



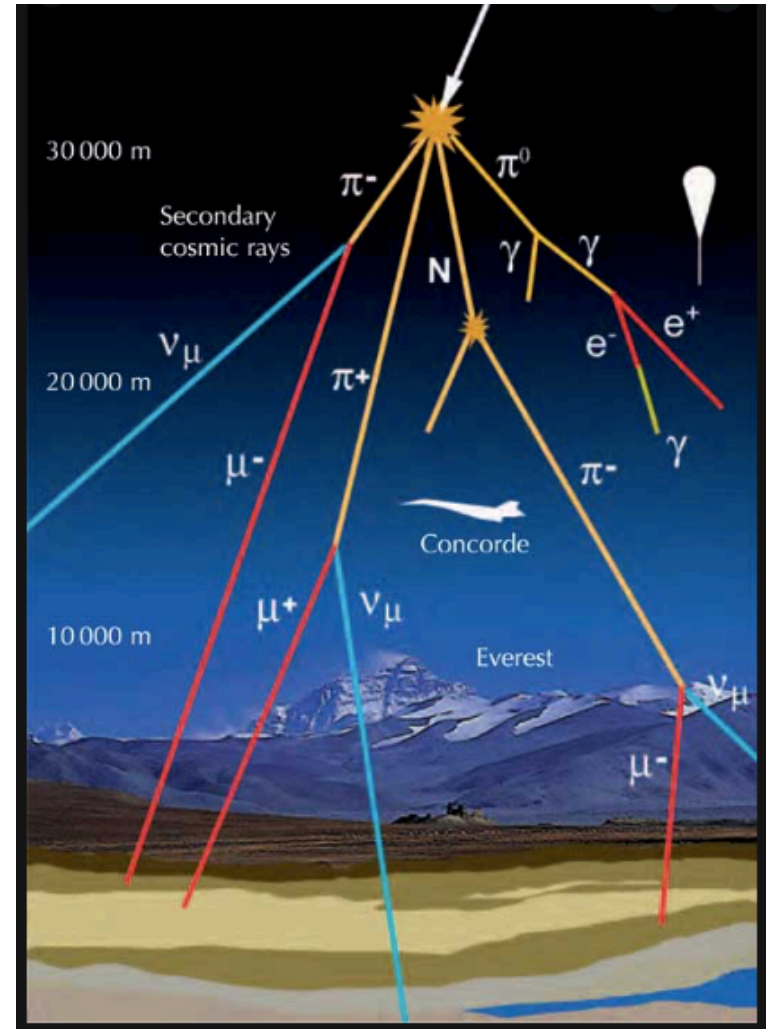
The GSU Neutron Detector

Carola Butler

Georgia State University

Introduction

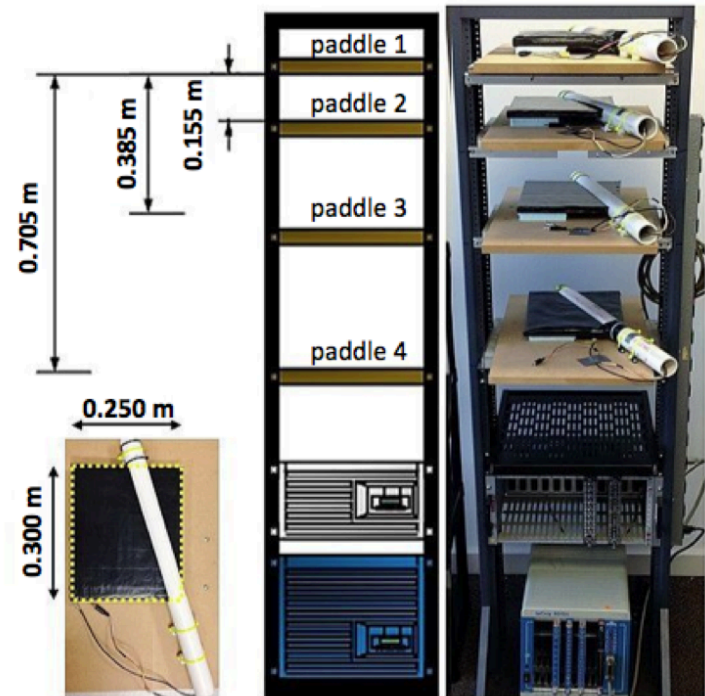
- It has long been known that the number of cosmic ray muons and neutrons reaching the surface of the earth depends on many conditions. We know this depends on:
- Location on earth (geomagnetic rigidity)
- Time of day
- Season of the year
- Solar conditions
 - Solar wind
 - Interplanetary magnetic field magnitude and polarity
 - Sunspot cycle
- Temperature
- Barometric pressure



At Georgia State we have a long history of building cosmic ray detectors for different studies



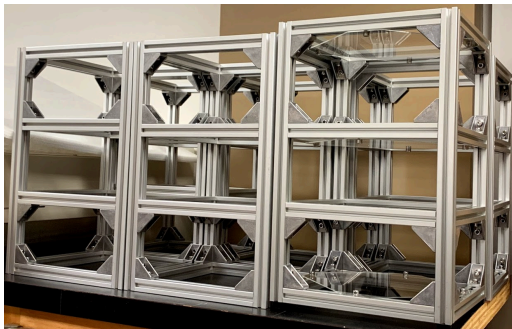
The 'Pot Detector' was originally built to measure muon lifetime. Built from scrap material and a cooking pot, it cost less than \$150. It was later converted to a cosmic ray counter. It has been running for 17 years



The four paddle detector is used to study cosmic rays counts using different acceptance angles,

