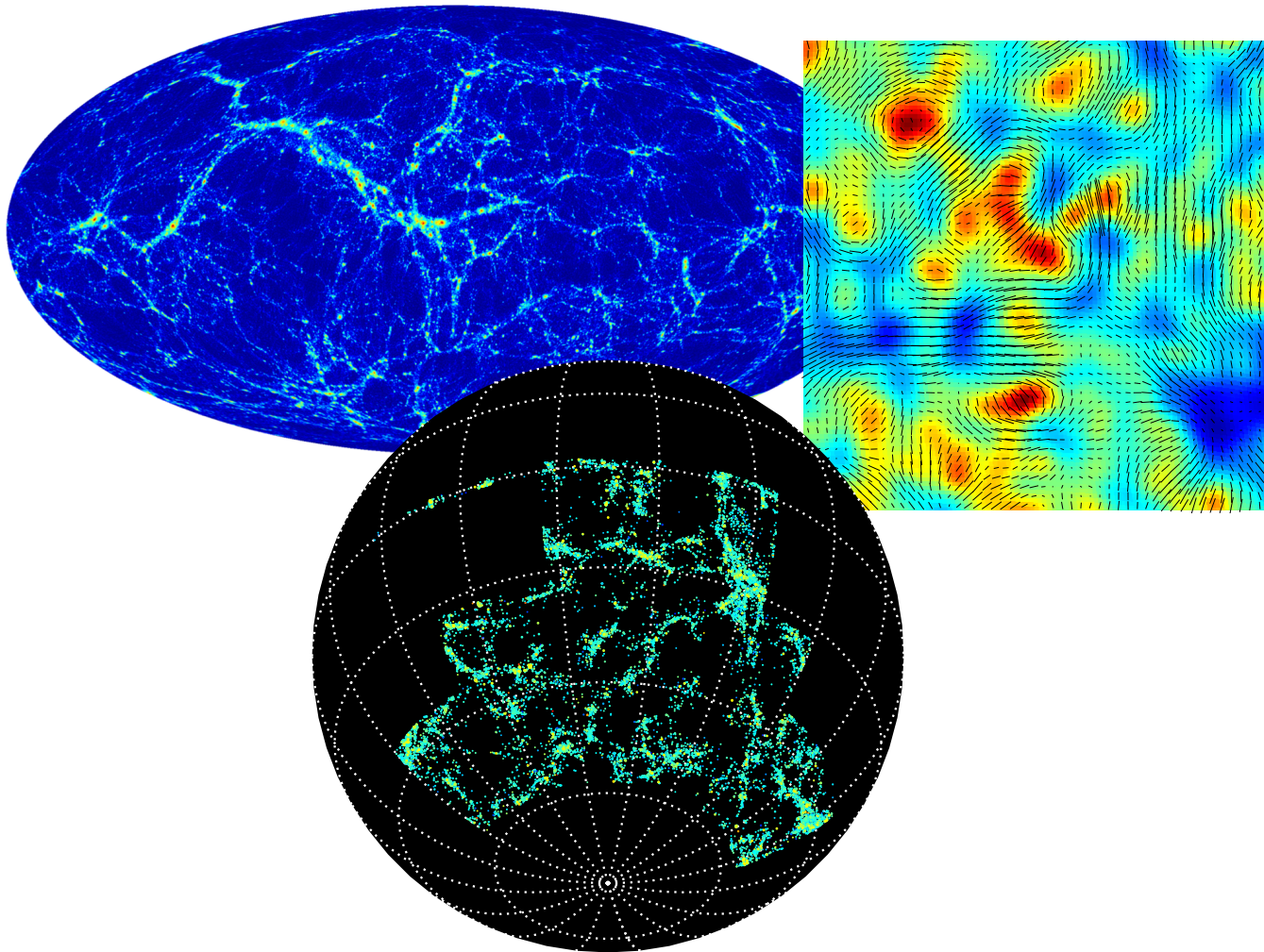


Simulated Galaxy Catalogs



Risa Wechsler

Stanford/SLAC/KIPAC

with Matt Becker,
Michael Busha
& DES sim working group

+

Yu Lu (SAMs)

+

Rachel Reddick

Eli Rykoff

(validation)

+

Hao-Yi Wu

Oliver Hahn

Sam Skillman

(cluster resims)

