

# Compton simulation preliminary results

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# Simulation setup

- Using our default Compton generator
  - <https://gitlab.com/eic/mceg/comptonRad>
- The generator gives out the 4-momentum for both the photon and electron together with 4 weight factors: unpolarized and polarized tree level cross sections (eq 26,27 of paper) and the order alpha corrections (not used for the following analysis)
- To obtain the average analyzing power for a particular configuration (or average over any number of bins) we need to weight by cross section
- An average FOM can also be calculated weighting by sqrt(N)

$$W_1 = \frac{1}{2\rho^{(n)}(x)} \left[ \frac{d^n \sigma^{(0)}}{dx^n}(s, -) + \frac{d^n \sigma^{(0)}}{dx^n}(s, +) \right] \quad \text{unpolarized xsection} \quad (26)$$

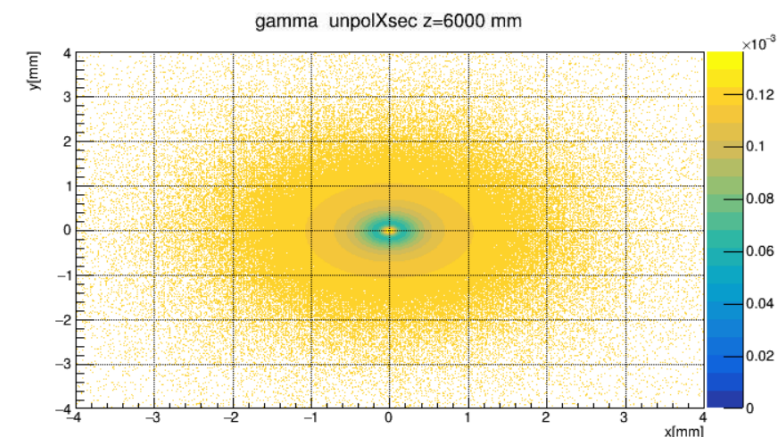
$$W_2 = \frac{1}{2\rho^{(n)}(x)} \left[ \frac{d^n \sigma^{(0)}}{dx^n}(s, -) - \frac{d^n \sigma^{(0)}}{dx^n}(s, +) \right] \quad \text{polarized xsection} \quad (27)$$

$$A_N = \frac{\sigma^- - \sigma^+}{\sigma^- + \sigma^+} = \frac{\sigma^p}{\sigma^u} \equiv \frac{W_1}{W_2}$$

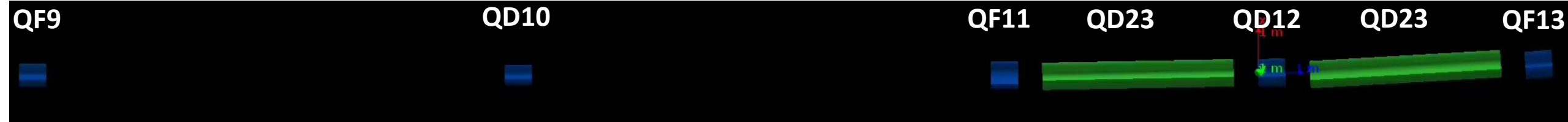
$$\langle A_N \rangle = \frac{\sum_i A_{N,i} \cdot \sigma_i^u}{\sum_i \sigma_i^u} = \frac{\sum_i \sigma_i^p}{\sum_i \sigma_i^u}$$

$$\langle FOM \rangle = \frac{\sum_i A_{N,i} \cdot \sqrt{\sigma_i^u} \cdot \sigma_i^u}{\sum_i \sigma_i^u} = \frac{\sum_i \sigma_i^p \cdot \sqrt{\sigma_i^u}}{\sum_i \sigma_i^u}$$

$$\langle E \cdot A_N \rangle = \frac{\sum_i E_i \cdot \sigma_i^p}{\sum_i E_i \cdot \sigma_i^u}$$



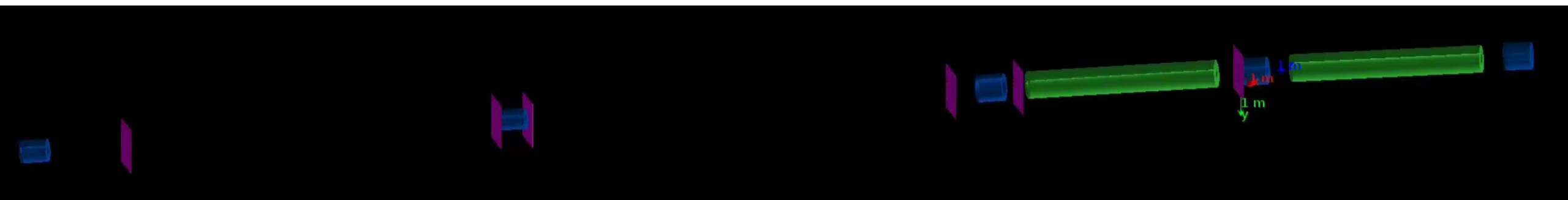
# Simulation setup



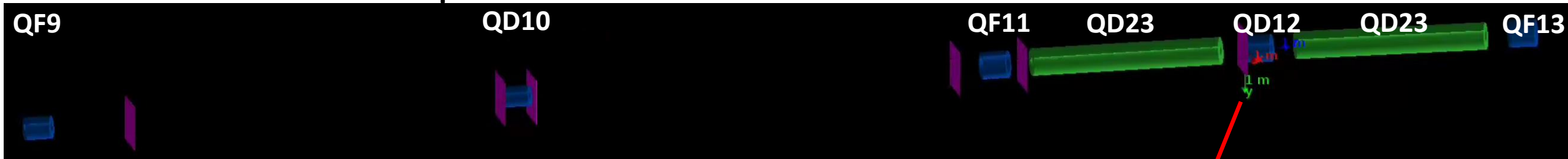
- Implemented IP12 magnet configuration

- <https://github.com/cipriangal/fun4all> Compton
- Particles going in  $-z$  direction with the origin at the exit of QD12
- The lunch angle is the same as the QD12 angle for both the electron and photon
  - Confirmed that with this lunch angle an 18GeV electron goes through the middle of all magnet apertures
- Look at the “truth” information at different planes before and after magnetic elements

|      | X [m] | Y [m] | Z [m]  | theta |
|------|-------|-------|--------|-------|
| QF13 | 0.17  | 0.00  | 6.19   | 38.40 |
| DB23 | 0.07  | 0.00  | 3.25   | 28.80 |
| QD12 | 0.00  | 0.00  | 0.30   | 19.20 |
| DB23 | -0.05 | 0.00  | -2.65  | 9.60  |
| QF11 | -0.06 | 0.00  | -5.59  | 0.00  |
| QD10 | -0.06 | 0.00  | -16.32 | 0.00  |
| QF9  | -0.06 | 0.00  | -27.04 | 0.00  |



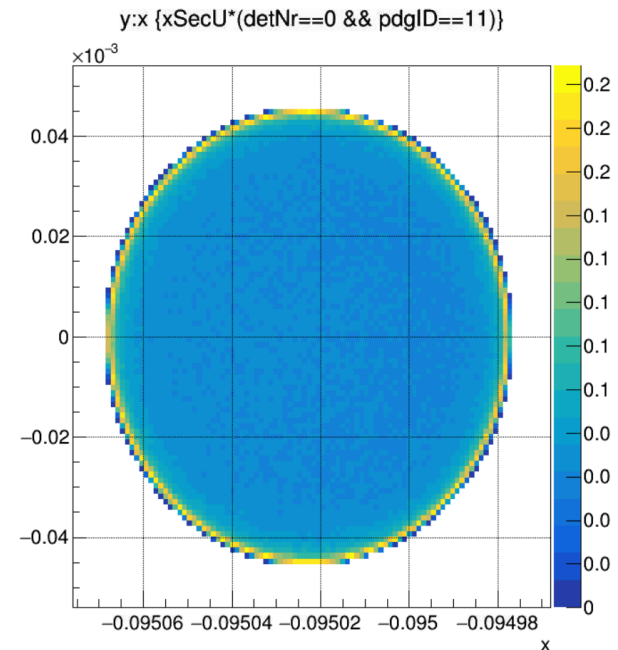
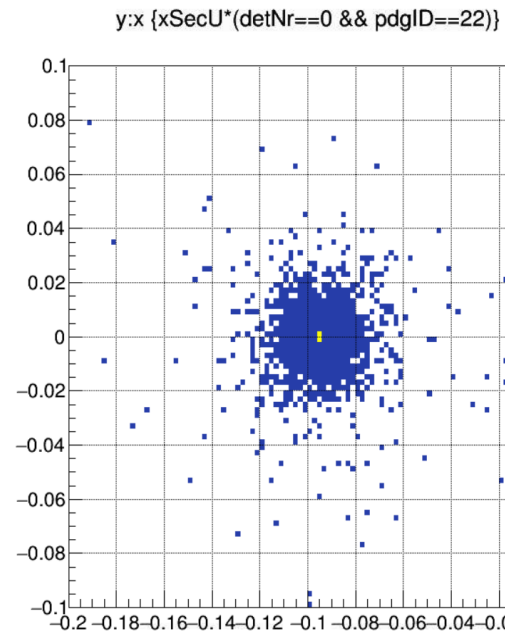
# Particle envelopes



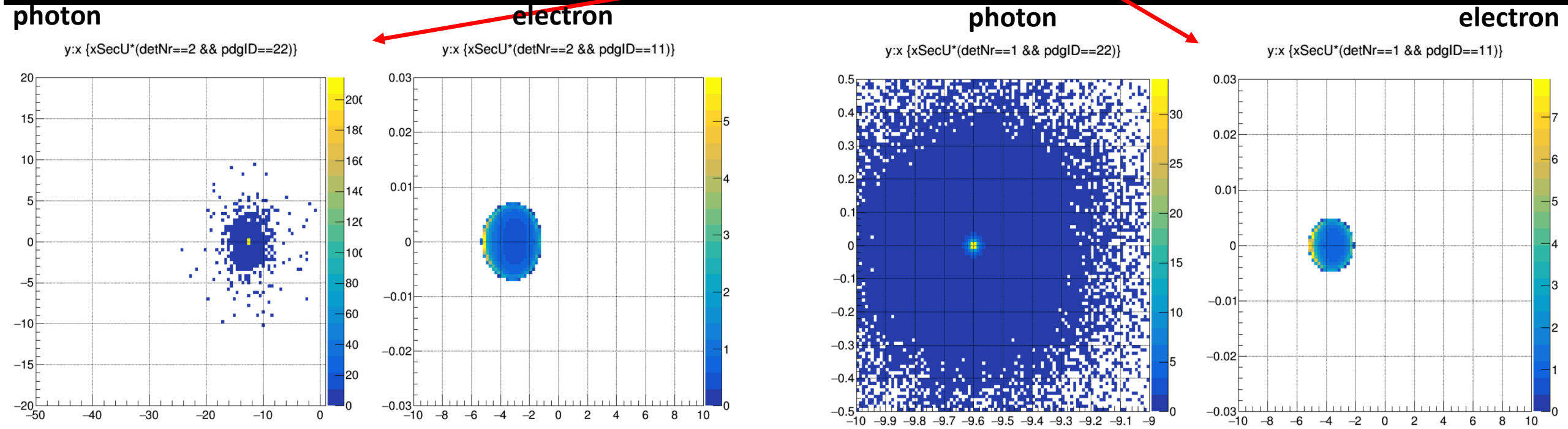
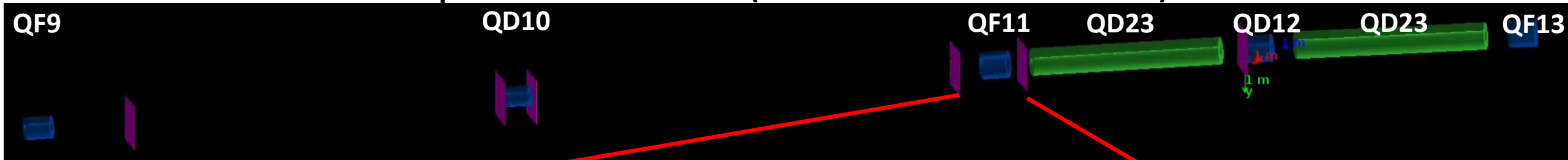
- All the axes are in cm
- The first “truth” plane is 6 cm downstream of the vertex point

photon

electron

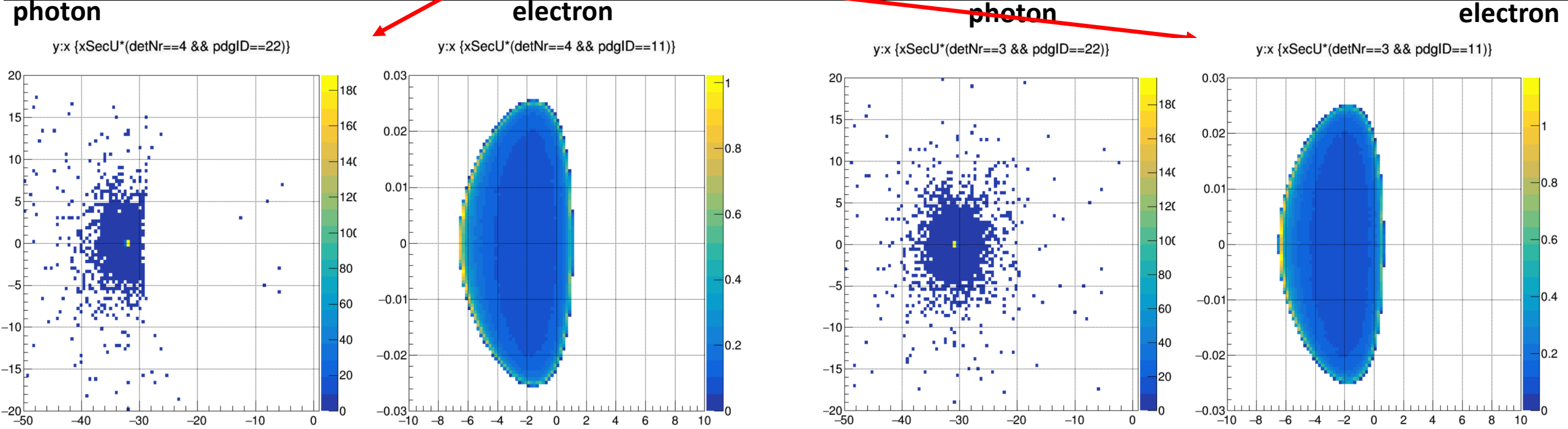


# Particle envelopes: QF11 ( $R_{min} = 20$ cm)



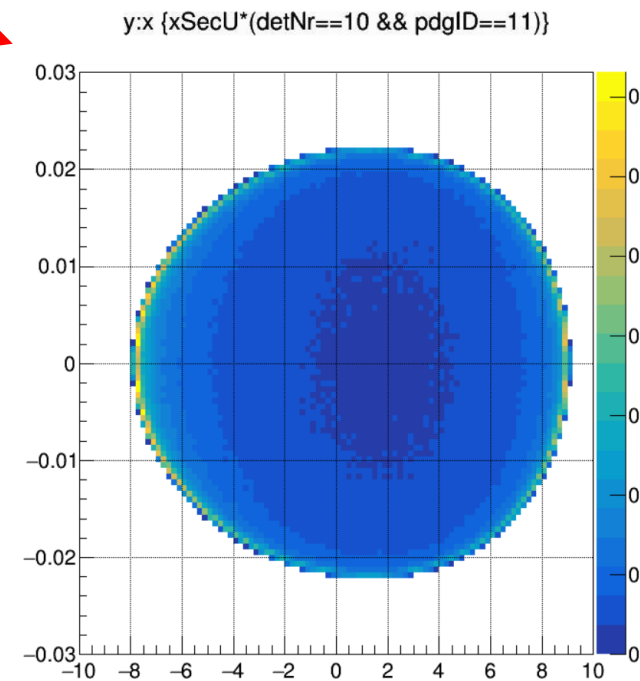
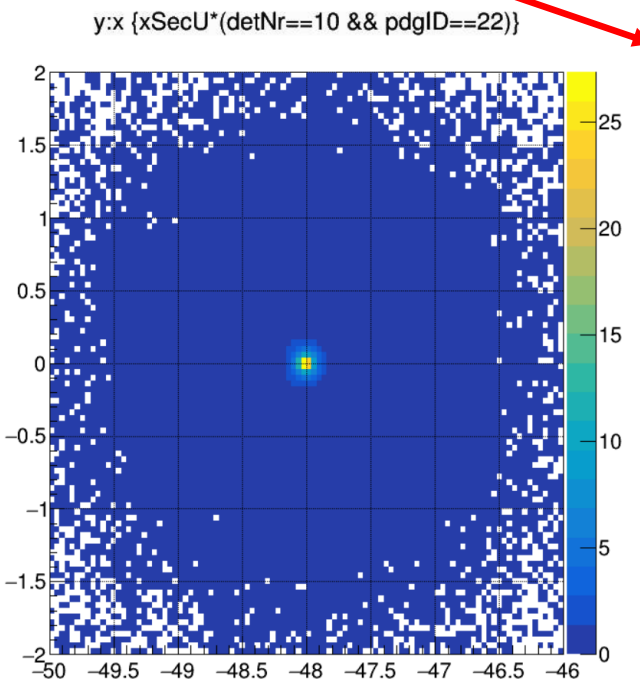
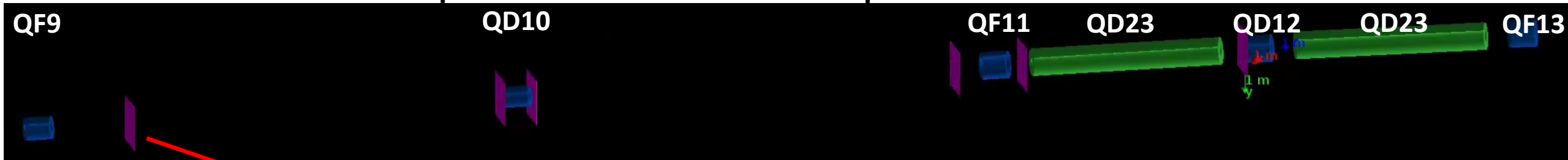
- I increased the bore of the quad to look downstream, but we can see we would start cutting into the photon envelope if  $R_{min} < 10$  cm (quad center is  $x = -6$  cm)