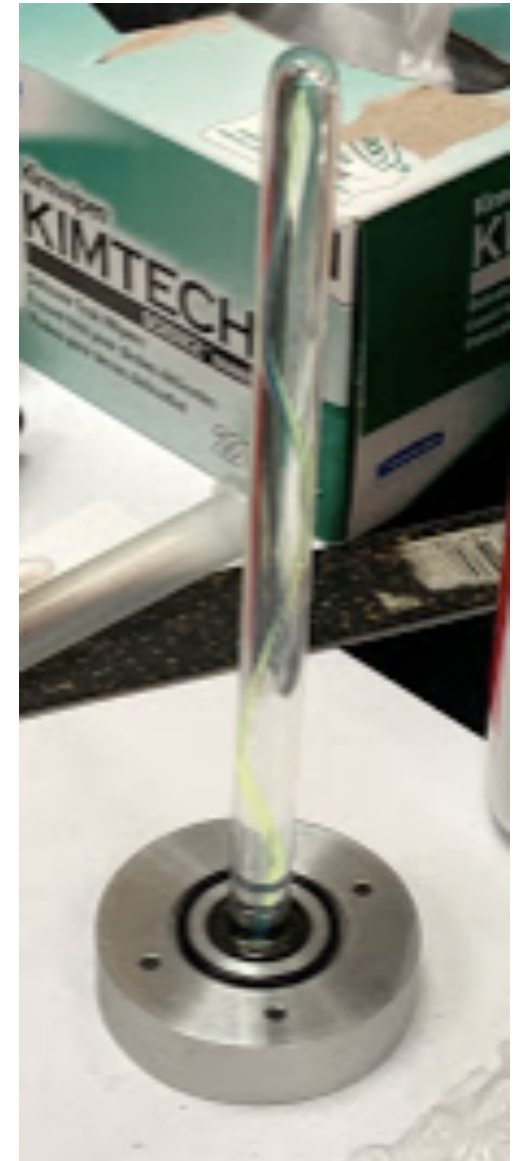
The Muon Telescope

Our latest project is to build muon telescopes that will use SiPm technology and will be portable and relatively inexpensive. The frame will hold plastic scintillator tiles and the acceptance angle can be varied.



When we started building our muon telescopes we wanted a neutron detector that would fit on the frame. The scintillator tiles could be used to help differentiate between charged particles and neutrons. The neutron detector had to be compact, reasonably light and use SiPM's rather than PM tubes.

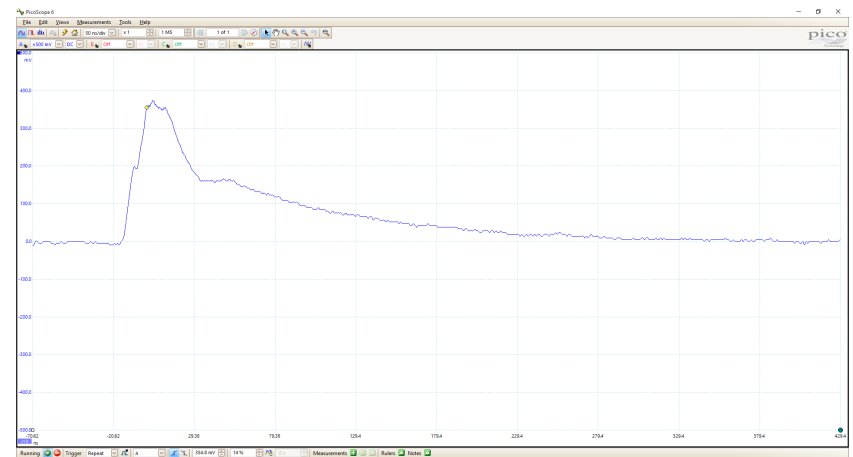
- When we started building the muon telescopes we used the designs used by sPHENIX- plastic scintillator plates with embedded wavelength shifting fibers coupled to SiPM's.
- The only practical way to detect neutrons is to use containers full of liquid scintillator.
- The problem was then to find a way to couple the SiPM to a tank full of scintillation liquid.
- Some people have tried immersing the fibers in liquid scintillator, but this has two problems. Some liquids are very chemically aggressive and the fibers are not rigid and would drift around in the tank.
- Putting the fibers into glass tubes was the solution.
- For optical coupling the tube is filled with optical grease.



The Prototype



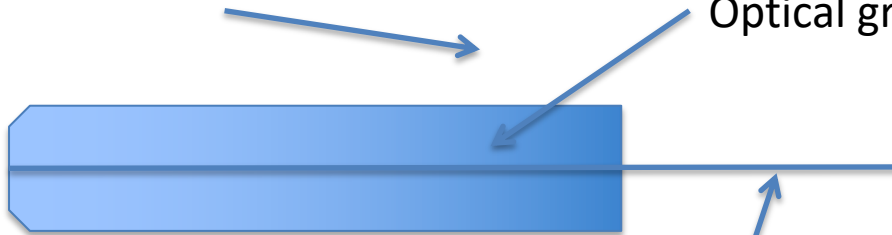
The prototype was made from a water bottle a student left in the lab and never came back for. When it was connected to an SiPM it gave a beautiful signal.



