



Reweighting the quark Sivers function with STAR jet data

Carlo Flore

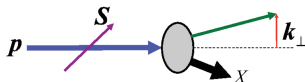
Laboratoire de Physique des 2 Infinis
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**QCD Evolution 2021
(virtual meeting)**

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M. Boglione, U. D'Alesio, CF,
J.O. Gonzalez-Hernandez, F. Murgia, A. Prokudin,
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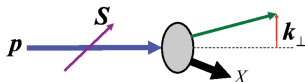
Introduction - the Sivers function



D.W. Sivers, PRD 41 (1990) 83 & PRD 43 (1991) 261

$$\begin{aligned}
 & f_{a/p\uparrow}(x, \mathbf{k}_\perp) - f_{a/p\uparrow}(x, -\mathbf{k}_\perp) \\
 &= \Delta^N f_{a/p\uparrow}(x, k_\perp^2) \frac{(\hat{\mathbf{p}} \times \mathbf{k}_\perp) \cdot \mathbf{S}}{|\mathbf{k}_\perp|} \quad (\text{TO} - \text{CA}) \\
 &= -\frac{2|\mathbf{k}_\perp|}{M} f_{1T}^{\perp a}(x, k_\perp^2) \frac{(\hat{\mathbf{p}} \times \mathbf{k}_\perp) \cdot \mathbf{S}}{|\mathbf{k}_\perp|} \quad (\text{Amsterdam})
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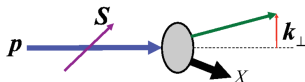


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- genuine TMD distribution

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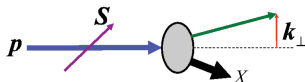


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- express correlation between proton transverse polarization and parton intrinsic $\mathbf{k}_{\perp}$

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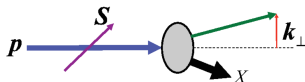
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- express correlation between proton transverse polarization and parton intrinsic $\mathbf{k}_\perp$
- extracted from SIDIS azimuthal asymmetries:

$$A_{UT}^{\sin(\phi_h - \phi_S)} \equiv \frac{F_{UT}^{\sin(\phi_h - \phi_S)}}{F_{UU,T}} = \frac{c[f_{1T}^{\perp a} D_1^q]}{c[f_1^q D_1^q]}$$

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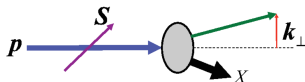


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- in the GPM, Sivers effect generates single spin asymmetries (SSAs) in $p^{\uparrow} p \rightarrow \text{jet } X$

Introduction - formalism

U. D'Alesio, F. Murgia PRD 70 (2004) 074009; M. Anselmino *et al.*, PRD 73 (2006) 014020;
L. Gamberg, Z.-B. Kang, PLB 696 (2011) 109; U. D'Alesio *et al.*, PRD 96 (2017) 036011 ...

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- Inclusive jet production in $p^\uparrow p$ collisions can be described in the **GPM**, where a **factorized formulation in terms of TMDs is assumed as a starting point**
- a **color gauge invariant formulation of GPM (CGI-GPM)** was developed, with inclusion of initial and final state interaction; **process dependence is recovered**
- A_N in $p^\uparrow p \rightarrow \text{jet } X$:

$$A_N \equiv \frac{d\sigma^\uparrow - d\sigma^\downarrow}{d\sigma^\uparrow + d\sigma^\downarrow} \equiv \frac{d\Delta\sigma}{2d\sigma},$$

with

$$d\Delta\sigma^{\text{CGI-GPM}} = \frac{2\alpha_s^2}{s} \sum_{a,b,c,d} \int \frac{dx_a dx_b}{x_a x_b} d^2\mathbf{k}_{\perp a} d^2\mathbf{k}_{\perp b} \left(-\frac{k_{\perp a}}{M_p} \right) f_{1T}^{\perp a}(x_a, \mathbf{k}_{\perp a}) \cos \varphi_a \\ \times f_{b/p}(x_b, \mathbf{k}_{\perp b}) H_{ab \rightarrow cd}^{\text{Inc}} \delta(\hat{s} + \hat{t} + \hat{u})$$

and

$$d\sigma = \frac{\alpha_s^2}{s} \sum_{a,b,c,d} \int \frac{dx_a dx_b}{x_a x_b} d^2\mathbf{k}_{\perp a} d^2\mathbf{k}_{\perp b} f_{a/p}(x_a, \mathbf{k}_{\perp a}) f_{b/p}(x_b, \mathbf{k}_{\perp b}) H_{ab \rightarrow cd}^U \delta(\hat{s} + \hat{t} + \hat{u}).$$

$$[\text{note: } d\Delta\sigma^{\text{GPM}} = d\Delta\sigma^{\text{CGI-GPM}} [H_{ab \rightarrow cd}^{\text{Inc}} \rightarrow H_{ab \rightarrow cd}^U]]$$