# Particle envelopes: QD10 ( $R_{min}=8.5\text{cm}$ $R_{max}=22.5\text{cm}$ )



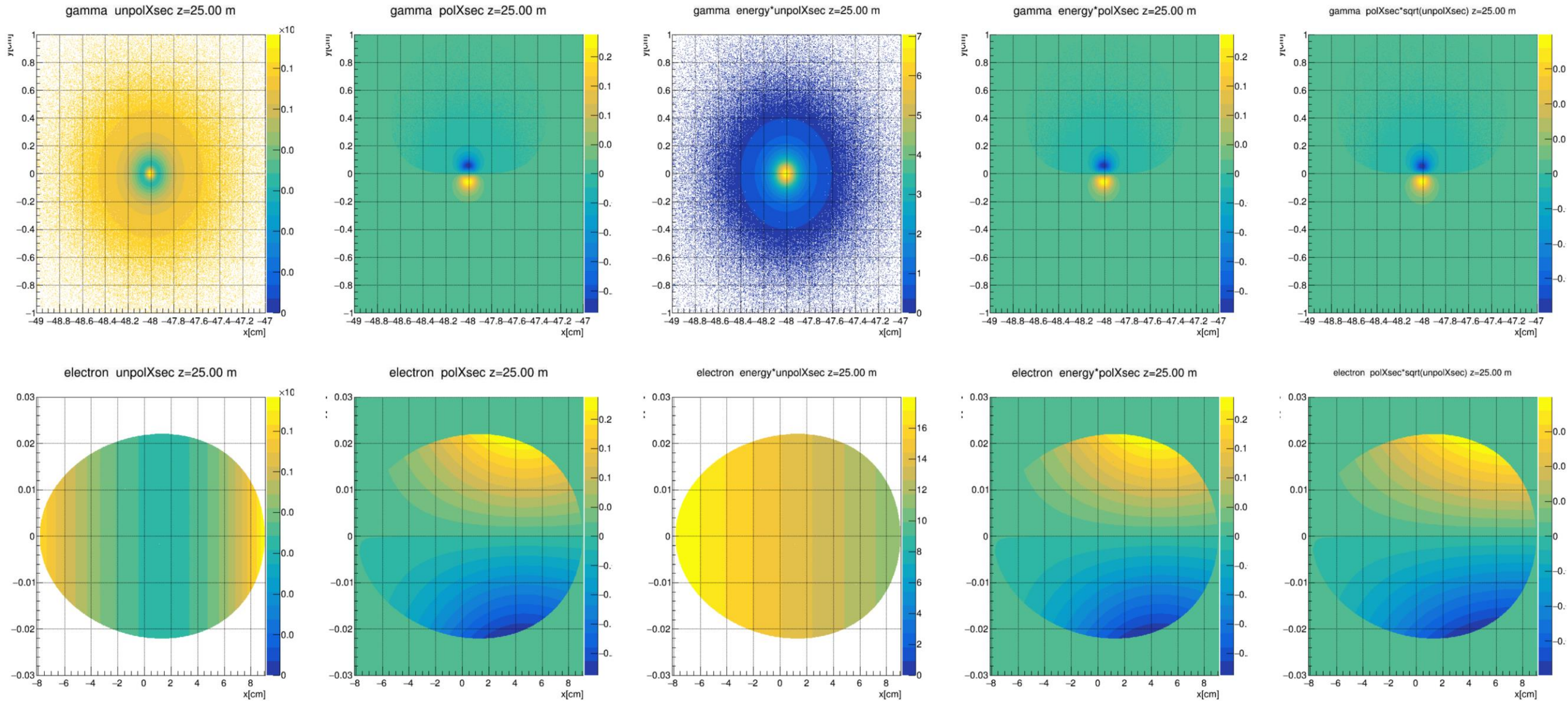
- Even with the restrictive magnet sizes we cut into the photon envelope with the QD10 outer size

# Particle envelopes: “detector plane”



- For QF9 with  $R_{min}=10\text{cm}$  and  $R_{max}=25\text{cm}$  we will have significant showering from the electron envelope
- This will produce a significant charged background for the electron detector tracker
- To place a calorimeter for the photon detector we only have about 42cm clearance from the center of the beamline

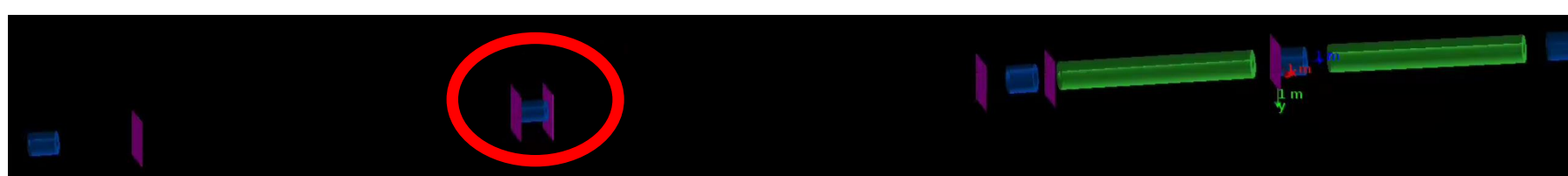
# Detector plane: nominal configuration



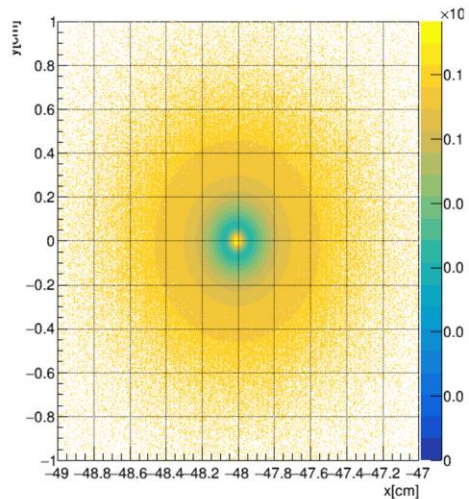


# Without QD10

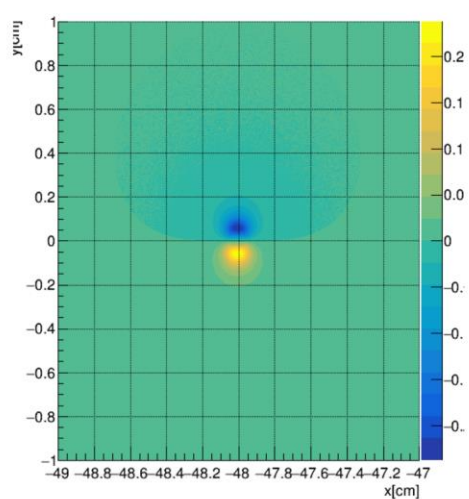
Nominal config



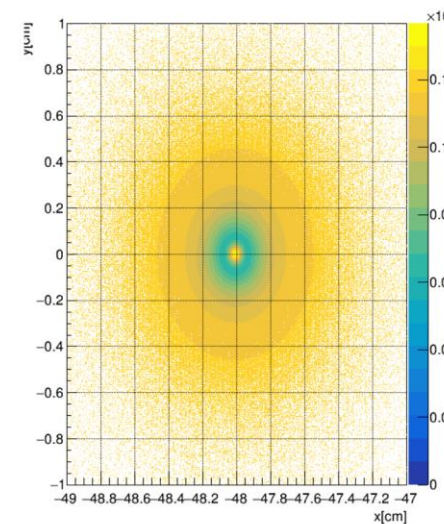
gamma unpolXsec z=25.00 m



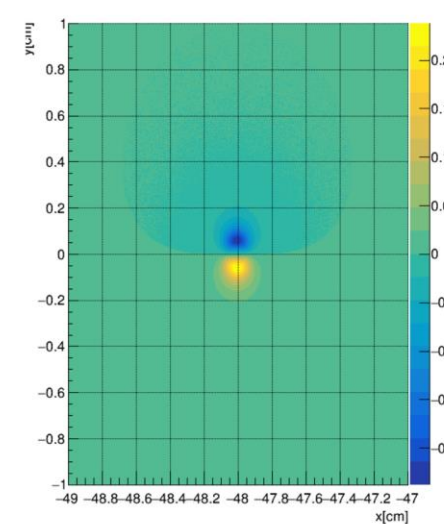
gamma polXsec z=25.00 m



gamma unpolXsec z=25.00 m

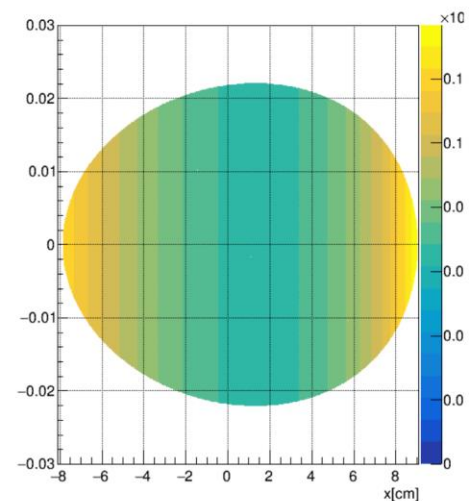


gamma polXsec z=25.00 m

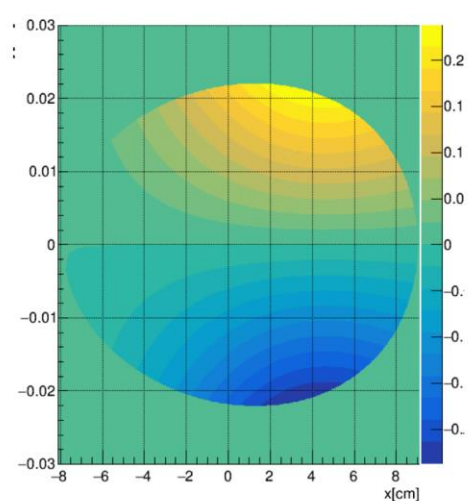


- Without the vertical focusing the scattered electron envelope would almost double by the time it gets to the detector plane

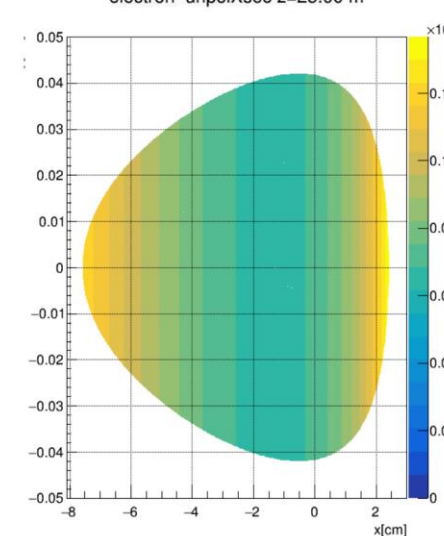
electron unpolXsec z=25.00 m



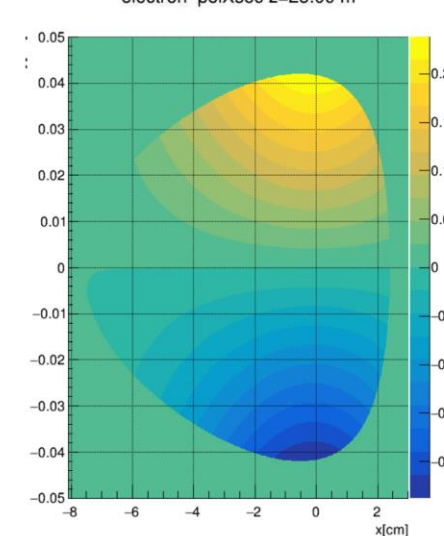
electron polXsec z=25.00 m



electron unpolXsec z=25.00 m



electron polXsec z=25.00 m



# Vertex smearing

```
eicfile->set_vertex_distribution_mean(0,0,0,0); //cm, ns
eicfile->set_vertex_distribution_width(4.3e-3,3.3e-3,0.9,0); //cm, ns
```

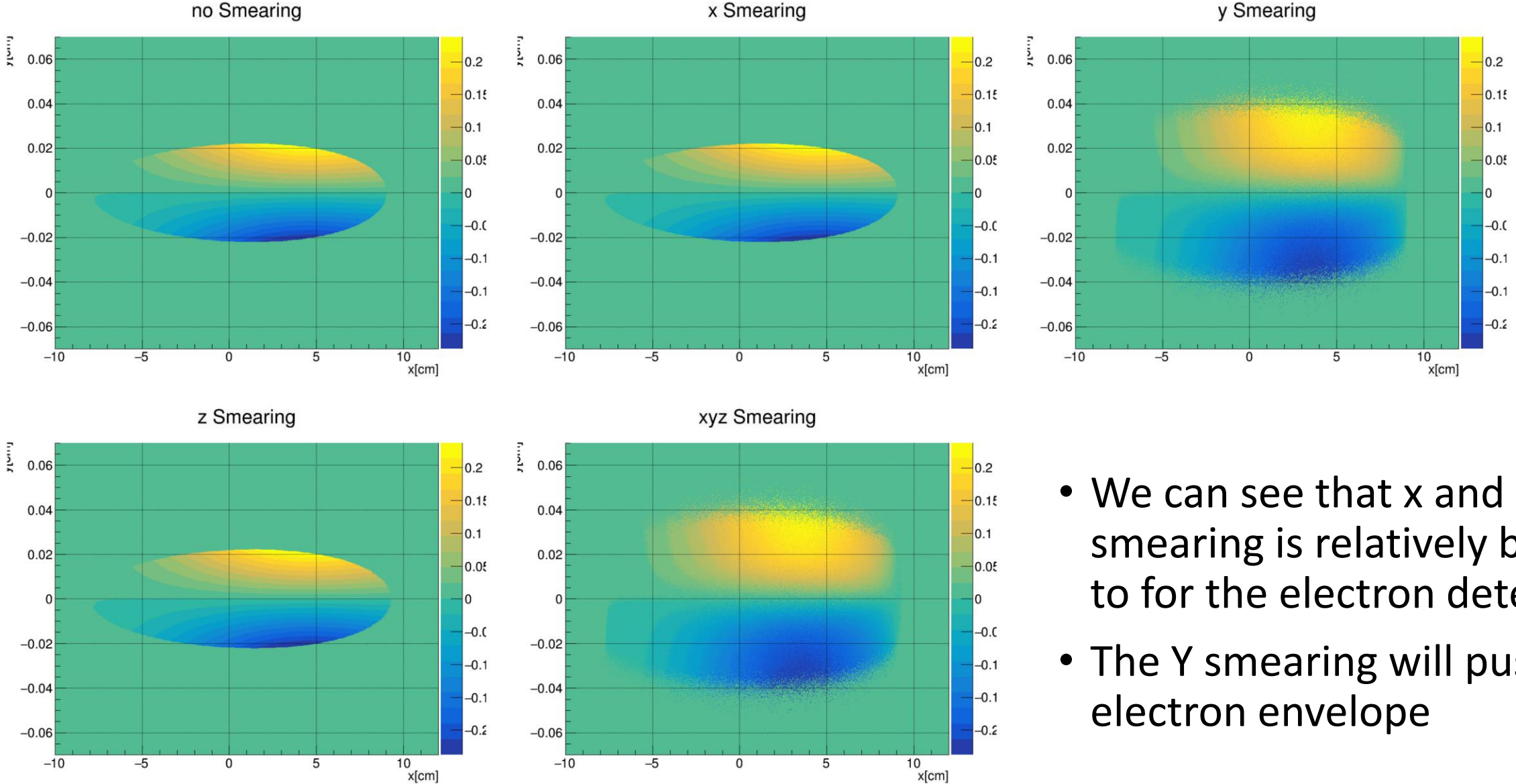
Here is the twiss parameters for IR12. "BETX", "BETY" is the beta function. "ALFX", "ALFY" is the slope of the beta function.

