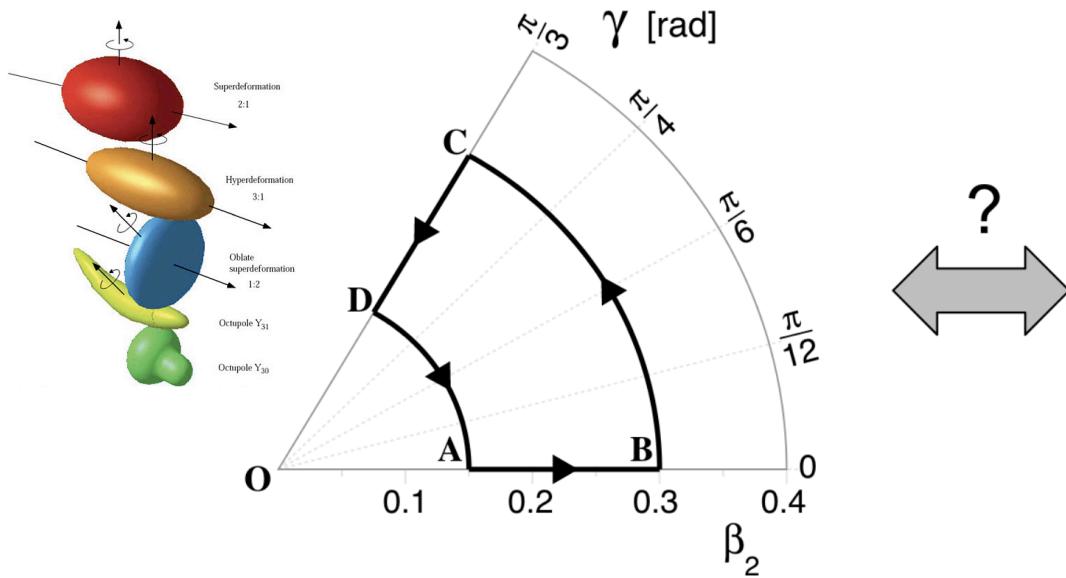


Imaging nuclear structure in heavy-ion collisions

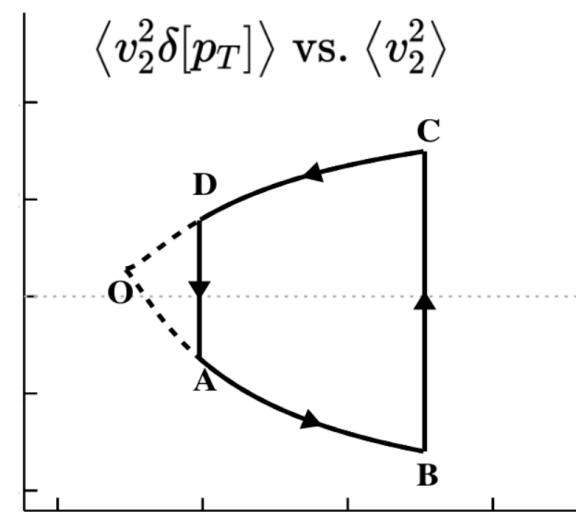
Jiangyong Jia

Collaborators: Giuliano Giacalone and Chunjian Zhang

Nuclear structure

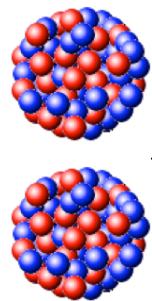


High-energy collisions



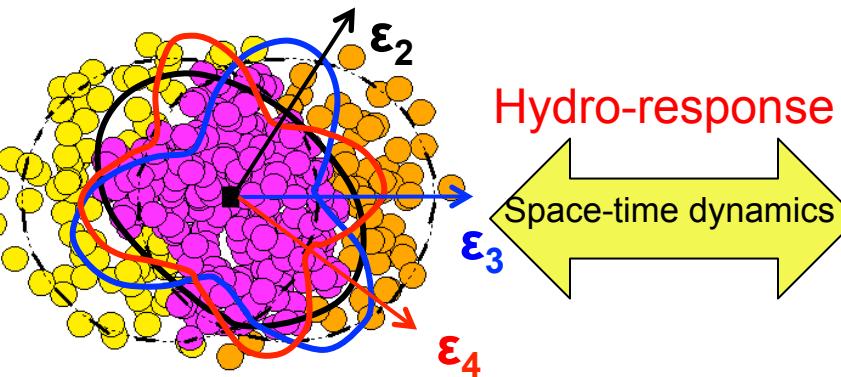
Hydrodynamic response to initial state

Nuclear Structure


 ρ_0

Imaging?

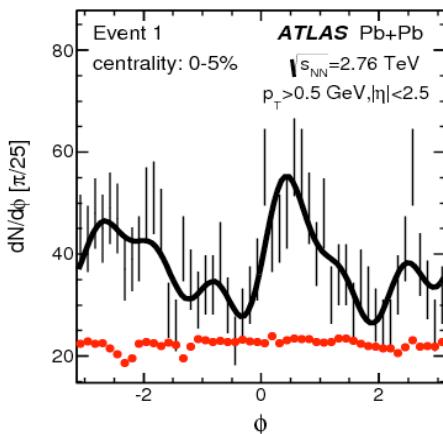
Initial State



Hydro-response

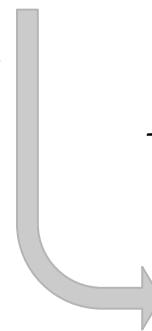
Space-time dynamics

Final Particle flow



$$\frac{1 + e^{(r - R_0(1 + \sum_n \beta_n Y_n^0(\theta, \phi)) / a_0)}}{1 + e^{(r - R_0(1 + \sum_n \beta_n Y_n^0(\theta, \phi)) / a_0)}}$$

- $\beta_2 \rightarrow$ Quadrupole deformation
- $\beta_3 \rightarrow$ Octupole deformation
- $a_0 \rightarrow$ Surface diffuseness
- $R_0 \rightarrow$ Nuclear size



Initial Size

$$R_\perp \propto \langle r_\perp^2 \rangle, \quad \mathcal{E}_n \propto \langle r_\perp^n e^{in\phi} \rangle$$

$$R_0 \quad a_0 \quad \beta_n$$

Initial Shape

$$??$$

Radial Flow

$$\frac{d^2 N}{d\phi dp_T} = N(p_T) \left(\sum_n V_n e^{-in\phi} \right)$$

Harmonic Flow

arXiv:1206.1905

High energy: approx. linear response in each event:

$$\frac{\delta[p_T]}{[p_T]} \propto -\frac{\delta R_\perp}{R_\perp} \quad V_n \propto \mathcal{E}_n$$

Influence of shape fluctuations in relativistic heavy ion collisions

A. Rosenhauer, H. Stöcker, J. A. Maruhn, and W. Greiner

Phys. Rev. C **34**, 185 – Published 1 July 1986

Article

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Uranium on uranium collisions at relativistic energies

Bao-An Li

Phys. Rev. C **61**, 021903(R) – Published 12 January 2000

Article

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High energy collisions of strongly deformed nuclei: An old idea with a new twist

E. V. Shuryak

Phys. Rev. C **61**, 034905 – Published 22 February 2000

Article

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Choice of colliding beams to study deformation effects in relativistic heavy ion collisions

S. Das Gupta and C. Gale

Phys. Rev. C **62**, 031901(R) – Published 23 August 2000

Article

References

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Focus particle production in UCC tip-tip vs body-body, start consider effect of collective flow

Elliptic flow in central collisions of deformed nuclei

P. Filip 

Physics of Atomic Nuclei 71, 1609–1618 (2008) | [Cite this article](#)

51 Accesses | 13 Citations | [Metrics](#)

Modern concept of elliptic flow and eccentricities, still focus on UCC region

Anisotropic Flow and Jet Quenching in Ultrarelativistic U + U Collisions

Ulrich Heinz and Anthony Kuhlman

Phys. Rev. Lett. 94, 132301 – Published 6 April 2005

Article

References

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Initial eccentricity in deformed $^{197}\text{Au} + ^{197}\text{Au}$ and $^{238}\text{U} + ^{238}\text{U}$ collisions at $\sqrt{s_{NN}} = 200$ GeV at the BNL Relativistic Heavy Ion Collider

Peter Filip, Richard Lednicky, Hiroshi Masui, and Nu Xu

Phys. Rev. C 80, 054903 – Published 5 November 2009

Article

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Export Citation

Parameterization of deformed nuclei for Glauber modeling in relativistic heavy ion collisions

Q.Y. Shou ^{a, b}  , Y.G. Ma ^a, P. Sorensen ^c, A.H. Tang ^c, F. Videbæk ^c, H. Wang ^c

Collision geometry and flow in uranium + uranium collisions

Andy Goldschmidt, Zhi Qiu, Chun Shen, and Ulrich Heinz
Phys. Rev. C **92**, 044903 – Published 7 October 2015

Article

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Spectra and elliptic flow of thermal photons from full-overlap U+U collisions at energies available at the BNL Relativistic Heavy Ion Collider

Pingal Dasgupta, Rupa Chatterjee, and Dinesh K. Srivastava
Phys. Rev. C **95**, 064907 – Published 15 June 2017

Article

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Hydrodynamic predictions for 5.44 TeV Xe+Xe collisions

Giuliano Giacalone, Jacquelyn Noronha-Hostler, Matthew Luzum, and Jean-Yves Ollitrault
Phys. Rev. C **97**, 034904 – Published 6 March 2018

Article

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Possible octupole deformation of ^{208}Pb and the ultracentral v_2 to v_3 puzzle

P. Carzon, S. Rao, M. Luzum, M. Sievert, and J. Noronha-Hostler
Phys. Rev. C **102**, 054905 – Published 9 November 2020

Article

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Sophisticated e-by-e hydrodynamic model comparison including deformations.