The reweighting method

W.T. Giele, S. Keller PRD 58 (1998) 094023; R.D. Ball *et al.*, NPB 849 (2011) 112;
N. Sato, J. Owens, H. Prosper, PRD 89 (2014) 114020; H. Paukkunen, P. Zurita, JHEP 12 (2014) 100

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Never applied to a fit with actual data...

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The reweighting method for TMDs

M. Boglione, U. D'Alesio, CF, J.O. Gonzalez-Hernandez, F. Murgia, A. Prokudin, PLB 815 (2021) 136135

BAYESIAN REWEIGHTING:

- consider a model for a TMD depending on $\mathbf{a} = \{a_1, \dots, a_n\}$ parameters with prior probability distribution $\pi(\mathbf{a})$
[if no prior knowledge exists, $\pi(\mathbf{a})$ is flat]

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- for a set of data $\mathbf{y} = y_1, \dots, y_{N_{\text{dat}}}$, χ^2 is defined as

$$\chi^2[\mathbf{a}, \mathbf{y}] = \sum_{i,j=1}^{N_{\text{dat}}} (y_i[\mathbf{a}] - y_j) C_{ij}^{-1} (y_j[\mathbf{a}] - y_i)$$

[if only uncorrelated uncertainties: $\chi^2[\mathbf{a}, \mathbf{y}] = \sum_{i=1}^{N_{\text{dat}}} (y_i[\mathbf{a}] - y_i)^2 / \sigma_i^2$]

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- parameter sets $\mathbf{a}_k$ produced using Monte Carlo (MC) procedures

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BAYESIAN REWEIGHTING (continued):

- **Posterior density** through **Bayes theorem**:

$$\mathcal{P}(\mathbf{a}|\mathbf{y}) = \frac{\mathcal{L}(\mathbf{y}|\mathbf{a}) \pi(\mathbf{a})}{Z}$$

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- expectation value and variance for an observable (symmetric)

$$\mathbb{E}[\mathcal{O}] \simeq \sum_k w_k \mathcal{O}(\mathbf{a}_k) \quad \mathbb{V}[\mathcal{O}] \simeq \sum_k w_k (\mathcal{O}(\mathbf{a}_k) - \mathbb{E}[\mathcal{O}])^2$$

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- **asymmetric errors** calculated using **rv_discrete** function of Python SciPy;
median as central value, uncertainties at **2 σ CL**

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Let's reweight the Sivers function with STAR data!

Reweighting the Sivers function (I)

M. Boglione, U. D'Alesio, CF, J.O. Gonzalez-Hernandez, JHEP 07 (2018) 148

- quark Sivers function parametrization:

$$\Delta^N f_{q/p^\uparrow}(x, k_\perp) = \frac{4M_p k_\perp}{\langle k_\perp^2 \rangle_S} \Delta^N f_{q/p^\uparrow}^{(1)}(x) \frac{e^{-k_\perp^2 / \langle k_\perp^2 \rangle_S}}{\pi \langle k_\perp^2 \rangle_S}$$

where $q = u, d$, and where $\Delta^N f_{q/p^\uparrow}^{(1)}(x)$ is the Sivers first $k_\perp$ -moment:

$$\begin{aligned} \Delta^N f_{q/p^\uparrow}^{(1)}(x) &= \int d^2 \mathbf{k}_\perp \frac{k_\perp}{4M_p} \Delta^N f_{q/p^\uparrow}(x, k_\perp) \equiv -f_{1T}^{\perp(1)q}(x) \\ &= N_q (1-x)^{\beta_q} \end{aligned}$$

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- 220 data points from SIDIS data, up to $x \lesssim 0.3$

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- quark Sivers function **parametrization**:

$$\Delta^N f_{q/p^\uparrow}(x, k_\perp) = \frac{4M_p k_\perp}{\langle k_\perp^2 \rangle_S} \Delta^N f_{q/p^\uparrow}^{(1)}(x) \frac{e^{-k_\perp^2 / \langle k_\perp^2 \rangle_S}}{\pi \langle k_\perp^2 \rangle_S}$$

where $q = u, d$, and where $\Delta^N f_{q/p^\uparrow}^{(1)}(x)$ is the **Sivers first $k_\perp$ -moment**:

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- 220 data points from SIDIS data, up to $x \lesssim 0.3$
- we adopt $\Delta\chi^2$ for 5 params at 2σ CL, i.e. $\Delta\chi^2 = 11.31$

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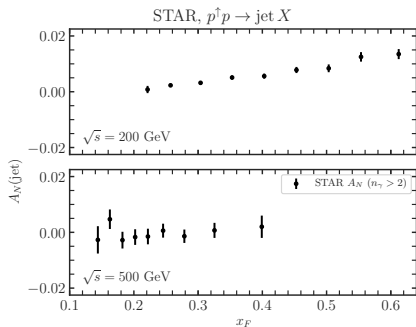
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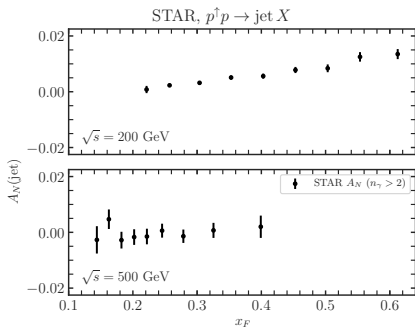
Reweighting the Sivers function (I)

M. Boglione, U. D'Alesio, CF, J.O. Gonzalez-Hernandez, F. Murgia, A. Prokudin, PLB 815 (2021) 136135;
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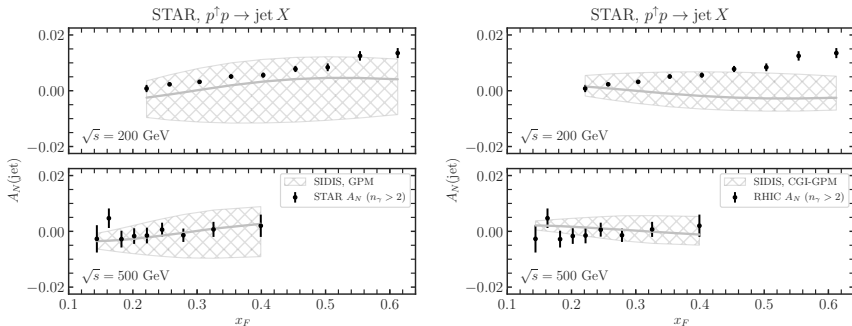
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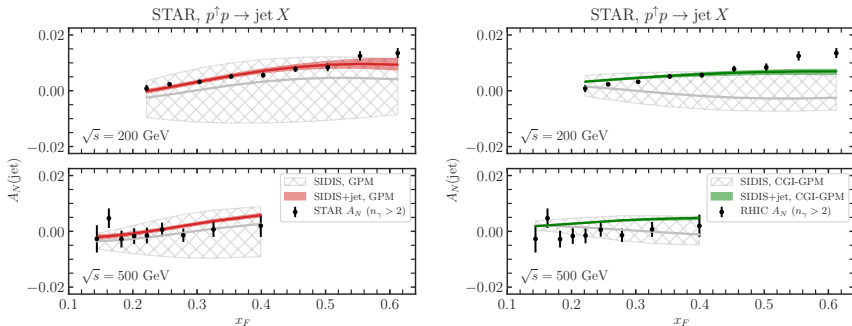
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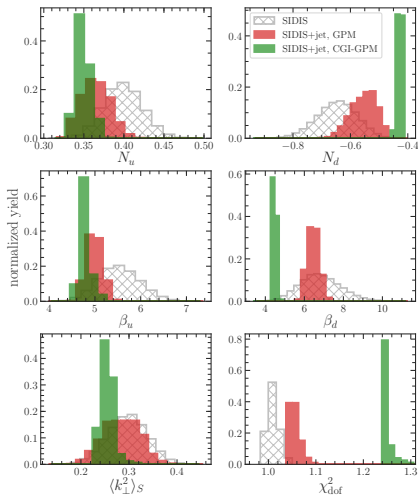
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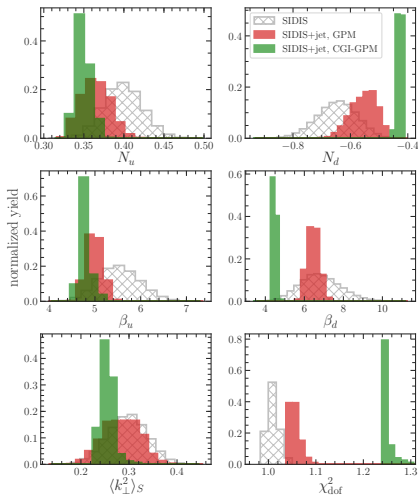
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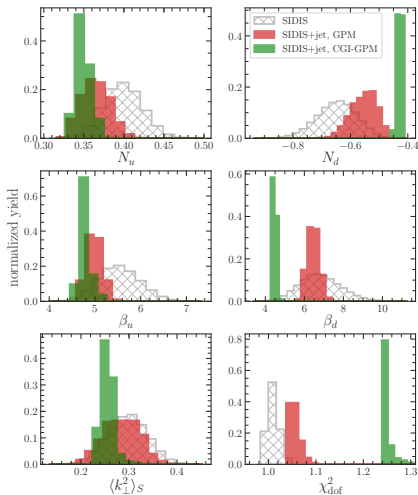