| * NAME   | S           | L           | BETX        | ALFX          | BETY        | ALFY           |
|--|-------------|-------------|-------------|---------------|-------------|----------------|
| "QF13"   | 1823.294259 | 0.6         | 4.669745436 | 0.310750451   | 23.49304068 | -0.7668571524  |
| "DB23"   | 1826.241095 | 4.223670697 | 3.777088439 | -0.2494192314 | 37.53084031 | -2.797041129   |
| "QD12"   | 1829.18793  | 0.6         | 7.833576927 | -1.577181773  | 54.98601912 | 0.06492603105  |
| ***** IP of the Compton polarimeter here ***** |             |             |             |               |             |                |
| "DB23"   | 1832.134765 | 4.223670697 | 25.91330713 | -4.11259067   | 36.90923362 | 2.876766807    |
| "QF11"   | 1835.0816   | 0.6         | 54.98959915 | -2.21962997   | 22.61170074 | 0.5441235163   |
| "QD10"   | 1845.805117 | 0.6         | 27.78894174 | 0.01007799586 | 54.99548409 | -0.1719185813  |
| "QF9"  | 1856.528634 | 0.6         | 54.99503024 | 0.2700717205  | 27.09541968 | 0.004988464914 |

$$\sigma_i = \sqrt{\epsilon_i * \beta_i}$$

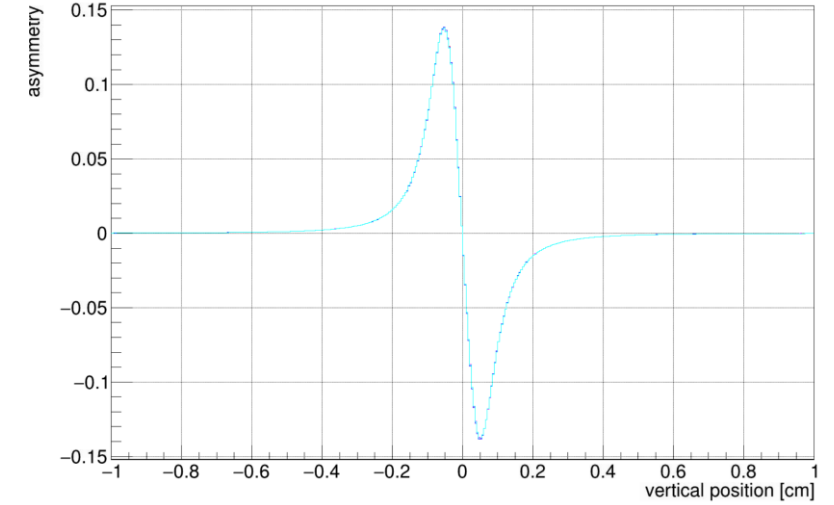
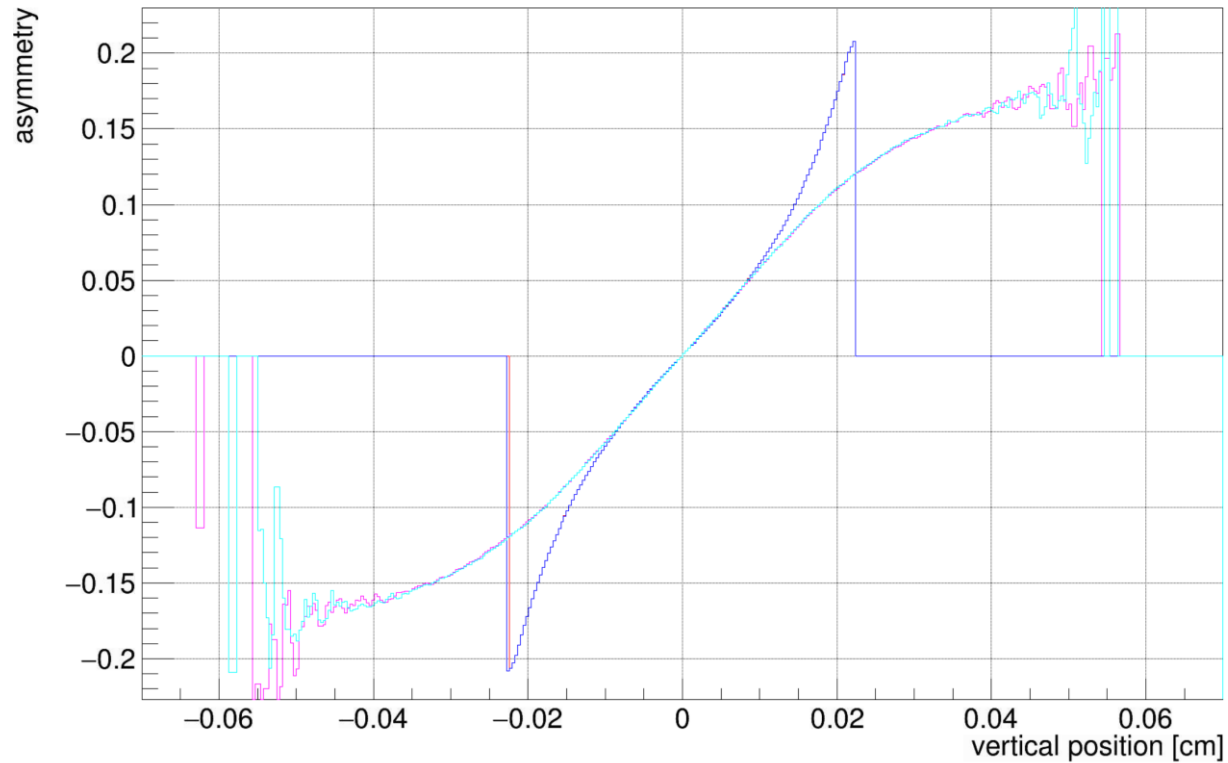
| Species   | proton   | electron |
|---|----------|----------|
| Energy [GeV]  | 275      | 18       |
| CM energy [GeV]   | 140.7    |          |
| Bunch intensity [10 <sup>10</sup> ]                               | 20.5     | 6.2      |
| No. of bunches  | 290      |          |
| Beam current [A]  | 0.74     | 0.227    |
| RMS norm. emit., h/v [μm]   | 4.6/0.75 | 845/72   |
| RMS emittance, h/v [nm]   | 16/2.6   | 24/2.0   |
| β*, h/v [cm]  | 90/4.0   | 59/5.0   |
| IP RMS beam size, h/v [μm]  | 119/10   |          |
| K <sub>x</sub>  | 11.8     |          |
| RMS Δθ, h/v [μrad]  | 132/253  | 202/202  |
| BB parameter, h/v [10 <sup>-3</sup> ]                             | 3/2      | 100/100  |
| RMS long. emittance [10 <sup>-3</sup> , eV·sec]                   | 36       |          |
| RMS bunch length [cm]   | 6        | 0.9      |
| RMS Δp/p [10 <sup>-4</sup> ]                                      | 6.8      | 10.9     |
| Max. space charge   | 0.006    | neglig.  |
| Piwinski angle [rad]  | 5.6      | 0.8      |
| Long. IBS time [h]  | 2.1      |          |
| Transv. IBS time [h]  | 2        |          |
| Hourglass factor H  | 0.86     |          |
| Luminosity [10 <sup>33</sup> cm <sup>-2</sup> sec <sup>-1</sup> ] | 1.65     |          |

# Vertex smearing

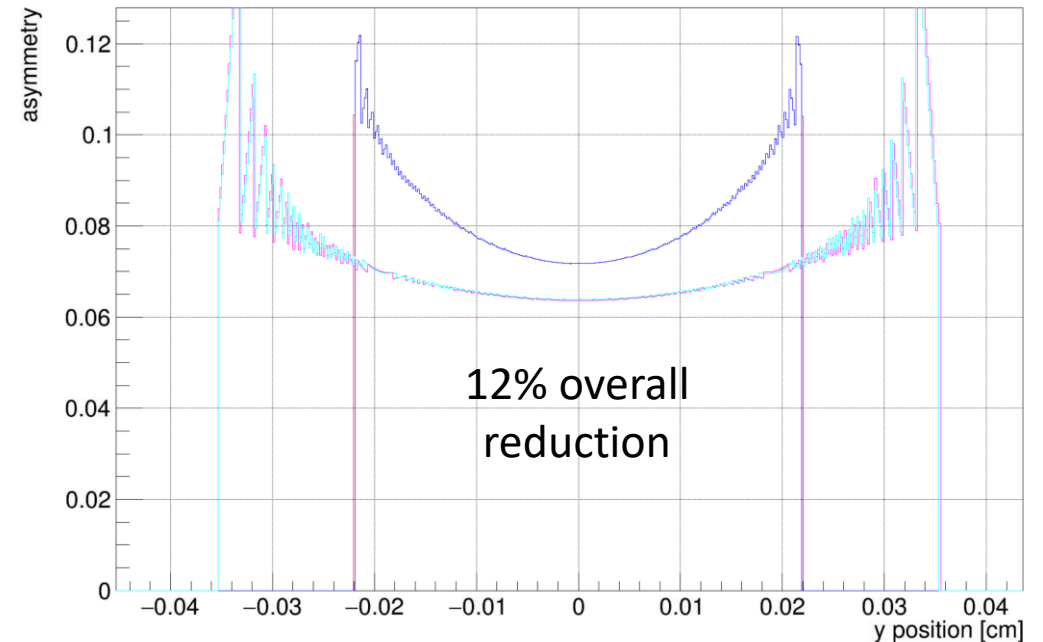


- We can see that x and z smearing is relatively benign to for the electron detector
- The Y smearing will push the electron envelope

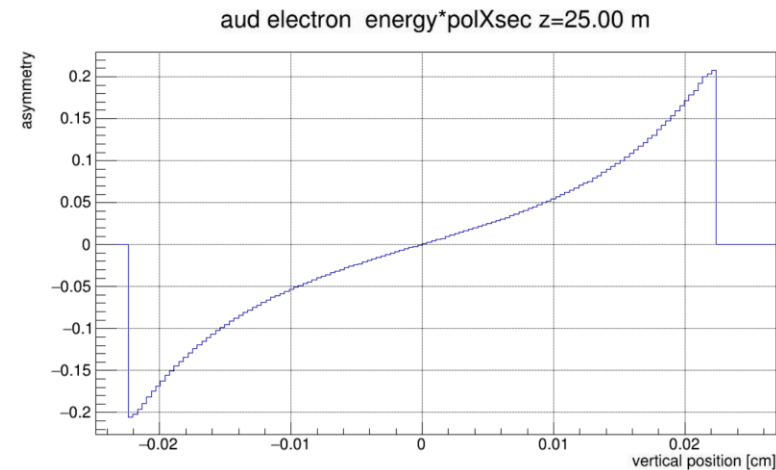
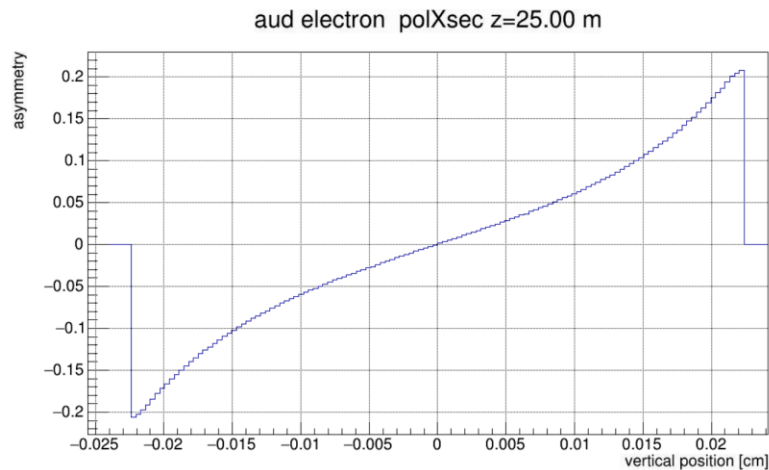
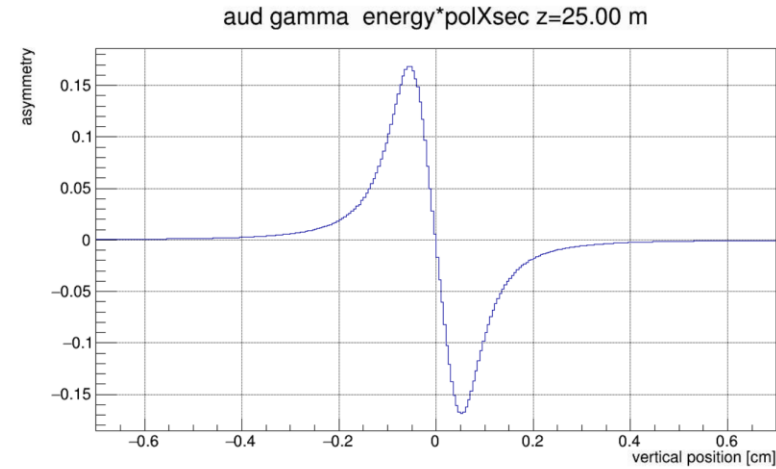
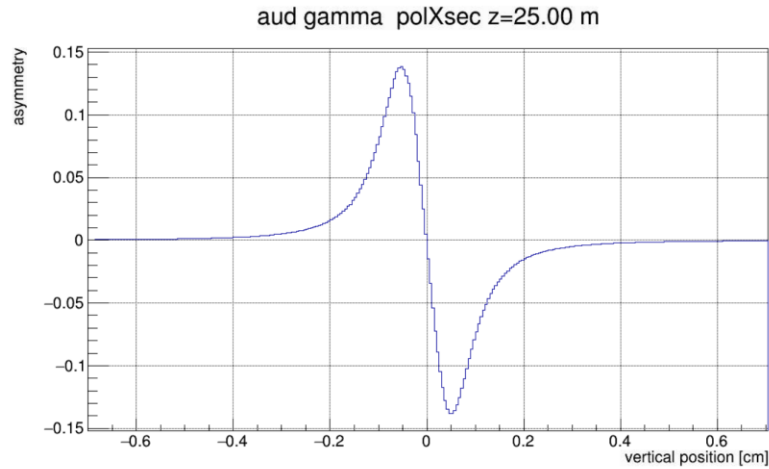
# Vertex smearing: vertical asymmetry



- The smearing will decrease the average measured asymmetry
- Plots in the following slides will be without any smearing



# Vertical asymmetry: detector plane

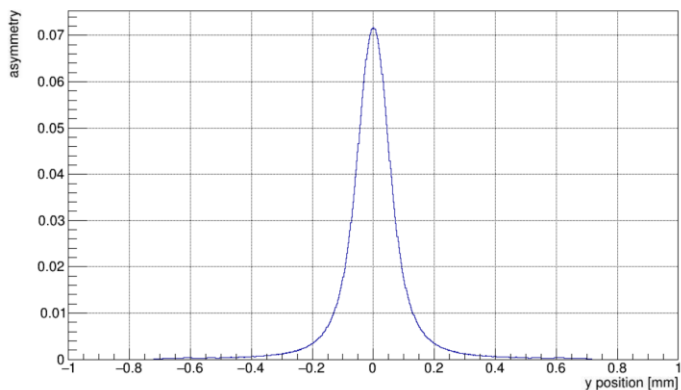


- We can see that the analyzing power goes up to  $\sim 14\%$  for the photon and  $\sim 20\%$  for the electron