Observing the Deformation of Nuclei with Relativistic Nuclear Collisions

Giuliano Giacalone

Phys. Rev. Lett. **124**, 202301 – Published 19 May 2020

v_2^2 - p_T correlation

Article

References

Citing Articles (8)

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Constraining the quadrupole deformation of atomic nuclei with relativistic nuclear collisions

Giuliano Giacalone

Phys. Rev. C **102**, 024901 – Published 3 August 2020

Article

References

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Export Citation

Multiphase transport model predictions of isobaric collisions with nuclear structure from density functional theory

Radial structure

Hanlin Li, Hao-jie Xu, Jie Zhao, Zi-Wei Lin, Hanzhong Zhang, Xiaobao Wang, Caiwan Shen, and Fuqiang Wang

Phys. Rev. C **98**, 054907 – Published 26 November 2018

Probing the Neutron Skin with Ultrarelativistic Isobaric Collisions

Hanlin Li, Hao-jie Xu, Ying Zhou, Xiaobao Wang, Jie Zhao, Lie-Wen Chen, and Fuqiang Wang

Phys. Rev. Lett. **125**, 222301 – Published 23 November 2020

Article

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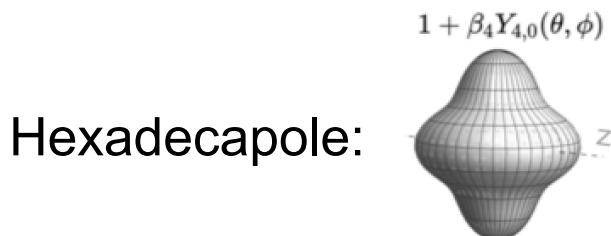
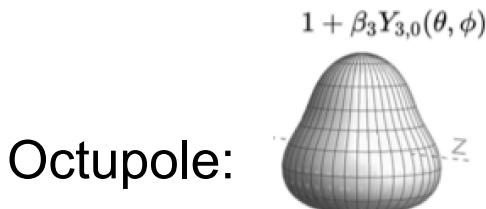
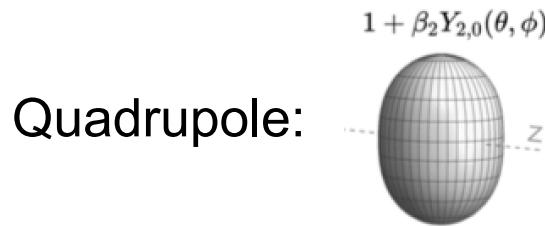
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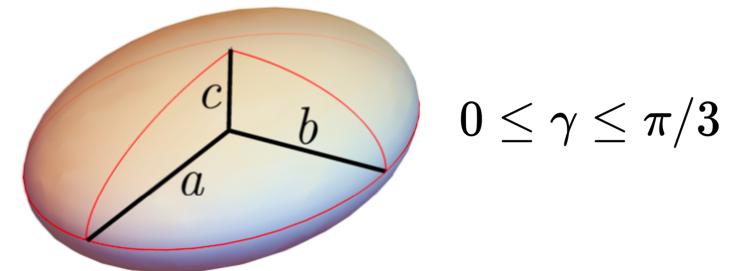
Shape of nuclei

$$\rho(r, \theta, \phi) = \frac{\rho_0}{1 + e^{(r - R(\theta, \phi))/a_0}}$$

$$R(\theta, \phi) = R_0 \left(1 + \beta_2 [\cos \gamma Y_{2,0} + \sin \gamma Y_{2,2}] + \beta_3 \sum_{m=-3}^3 \alpha_{3,m} Y_{3,m} + \beta_4 \sum_{m=-4}^4 \alpha_{4,m} Y_{4,m} \right)$$



Triaxial spheroid: $a \neq b \neq c$.

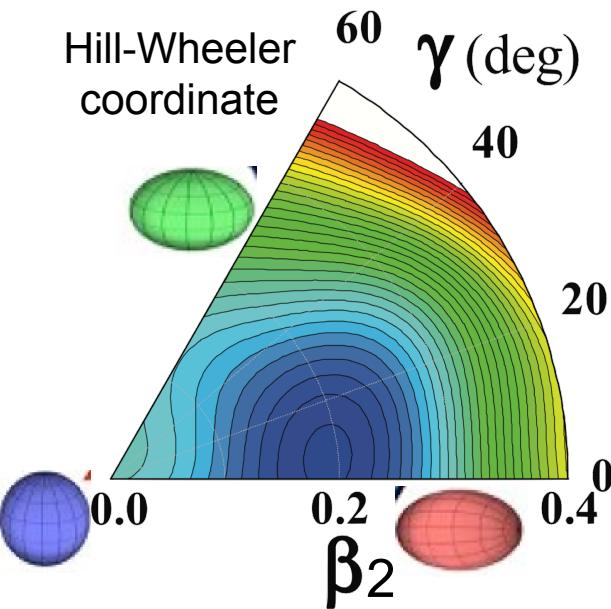
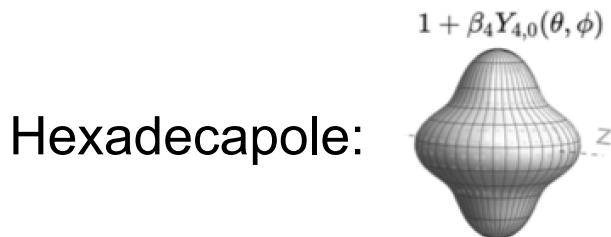
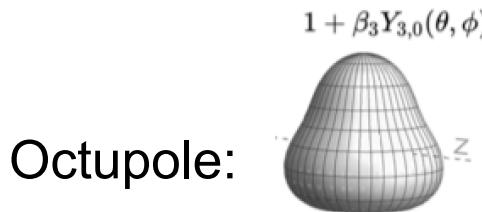
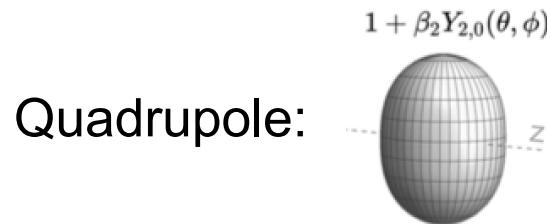


Prolate: $a=b < c \rightarrow \beta_2, \gamma=0$
 Oblate: $a < b=c \rightarrow \beta_2, \gamma=\pi/3$ or $-\beta_2, \gamma=0$

Shape of nuclei

$$\rho(r, \theta, \phi) = \frac{\rho_0}{1 + e^{(r - R(\theta, \phi))/a_0}}$$

$$R(\theta, \phi) = R_0 \left(1 + \beta_2 [\cos \gamma Y_{2,0} + \sin \gamma Y_{2,2}] + \beta_3 \sum_{m=-3}^3 \alpha_{3,m} Y_{3,m} + \beta_4 \sum_{m=-4}^4 \alpha_{4,m} Y_{4,m} \right)$$

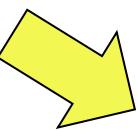
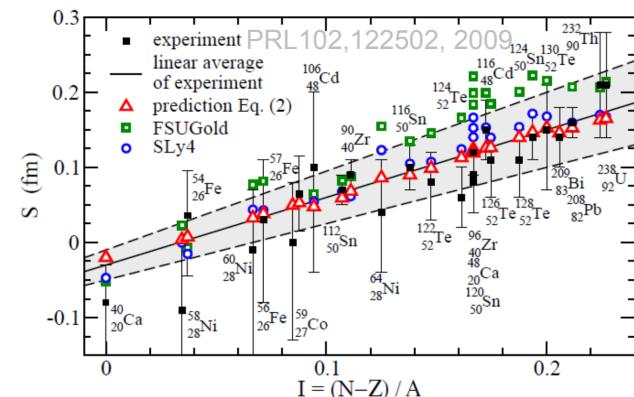
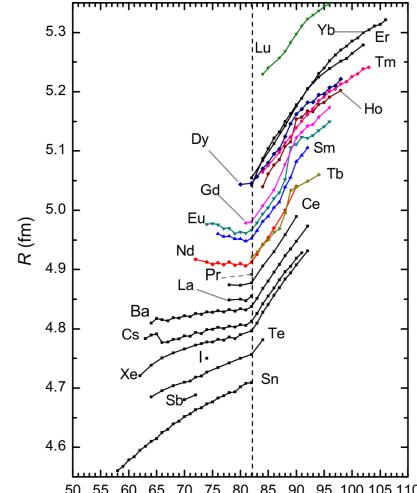
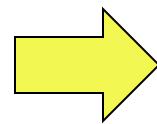
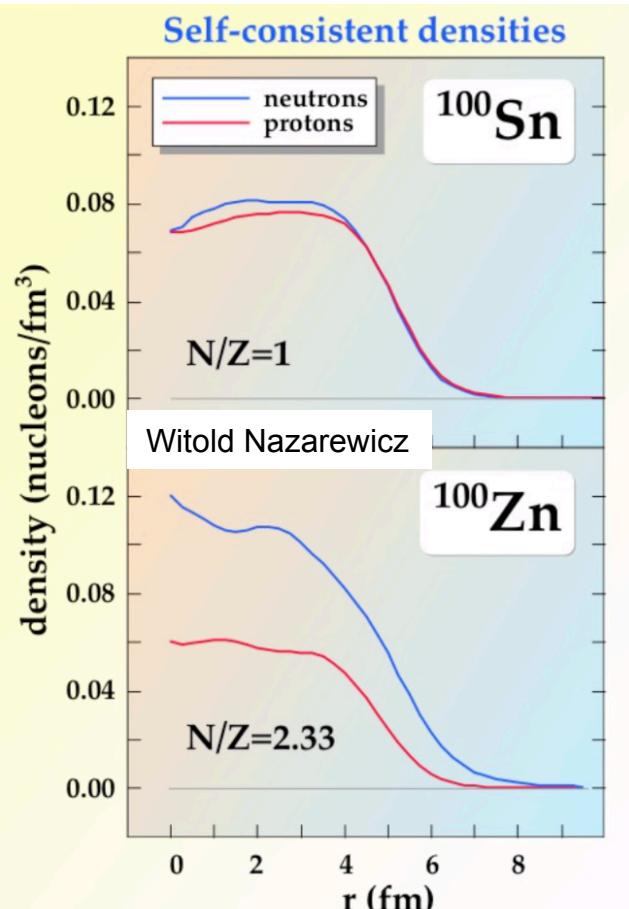


