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# 相对论同量异位素核碰撞中的

# 核结构研究

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## 湖州师范学院

原子核结构与中高能重离子碰撞交叉学科理论讲习班,

湖州, 2021年7月9-24日

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## I. 相对论重离子碰撞简介

## 相对论重离子碰撞





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相对论重离子碰撞的特点

- •极高的能量(RHIC, v=0.99995c)
- •极短的碰撞时间 fm/c~10-23s
- 低重子数密度
- •极高温度(QGP,夸克胶子等离子体)
- •极强的磁场
- •极大的涡旋

## 相对论重离子碰撞的观测量

- •粒子产额(统计模型;反物质产生)
- •集体流(各向异性流及其关联; QGP性质)
- 电荷关联 (手征磁效应等)
- •高阶矩(QCD相变)
- 电磁探针、重味探针、喷注淬火等等

## 相对论重离子碰撞的常用模型

- •相对论流体力学模型
- AMPT模型
- UrQMD模型
- Hijing模型

### • • • • • •





## 初始几何特性





$$\rho(r) = \rho_0 \frac{1 + w \left(\frac{r}{R}\right)^2}{1 + e^{(r-R)/a}}$$



Nucleus	R [fm]	a [fm]	w [fm]
<sup>2</sup> H	0.01	0.5882	0
<sup>16</sup> O	2.608	0.513	-0.51
<sup>28</sup> Si	3.34	0.580	-0.233
$^{32}S$	2.54	2.191	0.16
$^{40}$ Ca	3.766	0.586	-0.161
<sup>58</sup> Ni	4.309	0.517	-0.1308
<sup>62</sup> Cu	4.2	0.596	0
<sup>186</sup> W	6.58	0.480	0
<sup>197</sup> Au	6.38	0.535	0
$ ^{207}$ Pb <sup>a</sup>	6.62	0.546	0
<sup>238</sup> U	6.81	0.6	0





Michael L. Miller, Klaus Reygers, Stephen J. Sanders, Peter Steinberg, Ann.Rev.Nucl.Part.Sci. 57 (2007) 205-243

## Optical Glauber model

• Thickness function

$$T(s_T) = \int_{-\infty}^{\infty} \rho(s_T, z) dz$$



- The probability of a nucleon in A interact with a nucleon in B is  $p_{AB}(b) = \sigma_{in}^{NN} \frac{\int dx dy T_A(x + b/2, y) T_B(x - b/2, y)}{AB}$
- For a given nucleon in A, it interact with a nucleon in B is  $p_A(x,y;b) = \sigma_{in}^{NN} \frac{T_B(x-b/2,y)}{B}$

## A+B cross section

• The probability of having n binary collision is

$$P(n) = \binom{AB}{n} p_{AB}^n (1 - p_{AB})^{AB - n}$$



• The total cross section of an interaction between A and B is

$$\frac{d^2 \sigma_{in}^{A+B}}{db^2} = \sum_{\substack{i=1\\i=1}}^{AB} P(n;b) = 1 - (1 - p_{AB})^{AB}$$
$$\approx 1 - e^{-\sigma_{in}^{NN} \int dx dy T_A(x+b/2,y) T_B(x-b/2,y)}$$

Number of binary collisons

$$N_{coll}(b) = \sum_{i=1}^{AB} nP(n; b) = \sigma_{in}^{NN} \int dx dy T_A(x + b/2, y) T_B(x - b/2, y)$$

## Number of participant(wounded nucleons)

$$N_{part}(b) = \int dx dy T_A(x + b/2, y) \left\{ 1 - \left[ 1 - \sigma_{in}^{NN} \frac{T_B(x - b/2, y)}{B} \right]^B \right\}$$
  
+  $\int dx dy T_B(x - b/2, y) \left\{ 1 - \left[ 1 - \sigma_{in}^{NN} \frac{T_A(x + b/2, y)}{A} \right]^A \right\}$ 

### PHYSICAL REVIEW C 94, 054903 (2016)

### Cumulants of multiplicity distributions in most-central heavy-ion collisions

Hao-jie Xu\*

Department of Physics and State Key Laboratory of Nuclear Physics and Technology, Peking University, Beijing 100871, China (Received 17 April 2016; revised manuscript received 11 August 2016; published 11 November 2016)