The Light Collection Tube

Bicron 505 $n = 1.505$

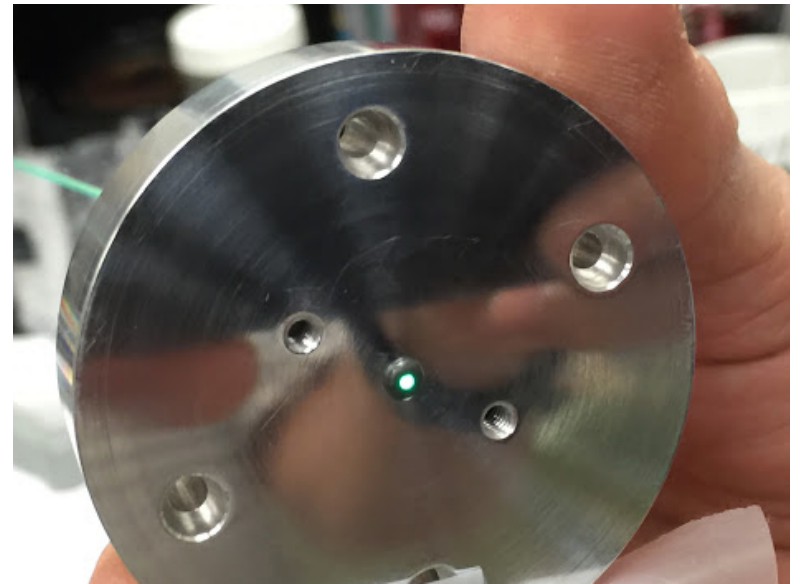
Optical grease, $n=1.465$



Borosilicate glass,
 $n=1.517$

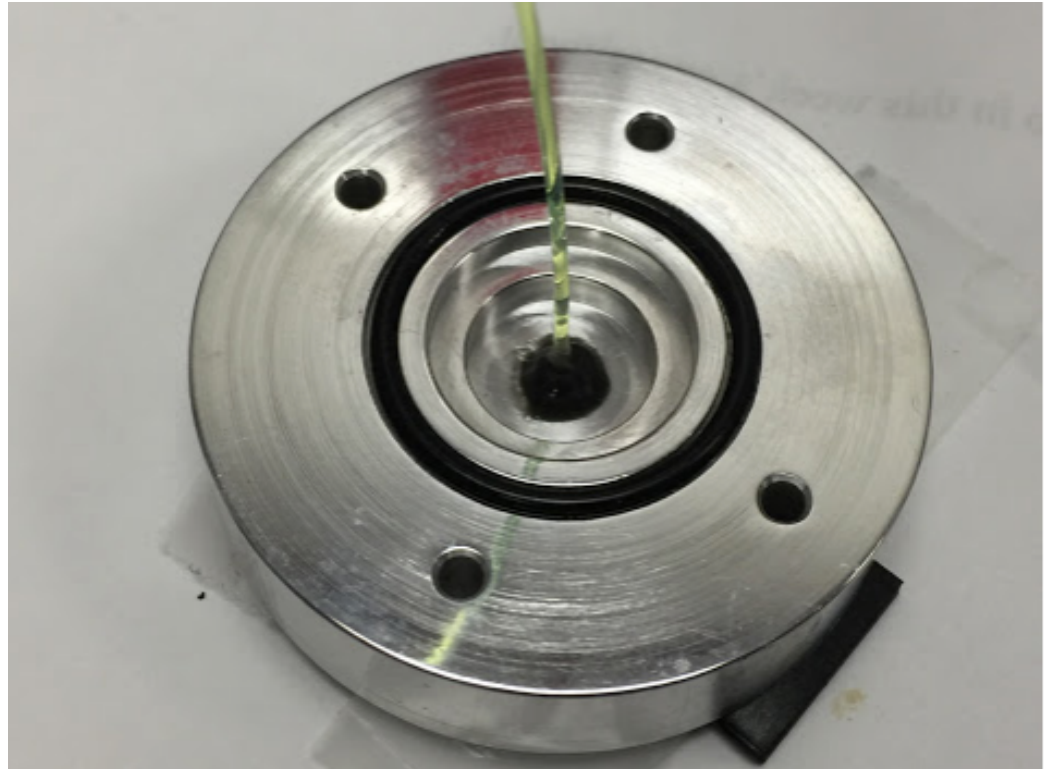
Fiber, $n=1.59$
Cladding, $n=1.49$

The differences in refractive indices is fairly small.



The End Cap

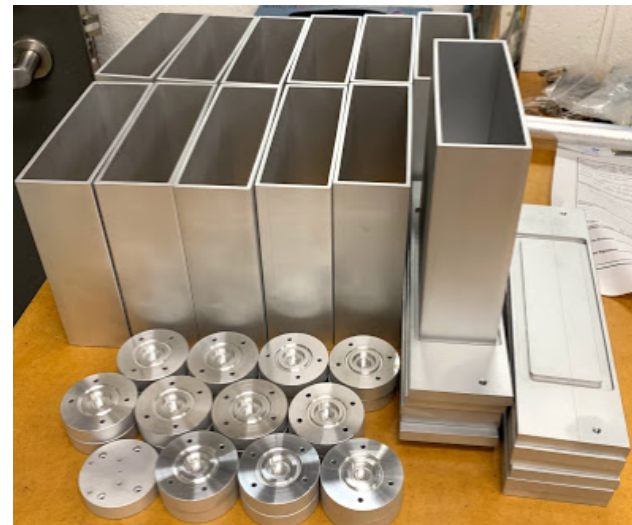
The most important part of the assembly is the end cap. It has to both hold the fiber and the grease tube, prevent the tank from leaking and provide a place to securely attach the SiPM holder.