Radial structure of nuclei

$$\rho = \frac{\rho_0}{1 + e^{(r - R_0(1 + \sum_n \beta_n Y_n^0(\theta, \phi))) / a_0}}$$

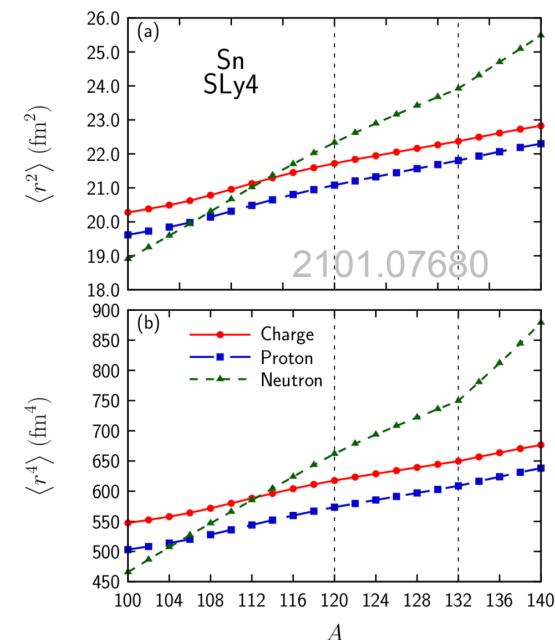
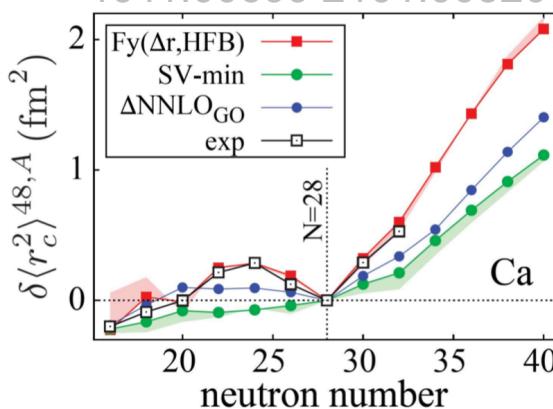
R_0

$\Delta r_{np} = R_n - R_p$



Higher radial moments

1911.00699 2101.00320

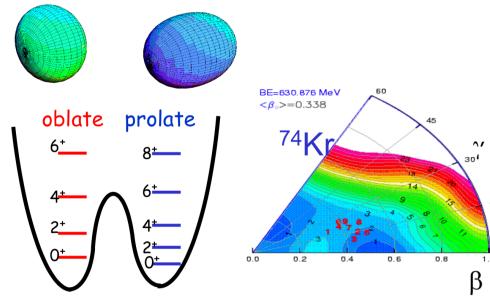
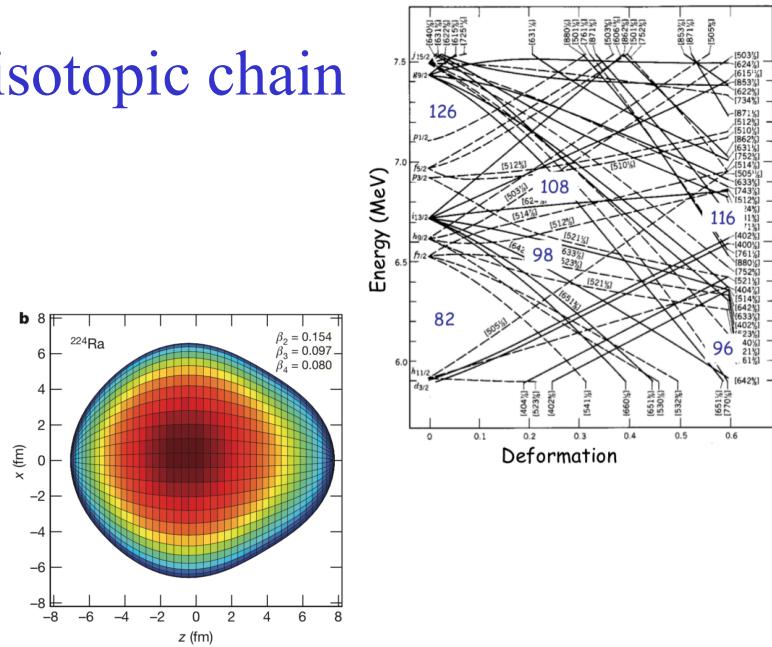


Some topics in nuclear structure

- How the shape/size/skin evolves along isotopic chain
 - Strong test on nuclear structure model

- Octuple (pear-shaped) deformation
 - Octupole correlation or static deformation
 - Strong test on EDM effects

- Triaxiality : infers from γ -band, Chiral and Wobbling bands.
Have large uncertainties.
 - shape coexistence

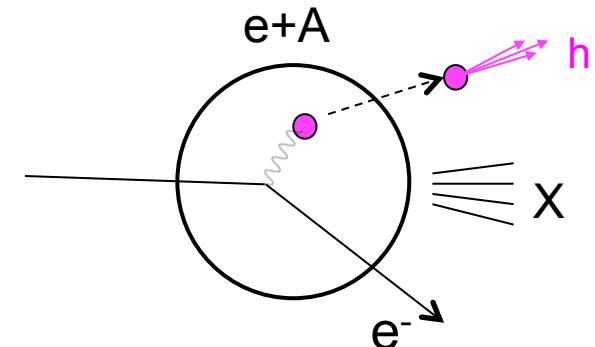
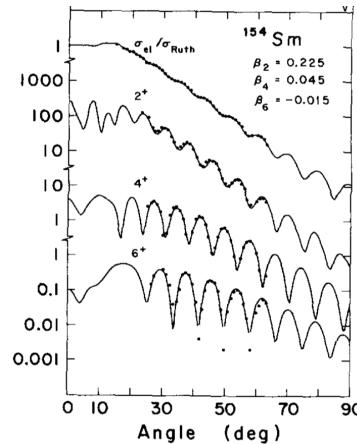
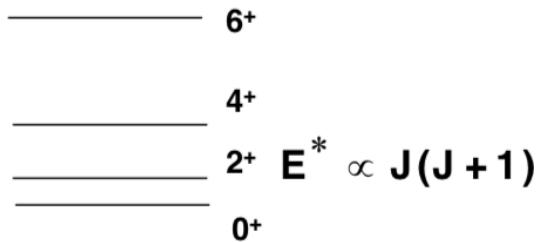


Use shape imaging in heavy-ion collision to help?

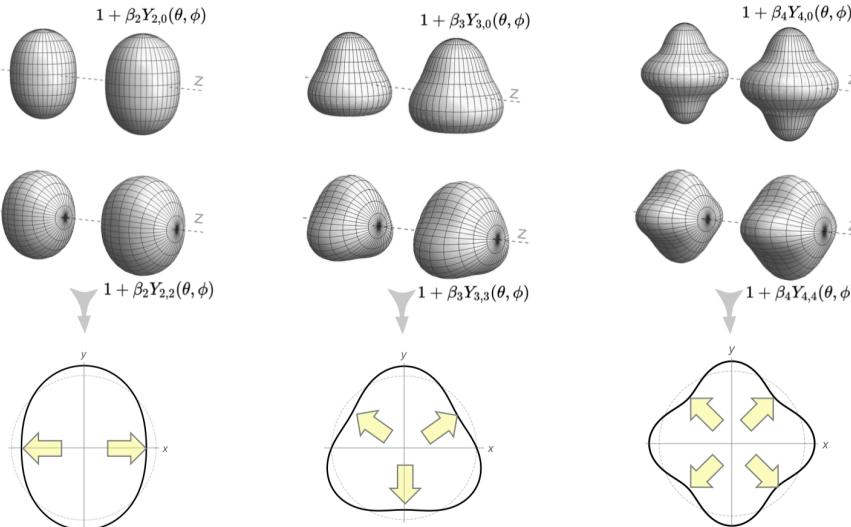
Nuclear structure vs HI method

- Shape from $B(En)$, radial profile from $e+A$ or ion-A scattering

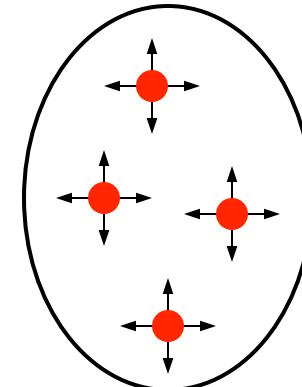
«rotational» spectrum



- Probe entire mass distribution: multi-point correlations



Flow response to probe the nuclear structure



$$S(\mathbf{s}_1, \mathbf{s}_2) \equiv \langle \delta\rho(\mathbf{s}_1)\delta\rho(\mathbf{s}_2) \rangle \\ = \langle \rho(\mathbf{s}_1)\rho(\mathbf{s}_2) \rangle - \langle \rho(\mathbf{s}_1) \rangle \langle \rho(\mathbf{s}_2) \rangle.$$