Strategies

- high resolution: associate all galaxies with resolved halos and subhalos.
 - assign luminosities using abundance matching + galaxy properties based on environment
 - extensive testing against data from SDSS at low z , including correlation functions, group statistics, galaxy-galaxy lensing, etc.
 - need very high resolution, e.g. \sim kpc force resolution and $1e8$ mass resolution to resolve $M_r = -19$ galaxies.
 - currently have/creating catalogs based on various boxes with ~ 150 -600 Mpc
 - SAM models on the same merger trees using model of Yu Lu, further development informed by empirical results
- medium resolution: minimum needed, in order to produce multiple sky surveys in many cosmologies
 - associate all galaxies with dark matter overdensities + central galaxies where halos are well resolved
 - well developed pipeline:
 - simulation lightcone
 - galaxy luminosities
 - SEDs for galaxies
 - shear at every galaxy position (current version, 6.2" resolution)
 - galaxies lensed / sheared & magnified
 - photometry in many bands, including LSST bands and DES, SDSS, VISTA, CFHTLS, NDWFS, DEEP, WISE, IRAC
 - photometric errors & photometric redshifts

Large area simulations BCC-Aardvark/Buzzard-v1.0

■ currently available simulations:

- LCDM cosmology; N-body lightcones to $z \sim 2$ (based on 3 sim boxes with 2048^3 particles)
- one additional cosmology (currently blind parameters for DES Blind Cosmology Challenge)
- halo finding from rockstar, includes multiple mass def., concentrations, etc.
- ~ 1 billion galaxies added using addgals, over $1/4$ sky (10313 sq. degrees), complete to $i \sim 25$
- photometry in many bands, including LSST bands and DES, SDSS (DR8+S82), VISTA (VHS +VIKING), CFHTLS, NDWFS, DEEP, WISE, IRAC

■ shear on the full quarter of sky using CALCLENS; currently with 6.2" resolution

■ extensive development and testing with SDSS data and other higher redshift data, including early DES data; designed to go to full DES depth

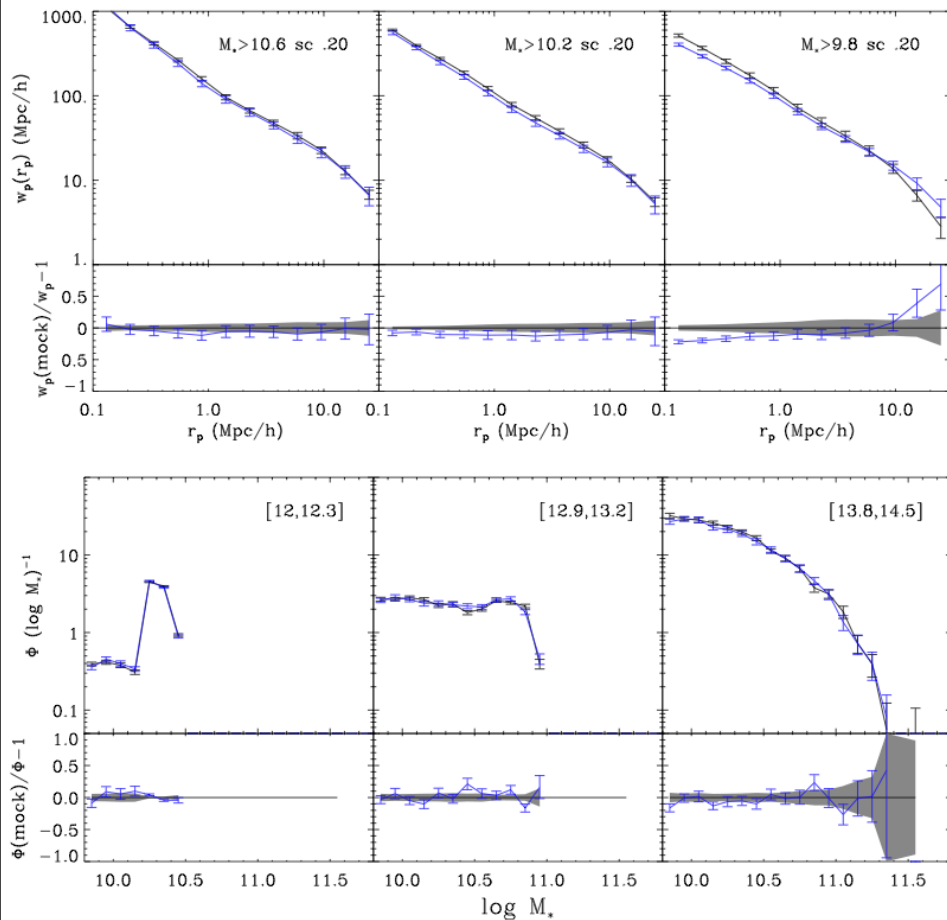
■ should contain all of the galaxies in the LSST “gold sample”

■ stars included now, quasar model very soon

■ simulations currently in hand: five cosmologies; 10 $1/4$ skies in one cosmology.