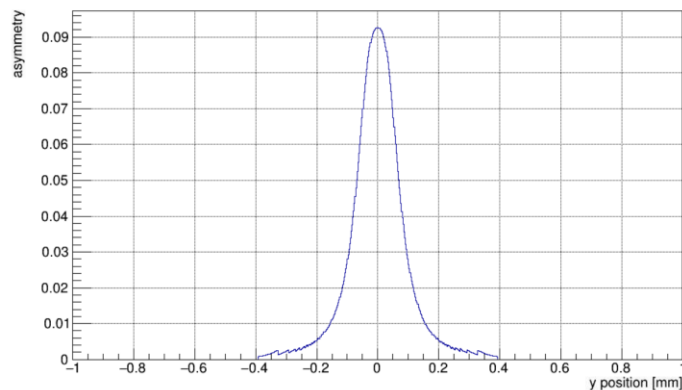
# AUD and position offsets (detector plane)

Vertical axis:  $(\langle A \rangle_{\text{up}} - \langle A \rangle_{\text{down}}) / 2$

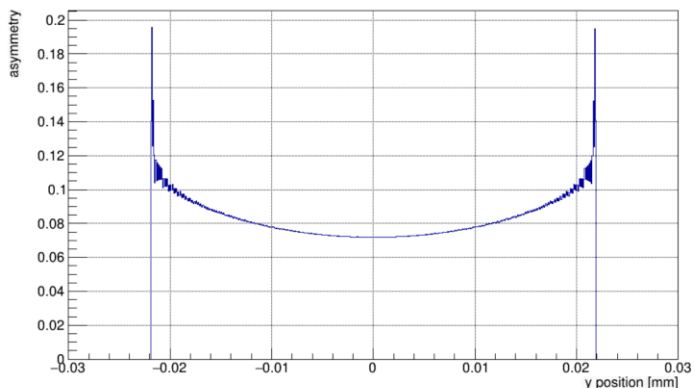
aud gamma polXsec z=25.00 m



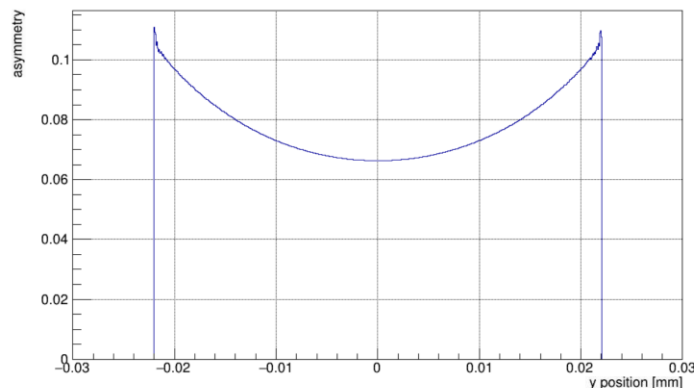
aud gamma energy\*polXsec z=25.00 m



aud electron polXsec z=25.00 m



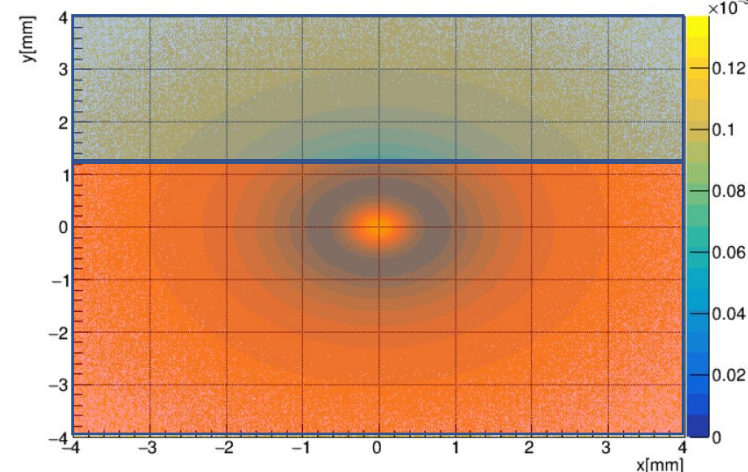
aud electron energy\*polXsec z=25.00 m



$\langle A \rangle_{\text{up}}$



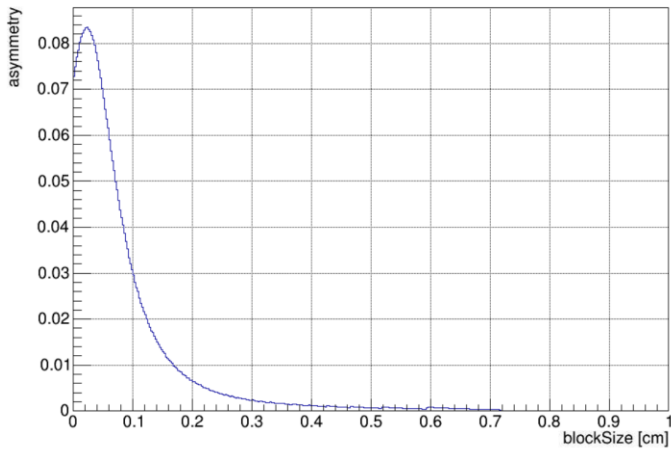
$\langle A \rangle_{\text{down}}$



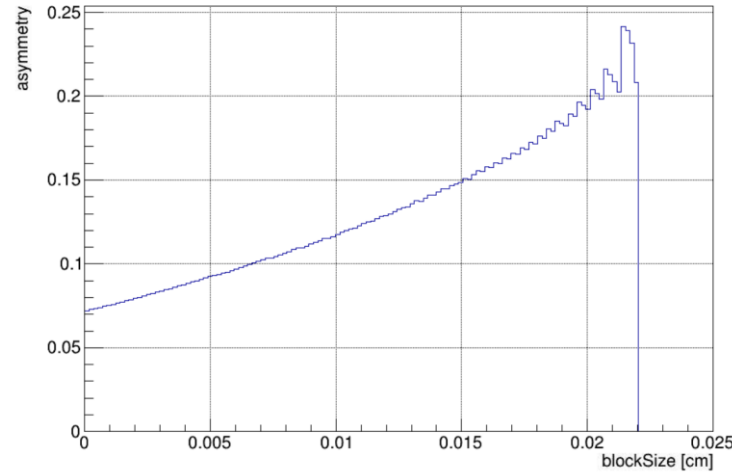
- For a simple “2-side” analysis both the photon and electron effective analyzing power is a  $\sim 7\%$  for the counting analysis
  - An energy weighted analysis would increase the analyzing power for the photon by about 2% and decrease the electron analyzing power by about 1%

# AUD block measurements – with Bfields

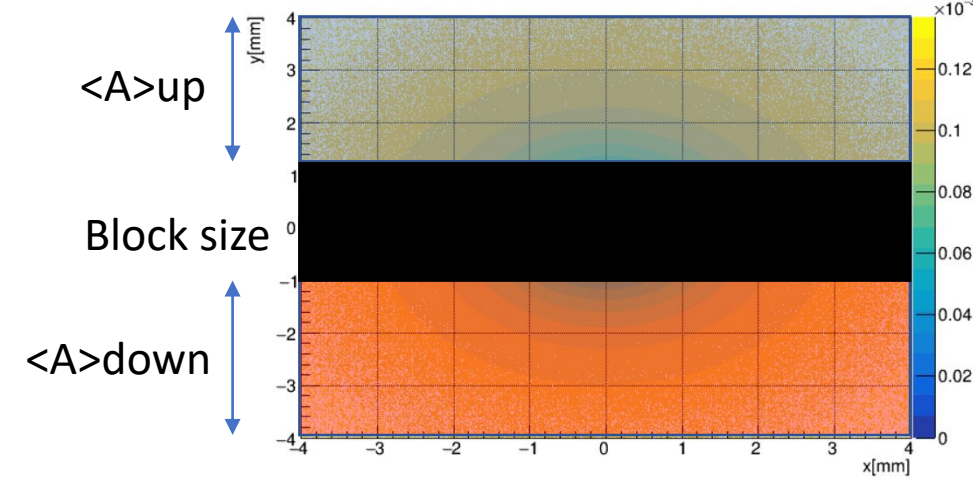
aud center=0.00 gamma polXsec z=25.00 m



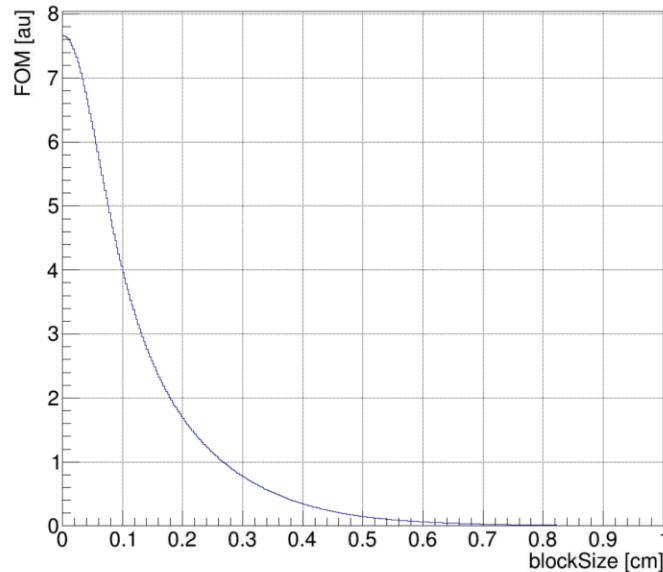
aud center=0.00 electron polXsec z=25.00 m



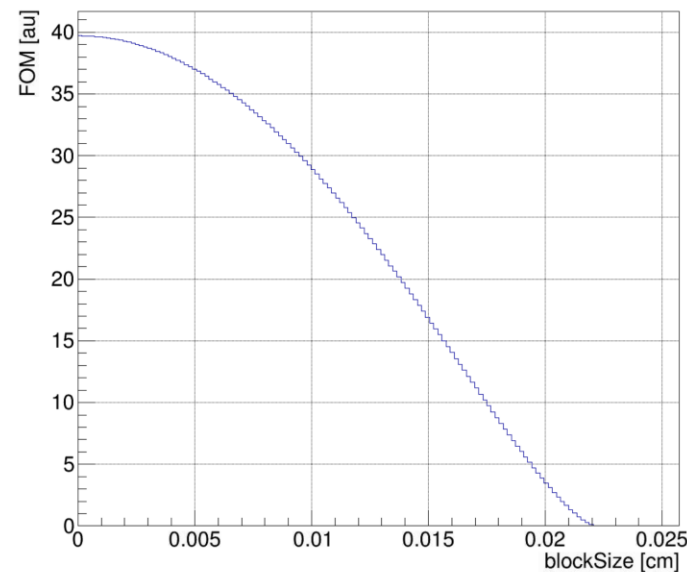
gamma unpolXsec z=25000 mm



FOM center=0.00 gamma polXsec\*sqrt(unpolXsec) z=25.00 m



FOM center=0.00 electron polXsec\*sqrt(unpolXsec) z=25.00 m

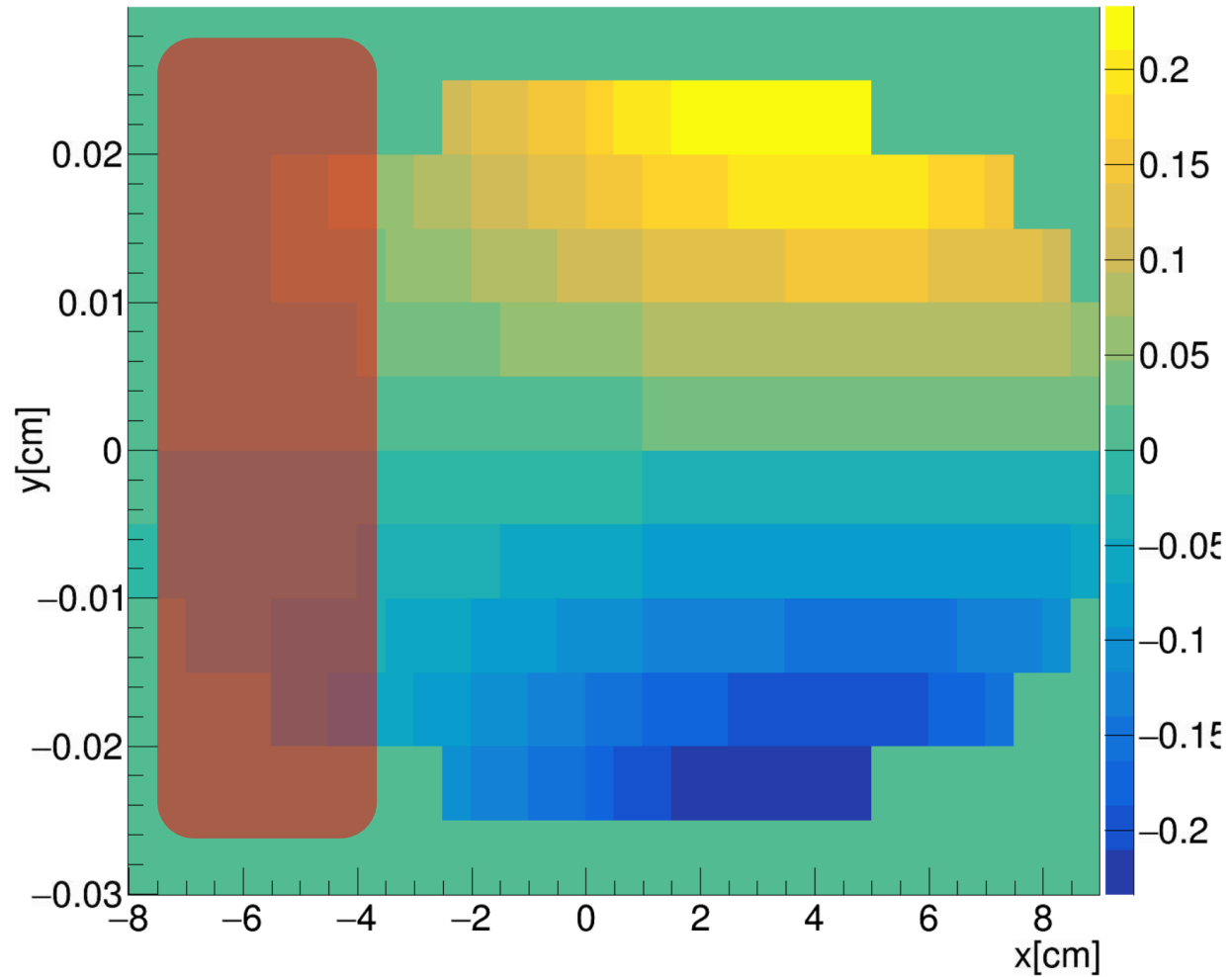
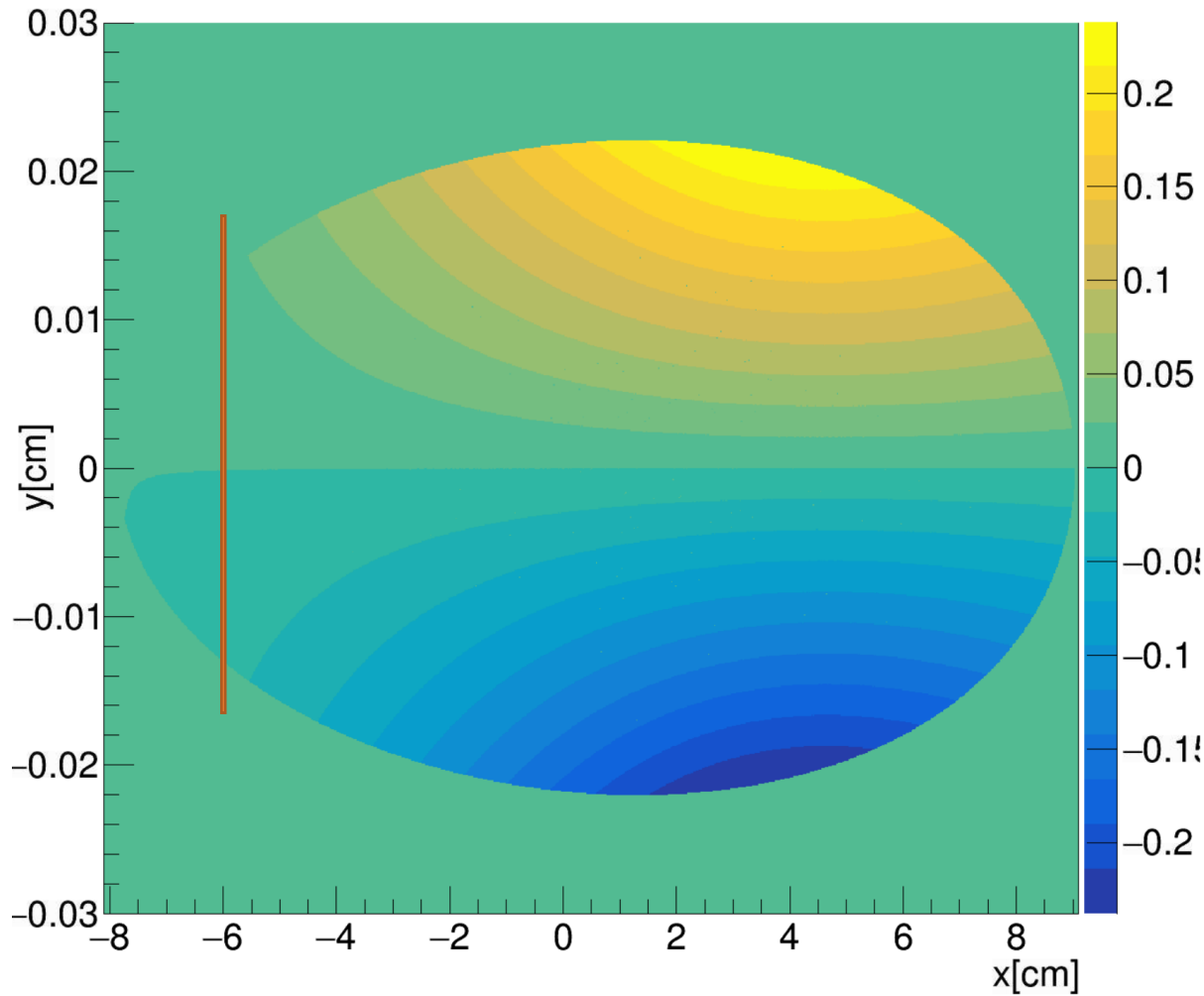