I investigate the volume corrections on cumulants of total charge distributions and net proton distributions. The required volume information is generated by an optical Glauber model. I find that the corrected statistical expectations of multiplicity distributions mimic the negative binomial distributions at noncentral collisions, and

they tend to approach the Poisson ones at most-central collis the volume corrections. However, net proton distributions at to the external volume fluctuations at most-central collisions, volume distributions in event-by-event multiplicity fluctuation

优点: 解析公式,不占计算资源 缺点: 没有涨落



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### Glauber modeling in high energy nuclear collisions

Michael L. Miller (MIT, LNS), Klaus Reygers (Munster U.), Stephen J. Sanders (Kansas U.), Peter Steinberg (Brookhaven) (Jan, 2007) Published in: *Ann.Rev.Nucl.Part.Sci.* 57 (2007) 205-243 • e-Print: nucl-ex/0701025 [nucl-ex]

🔓 pdf 🥜 DOI 🖃 cite

### The PHOBOS Glauber Monte Carlo

B. Alver (MIT), M. Baker (Brookhaven), C. Loizides (MIT), P. Steinberg (Brookhaven) (May, 2008) e-Print: 0805.4411 [nucl-ex]

🛱 pdf 🖃 cite



➔ 1,467 citations

#1

#3



The effect of triangular flow on di-hadron azimuthal correlations in relativistic heavy ion collisions Jun Xu (Texas A-M), Che Ming Ko (Texas A-M, Cyclotron Inst. and Texas A-M) (Nov, 2010)				
Published in: Phys.Rev.C 83 (2011) 021903 • e-Print: 1011.3750 [nucl-th]				
∄ pdf ∂ DOI ⊑ cite				
Elliptic and triangular flow in event-by-event (3+1)D viscous hydrodynamics	#2			
Bjorn Schenke (McGill U.), Sangyong Jeon (McGill U.), Charles Gale (McGill U.) (Sep. 2010)				
Published in: Phys.Rev.Lett. 106 (2011) 042301 • e-Print: 1009.3244 [hep-ph]				
₽ pdf	€ 611 citations			
Triangular flow in event-by-event ideal hydrodynamics in Au+Au collisions at $\sqrt{s_{ m NN}}=200A$ GeV	#3			
Hannah Petersen (Duke U.), Guang-You Qin (Duke U.), Steffen A. Bass (Duke U.), Berndt Muller (Duke U.) (Aug, 2010)				
Published in: Phys.Rev.C 82 (2010) 041901 • e-Print: 1008.0625 [nucl-th]				
₽ pdf ∂ DOI Ξ cite	➔ 202 citations			
Triangular flow in hydrodynamics and transport theory	#4			
Burak Han Alver (MIT, LNS), Clement Gombeaud (Saclay, SPhT), Matthew Luzum (Saclay, SPhT), Jean-Yves Ollitrault (Saclay, SPhT)	(Jul, 2010)			
Published in: Phys.Rev.C 82 (2010) 034913 • e-Print: 1007.5469 [nucl-th]				
₽ pdf @ DOI	➔ 345 citations			
Collision geometry fluctuations and triangular flow in heavy-ion collisions	#5			
B. Alver (MIT), G. Roland (MIT) (Mar, 2010)				
Published in: Phys.Rev.C 81 (2010) 054905, Phys.Rev.C 82 (2010) 039903 (erratum) • e-Print: 1003.0194 [nucl-th]				

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🔓 pdf 🕜 DOI 🖃 cite

蒙特卡洛Glauber模型的一些细节

• 如何判断碰撞发生

Step function

Gaussian function

PHYSICAL REVIEW C 84, 064913 (2011)

## Wounded-nucleon model with realistic nucleon-nucleon collision profile and observables in relativistic heavy-ion collisions

Maciej Rybczyński<sup>1,\*</sup> and Wojciech Broniowski<sup>2,1,†</sup> <sup>1</sup>Institute of Physics, Jan Kochanowski University, PL-25406 Kielce, Poland <sup>2</sup>The H. Niewodniczański Institute of Nuclear Physics, Polish Academy of Sciences, PL-31342 Kraków, Poland (Received 1 September 2011; published 22 December 2011)

$$p(b) = Ae^{-\pi Ab^2/\sigma_{\rm inel}}.$$

 $d \leq \sqrt{\sigma_{\text{inel}}^{\text{NN}}/\pi}$ 

• Recenter(1)