Evidence of deformation in U+U vs Au+Au¹²

<https://indico.cern.ch/event/854124/contributions/4135480/>

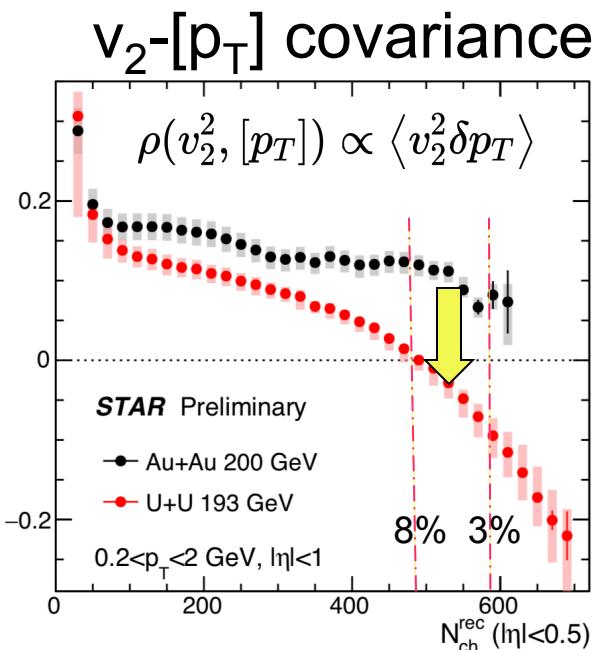
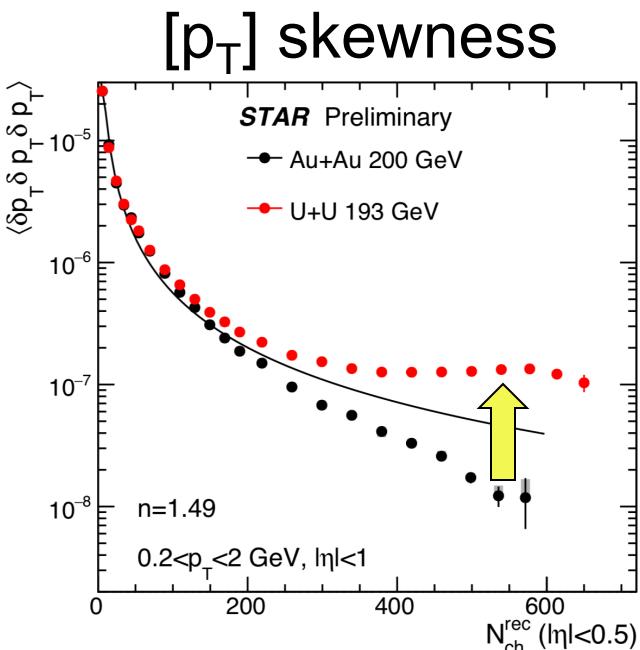
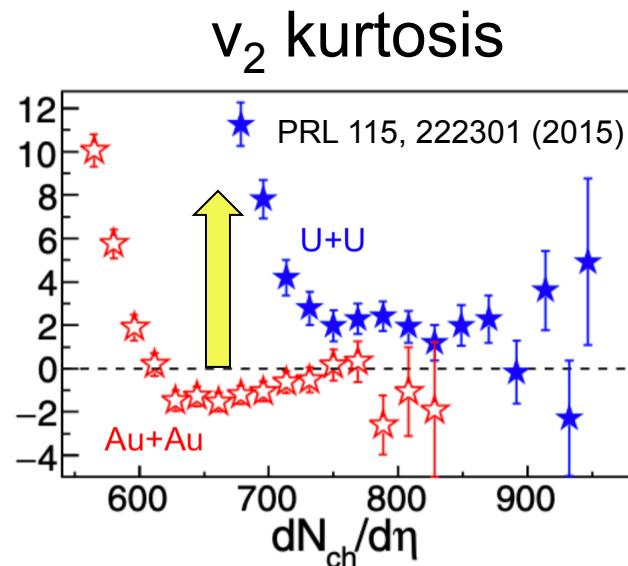
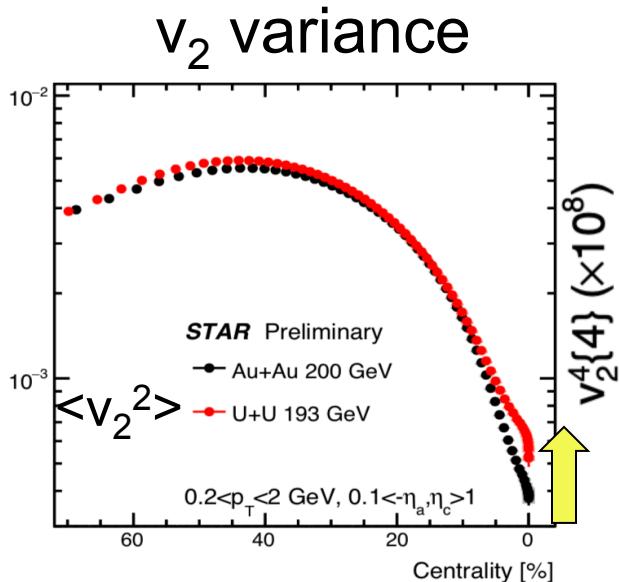
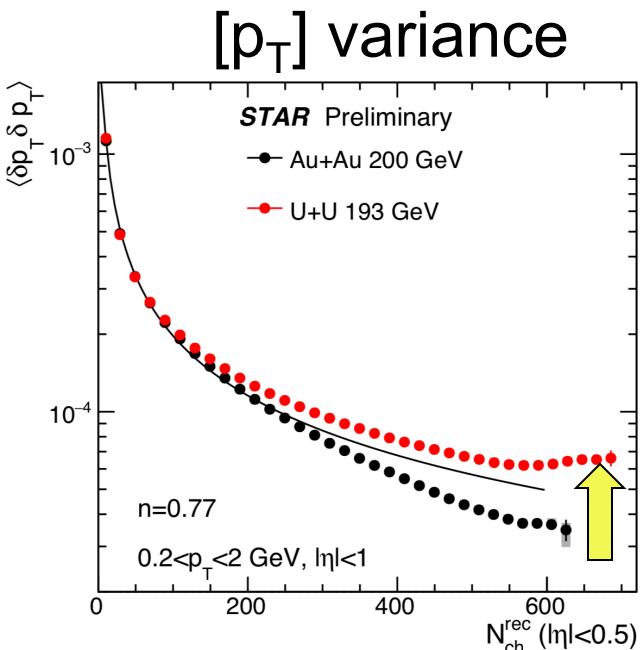
Collisions at $\sqrt{s_{NN}}=193\text{-}200 \text{ GeV}$

Large deformation in ^{238}U
relative to ^{197}Au strongly
influence flow signals

$$\beta_{2\text{U}} \sim 0.28$$

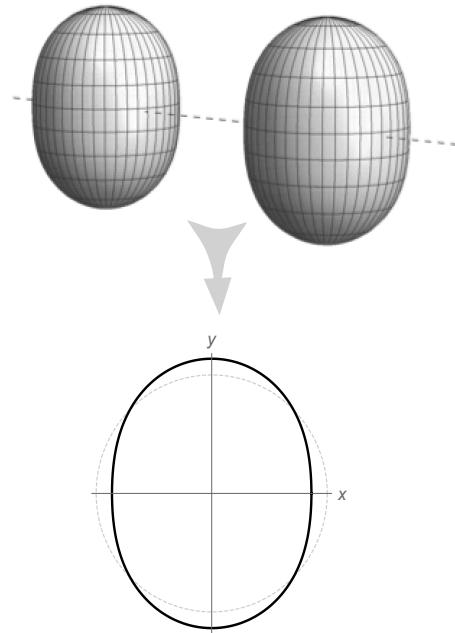
$$\beta_{2\text{Au}} \sim -0.13?$$

Working to turn these into
a quantitative tool !



Shape of the initial state in HI

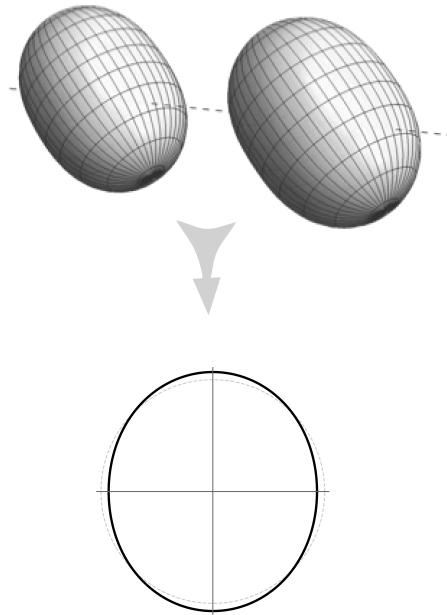
Body-Body



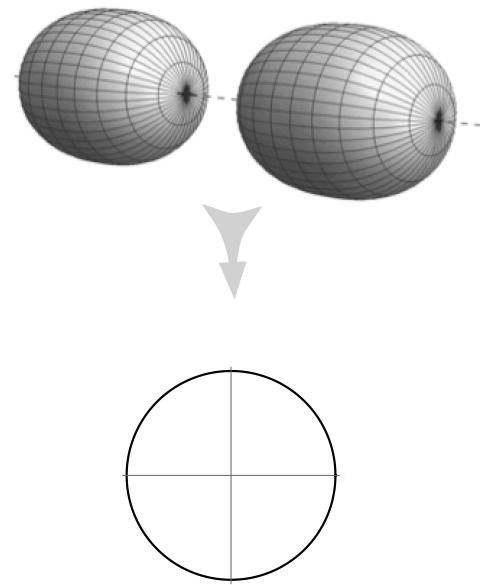
$$\varepsilon_2 \sim 0.95\beta_2$$

$$\mathcal{E}_n = \epsilon_n e^{in\Phi} \propto \langle r_\perp^n e^{in\phi} \rangle$$