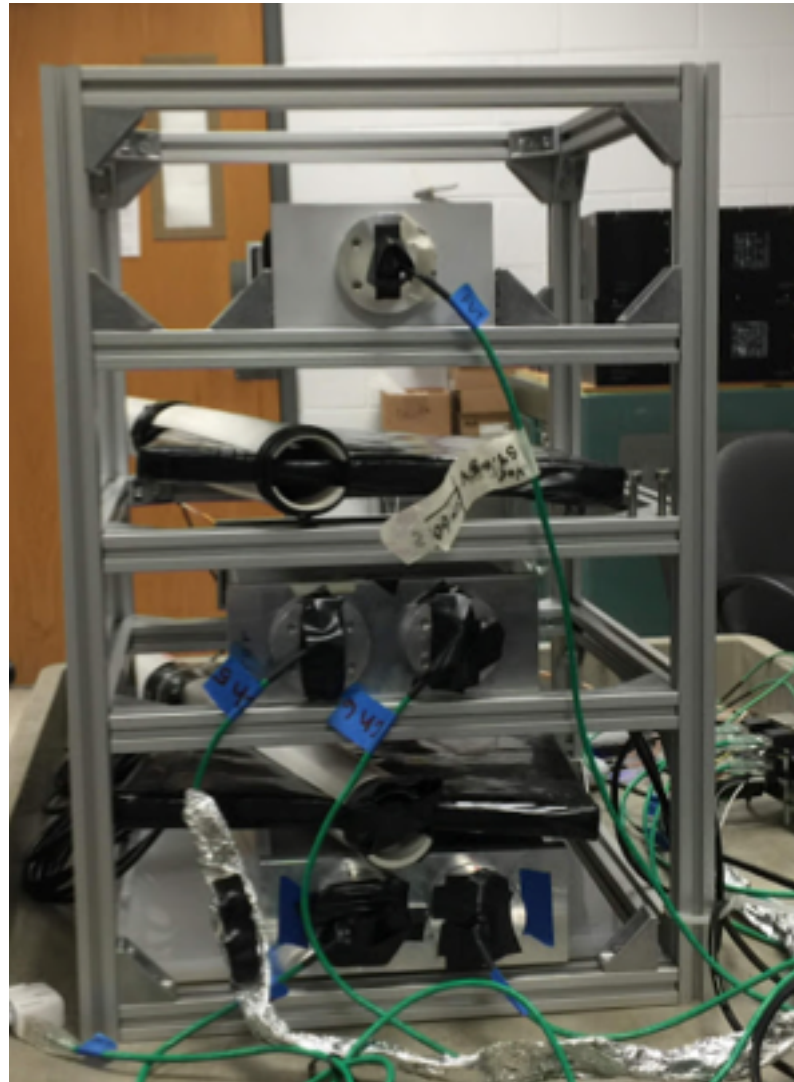
The Mark II-Mark V



In principle any container would work. Our detectors are made from 6061 aluminum which is inexpensive, comes in many sizes, and is easy to machine. This is 6" x 2" standard tubing. We cement the parts together using Araldite 2014 epoxy which is very strong, not electrically conductive and chemical resistant.



The Mark II model had only one tube and SiPM. However, SiPM's are very noisy (about 300,000 dark counts/second) so all our subsequent models have two SiPM's and rely on coincidence to declare a hit.