- While I made the claim last time there could be some gain from using two detector halves if we look at the FOM we see that is not really the case

# Ideal vs detector segmentation

electron polXsec z=25.00 m

average asymmetry



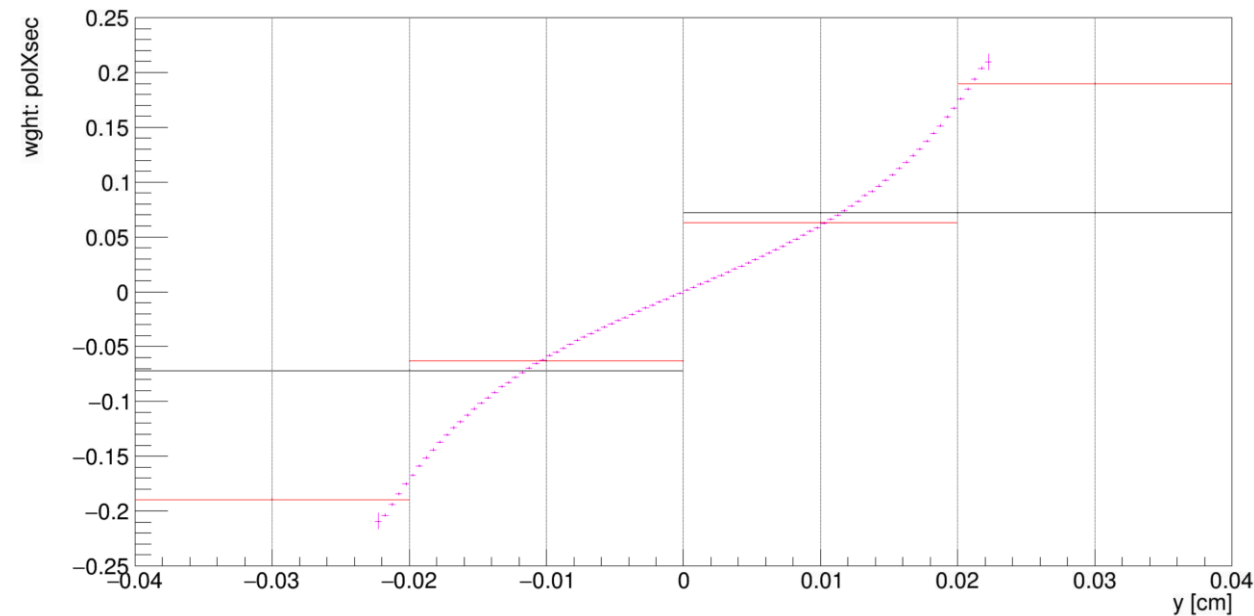
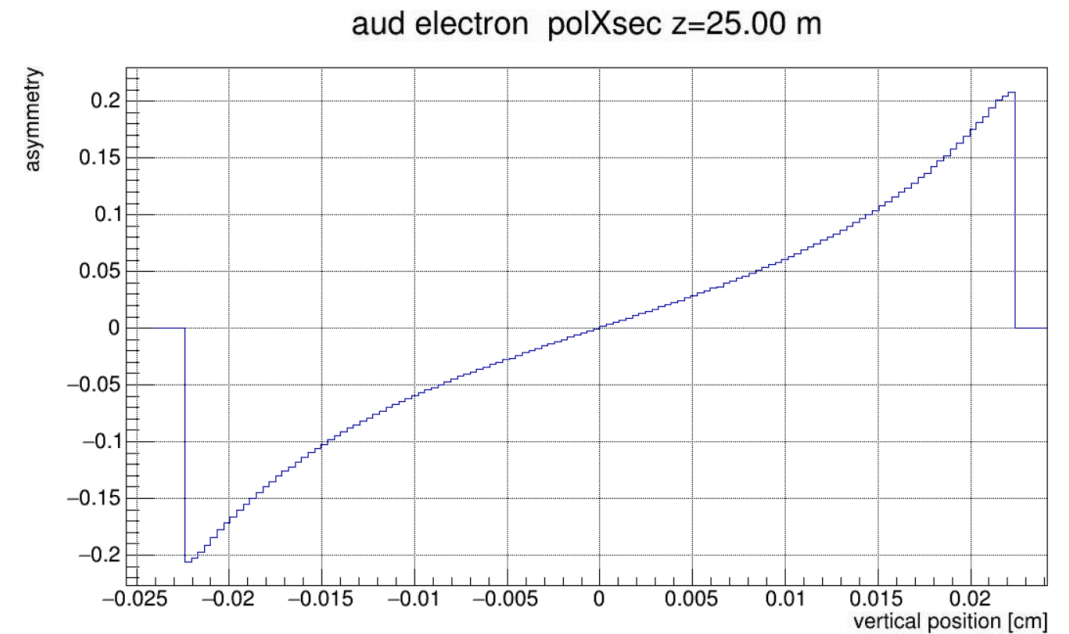
- For the “detector” analysis the left side should be removed (not done in the following)



# Electron detector segmentation

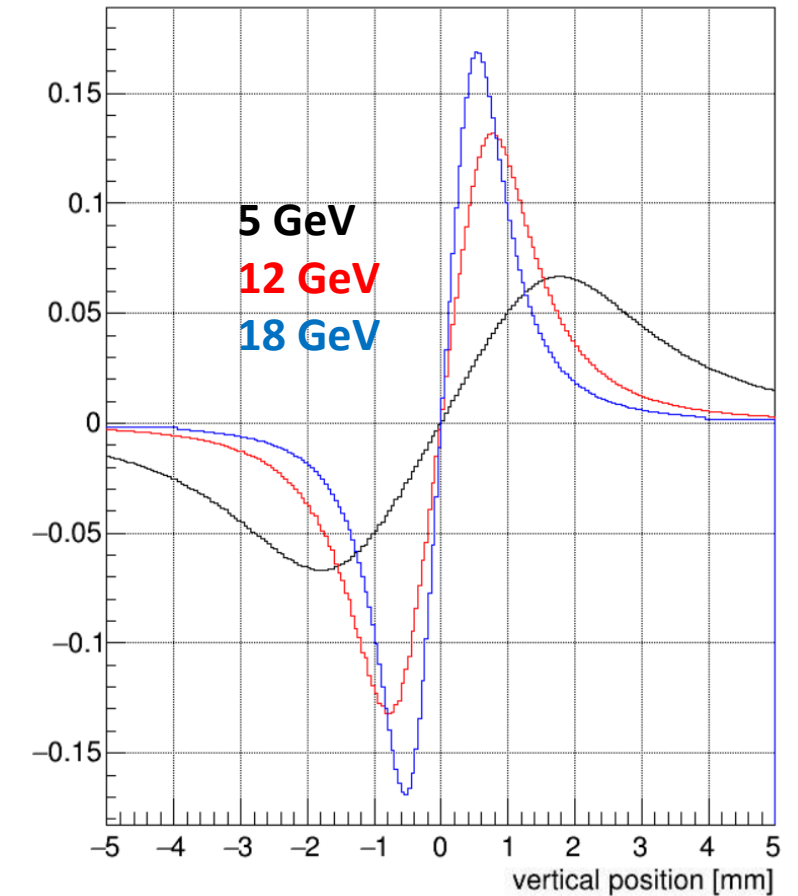
| segmentation<br>[um] | norm  | uncertainty |
|----------------------|-------|-------------|
| 400                  | 30.53 | 0.06        |
| 200                  | 75.71 | 0.16        |
| 100                  | 73.74 | 0.14        |
| 50                   | 73.43 | 0.13        |
| 10                   | 73.01 | 0.11        |
| 5                    | 73.00 | 0.11        |

- Similar to the analysis Dave showed I vertically “segmented” and normalized (0.73) the asymmetry and fit it to the expected simulation asymmetry
- We can see that event at 50um we have a shift in the “extracted” polarization



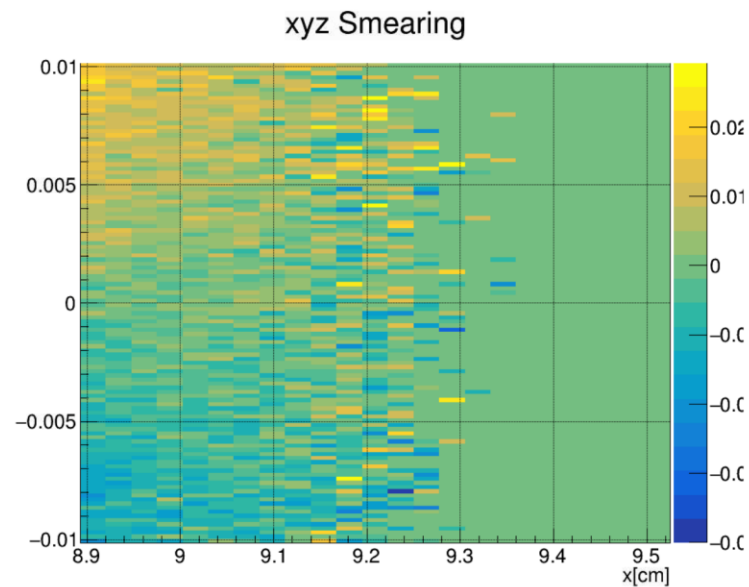
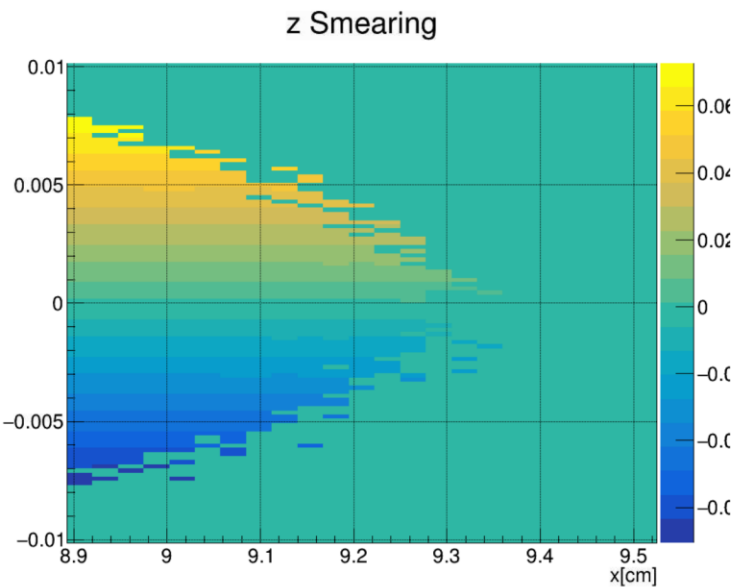
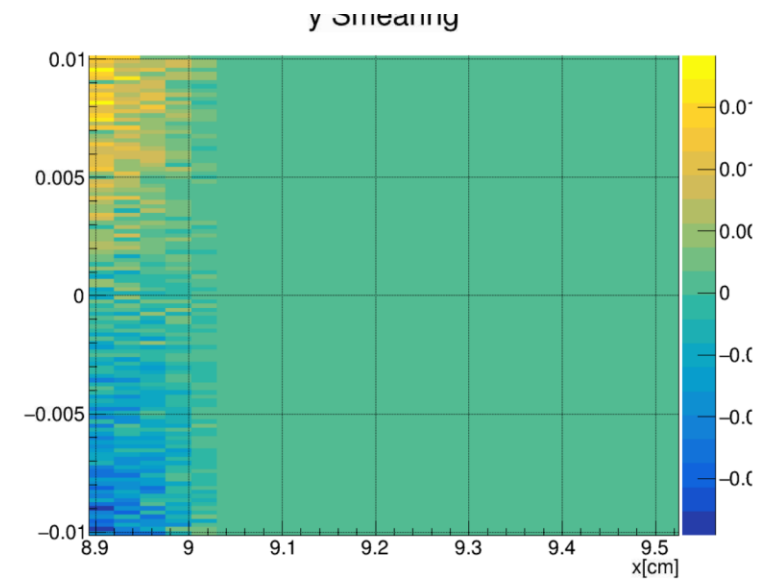
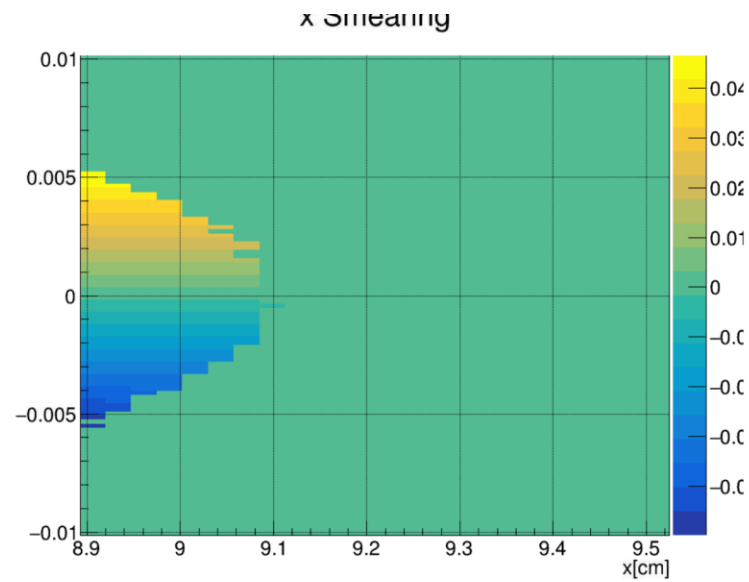
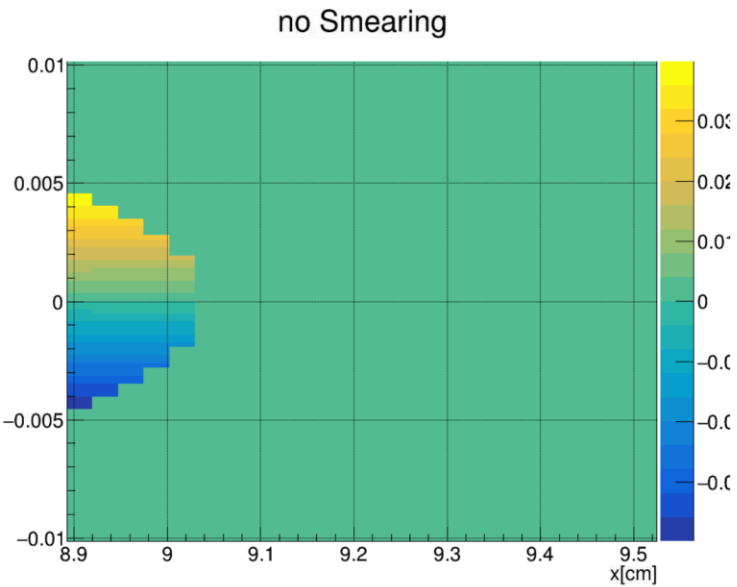
# To do