Tip-Tip



$$\varepsilon_2 \sim 0.48\beta_2$$



$$\varepsilon_2 \sim 0$$

shape of overlap = shape of nucleon dist. projected along Euler angle $\Omega=\varphi\theta\psi$

Parametric dependence

- ε_n has the form $\varepsilon_n = \varepsilon_{n;0} + \sum_{m=2}^4 \underbrace{p_{n;m}(\Omega_1, \Omega_2)}_{\text{undeformed}} \beta_m + \mathcal{O}(\beta^2)$
 γ only appear here, in the form of $\cos 3\gamma, \cos 6\gamma, \dots$
- $R_\perp^2 = \langle x^2 \rangle + \langle y^2 \rangle$ has the form $\delta d_\perp/d_\perp = \delta_d + \sum_{m=2}^4 p_{0;m}(\Omega_1, \Omega_2) \beta_m + \mathcal{O}(\beta^2)$
 $d_\perp \equiv 1/R_\perp$
- Two particle correlation

$$\langle \varepsilon_n^2 \rangle \approx \langle \varepsilon_{n;0}^2 \rangle + \sum_m \langle p_{n;m} p_{n;m}^* \rangle \beta_m^2 \quad \left\langle \left(\frac{\delta d_\perp}{d_\perp} \right)^2 \right\rangle \approx \langle \delta_d^2 \rangle + \sum_m \langle p_{0;m}^2 \rangle \beta_m^2$$

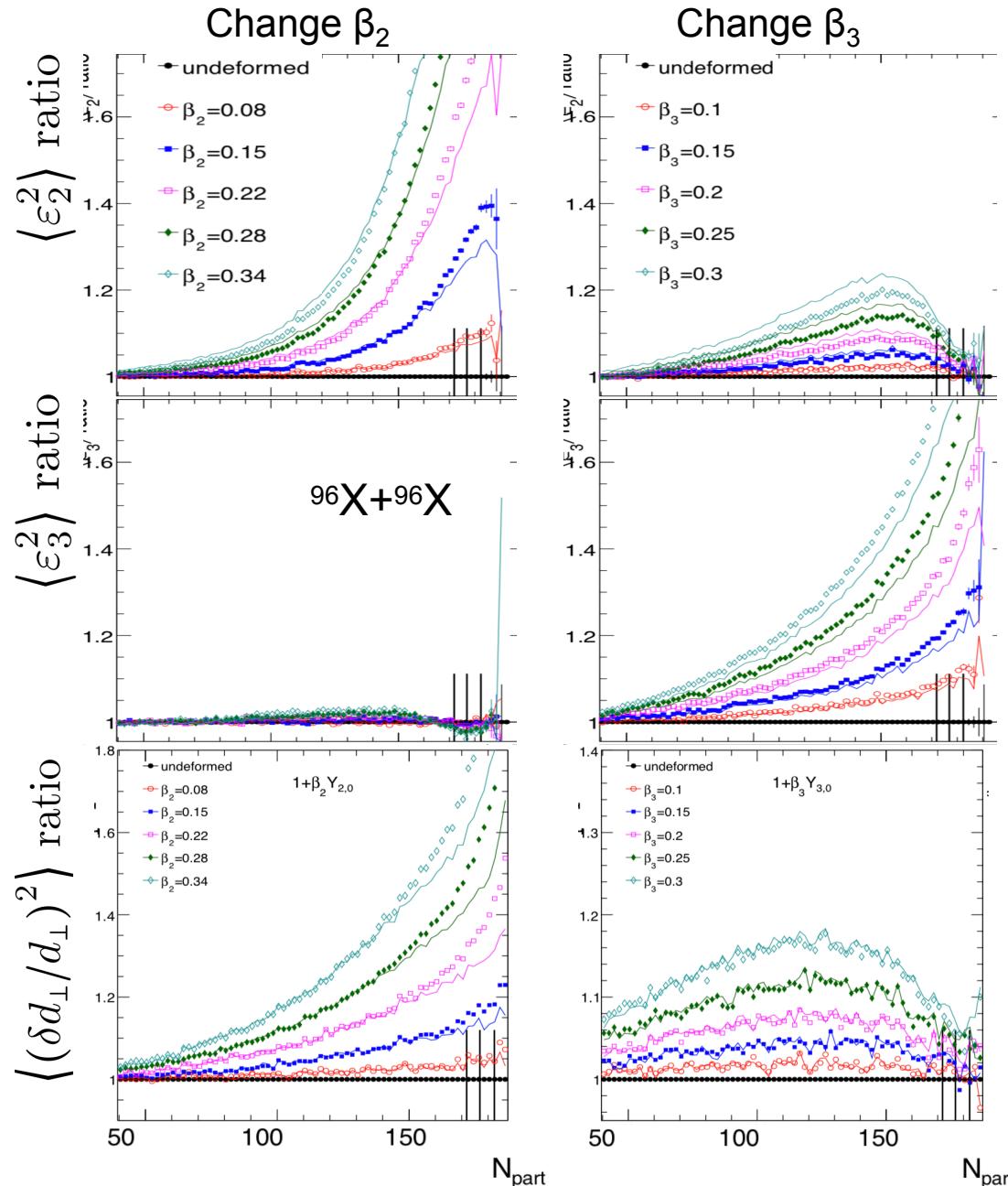
- Consider also the influence of R_0 and a $\frac{\rho_0}{1 + e^{(r - \mathbf{R}_0(1 + \sum_n \beta_n Y_n^0(\theta, \phi))) / \mathbf{a}_0}}$

$$\langle \varepsilon_n^2 \rangle \approx b_0 + b_1 \beta_2^2 + b_2 \beta_3^2 + b_3 (R_0 - R_{0,\text{ref}}) + b_4 (a - a_{\text{ref}})$$

- Again linear response to relate to final state: $v_n \propto \varepsilon_n \frac{\delta[p_T]}{[p_T]} \propto \frac{\delta d_\perp}{d_\perp}$

Parametric dependence

See 2106.08768

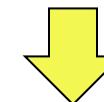


medium size system:

$$\varepsilon_2^2 = a'_2 + b'_2 \beta_2^2 + b'_{2,3} \beta_3^2$$

$$\varepsilon_3^2 = a'_3 + b'_3 \beta_3^2$$

$$(\delta d_\perp/d_\perp)^2 = a'_0 + b'_0 \beta_2^2 + b'_{0,3} \beta_3^2$$



$$v_2^2 = a_2 + b_2 \beta_2^2 + b_{2,3} \beta_3^2$$

$$v_3^2 = a_3 + b_3 \beta_3^2$$

$$(\delta p_T/p_T)^2 = a_0 + b_0 \beta_2^2 + b_{0,3} \beta_3^2$$

Isobar collisions as precision tool

- Unique running mode of RHIC and STAR to minimize the detector systematics
 - 0.4% precision is achieved in ratio of many observables between two isobar systems → precision imaging tool

A key question for any
HI observable O

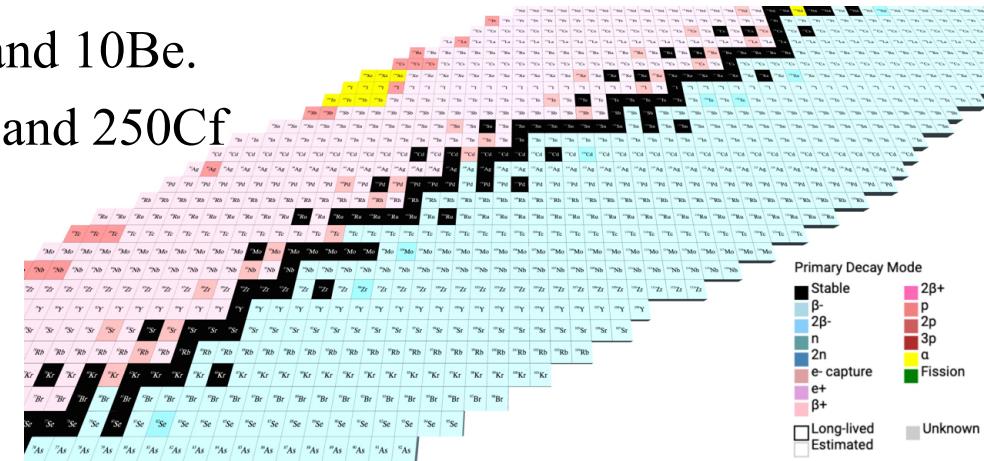
$$\frac{O_{X+X}}{O_{Y+Y}} = ? \quad 1$$

${}^A_X + {}^A_X$ vs ${}^A_Y + {}^A_Y$

Deviation from 1 must has its origin in the nuclear structure, which is reflected by the initial state and then survives the final state. A precision tool to study initial state and final state responses

- Many such pairs of isobars in the nuclear chart.
 - Small system isobar such as ${}^{10}\text{B}$ and ${}^{10}\text{Be}$.
 - Large system isobar up to ${}^{250}\text{Cm}$ and ${}^{250}\text{Cf}$

A new handle to probe
heavy ion physics



Isobar collisions as precision tool

- Unique running mode of RHIC and STAR to minimize systematics
 - 0.4% precision is achieved in ratio of many observables between two isobar systems → precision imaging tool

