_{\mu\bar{\tau}}^Z  < 1.7 \times 10^{-18}$	$ \tilde{g}_{\mu\bar{\tau}}^{ZT}  < 4.4 \times 10^{-16}$	$ \tilde{g}_{\mu\bar{\tau}}^{ZZ}  < 1.5 \times 10^{-15}$

Limits at 90% C.L. for the 15 independent SME coefficients that produce neutrino-antineutrino oscillations. The coefficients for CPT-conserving Lorentz violation  $\tilde{H}_{c\bar{d}}^Z$  are given in units of GeV and the coefficients for CPT-violating Lorentz violation  $\tilde{g}_{c\bar{d}}^{\alpha\beta}$  are dimensionless.

# Last words

- As a community, we have lots of neutrinos and antineutrinos in the book. Let's spend some (perhaps small) fraction of our time (thoroughly) looking for the unexpected.
- Maybe our ~~proton decay~~  $\theta_{13}$  experiment(s) will discover ~~neutrino oscillations~~ Lorentz violation. ~~Supernova~~ Sterile neutrinos wouldn't be so bad either.



# Double Chooz



**Brazil**

CBPF  
UNICAMP  
UFABC



**France**

APC  
CEA/DSM/IRFU:  
SPP  
SPhN  
SEDI  
SIS  
SENAC  
CNRS/IN2P3:  
Subatech  
IPHC  
ULB/VUB



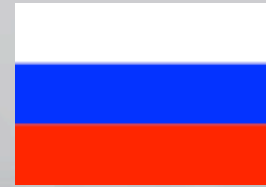
**Germany**

EKU Tübingen  
MPIK Heidelberg  
RWTH Aachen  
TU München  
U. Hamburg



**Japan**

Tohoku U.  
Tokyo Inst. Tech.  
Tokyo Metro. U.  
Niigata U.  
Kobe U.  
Tohoku Gakuin U.  
Hiroshima Inst  
Tech.