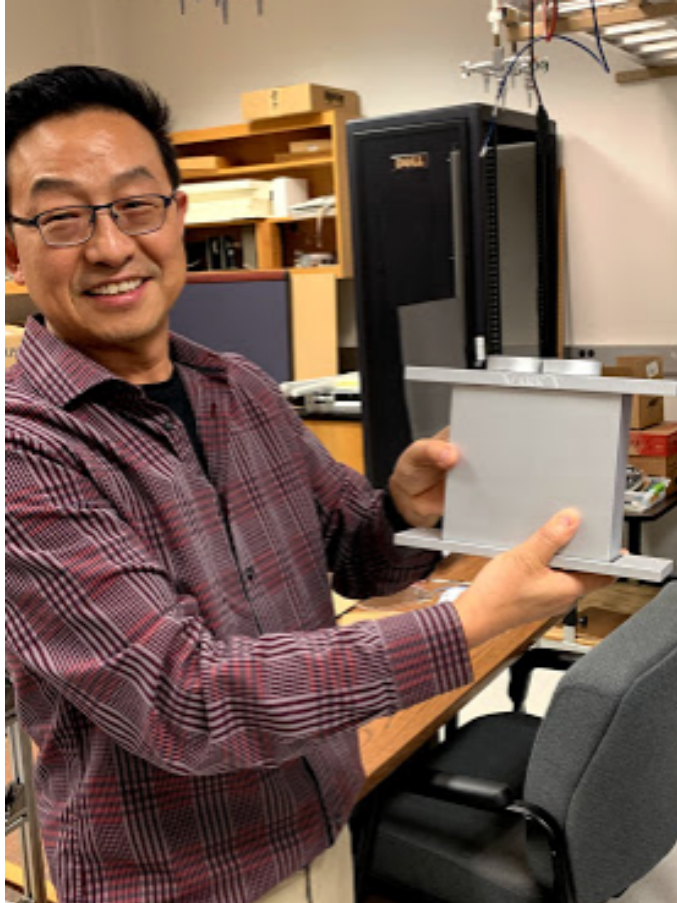


Mark II

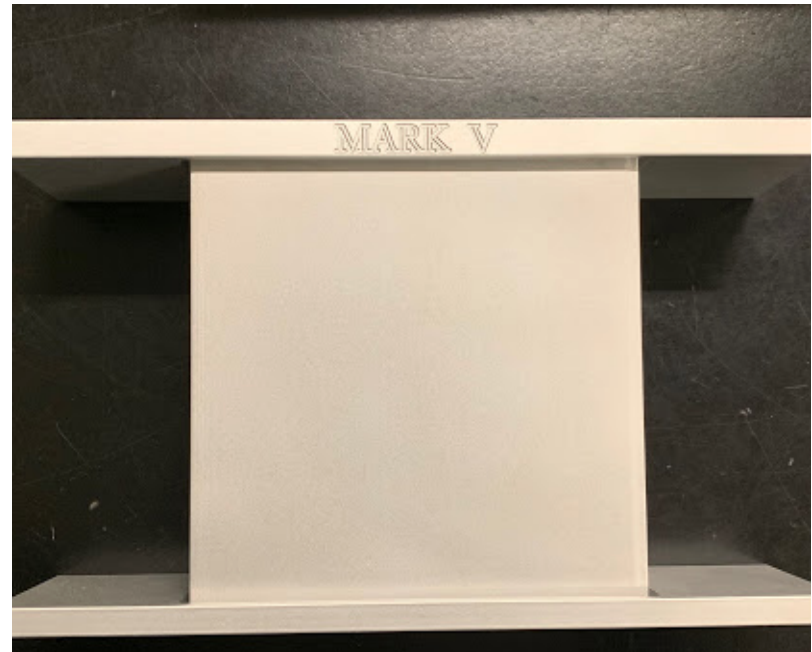
Mark III

Mark IV

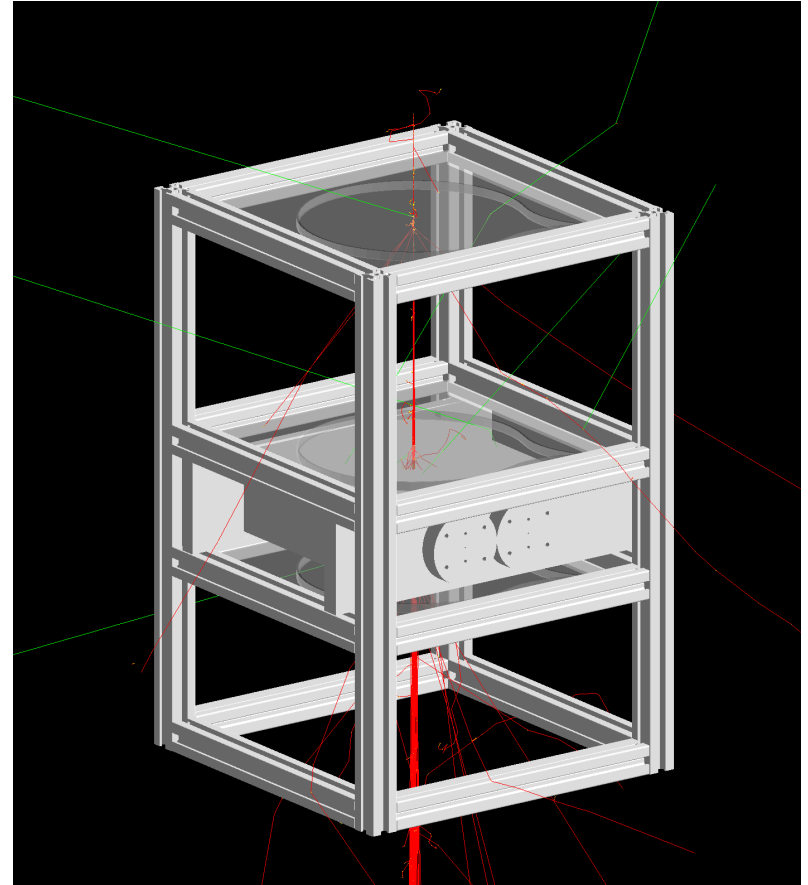
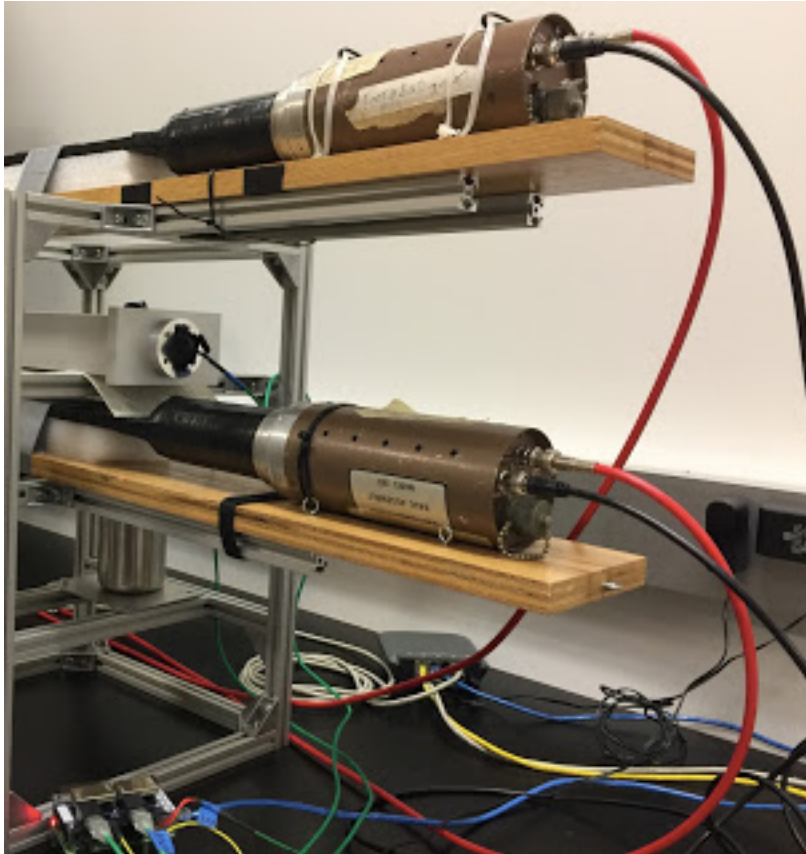
Current design



The Mark V is six by six by two inches and has two end caps. The flanges on the side let us bolt it securely to the muon telescope frame.

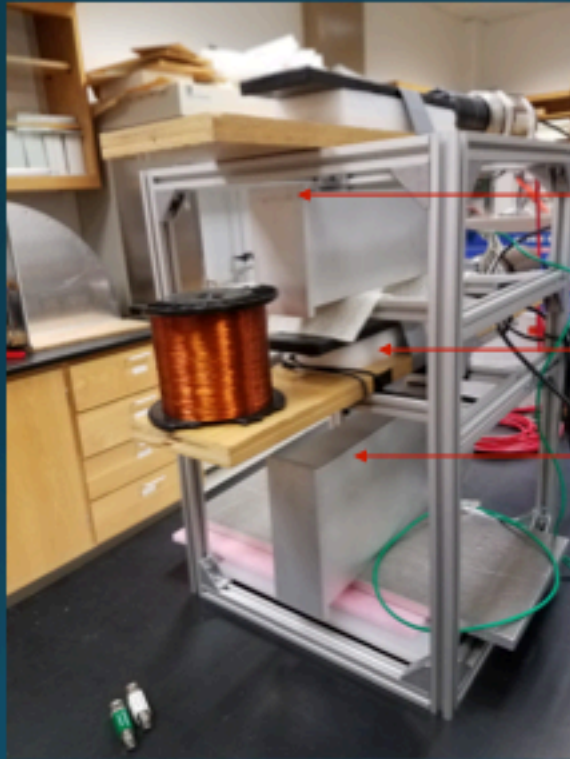


Performance



Test set up by Montgomery Steele

Initial Detection of Cosmic Ray Neutrons



Neutron detector

Muon Detector

Neutron detector

One plastic detector placed between two liquid scintillator neutron detectors.

Cosmic rays must pass through both neutron detectors

Then if the muon detector sees nothing, the cosmic ray is a neutron.