Eccentricity (Point-like nucleons)

$$\psi_n \{\text{part}\} = \frac{1}{n} \left[ \arctan \frac{\langle r^n \sin(n\varphi) \rangle}{\langle r^n \cos(n\varphi) \rangle} + \pi \right] \text{Recenter(2)}$$
$$\varepsilon_n \{\text{part}\} = \frac{\sqrt{\langle r^n \cos(n\varphi) \rangle^2 + \langle r^n \sin(n\varphi) \rangle^2}}{\langle r^n \rangle},$$

 $\varepsilon_n(b), \varepsilon_n(N_{part})$ 

$$\varepsilon_n(N_{ch}),$$
  

$$N_{ch} = n_{pp}(xN_{coll} + (1 - x)N_{part}/2)$$



## Energy (entropy) density

Two-component Glauber

$$s_0(\boldsymbol{x}_{\perp}) = \frac{dS}{\tau_0 dx dy d\eta_s} \bigg|_{\eta_s = 0}$$
  
=  $\frac{C}{\tau_0} \left( \frac{1 - \delta}{2} \frac{dN_{\text{part}}}{d^2 x_{\perp}} + \delta \frac{dN_{\text{coll}}}{d^2 x_{\perp}} \right).$ 

### Color Glass condensate (CGC)

$$s_0(\mathbf{x}_{\perp}) = 3.6n_g$$
  
=  $3.6 \left. \frac{dN_g}{\tau_0 d^2 x_{\perp} d\eta_s} \right|_{y=\eta_s=0}$ .

$$\begin{aligned} \frac{dN_g}{d^2x_\perp dy} &= \frac{2\pi^2}{C_F} \int \frac{d^2p_T}{p_T^2} \int^{p_T} \frac{d^2k_T}{4} \alpha_s(Q^2) \\ &\times \phi_A(x_1, (\boldsymbol{p}_T + \boldsymbol{k}_T)^2/4; \boldsymbol{x}_\perp) \\ &\times \phi_B(x_2, (\boldsymbol{p}_T - \boldsymbol{k}_T)^2/4; \boldsymbol{x}_\perp), \end{aligned}$$

$$Q_{s,A}^2(x; \boldsymbol{x}_\perp) = Q_{s,0}^2 \frac{T_A(\boldsymbol{x}_\perp)}{T_{A,0}} \left(\frac{x_0}{x}\right)^{\lambda}$$

## Energy (entropy) density

### Trento model



### PHYSICAL REVIEW C 92, 011901(R) (2015)

### Alternative ansatz to wounded nucleon and binary collision scaling in high-energy nuclear collisions

J. Scott Moreland, Jonah E. Bernhard, and Steffen A. Bass Department of Physics, Duke University, Durham, North Carolina 27708-0305, USA (Received 23 December 2014; revised manuscript received 11 May 2015; published 6 July 2015)

We introduce a new parametric initial-condition model for high-energy nuclear collisions based on eikonal entropy deposition via a "reduced-thickness" function. The model simultaneously describes experimental protonproton, proton-nucleus, and nucleus-nucleus multiplicity distributions and generates nucleus-nucleus eccentricity harmonics consistent with experimental flow constraints. In addition, the model is compatible with ultracentral uranium-uranium data unlike existing models that include binary collision terms.

$$f = T_R(p; T_A, T_B) \equiv \left(\frac{T_A^p + T_B^p}{2}\right)^{1/p},$$

$$T_R = \begin{cases} \max(T_A, T_B), & p \to +\infty \\ (T_A + T_B)/2, & p = +1 \text{ (arithmetic)} \\ \sqrt{T_A T_B}, & p = 0 \text{ (geometric)} \\ 2T_A T_B/(T_A + T_B), & p = -1 \text{ (harmonic)} \\ \min(T_A, T_B), & p \to -\infty. \end{cases}$$

## Collision geometry and anisotropic flow



With the WS densities, we have made very successful connections between the final flow observable and the initial collision geometry.

Prefect fluid - strong coupling QGP (sQGP)



$$\frac{dN}{d\phi} = N(1 + 2\sum_{n} \nu_{n} \cos\left[n(\phi - \Psi_{n})\right]))$$

STAR, PRC 77, 054901 (2018)

## Q vector

### PHYSICAL REVIEW C 83, 044913 (2011)

### Flow analysis with cumulants: Direct calculations

Ante Bilandzic,<sup>1,2</sup> Raimond Snellings,<sup>2</sup> and Sergei Voloshin<sup>3</sup> <sup>1</sup>Nikhef, Science Park 105, NL-1098 XG Amsterdam, The Netherlands <sup>2</sup>Utrecht University, P.O. Box 80000, NL-3508 TA Utrecht, The Netherlands <sup>3</sup>Wayne State University, 666 West Hancock Street, Detroit, Michigan 48201, USA (Received 6 October 2010; published 26 April 2011)