■ allows science analysis related to clusters, weak lensing, LSS, photometric redshifts, spectroscopic followup design, etc.

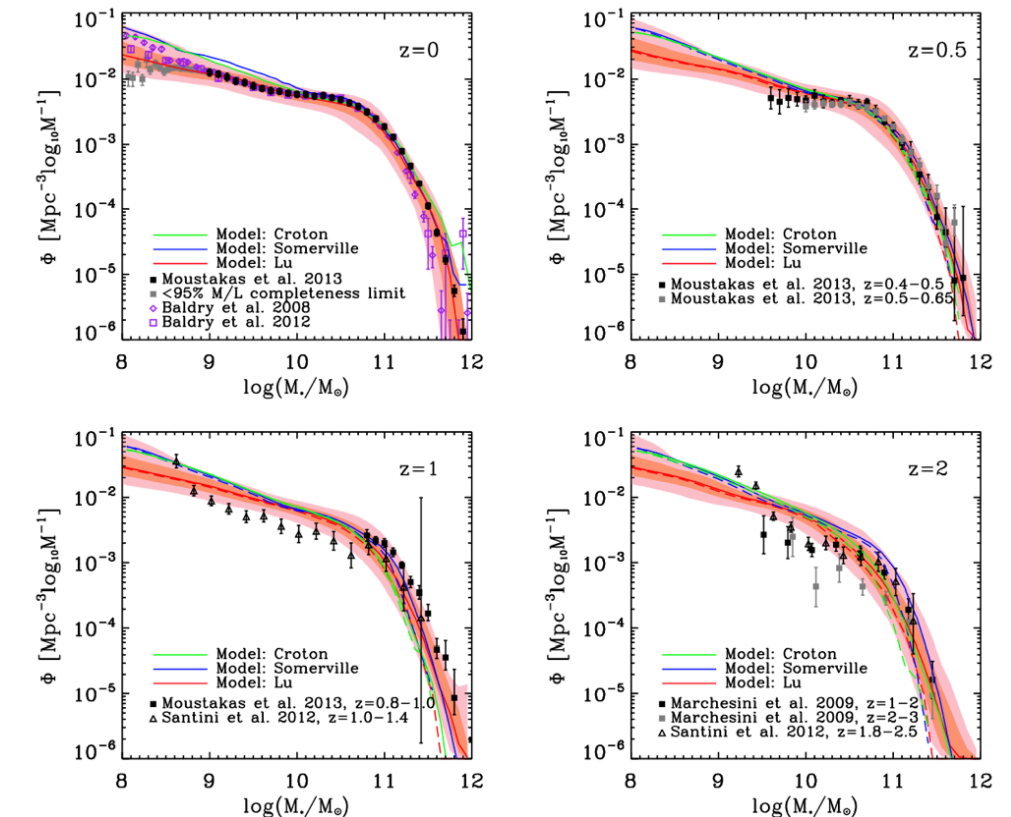
High resolution / abundance matching



- luminosities assigned to resolved halos and subhalos
- currently testing algorithms and resolution requirements
- colors and additional properties added based on selecting from real galaxies with similar local density & luminosity
- doing comparisons now between this method and other color methods based on same base catalogs and lum assignment (e.g. Hearin & Watson)
- Status: plan to produce fully lensed 100-200 sq. degree catalog using this method early 2014.
- Future plans: in 2014 large high res box (INCITE proposal PI: Warren) will allow similar method in Gpc volume

Semi-analytic model

- running Lu SAM on same merger trees as the previous version.
- Bayesian SAM which can be constrained to various observables.
- actively developing the model, including constraints with CANDELS, DES data



CALCLENS: Curved-sky gravitational Lensing for Cosmological Light cone simulations

See Becker 2013

CALCLENS is a multiple-plane ray tracing algorithm designed to add weak lensing signals to mock catalogs from N-body light cones.