- Redo analysis with smeared distributions and potential detector area
- Fold in luminosity and get a time estimate for 1% measurements
- Repeat for other energies
- Look at IP6



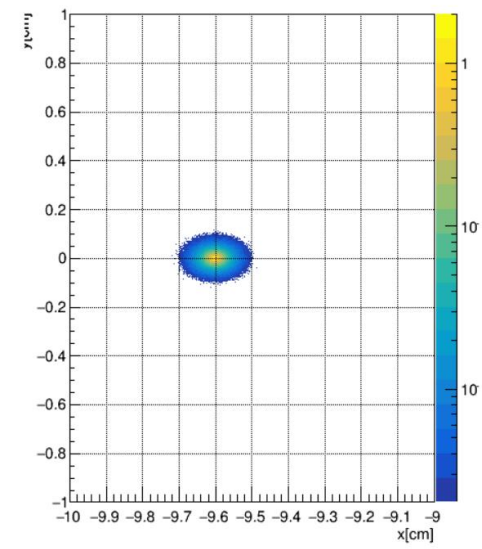
# Backup

# Vertex smearing

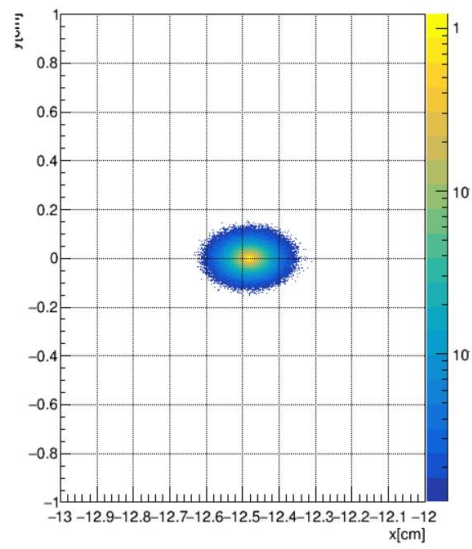


# Xsection

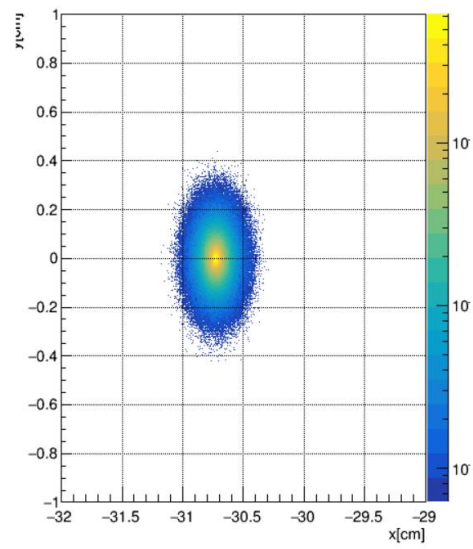
gamma unpolXsec z=5.00 m



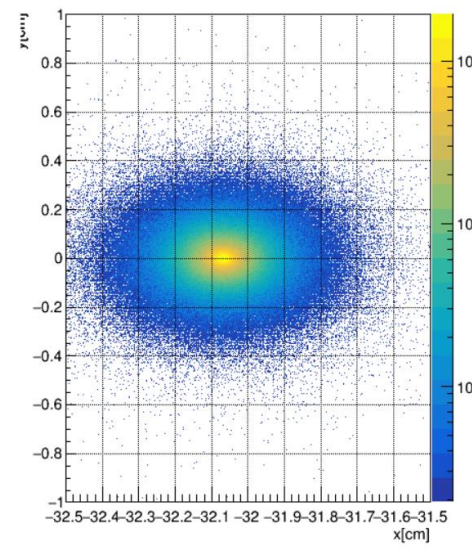
gamma unpolXsec z=6.50 m



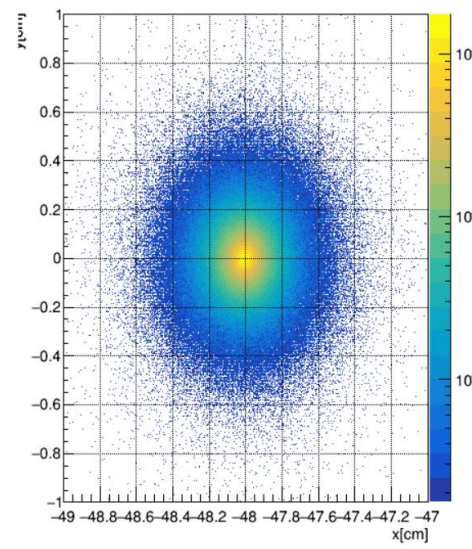
gamma unpolXsec z=16.00 m



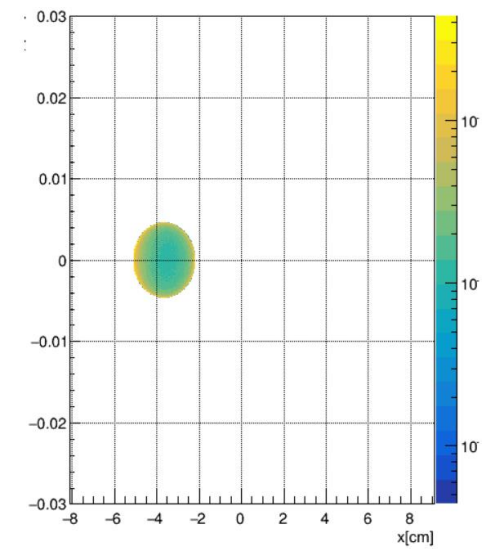
gamma unpolXsec z=16.70 m



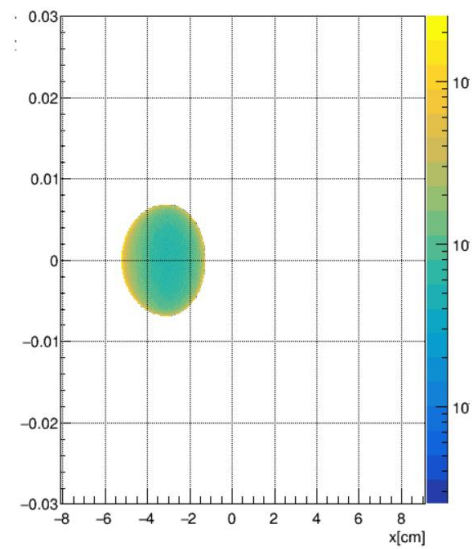
gamma unpolXsec z=25.00 m



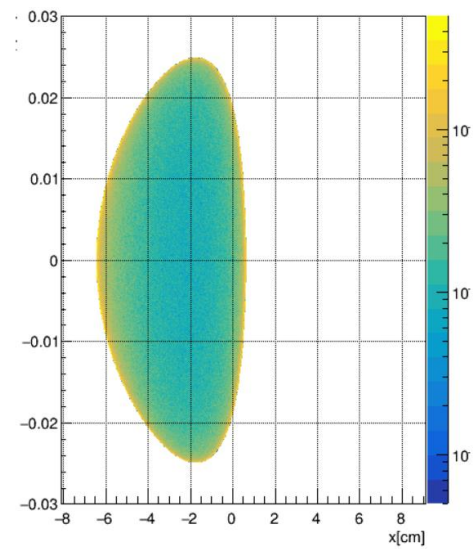
electron unpolXsec z=5.00 m



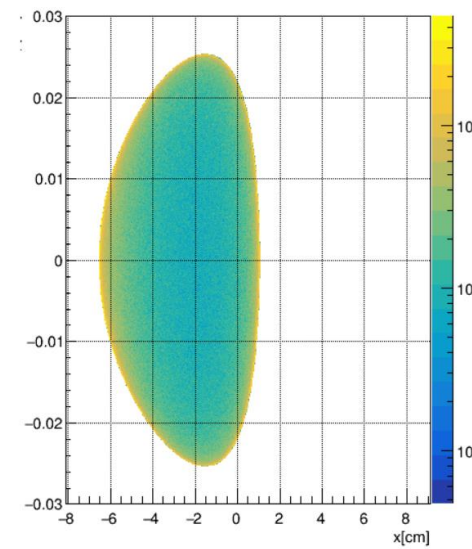
electron unpolXsec z=6.50 m



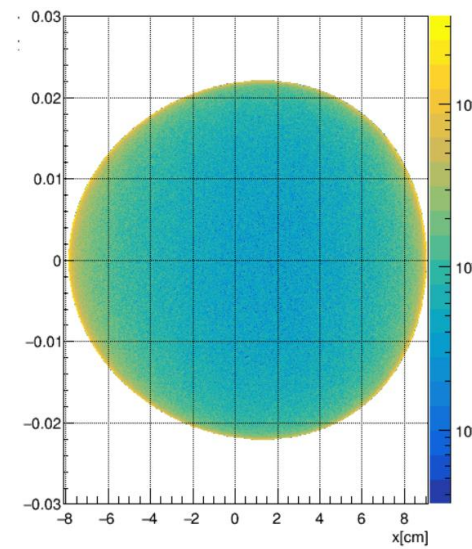
electron unpolXsec z=16.00 m



electron unpolXsec z=16.70 m

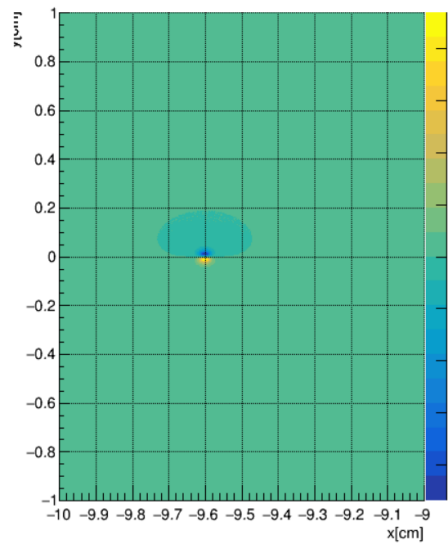


electron unpolXsec z=25.00 m

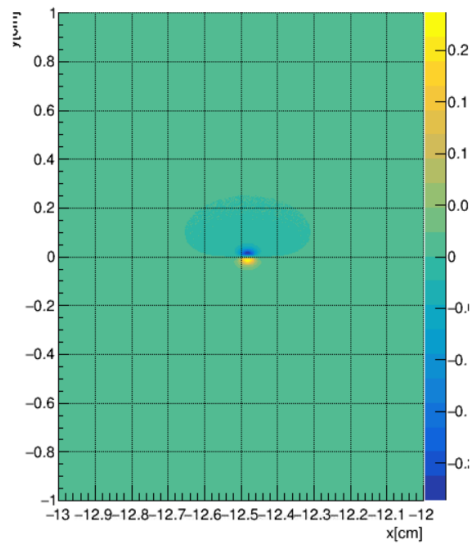


# Asym

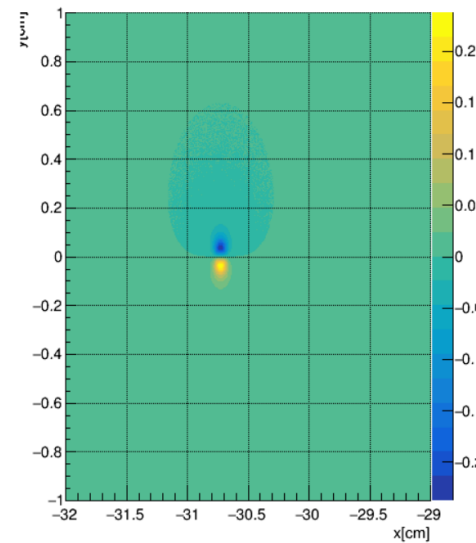
gamma polXsec z=5.00 m



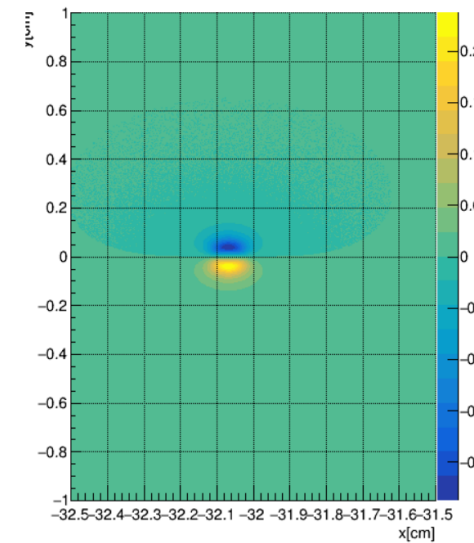
gamma polXsec z=6.50 m



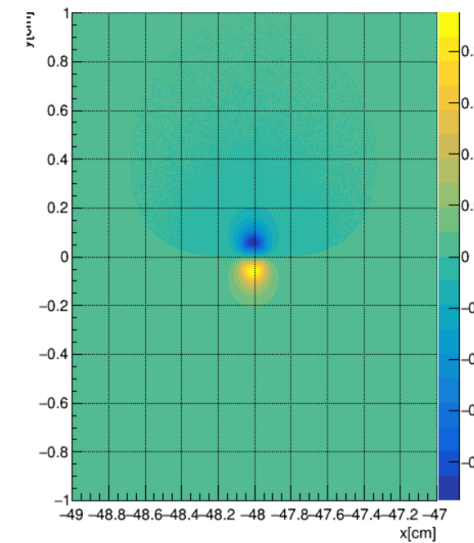
gamma polXsec z=16.00 m



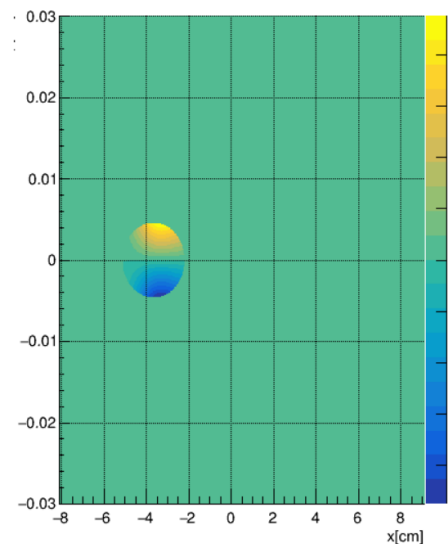