Anisotropic flow measurements in heavy-ion collisions provide important information on the properties of hot and dense matter. These measurements are based on analysis of azimuthal correlations and might be biased by contributions from correlations that are not related to the initial geometry, so-called nonflow. To improve anisotropic flow measurements, advanced methods based on multiparticle correlations (cumulants) have been developed to suppress nonflow contribution. These multiparticle correlations can be calculated by looping over all possible multiplets, however, this quickly becomes prohibitively CPU intensive. Therefore, the most used technique for cumulant calculations is based on generating functions. This method involves approximations, and has its own biases, which complicates the interpretation of the results. In this paper we present a new exact method for direct calculations of multiparticle cumulants using moments of the flow vectors.

$$Q_n \equiv \sum_{i=1}^M e^{in\phi_i},$$

$$\langle 2 \rangle = \frac{|Q_n|^2 - M}{M(M-1)}.$$

$$\langle 4 \rangle = \frac{|Q_n|^4 + |Q_{2n}|^2 - 2 \cdot \operatorname{Re}[Q_{2n}Q_n^*Q_n^*]}{M(M-1)(M-2)(M-3)} -2\frac{2(M-2) \cdot |Q_n|^2 - M(M-3)}{M(M-1)(M-2)(M-3)}.$$

## Hydrodynamics and flow

 $\partial_{\mu}T^{\mu\nu}(x) = 0$  and  $\partial_{\mu}j^{\mu}(x) = 0$ ,



22 March 2001

PHYSICS LETTERS B

Physics Letters B 503 (2001) 58-64

www.elsevier.nl/locate/npe

### Radial and elliptic flow at RHIC: further predictions

P. Huovinen<sup>a</sup>, P.F. Kolb<sup>b,c</sup>, U. Heinz<sup>b</sup>, P.V. Ruuskanen<sup>d</sup>, S.A. Voloshin<sup>e</sup>

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 <sup>b</sup> Department of Physics, The Ohio State University, 174 West 18th Avenue, Columbus, OH 43210, USA
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 <sup>d</sup> Department of Physics, University of Jyväskylä, FIN-40341 Jyväskylä, Finland
 <sup>c</sup> Department of Physics and Astronomy, Wayne State University, 666 W. Hancock Street, Detroit, MI 48202, USA

Received 12 January 2001; accepted 30 January 2001 Editor: R. Gatto



Fig. 6. Simple source of four fireballs.

$$v_2(p_t) \approx \tanh\left(\frac{1}{2}\left(\frac{\kappa p_t - \lambda m_t}{T} + \mu\right)\right),$$

## flow fluctuation & correlations

### Collective flow in 2.76 A TeV and 5.02 A TeV Pb+Pb collisions

Wenbin Zhao (Peking U. and Peking U., SKLNPT), Hao-jie Xu (Peking U. and Peking U., SKLNPT), Huichao Song (CICQM, Beijing and Peking U., CHEP and Peking U., SKLNPT) (Mar 31, 2017)

#14

Published in: Eur.Phys.J.C 77 (2017) 9, 645 • e-Print: 1703.10792 [nucl-th]





Effect of deformation

$$\rho(r) = \frac{\rho_0}{1 + \exp[\frac{r - R(1 + \beta_2 Y_{20} + \beta_4 Y_{40})}{a}]}$$

Au:  $\beta_2 = -0.131$ ; U:  $\beta_2 = 0.28$ 

Knee structure of v2 distributions at most-central U+U collisions



S. Voloshin, PRL105, 172301 (2010)

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## From Glauber to Trento





STAR, PRL115, 222301 (2015)

S. Voloshin, PRL105, 172301 (2010)

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## Prediction with Trento



G. Giacalone, PRL124, 202301 (2020)

Effect of cluster correlations

 $^{12}_{6}C + ^{208}_{82}Pb$ 





W. Broniowski, E. Arriola, PRL112, 112501 (2014)

The  $\alpha$ -clustered and uniform 12C have very different predictions on v3, its event-by-event fluctuations, or the correlations of the v2 and v3

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## 利用系统扫描甄别 $\alpha$ -cluster结构