Features:

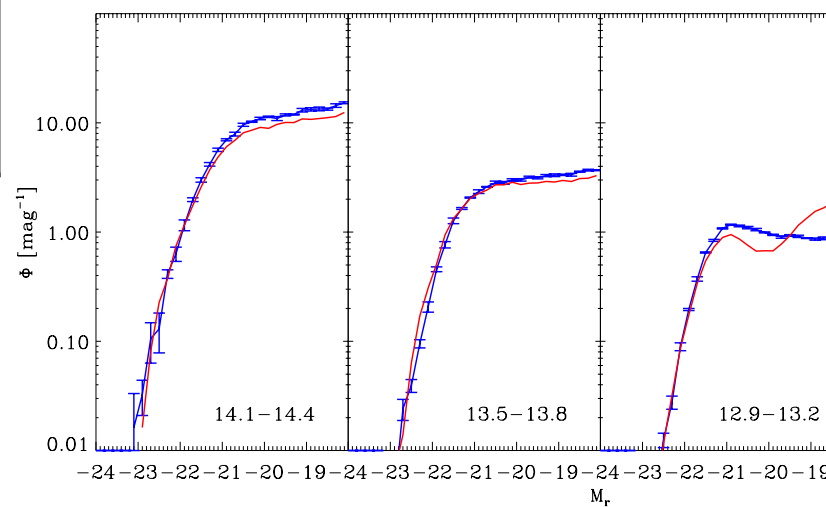
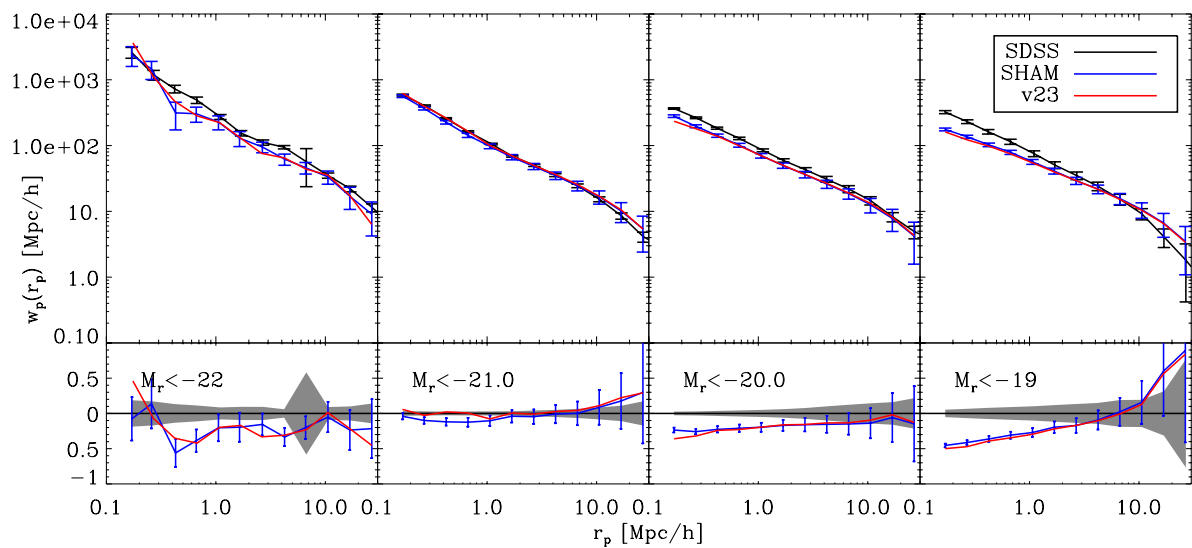
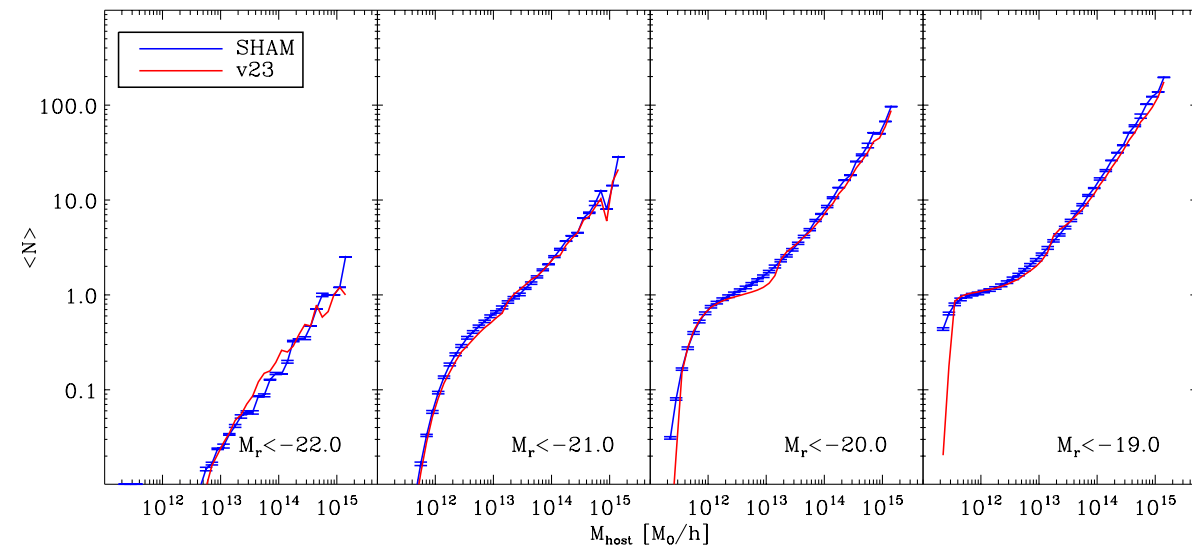
- works on the curved sky
- fast, approximate 2D Poisson solver
- works in the Limber approximation
- fully redshift dependent shear
- captures all of the magnification effects (i.e., finds galaxy images correctly)