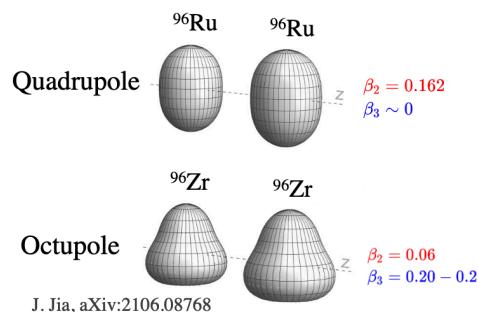
A key question for any
HI observable \mathcal{O}

$$\frac{\mathcal{O}_{X+X}}{\mathcal{O}_{Y+Y}} = ?$$

${}^A\text{X} + {}^A\text{X}$ vs ${}^A\text{Y} + {}^A\text{Y}$

Deviation from 1 must has its origin in the nuclear structure, which is reflected by the initial state and then survives the final state. A precision tool to study initial state and final state responses

■ Expectation



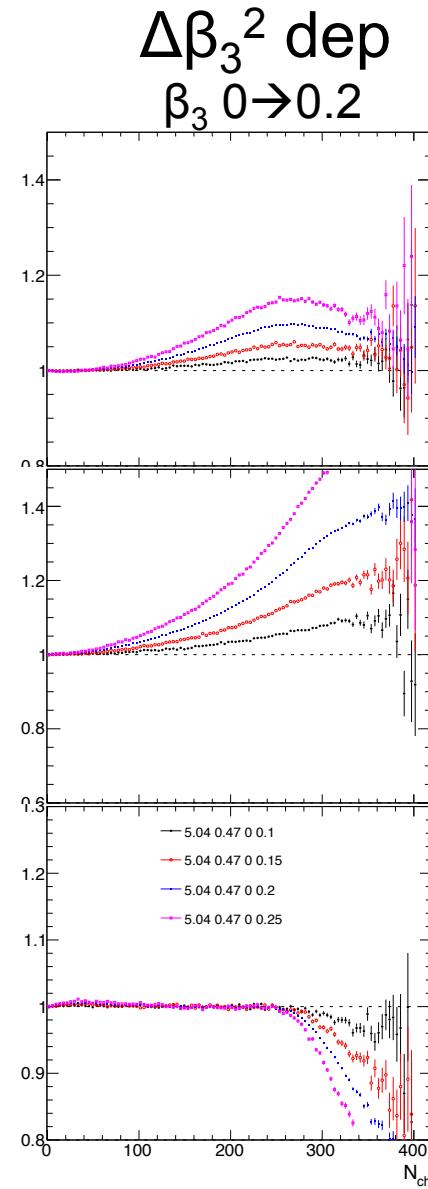
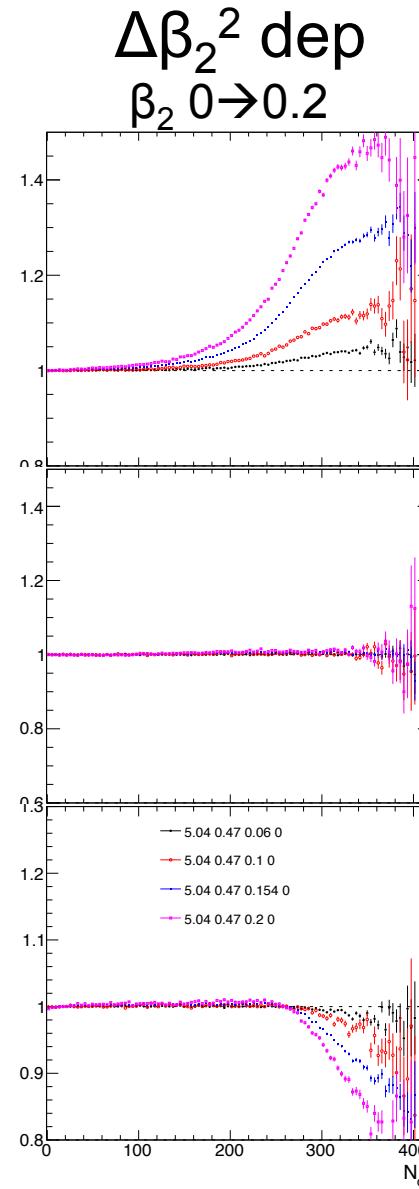
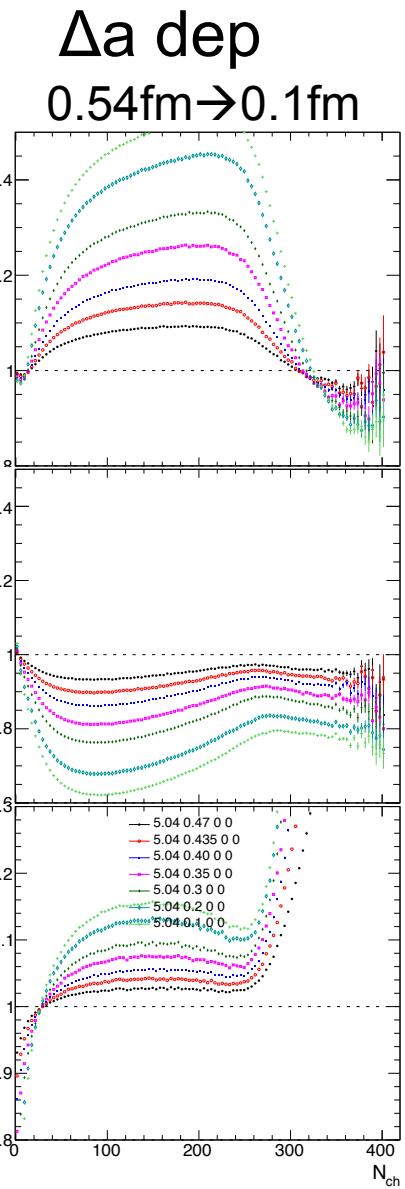
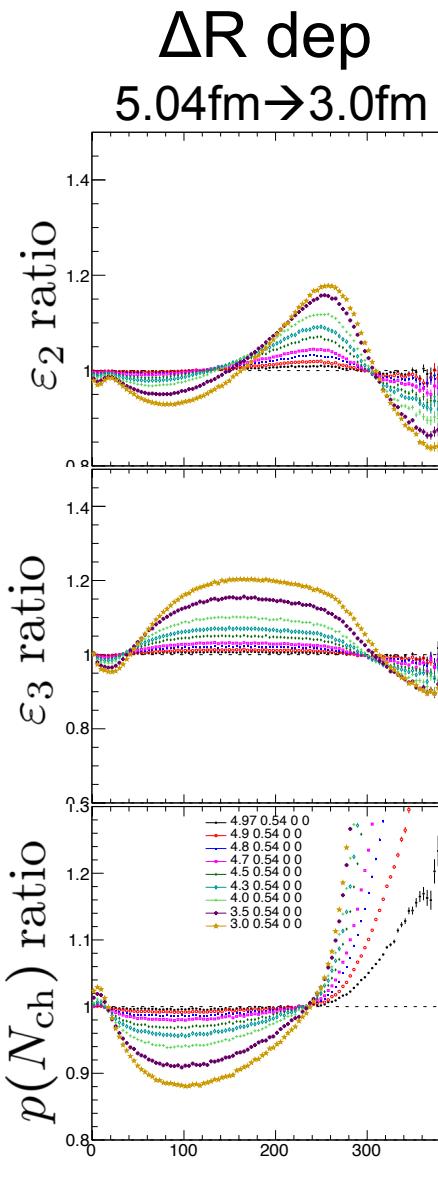
$$\mathcal{O} \approx b_0 + b_1 \beta_2^2 + b_2 \beta_3^2 + b_3 (R_0 - R_{0,\text{ref}}) + b_4 (a - a_{\text{ref}})$$

$$R_{\mathcal{O}} \equiv \frac{\mathcal{O}_{\text{Ru}}}{\mathcal{O}_{\text{Zr}}} \approx 1 + c_1 \Delta \beta_2^2 + c_2 \Delta \beta_3^2 + c_3 \Delta R_0 + c_4 \Delta a$$