gamma polXsec z=16.70 m



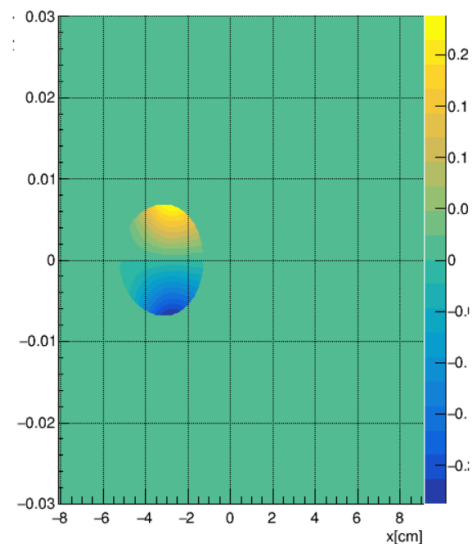
gamma polXsec z=25.00 m



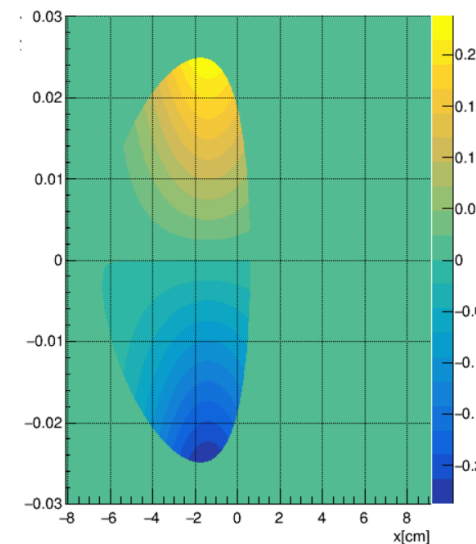
electron polXsec z=5.00 m



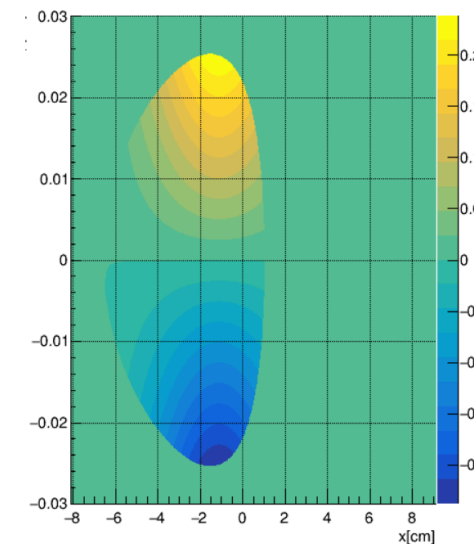
electron polXsec z=6.50 m



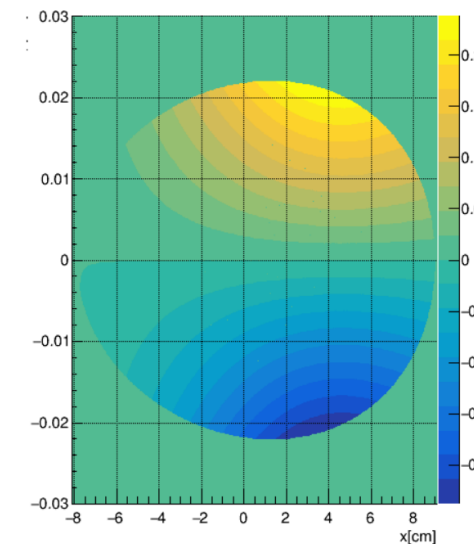
electron polXsec z=16.00 m



electron polXsec z=16.70 m

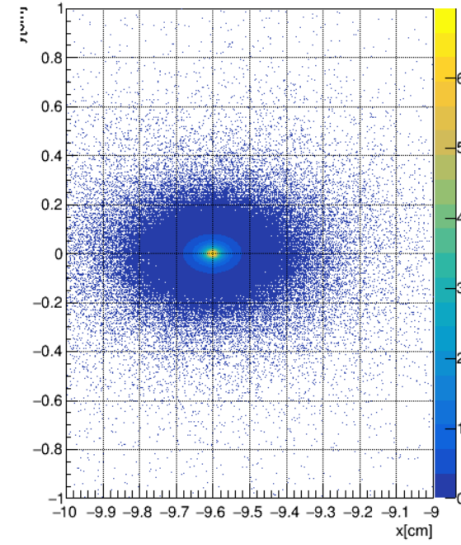


electron polXsec z=25.00 m

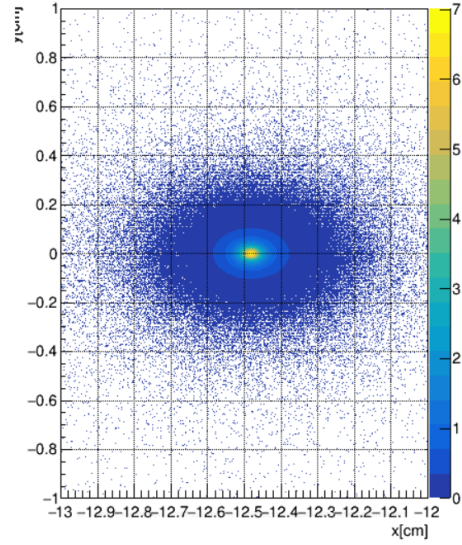


# <Energy>

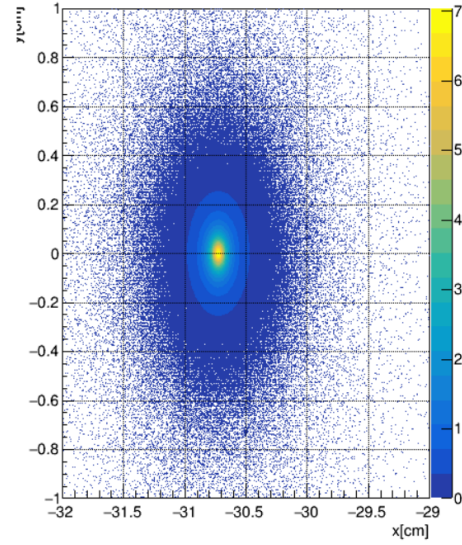
gamma energy\*unpolXsec z=5.00 m



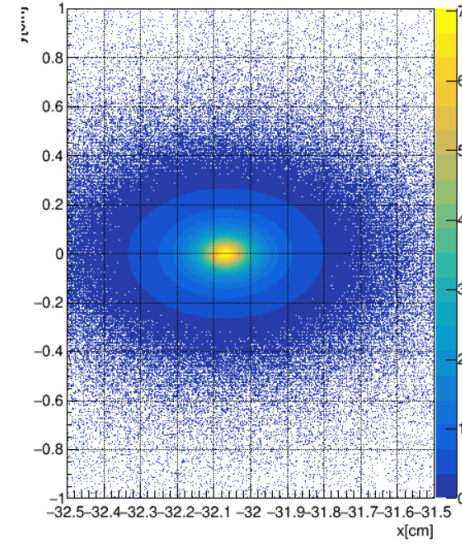
gamma energy\*unpolXsec z=6.50 m



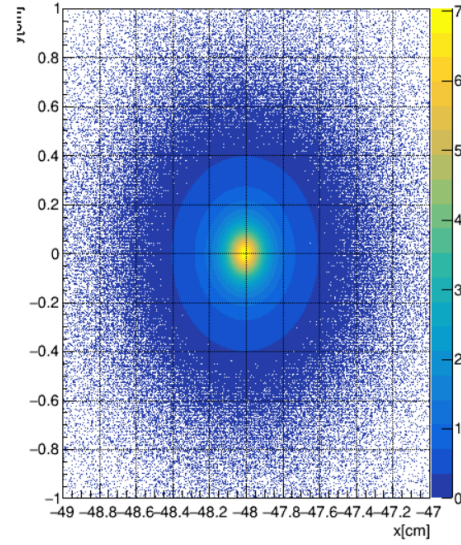
gamma energy\*unpolXsec z=16.00 m



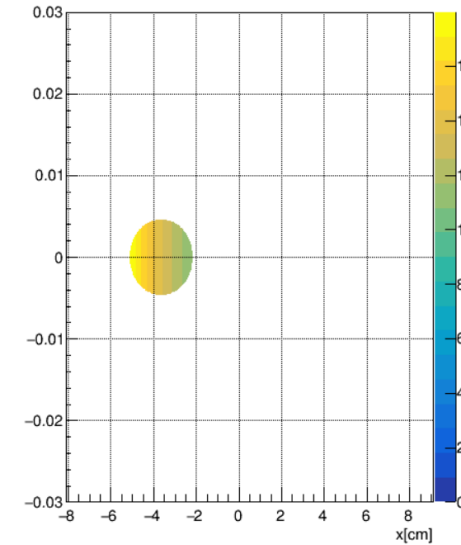
gamma energy\*unpolXsec z=16.70 m



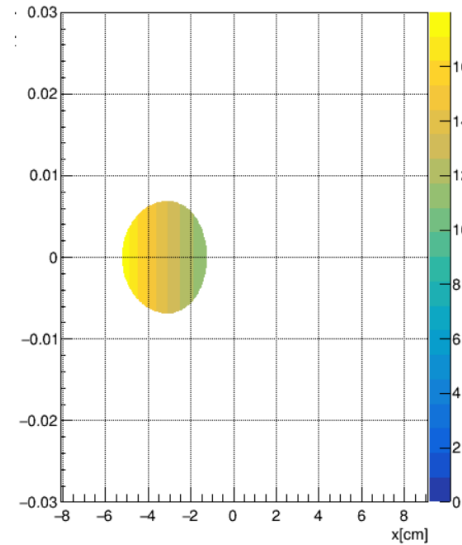
gamma energy\*unpolXsec z=25.00 m



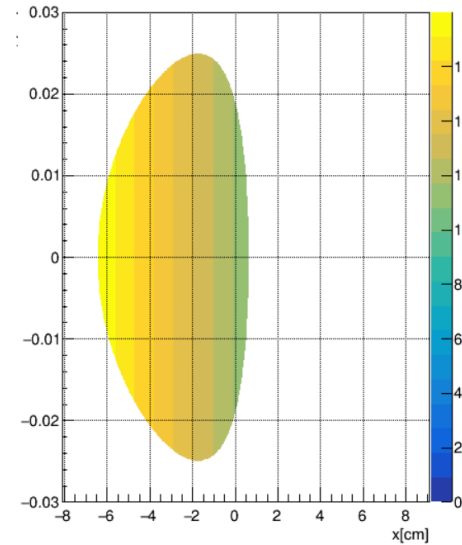
electron energy\*unpolXsec z=5.00 m



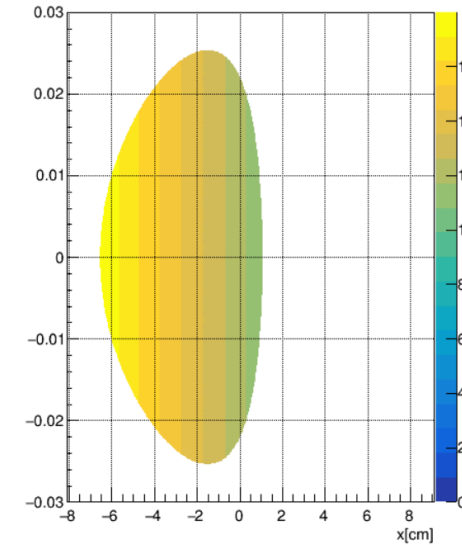
electron energy\*unpolXsec z=6.50 m



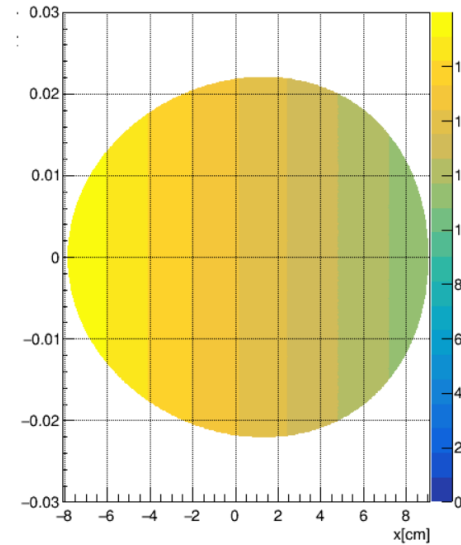
electron energy\*unpolXsec z=16.00 m



electron energy\*unpolXsec z=16.70 m

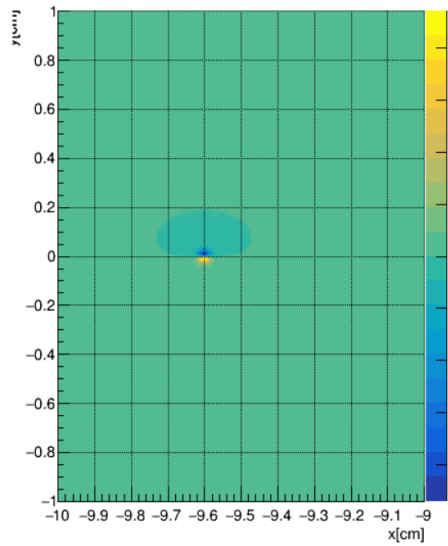


electron energy\*unpolXsec z=25.00 m

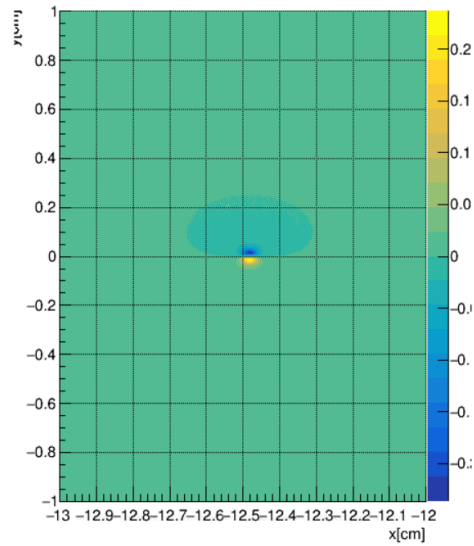