**Russia**

INR RAS  
IPC RAS  
RRC Kurchatov



**Spain**

CIEMAT-Madrid

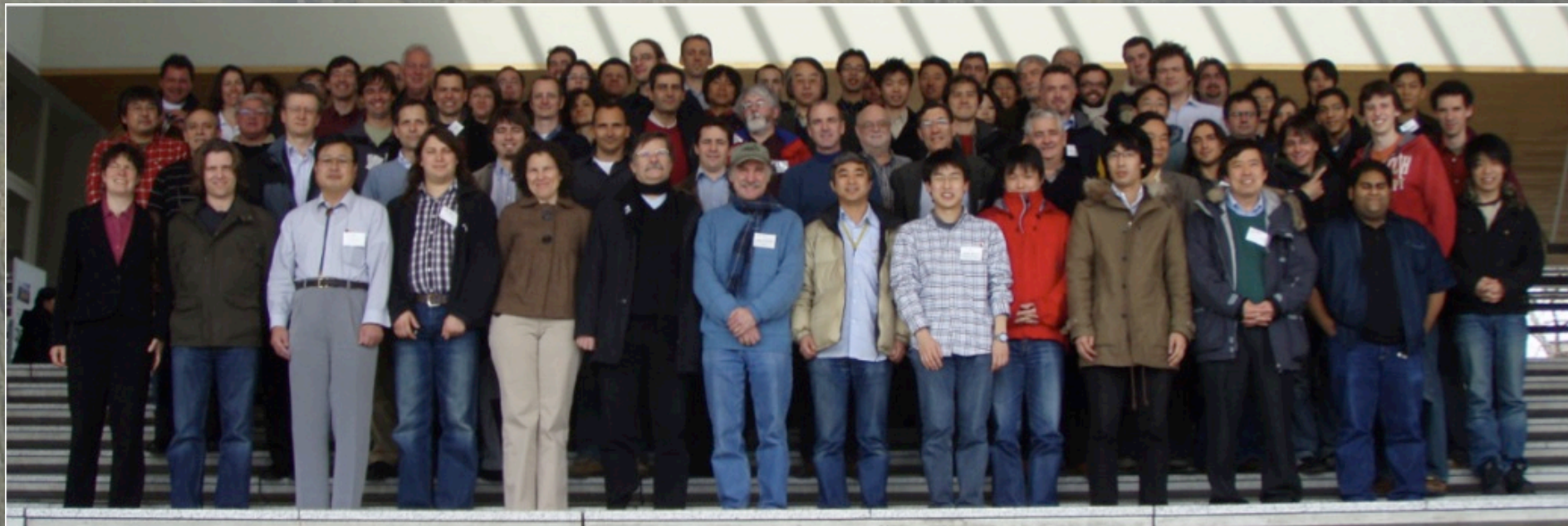


**USA**

U. Alabama  
ANL  
U. Chicago  
Columbia U.  
UCDavis  
Drexel U.  
IIT  
Kansas State  
LLNL  
MIT  
U. Notre Dame  
Sandia National  
Laboratories  
U. Tennessee

Spokesperson: H. de Kerret (IN2P3)  
Project Manager: Ch. Veyssière (CEA-Saclay)

Web Site: [www.doublechooz.org/](http://www.doublechooz.org/)





# $\theta_{13}$ and Lorentz violation

All measured time dependent parameters are consistent with zero. However, the time independent parameter is non-zero at 2.1 sigma.

---

$$P_{\bar{\nu}_e \rightarrow \bar{\nu}_e} \simeq 1 - \frac{L^2}{(\hbar c)^2} ((\mathcal{C})_{e\bar{\tau}}^2 + (\mathcal{C})_{e\bar{\mu}}^2)$$

A normalization-only fit finds  $(\mathcal{C})_{e\bar{\tau}}^2 + (\mathcal{C})_{e\bar{\mu}}^2 = 34.2 \pm 9.2$