Other “Features”:

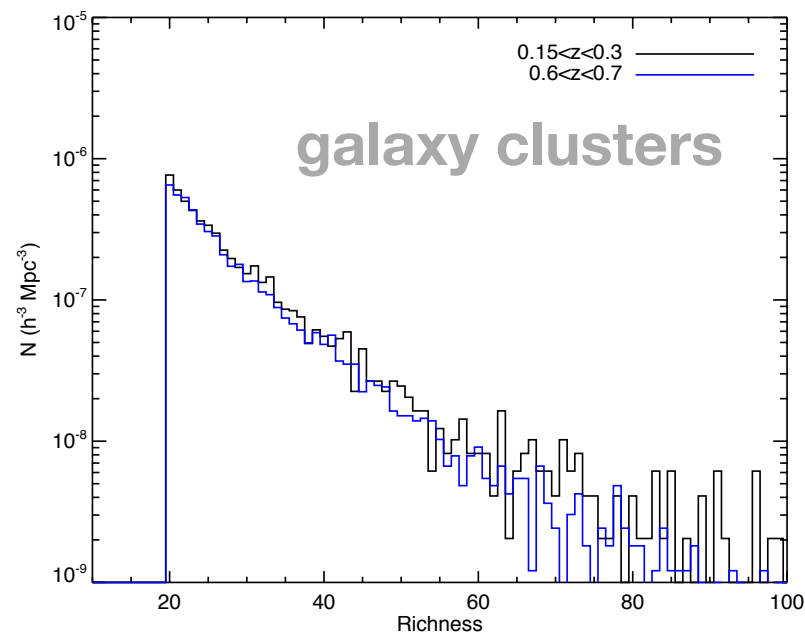
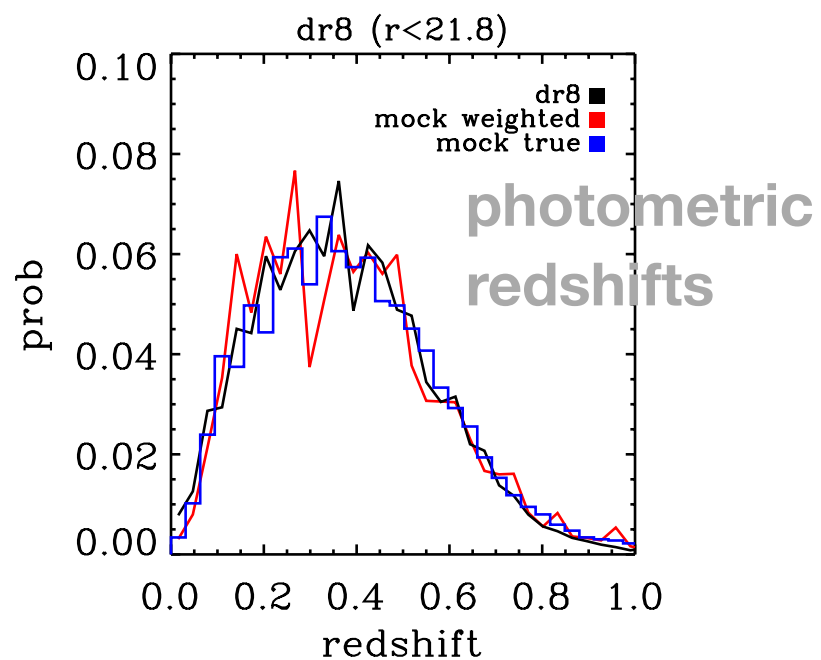
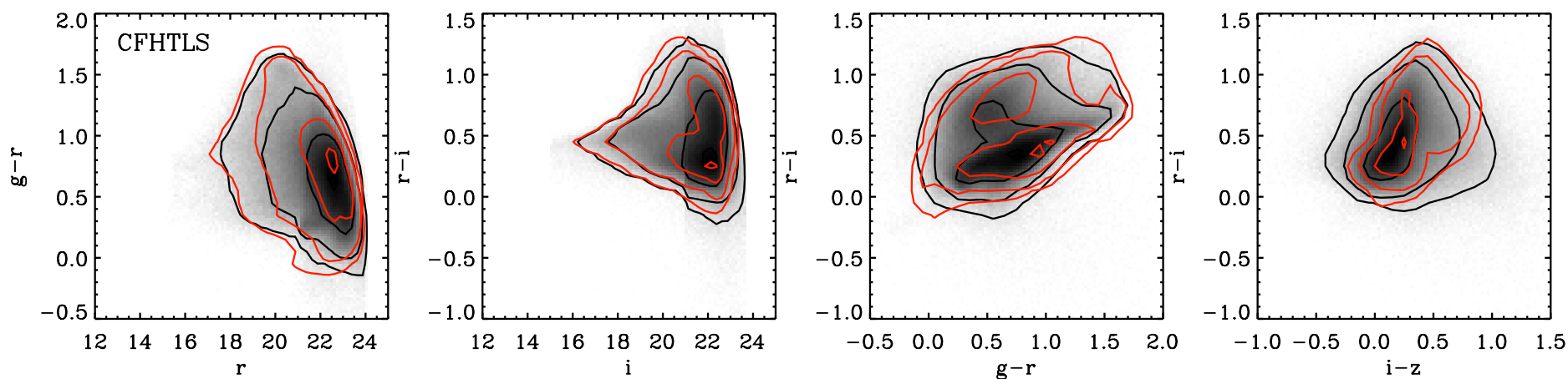
- approximate 2D Poisson solver
- works in the Limber approximation