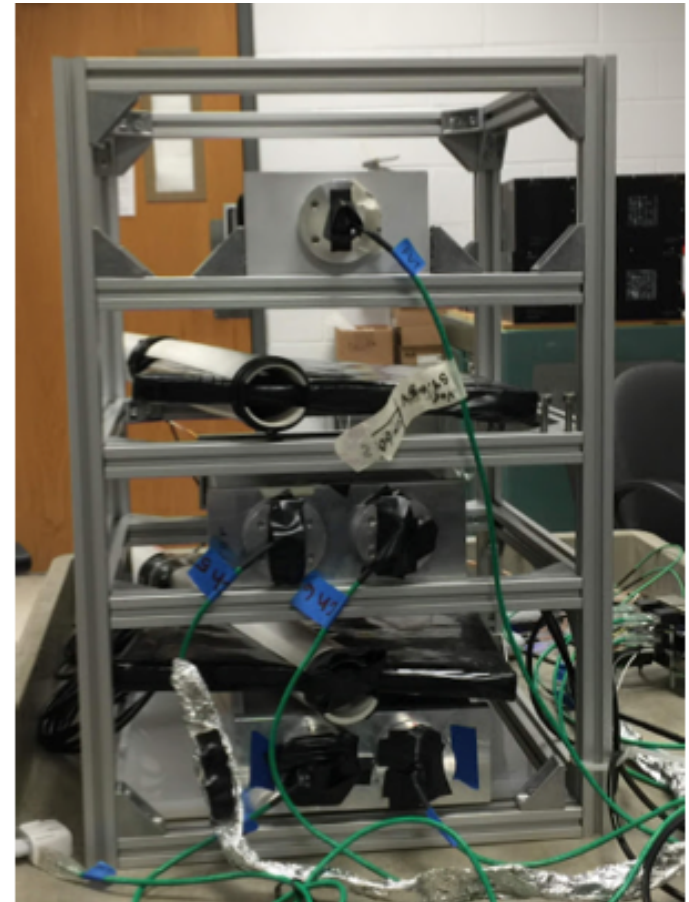
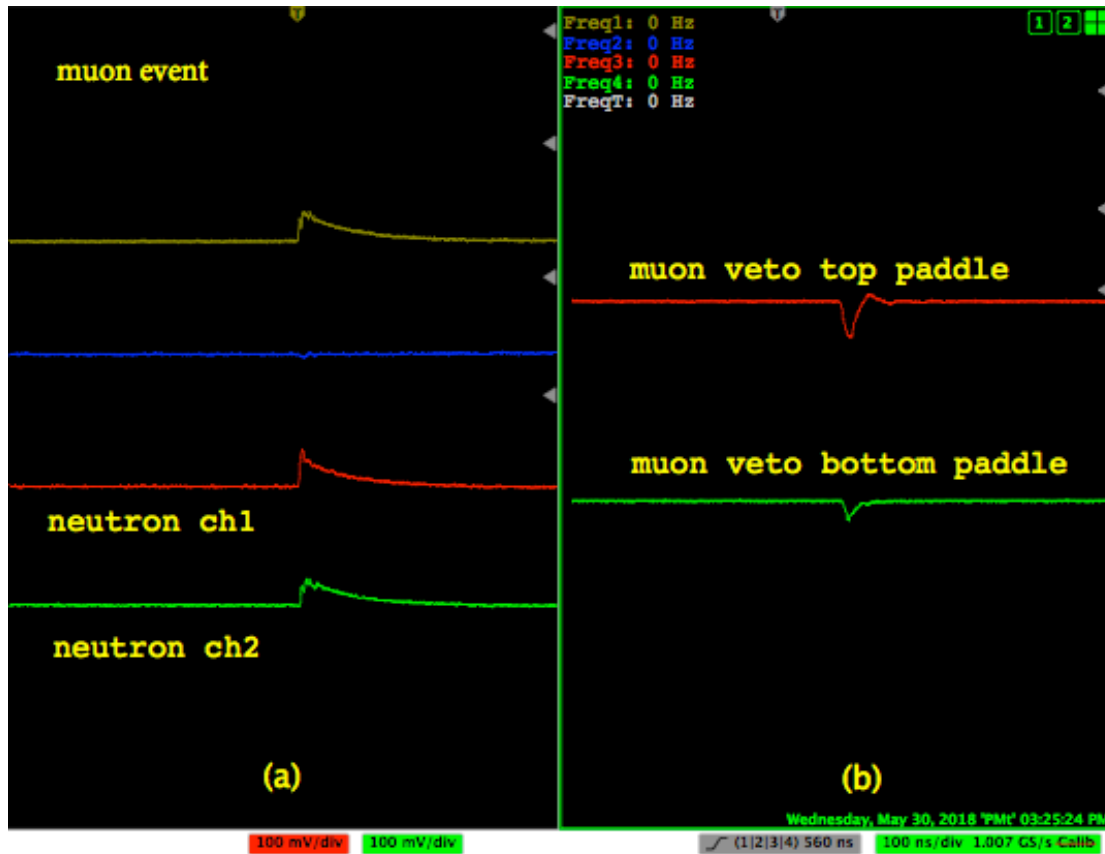
10/3/19