# $\langle \text{Energy} * \text{Asym} \rangle$

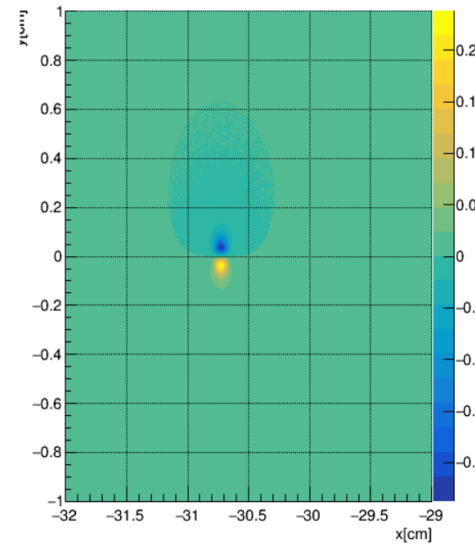
gamma energy\*polXsec z=5.00 m



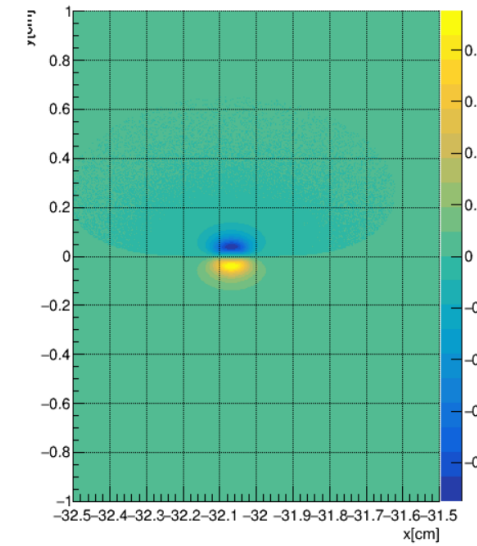
gamma energy\*polXsec z=6.50 m



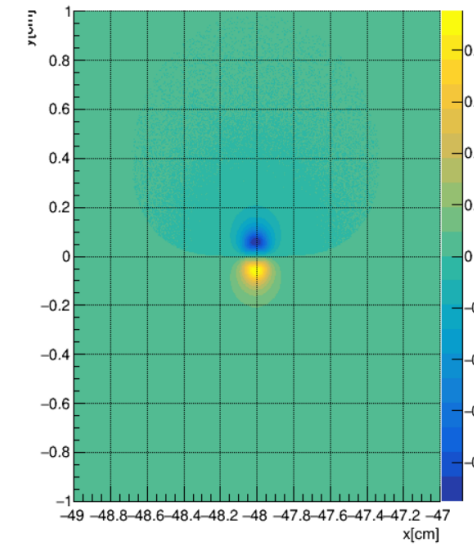
gamma energy\*polXsec z=16.00 m



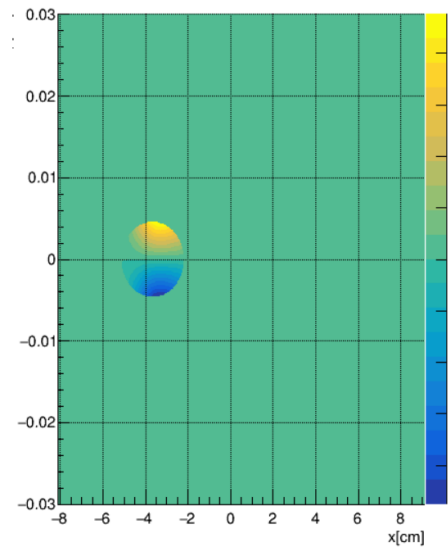
gamma energy\*polXsec z=16.70 m



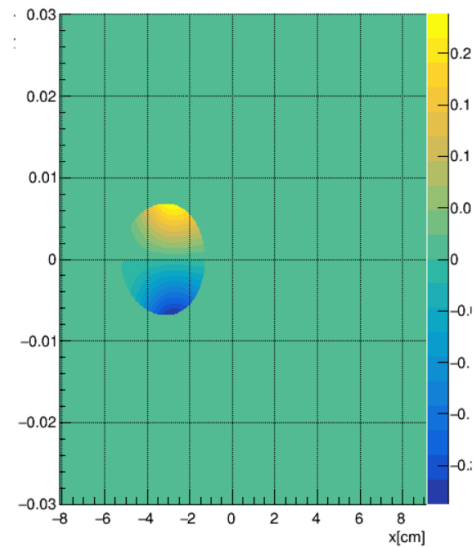
gamma energy\*polXsec z=25.00 m



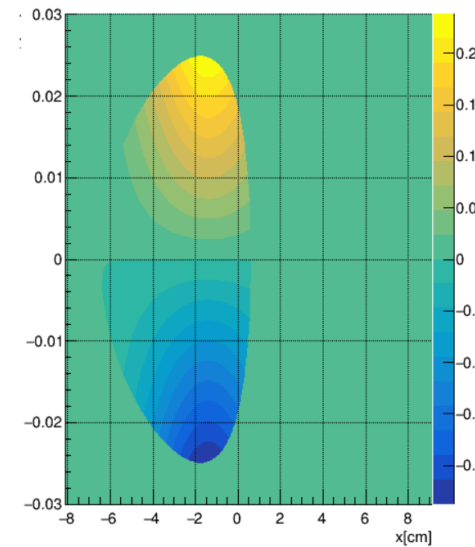
electron energy\*polXsec z=5.00 m



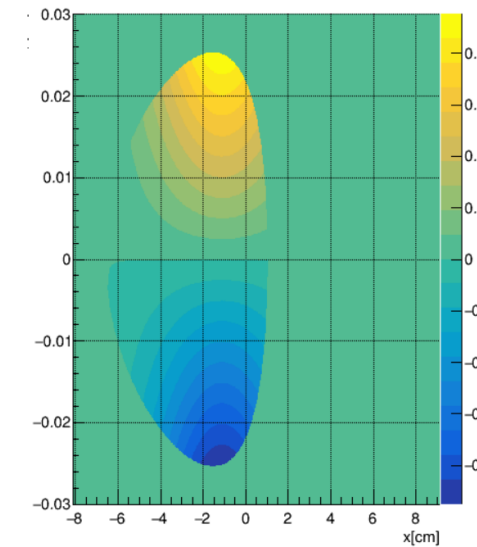
electron energy\*polXsec z=6.50 m



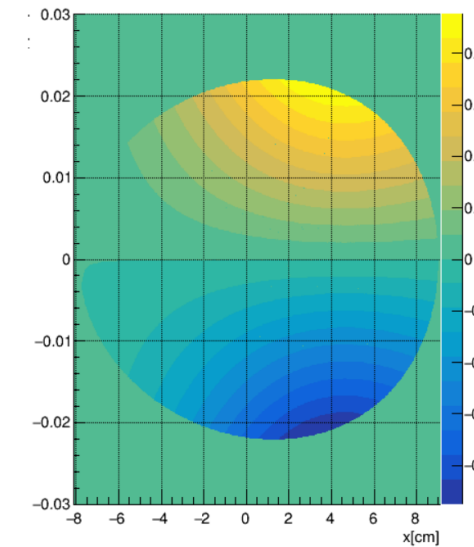
electron energy\*polXsec z=16.00 m



electron energy\*polXsec z=16.70 m

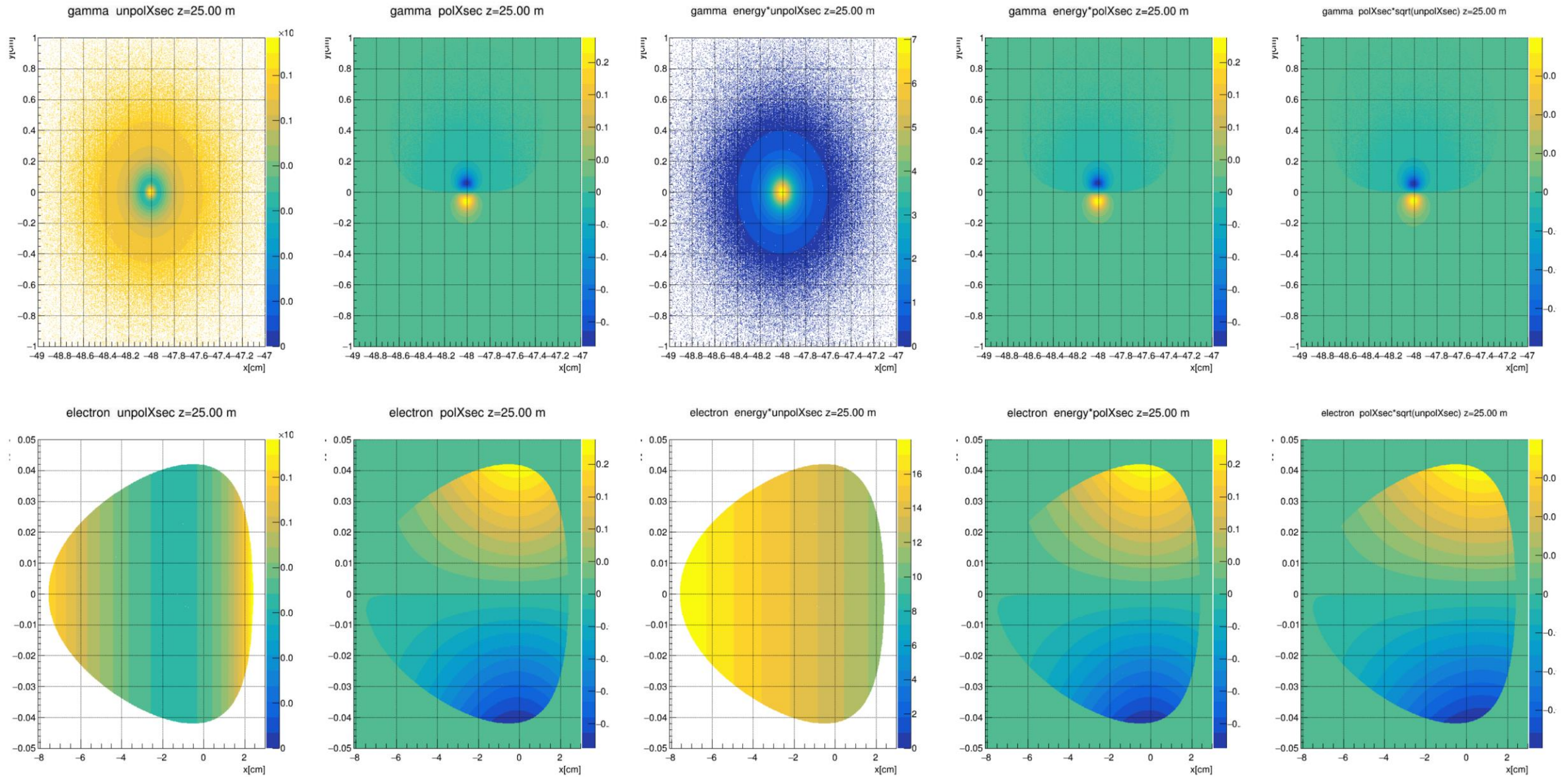


electron energy\*polXsec z=25.00 m



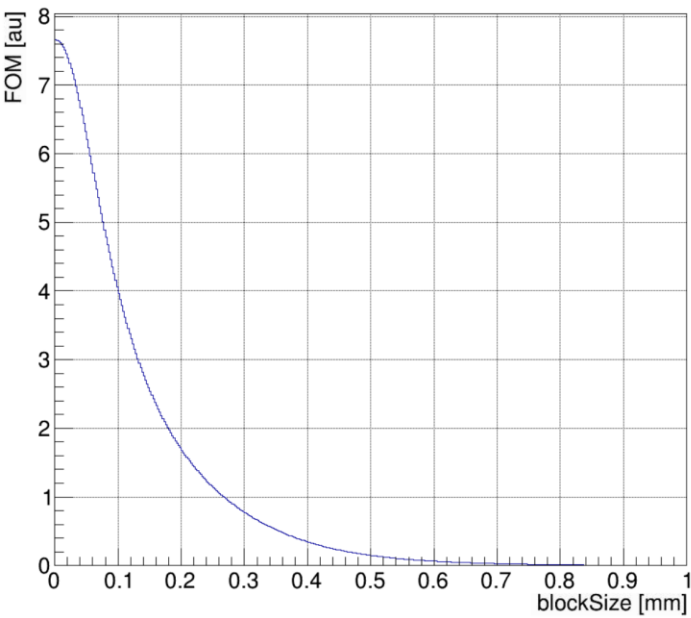


# 25m (det plane): noQD10

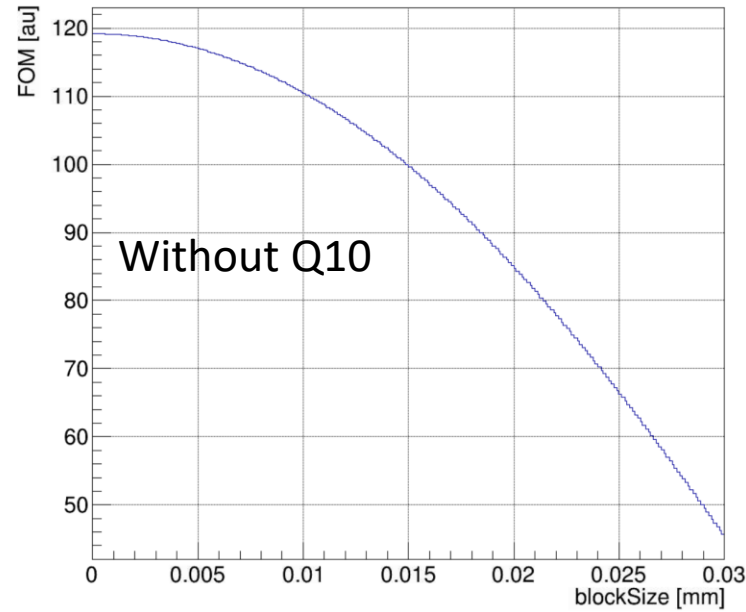


# AUD block measurements – with Bfields

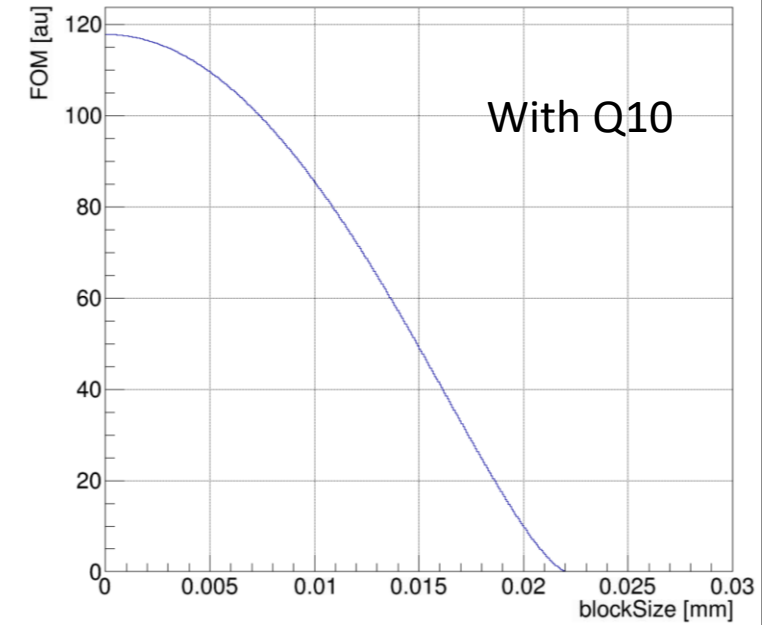
FOM center=0.00 gamma polXsec\*sqrt(unpolXsec) z=25.00 m



FOM center=0.00 electron polXsec\*sqrt(unpolXsec) z=25.00 m

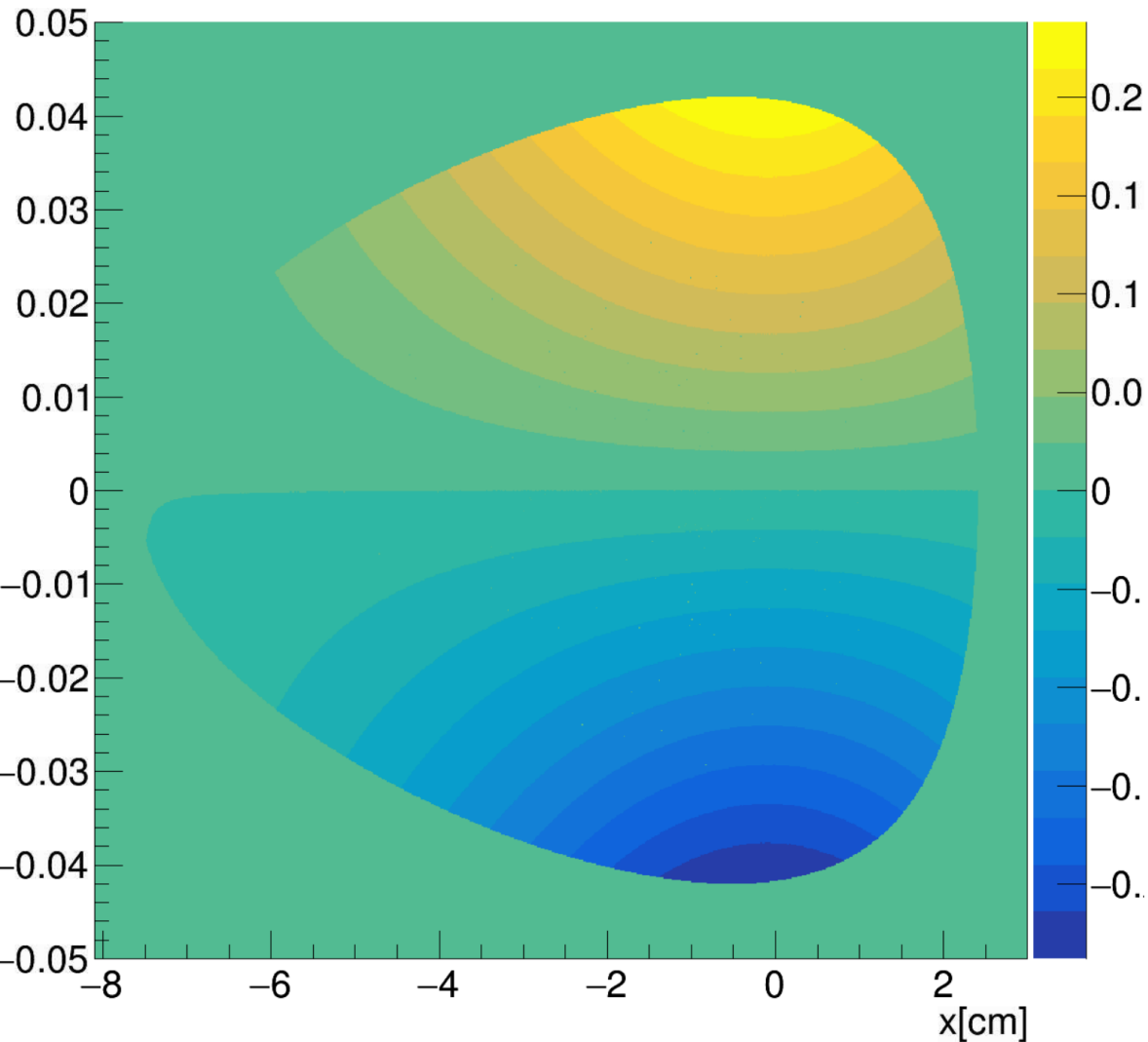


FOM center=0.00 electron polXsec\*sqrt(unpolXsec) z=25.00 m



# Ideal vs detector segmentation

electron polXsec z=25.00 m



average asymmetry

X= 5mm  
Y = 50um