abundance matching vs. addgals

with Michael Busha
& Rachel Reddick

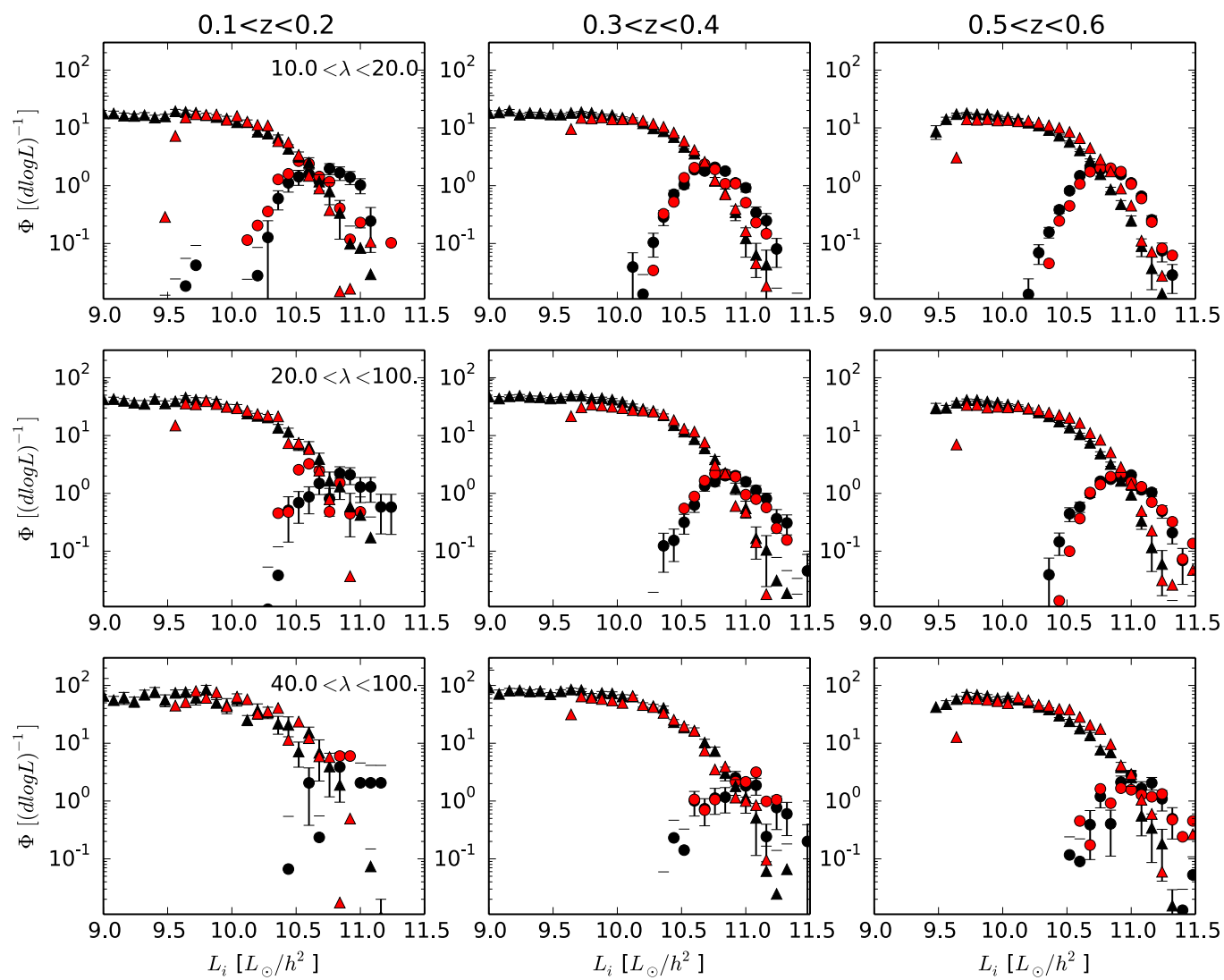


galaxy colors and luminosities

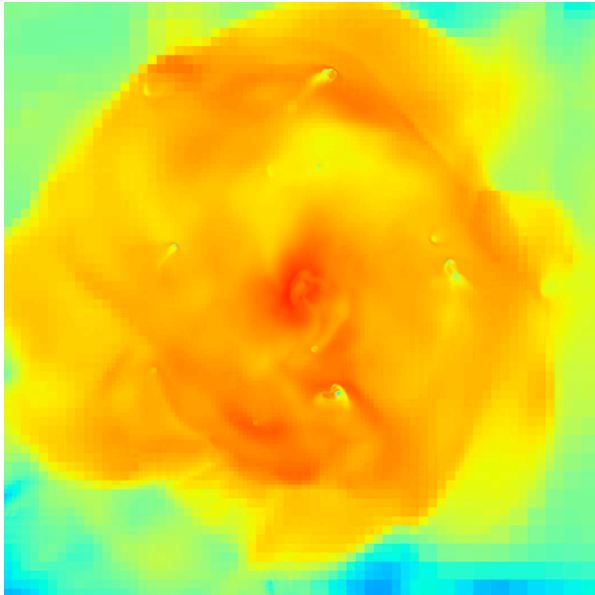


conditional luminosity function in clusters

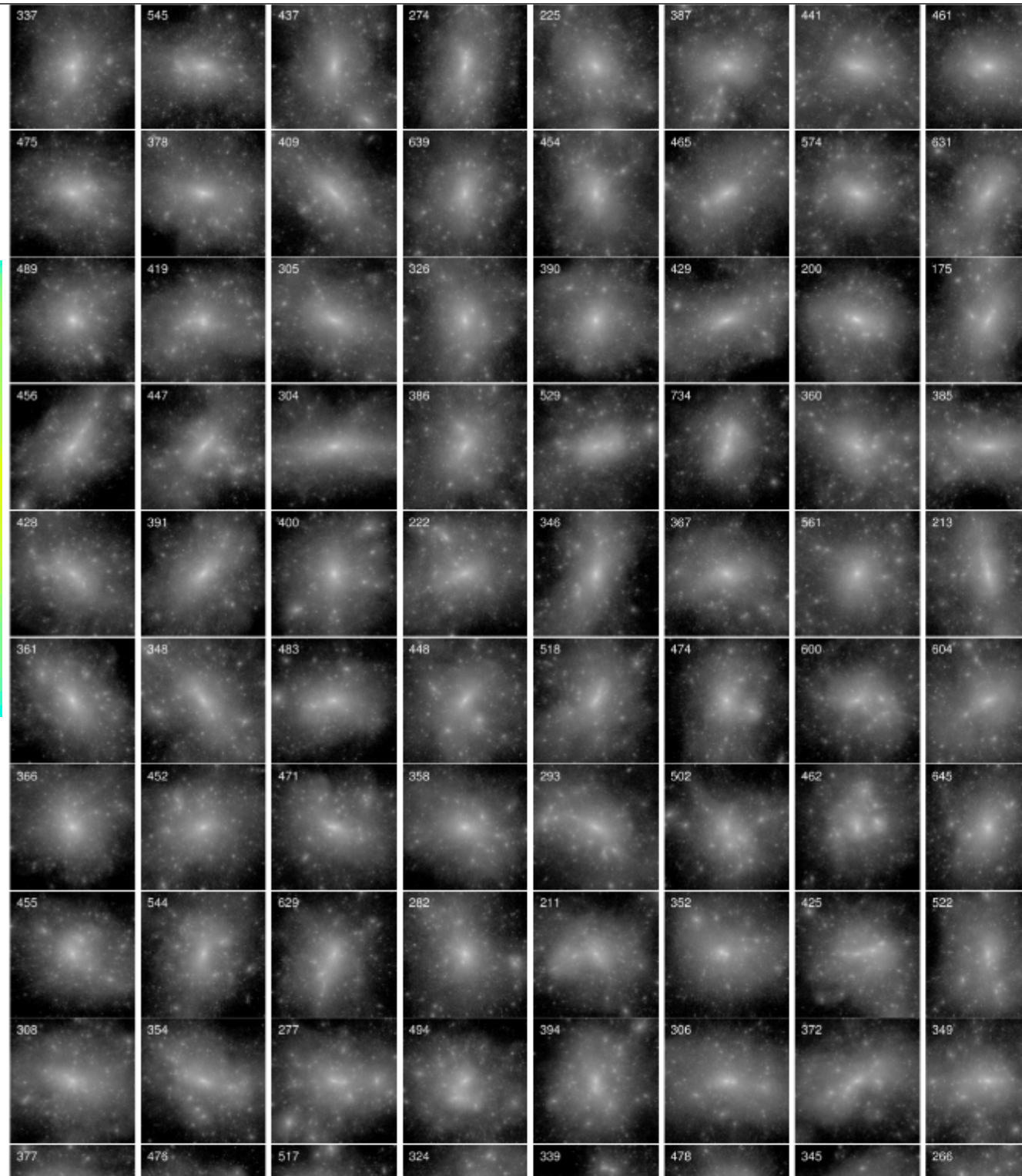
S82 vs mocks



cluster sims



Rhapsody resims of 100
clusters,
working now on hydro
(RAMSES & ENZO)



LSST-DESC summary

■ large area / low res

- quarter-sky, lensed galaxy catalogs on a lightcone, with LSST magnitudes to $i \sim 25$ and $z=2$ exist now in two cosmological models.
- expanding sample to cover several cosmologies, further developing galaxy model and testing against data (SDSS, DES), improving resolution, etc.

■ small area / high res

- high-resolution simulations which allow galaxies to be placed on resolved halos and subhalos available for smaller areas of sky (currently ~ 100 sq. degrees, with replication)
- populated using empirical (abundance matching) and semi-analytic methods
- mid-2014: apply these methods to new Gpc box with $1e12$ particles (approved INCITE w M. Warren)

■ working on integrating both sets of simulations with ImSim tools (with Debbie Bard)

■ for cluster work: sample of 100 high resolution cluster simulations (RHAPSODY), now under development with hydro

■ *please contact me (rwechsler@stanford.edu) if you would like to use them for DESC tasks!
please give us feedback on DESC priorities for accuracy, resolution, properties, volume!*