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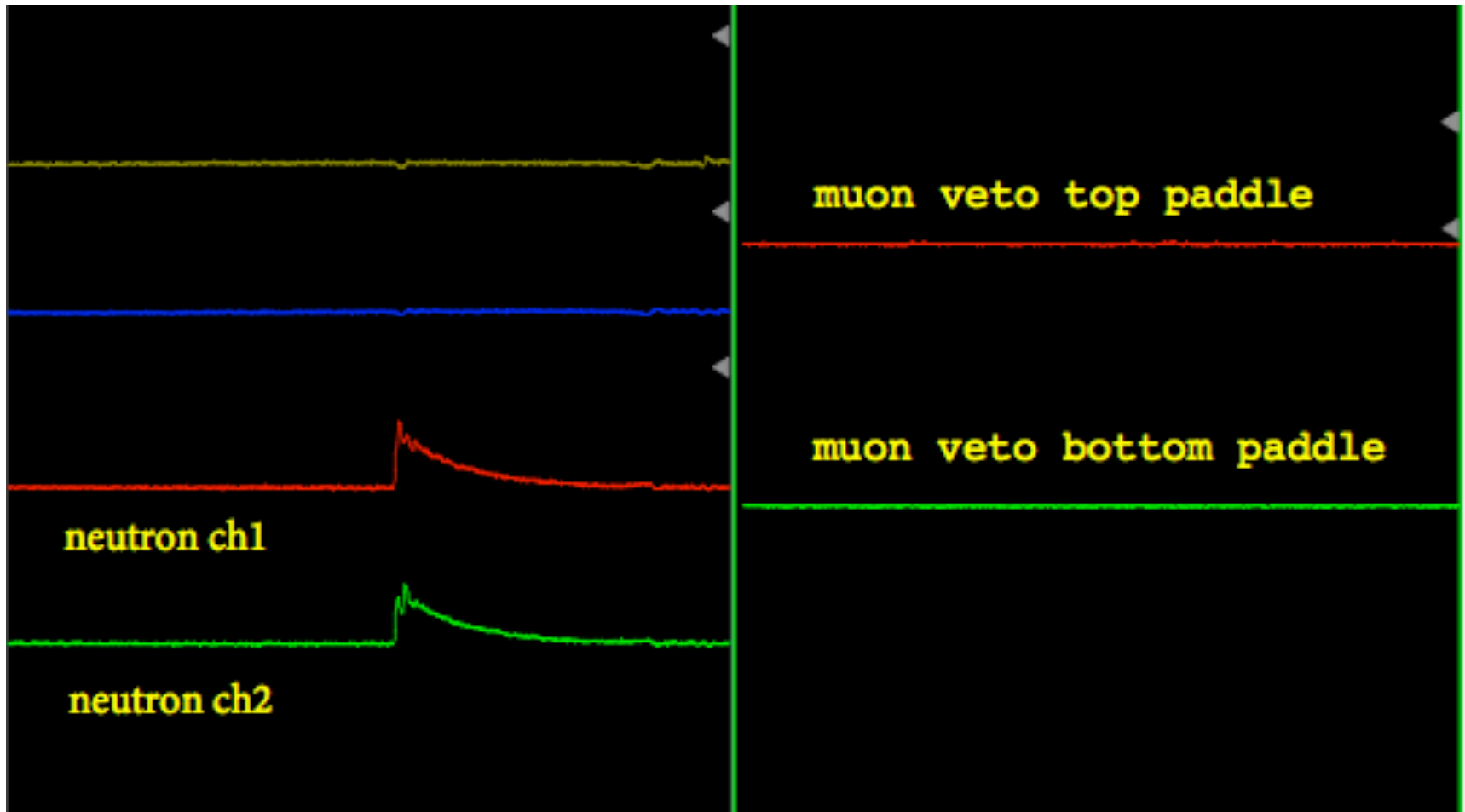
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Slide courtesy of Montgomery Steele

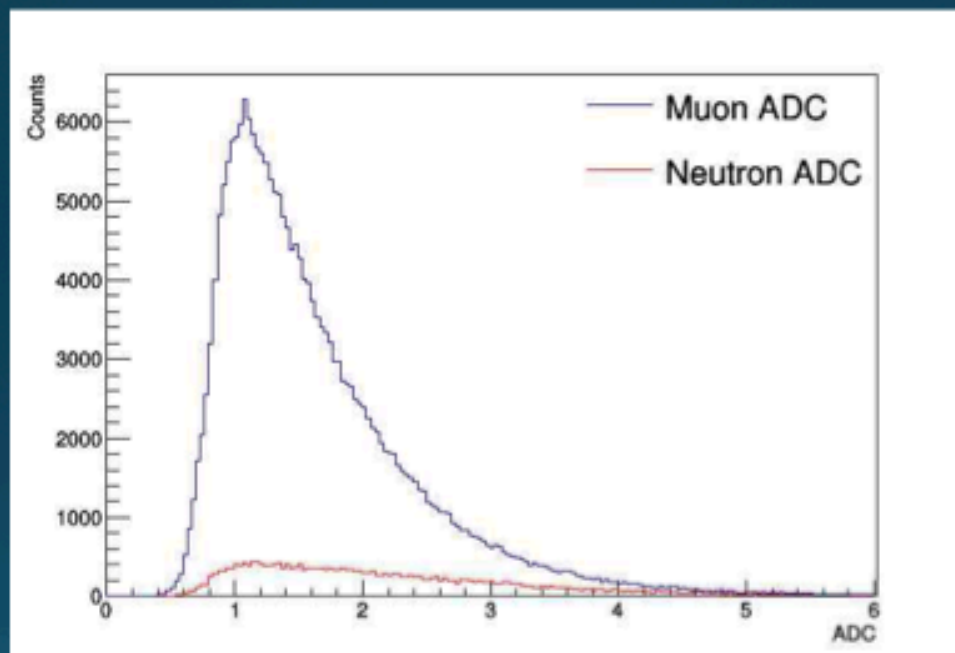
Muon Signals



Neutron Signal



Neutron detector ADC spectrum



As expected, cosmic rays comprise ~10% of our detected events!

10/3/19

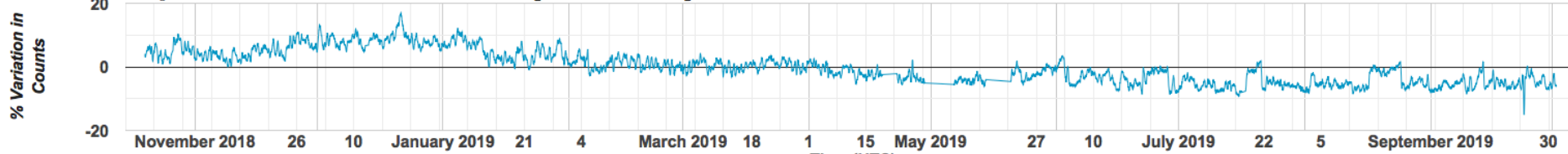
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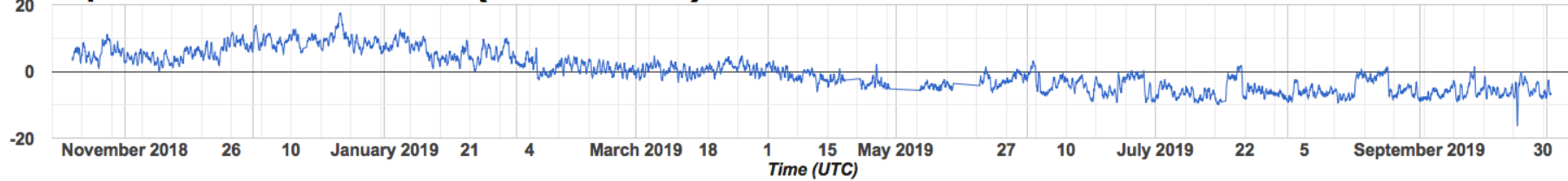
Slide courtesy of Montgomery Steele