✓ 非对称系统扫描, $v_3/v_2$ 网种构型具有明显的差别,WS构型非常平坦 ✓  $^{16}O + ^{197}$  Au中心度依赖,高多重数下 $v_3/v_2$ 的比,两种构型具有明显的差别 ✓ 对称系统扫描,明显看到四面体构型的 $^{16}O + ^{16}O$ 系统系的 $v_3/v_2$ 偏离系统学

Y.A. Li, S. Zhang, Y.G. Ma, Phys. Rev. C 102, 054907 (2020)

S. Zhang (张松), IMP, Fudan,

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## Chiral Magnetic Effect (CME)



- CME: a novel phenomenon predicted in HI collisions
  - Prerequisite: chiral imbalance+ magnetic field
  - Consequence: charge separation along B field
- Experimental search is challenging due to overwhelming background  $\rightarrow$  Isobar





*CME* measured in Au+Au collisions

STAR, arXiv:2106.09243



H. Xu, et.al, CPC42, 084103 (2018)

# III. 核结构对手征磁效应测量的影响

## Chiral magnetic effect



 $eB \sim m_{\pi}^2 @HIC$ 

D. Kharzeev, PPNP88, 1(2016)

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STAR, PRL103, 251601 (2009) ALICE, PRL110, 012301 (2013)

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Cluster decay + elliptic flow $(v_2)$ 

 $\langle \cos(\phi_{\alpha} + \phi_{\beta} - 2\Psi_{RP}) \rangle$   $\propto \quad \langle \cos(\phi_{\alpha} + \phi_{\beta} - 2\phi_{\rho} + 2\phi_{\rho} - 2\Psi_{RP}) \rangle$   $\simeq \quad \langle \cos(\phi_{\alpha} + \phi_{\beta} - 2\phi_{\rho}) \rangle \langle 2(\phi_{\rho} - |\Psi_{RP}) \rangle$   $= \quad \langle \cos(\phi_{\alpha} + \phi_{\beta} - 2\phi_{\rho}) \rangle v_{2}^{\rho}$ 

## Isobar collisions

The isobar collisions was proposed to measure the chiral magnetic effect. S. Voloshin, PRL105, 172301 (2010)



- Same eccentricities => flow background
- Different magnetic field => CME signals

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## DFT densities VS WS densities



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Instead of the WS densities, we use the densities obtained from the **density functional theory (DFT)** with parameter set SLy4.

### Static model: Monte Carlo Glauber model



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## AMPT simulations





AMPT model

J. Zhao (STAR), NPA982, 535(2019)

### **Background dominated**

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--- The CME signal, if exist, is very small





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## IV. 通过同量异位素核碰撞测量中子皮

## Neutron skin from antiprotonic atoms

VOLUME 87, NUMBER 8 PHYSICAL REVIEW LETTERS 20 AUGUST 2001

### **Neutron Density Distributions Deduced from Antiprotonic Atoms**



## Determine the neutron skin type

H. Xu, H. Li, X. Wang, C. Shen, F. Wang, PLB819, 136453 (2021)

	<sup>96</sup> Ru		<sup>96</sup> Zr	
	R	a	R	a
р	5.085	0.523	5.021	0.523
skin-type n	5.085	0.523	5.194	0.523
halo-type n	5.085	0.523	5.021	0.592



Haojie Xu, Huzhou University

Trento Model

- The Nch ratios bends up above unity at large Nch because the larter neutron skin thickness of Zr
- The shapes of the Ru+Ru/Zr+Zr ratios of the Nch distributions and eccentricity in mid-central collisions can further distinguish between skin-type and halo-type neutron densities, both having the same skin thickness.