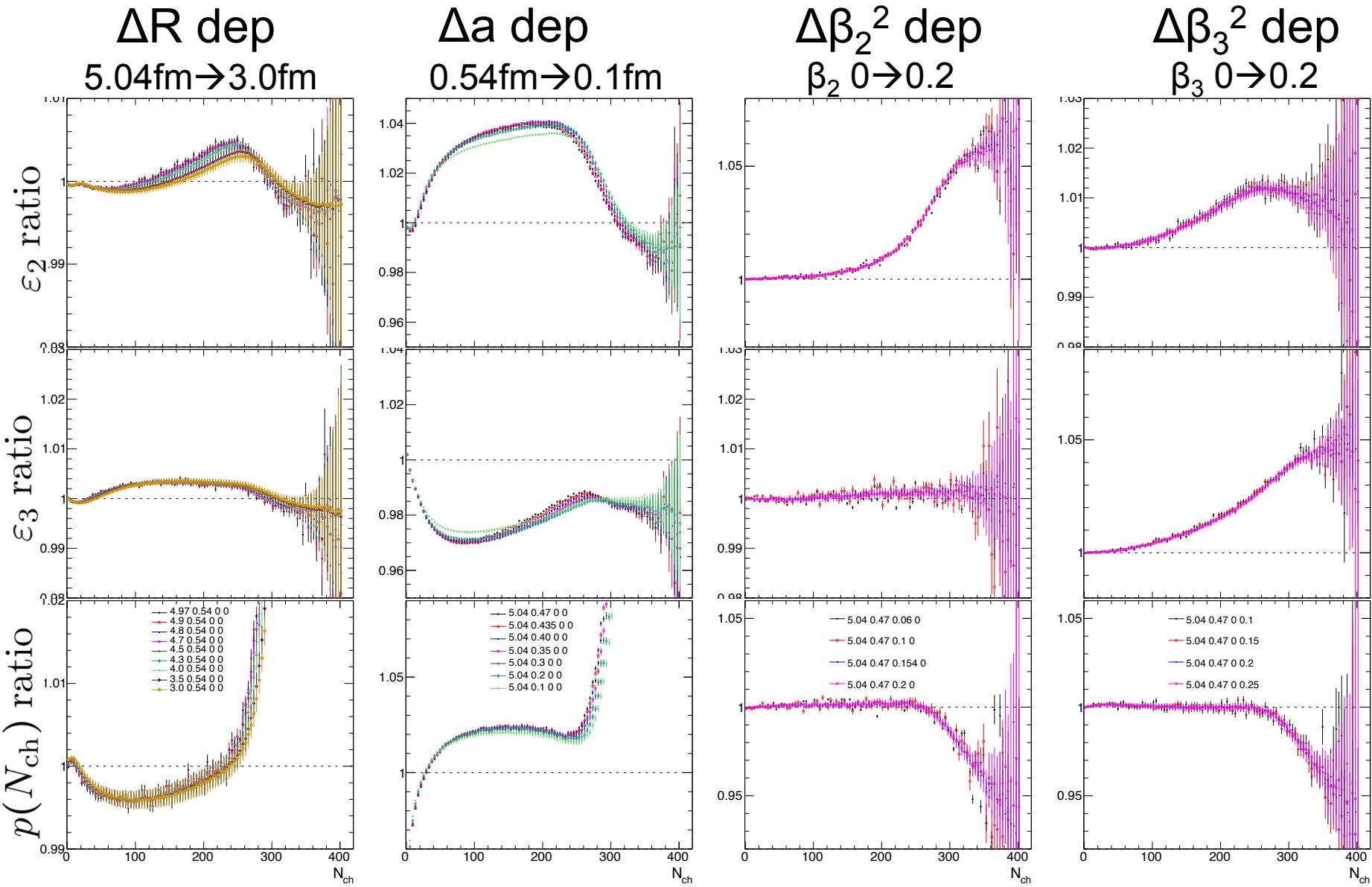
Species	β_2	β_3	a_0	R_0
Ru	0.162	0	0.46 fm	5.09 fm
Zr	0.06	0.20	0.52 fm	5.02 fm
difference	$\Delta \beta_2^2$	$\Delta \beta_3^2$	Δa_0	ΔR_0
	0.0226	-0.04	-0.06 fm	0.07 fm

Valid for most single- and two-particle observable: $v2, v3, p(N), \langle p_T \rangle, \langle \delta p_T^2 \rangle ..$

Glauber results: N_{ch} dep

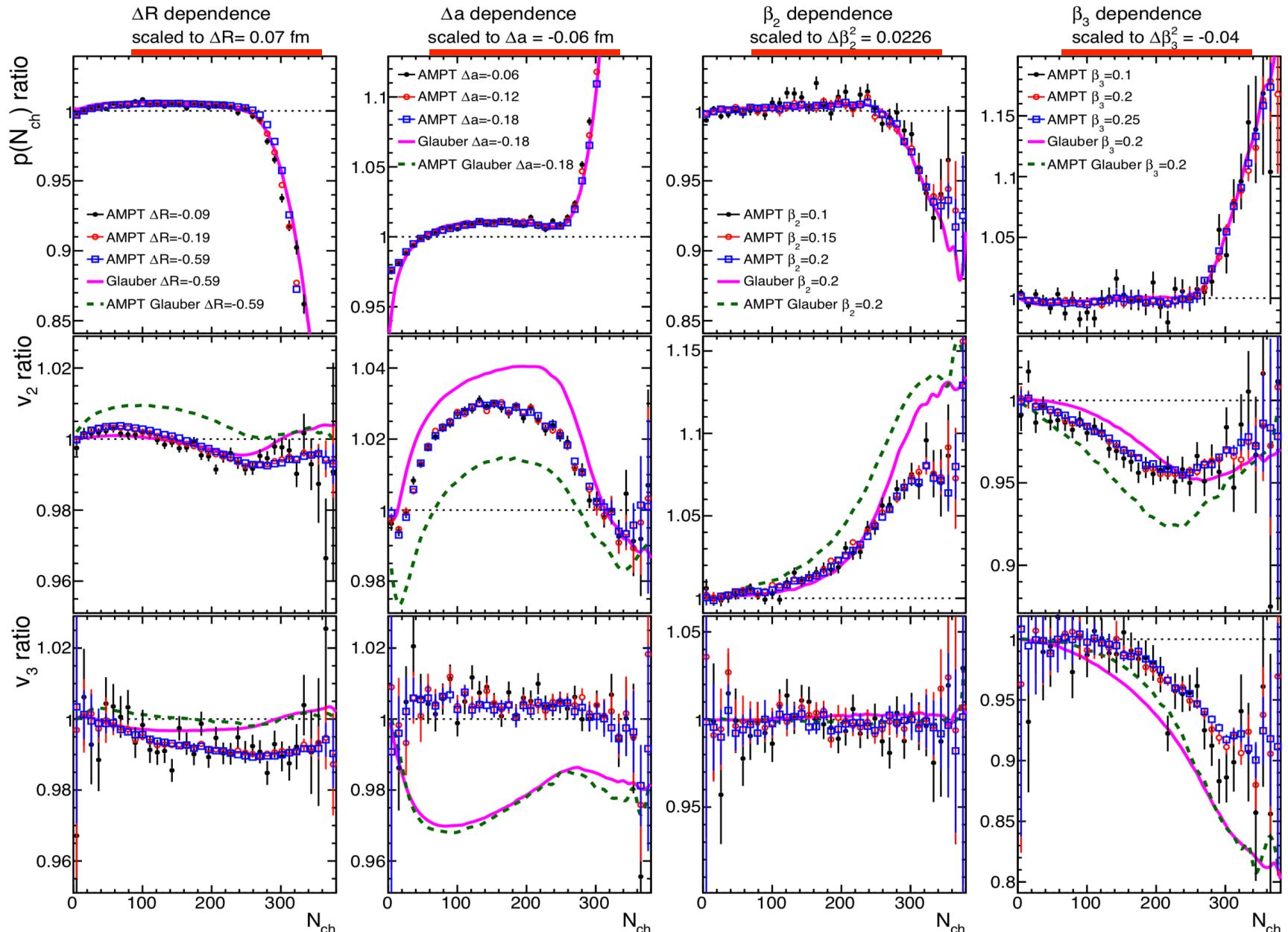


Glauber results: scaled



Verifies the relation: $1 + c_1 \Delta \beta_2^2 + c_2 \Delta \beta_3^2 + c_3 \Delta a + c_4 \Delta R$

AMPT results: scaled

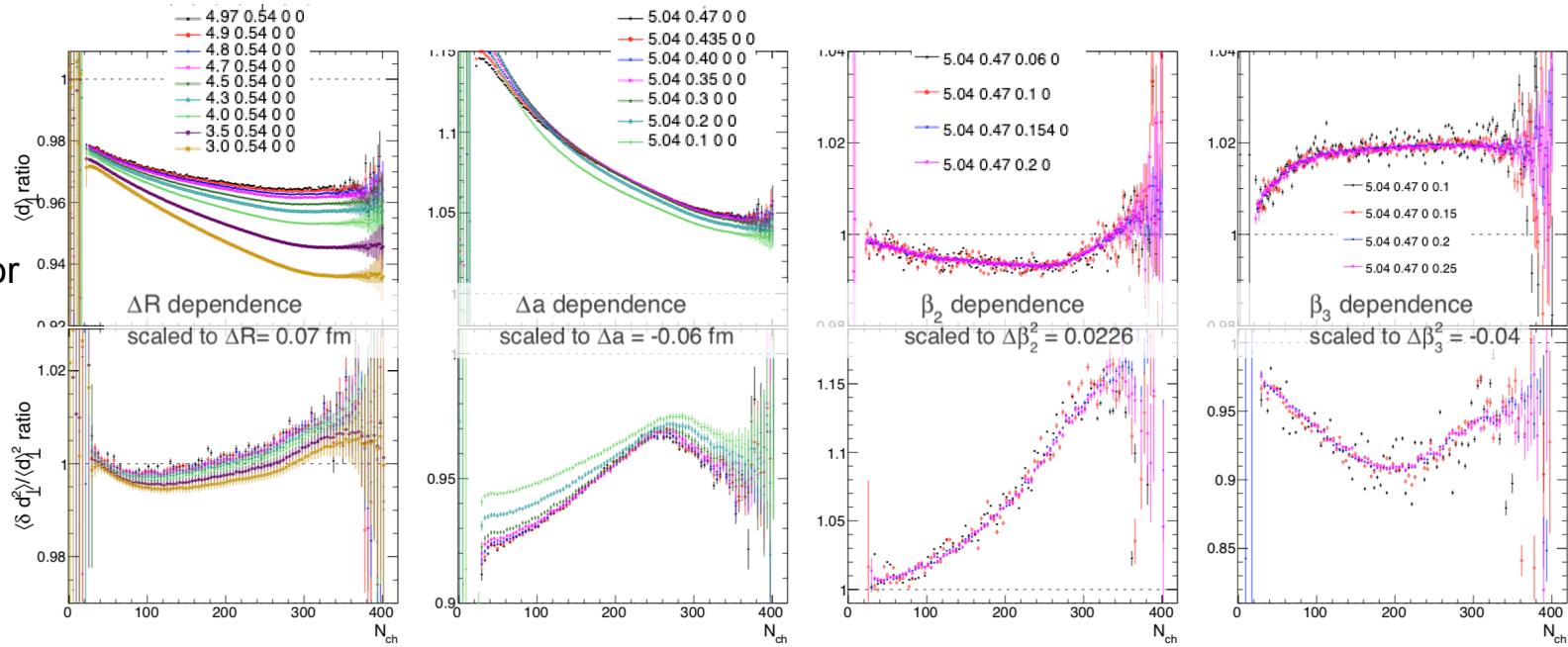


Verifies the relation: $1 + c_1 \Delta \beta_2^2 + c_2 \Delta \beta_3^2 + c_3 \Delta a + c_4 \Delta R$