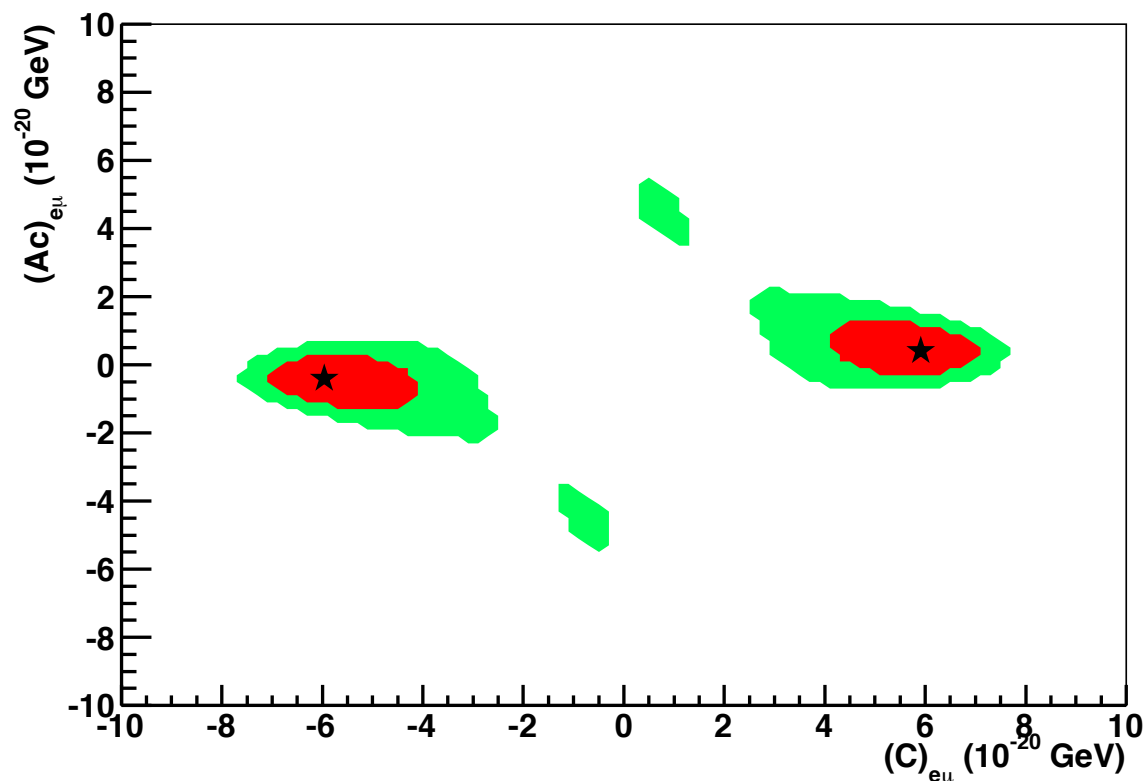
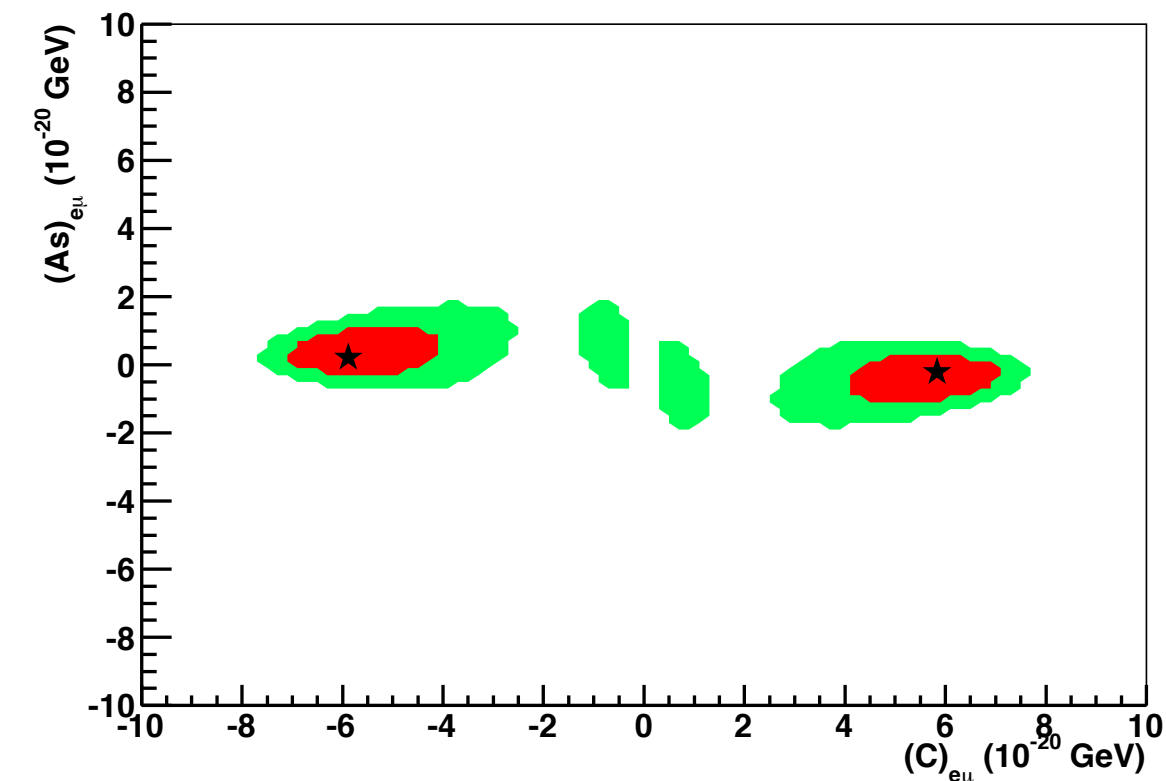
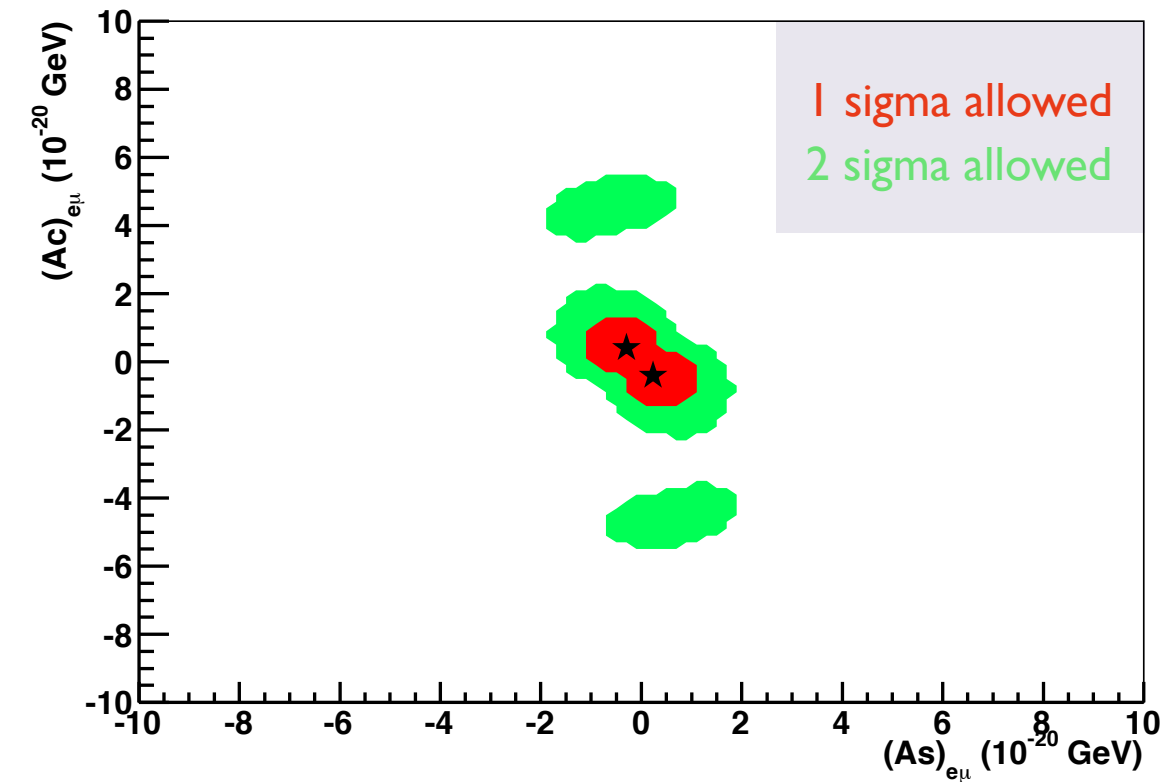
Generally we interpret the sidereal time independent disappearance observed (consistent with the Double Chooz second publication) as due to  $\theta_{13}$  in the 3 flavor neutrino oscillation framework rather than LV.

# Three parameter fit

$$P_{\bar{\nu}_e \rightarrow \bar{\nu}_e} \simeq 1 - \frac{L^2}{(\hbar c)^2} [ |(\mathcal{C})_{\bar{e}\bar{\mu}} + (\mathcal{A}_s)_{\bar{e}\bar{\mu}} \sin \omega_{\oplus} T_{\oplus} + (\mathcal{A}_c)_{\bar{e}\bar{\mu}} \cos \omega_{\oplus} T_{\oplus}|^2 ]$$

The time dependent parameters are consistent with zero.

The time independent parameter is non-zero at 2.1 sigma.



# Rate and (solar) time

IBD rate in solar time (227.9 live solar days)