eSHF	SHF
SHF: Standard Skyrme-Hartree-Fock (SHF) model	SLv4
eSHF: Extended SHF model $L(\rho_c)$ (MeV) 20.000 47.300 70.000	42.661
$F(\rho, \delta) = F_{\rho}(\rho) + F_{\rho}(\rho) \delta^{2} + O(\delta^{4}) \qquad \xrightarrow{E_{\text{sym}}(\rho_{c})(\text{MeV})} 26.650  26.650  26.650$	26.450
$L(p,0) = L_0(p) + L_{sym}(p)0 + O(0) \qquad \qquad$	0.15954
$\rho = \rho_n + \rho_p; \ \delta = \frac{r_n + \rho_p}{\rho}; \ \rho_c \simeq 0.11 fm^{-3}$ $t_0 (\text{MeV} \cdot \text{fm}^3)$ $-2063.0$ $-2037.3$ $-1855.3$	-2488.9
$\begin{bmatrix} dF & (0) \end{bmatrix} = \begin{bmatrix} t_1 (\text{MeV} \cdot \text{fm}^3) & 442.48 & 524.18 & 576.91 \\ t_1 (\text{MeV} \cdot \text{fm}^3) & 442.48 & 524.18 & 576.91 \\ t_2 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 524.68 & 576.91 \\ t_3 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 524.68 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 524.68 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 524.68 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 524.68 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 524.68 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 524.68 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 524.68 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 524.68 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 524.68 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 524.68 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 524.68 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 524.68 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 524.68 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 524.68 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 524.68 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 524.68 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 524.68 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 524.68 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 524.68 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 524.68 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 524.68 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 562.62 & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 576.91 \\ t_4 (\text{MeV} \cdot \text{fm}^3) & 576.91 \\ t_5 (\text{MeV} \cdot \text{fm}^3) & 576.91 \\ t_6 (\text{MeV} \cdot \text{fm}^3) & 576.91 \\ t_6 (\text{MeV} \cdot \text{fm}^3) & 576.91 \\ t_6 (\text{MeV} \cdot \text{fm}^3) & 576.72 \\ t_6 (\text{MeV} \cdot \text{fm}^3) & 576.91 \\ t_6 (\text{MeV} \cdot \text{fm}^3) & $	486.82
$L(\rho_c) = 3\rho_c \left[\frac{\alpha L_{sym}(\rho)}{1 - 100}\right] + \frac{1000}{1 - 100} + 10$	-54.640
$ \left[ \begin{array}{c} d\rho \\ d\rho \end{array} \right]_{\rho=\rho_{c}} \left[ \begin{array}{c} t_{3}(\text{MeV} \cdot \text{Im}^{-1}) \\ t_{4}(\text{MeV} \cdot \text{fm}^{5+3\beta}) \\ t_{5}(\text{MeV} \cdot \text{fm}^{$	13///.
Z. Zhang, PRC94, 064326(2016) $t_{t}(MeV \cdot fm^{5+3\gamma})$ $\begin{bmatrix} 1352.3 & -1013.7 & -1050.2 \\ 3037.5 & 2153.2 & -436.51 \end{bmatrix}$	_
$1 \qquad \qquad 1 \qquad $	0.83400
$v_{i,j} = t_0(1+x_0P_{\sigma})\delta(\mathbf{r}) + \frac{1}{c}t_3(1+x_3P_{\sigma})\rho^{\alpha}(\mathbf{R})\delta(\mathbf{r}) \qquad / \begin{array}{c} x_0 \\ x_1 \\ 1.3163 \\ 0.37275 \\ -0.51268 \end{array}$	-0.34400
$x_2 = -0.55463 - 0.55121 - 3.1558$	-1.00000
$+ \frac{1}{2}t_1(1+x_1P_{\sigma})[K'^2\delta(\mathbf{r})+\delta(\mathbf{r})K^2] \qquad \qquad$	1.35400
2 (x <sub>4</sub> ) 1.7600 0.29499 -1.5709	-
$+ t_2(1+x_2P_{\sigma})\mathbf{K}' \cdot \delta(\mathbf{r})\mathbf{K}$ / (x <sub>5</sub> ) -0.83852 -0.65206 -4.1683	
$\frac{1}{\alpha} = \frac{1}{\alpha} = \frac{1}$	0.16667
$+ \frac{1}{2} \iota_4 (1 + x_4 P_{\sigma}) [\mathbf{K} \ \delta(\mathbf{r}) \rho(\mathbf{K}) + \rho(\mathbf{K}) \delta(\mathbf{r}) \mathbf{K}] \qquad \beta \qquad 1 \qquad 1 \qquad 1$	-
+ $t_5(1+x_5P_{\sigma})\mathbf{K}' \cdot \rho(\mathbf{R})\delta(\mathbf{r})\mathbf{K}$ Extended $\gamma$ $\frac{\gamma}{1}$ $\frac{1}{1}$ $\frac{1}{1}$	
$= iW_0(\text{MeV} \cdot \text{fm}^3) = 92.759  100.14  113.61$	123.00
$+ i W_0(\boldsymbol{\sigma}_i + \boldsymbol{\sigma}_j) \cdot [\boldsymbol{\Lambda} \times \boldsymbol{\sigma}(\boldsymbol{r})\boldsymbol{\Lambda}], \qquad (4)$	11