Online monitoring

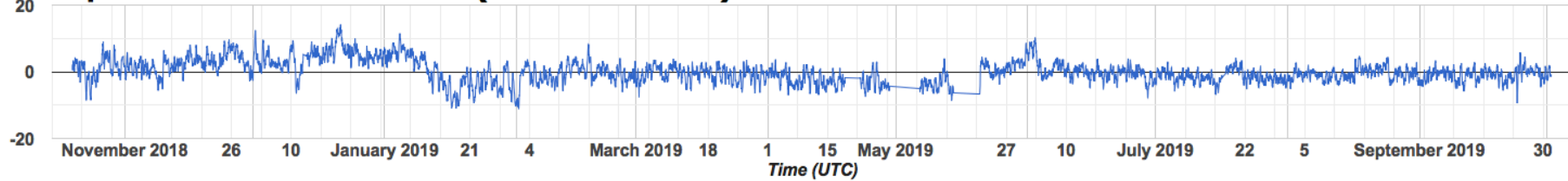
Liquid scintillator test stand (all counts)



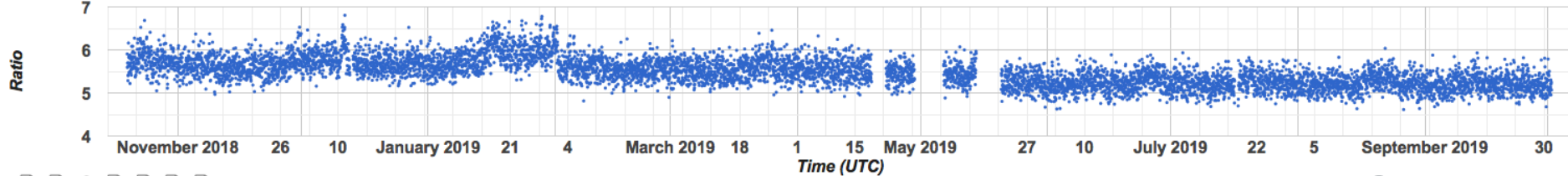
Liquid scintillator test stand (muon counts)



Liquid scintillator test stand (neutron counts)

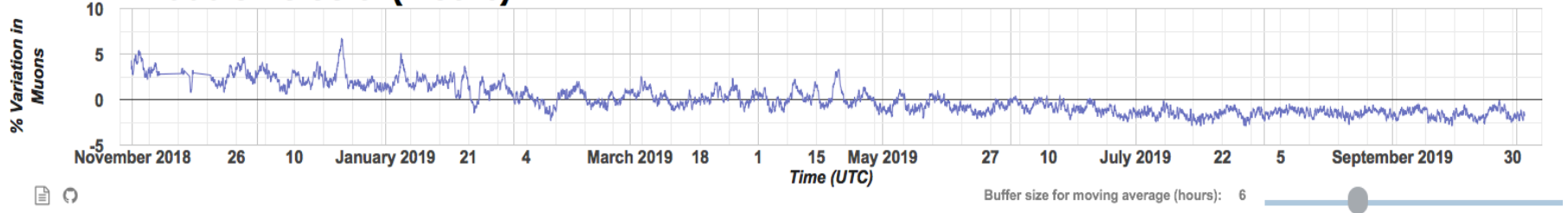


Ratio of Muon counts to Neutron counts



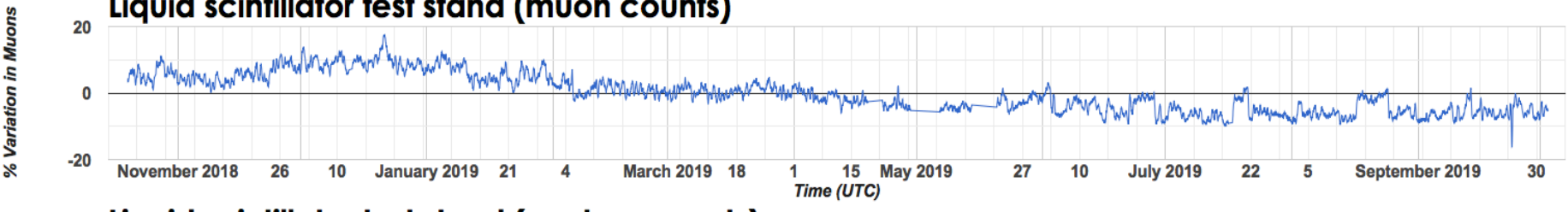
Comparison with Other Detectors

4-Paddle Detector (muons)



This data is taken from a cosmic ray detector test stand in Atlanta, GA. It is located in 1 Park Place, Rm. 706. The graph shows the % variation in cosmic rays detected per hour using a running average. The span of the running average can be adjusted using the slider. The detectors in this test stand use a scintillating plastic which can detect cosmic ray muon events.

Liquid scintillator test stand (muon counts)



<http://phynp6.phy-astr.gsu.edu/~cosmic/dynamicDisplay>

Future work

- Improve signal by adding more fibers to the grease tube
- Calibrate ADC using known neutron energies
- Build more detectors

Deep bows of Gratitude

Dr. Xiaochun He

Dr. Murad Sarsour

Dr. Megan Connors

For letting me play with their group

Peter Walker

Dwayne Torres

Sam Mayberry

*Our instrument shop people for their
endless patience and help*

Jonathan He

For help with drawings

Montgomery Steele

For testing and data analysis

Sawaiz Syed

For electronics and rendering help

The entire Nuclear and
Particle Physics Group –Past
and Present Members





And Thank You for Your Attention