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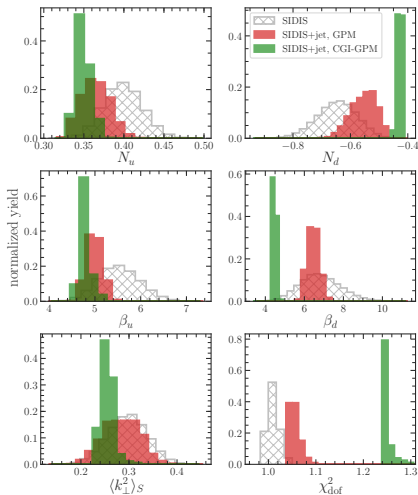


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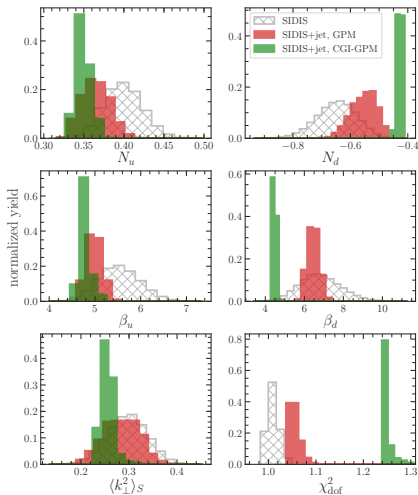


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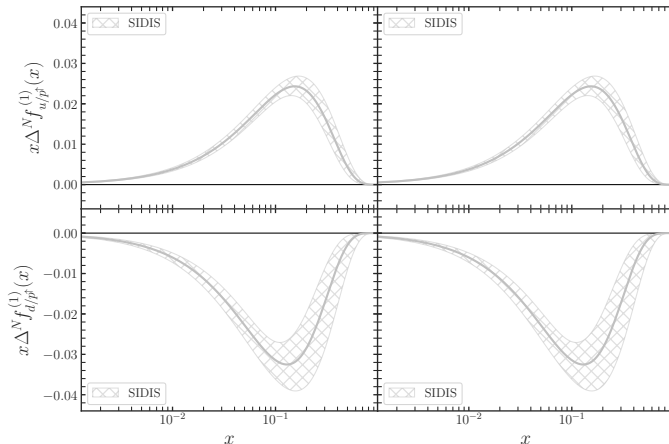


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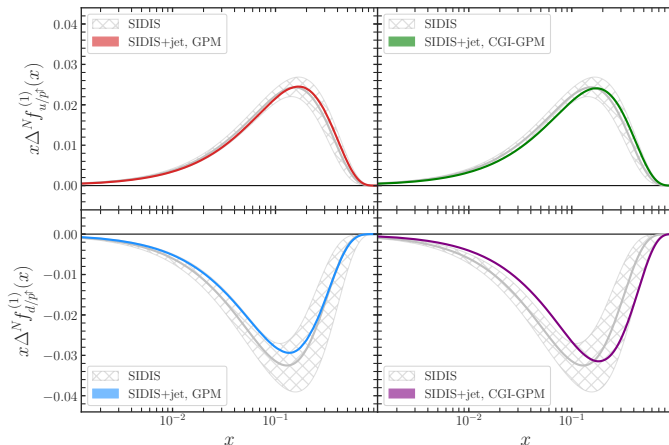
Reweighting the Sivers function (III)