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Additional density-dependent two-body forces to effectively simulate the momentum-dependent three-body forces

## Neutron skin thickness



The four interactions give similar proton rms, but the neutron radius increase with  $L(\rho_c)$ 

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Roca-Maza, PRL106, 252501 (2011)

L (MeV)

100

50

 $^{208}$ Pb

 $r_p \quad \Delta r_{\rm np} \mid \Delta r_{\rm np}$ 

150

 $^{96}$ Ru

## Heavy ion event generators

- Heavy ion jet interaction generator (Hijing)
- A Multi-Phase Transport model (AMPT)
- Default (String fragmentation)
- String melting
- Ultra relativistic Quantum Molecular Dynamics (UrQMD)

## Multiplicity distributions



 $dP/dN_{\rm ch} \propto -\text{Erf}[-(N_{\rm ch}/N_{1/2}-1)/w] + 1,$  (1)

The effect is hardly observable in a plot of the  $N_{ch}$  distributions themselves.

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## Ratios of N<sub>ch</sub> distributions



- The ratio of  $N_{ch}$  distributions highlight the differences but cumbersome to quantify
- To quantify the differences, we use the R observable of  $\langle N_{ch} \rangle$  at top 5% centrality.

## The R observable

R is a relative measure, **much of experimental effects** cancel:

- **1. Track inefficiency:** We use only 0-5% central collisions , where the tracking efficiency is constant to a good degree
- 2. Trigger inefficiency: the trigger inefficiency can be corrected in experiment. Even without correction, the uncertainty is about  $2 \times 10^{-4}$ , negligible small.

Question: The R observable is actually an isospin insensitive observable, why it have **rather weak model dependence**?

The particle production in relativistic heavy ion collisions is insensitive to the details of the QCD physics, which is in contrast to the hadronic observables in low-energy studies.





### • Glauber model

$$\frac{dN}{dy} \propto s \propto \frac{(1-x)}{2}n_{part} + xn_{coll}$$

 $n_{part}$ : number of participant;  $n_{coll}$ : number of binary collisions

x = 0.1(default), x = 0.2(extreme)

• Trento model (J. Moreland, et.al, PRC92, 011901(2015))

$$\frac{dN}{dy} \propto s \propto \left(\frac{T_A^p + T_B^p}{2}\right)^{1/p}; p = 0 \text{ and } k = 1.4$$

## Probe the neutron skin thickness



4 dynamic models + 2 static models

The R observable in isobaric collisions at ultra-relativistic energies provide a novel approach to determine the neutron skin thickness to a precision that may comparable to or even exceed those achieved by traditional low-energy nuclear experiments.

STAR isobar collisions (2018):

- More statistics: 6.3 billion isobar events
- Less systematical uncertainty

## Grazing isobaric collisions

H. Xu, H. Li, Y. Zhou, X. Wang, J. Zhao, L. Chen, F. Wang, arXiv:2105.04052 (2021)



We propose a direct measurement of the neutron skin by using net-charge multiplicities in ultra-peripheral (grazing) collisions of those isobars.

### DFT and WS densities

H. Xu, H. Li, X. Wang, C. Shen, F. Wang, PLB819, 136453 (2021)



 $\langle r \rangle_{\rm DFT} = \langle r \rangle_{\rm WS}$  $\langle r^2 \rangle_{DFT} = \langle r^2 \rangle_{WS}$ 

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## Effective of nuclei deformation

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H. Xu, H. Li, X. Wang, C. Shen, F. Wang, PLB819, 136453 (2021)

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Giuliano Giacalone, Jiangyong Jia, Vittorio Soma, arXiv:2102.08158 Giuliano Giacalone, Jiangyong Jia, Chunjian Zhang, arXiv: 2105.01638 Jiangyong Jia, Shenli Huang, Chunjiang Zhang, arXiv:2105.05713 Jiangyong Jia, arXiv:2106.08768

## V总结

#### PHYSICAL REVIEW LETTERS 125, 222301 (2020)

### arXiv: 1910.06170

#### Probing the Neutron Skin with Ultrarelativistic Isobaric Collisions

Hanlin Li<sup>0</sup>,<sup>1</sup> Hao-jie Xu<sup>0</sup>,<sup>2,\*</sup> Ying Zhou,<sup>3</sup> Xiaobao Wang,<sup>2</sup> Jie Zhao,<sup>4</sup> Lie-Wen Chen,<sup>3,†</sup> and Fuqiang Wang<sup>2,4,‡</sup> <sup>1</sup>College of Science, Wuhan University of Science and Technology, Wuhan, Hubei 430065, China <sup>2</sup>School of Science, Huzhou University, Huzhou, Zhejiang 313000, China <sup>3</sup>School of Physics and Astronomy and Shanghai Key Laboratory for Particle Physics and Cosmology, Shanghai Jiao Tong University, Shanghai 200240, China <sup>4</sup>Department of Physics and Astronomy, Purdue University, West Lafavette, Indiana 47907, USA

(Received 24 October 2019; accepted 15 October 2020; published 23 November 2020)

Particle production in ultrarelativistic heavy ion collisions depends on the details of the nucleon den distributions in the colliding nuclei. We demonstrate that the charged hadron multiplicity distribution isobaric collisions at ultrarelativistic energies provide a novel approach to determine the poorly knu neutron density distributions and thus the neutron skin thickness in finite nuclei, which can in turn stringent constraints on the nuclear symmetry energy.

DOI: 10.1103/PhysRevLett.125.222301





Physics Letters B 819 (2021) 136453

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### Determine the neutron skin type by relativistic isobaric collisions



Hao-jie Xu<sup>a,\*</sup>, Hanlin Li<sup>b</sup>, Xiaobao Wang<sup>a</sup>, Caiwan Shen<sup>a</sup>, Fugiang Wang<sup>a,c,\*</sup>

ABSTRACT

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### arXiv: 2103.05595

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The effects of neutron skin on the multiplicity ( $N_{ch}$ ) and eccentricity ( $\epsilon_2$ ) in relativistic  $\frac{96}{44}Ru + \frac{96}{44}Ru$  and  ${}_{40}^{96}$ Zr+ ${}_{40}^{96}$ Zr collisions at  $\sqrt{s_{_{NN}}} = 200$  GeV are investigated with the Trento model. It is found that the Ru+Ru/Zr+Zr ratios of the N<sub>ch</sub> distributions and  $\epsilon_2$  in mid-central collisions are exquisitely sensitive to the neutron skin type (skin vs. halo). The state-of-the-art calculations by energy density functional theory (DFT) favor the halo-type neutron skin and can soon be confronted by experimental data. It is demonstrated that the halo-type density can serve as a good surrogate for the DFT density, and thus can be efficiently employed to probe nuclear deformities by using elliptic flow data in central collisions. We provide hereby a proof-of-principle venue to simultaneously determine the neutron skin type, thickness, and nuclear deformity.

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### Hao-jie Xu<sup>\*</sup>,<sup>1</sup> Hanlin Li<sup>†</sup>,<sup>2</sup> Ying Zhou,<sup>3</sup> Xiaobao Wang,<sup>1</sup> Jie Zhao,<sup>4</sup> Lie-Wen Chen<sup>‡</sup>,<sup>3</sup> and Fuqiang Wang<sup>§1,4</sup> <sup>1</sup>School of Science, Huzhou University, Huzhou, Zhejiang 313000, China <sup>2</sup>College of Science, Wuhan University of Science and Technology, Wuhan, Hubei 430065, China <sup>3</sup> School of Physics and Astronomy, Shanghai Key Laboratory for Particle Physics and Cosmology, and Key Laboratory for Particle Astrophysics and Cosmology (MOE), Shanghai Jiao Tong University, Shanghai 200240, China <sup>4</sup>Department of Physics and Astronomy, Purdue University, West Lafayette, Indiana 47907, USA Neutron skin thickness ( $\Delta r_{np}$ ) of nuclei and the inferred nuclear symmetry energy are of critical importance to nuclear physics and astrophysics. It is traditionally measured by nuclear processes with significant theoretical uncertainties. We recently proposed an indirect measurement of the $\Delta r_{np}$ by charged hadron multiplicities in central isobaric collisions at relativistic energies, which are sensitive to nuclear densities. In this paper we propose a direct measurement of the $\Delta r_{nn}$ by using net-charge multiplicities in ultra-peripheral (grazing) collisions of those isobars, under the assumption that they are simple superimposition of nucleon-nucleon interactions. We illustrate this novel approach by the TRENTO and URQMD models. arXiv: 2105.04052 PACS numbers:

Measuring neutron skin by grazing isobaric collisions



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徐浩洁,湖州师范学院

V总结

相对论同量异 位素核碰撞是 研究核结构绝 佳平台!



# Thank you for your attention!