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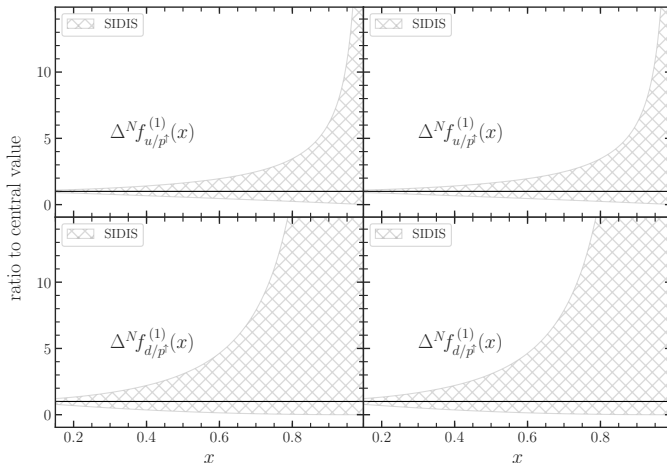
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drastic reduction of uncertainties

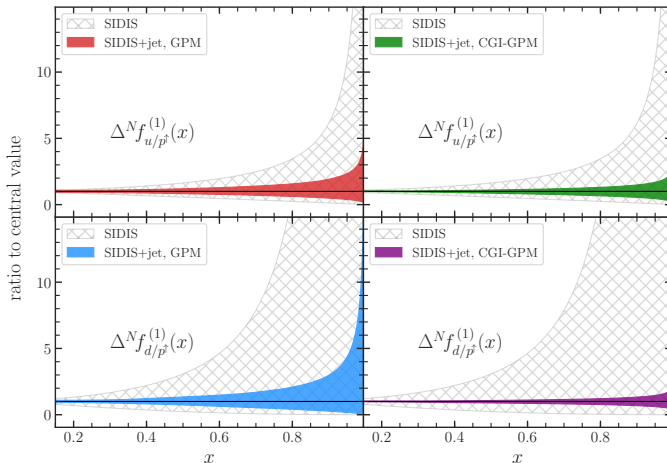
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STAR data allows to constrain the Sivers function at large x !

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- Forthcoming measurements from **COMPASS**, **JLab** and **the EIC** will play **a crucial role in unravelling the nucleon structure in its full complexity**

Thank you

Backup

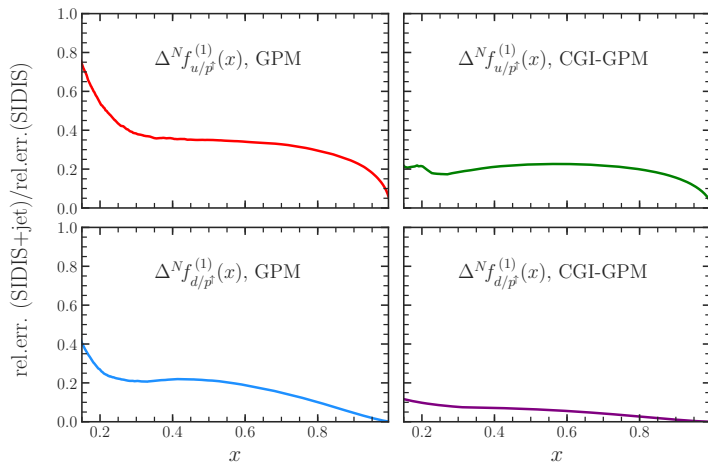
Reweighting the Sivers function - parameters

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	SIDIS	SIDIS+jet, GPM	SIDIS+jet, CGI-GPM
N_u	$0.40^{+0.05}_{-0.04}$	$0.36^{+0.04}_{-0.03}$	$0.35^{+0.02}_{-0.01}$
N_d	$-0.65^{+0.13}_{-0.15}$	$-0.55^{+0.07}_{-0.10}$	$-0.43^{+0.01}_{-0.02}$
β_u	$5.52^{+0.93}_{-0.83}$	$4.98^{+0.34}_{-0.30}$	$4.79^{+0.28}_{-0.19}$
β_d	$6.77^{+2.29}_{-1.85}$	$6.45^{+0.63}_{-0.52}$	$4.48^{+0.17}_{-0.13}$
$\langle k_{\perp}^2 \rangle_S$	$0.30^{+0.08}_{-0.08}$	$0.28^{+0.07}_{-0.07}$	$0.26^{+0.03}_{-0.02}$
χ^2_{dof}	$1.01^{+0.03}_{-0.02}$	$1.05^{+0.03}_{-0.01}$	$1.25^{+0.04}_{-0.01}$

Reweighting the Siverts function - relative errors

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Ratio of relative errors