Scaling approach to nuclear structure

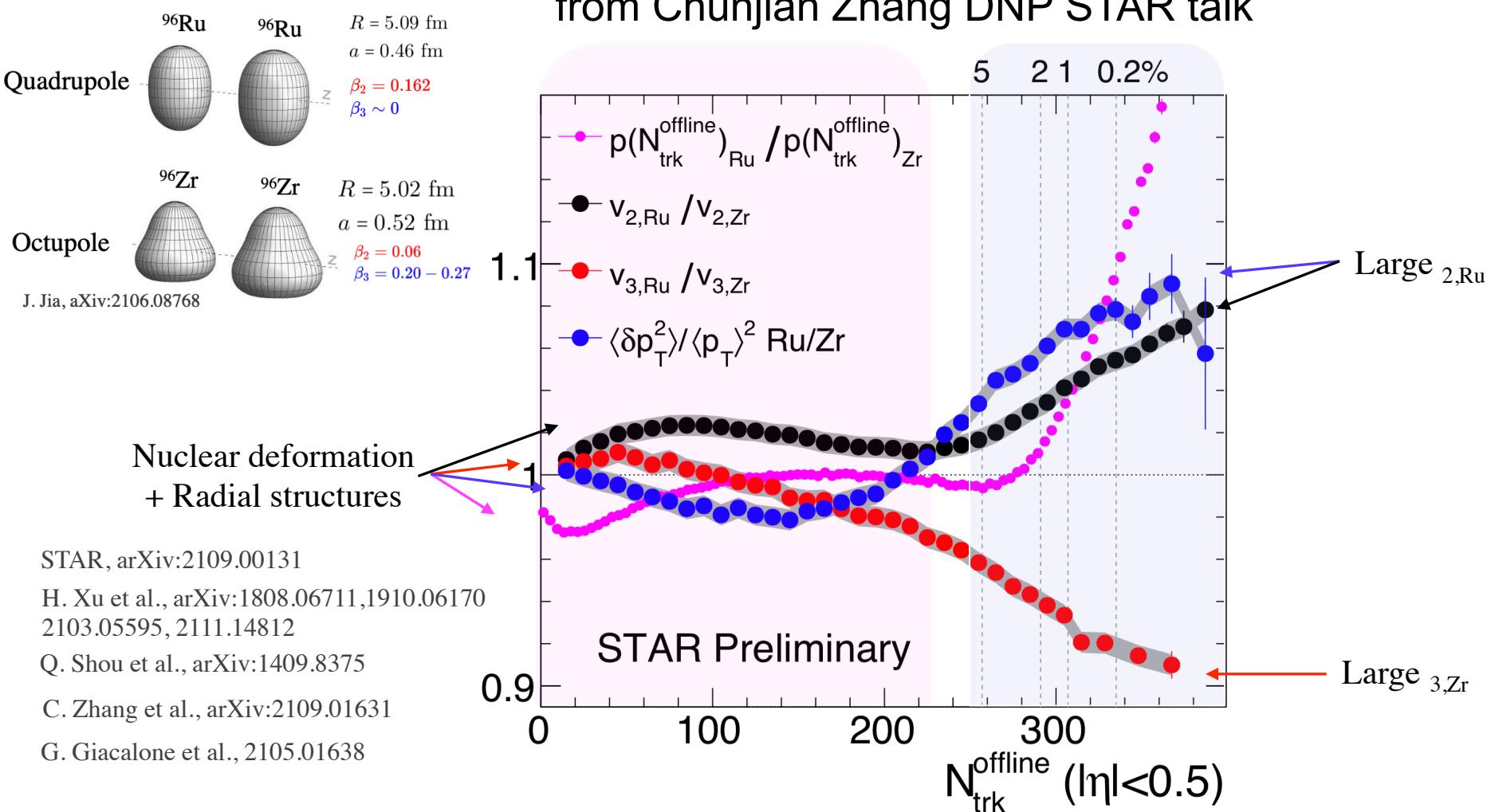
$$\frac{\mathcal{O}_{\text{Ru}}}{\mathcal{O}_{\text{Zr}}} \approx 1 + c_1 \Delta \beta_2^2 + c_2 \Delta \beta_3^2 + c_3 \Delta R_0 + c_4 \Delta a$$

Works well for most single- and two-particle observables:
 $v2, v3, p(N)$, but also $\langle p_T \rangle, \langle \delta p_T^2 \rangle \dots$



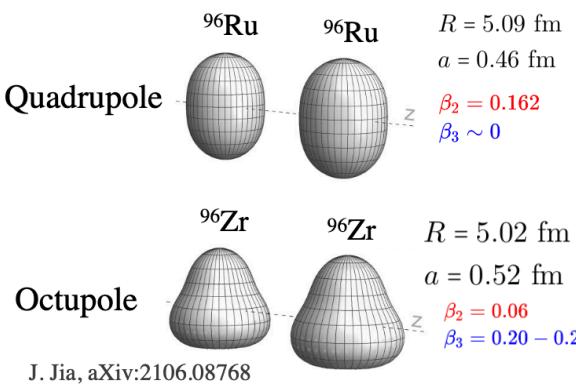
- Determine c_n once, and predict ratios for other parameter values.
- Constrain parameters via χ^2 analysis or Bayesian inference.
- Generalize to multi-particle observables...

Compare with isobar data

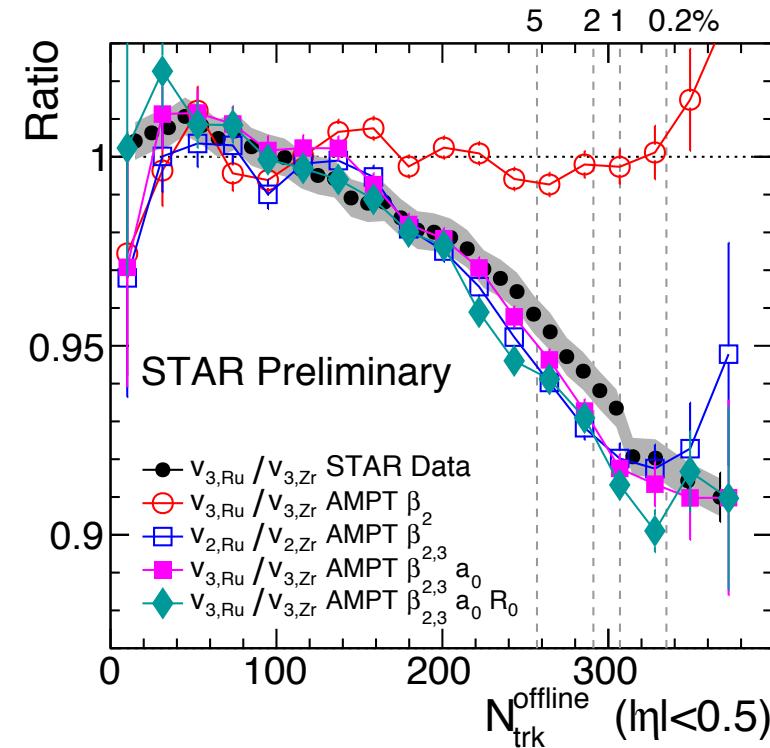
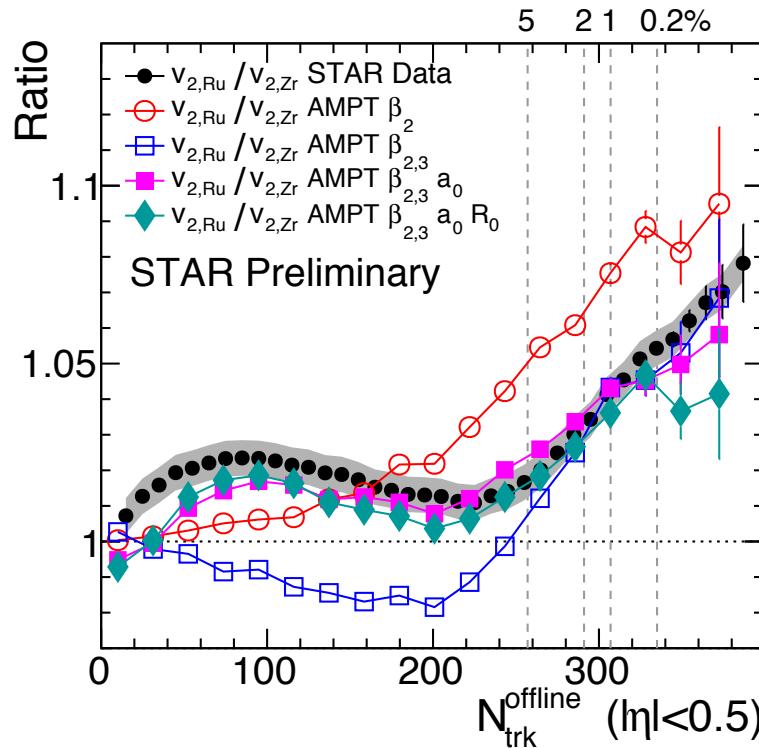


Use these ratios to probe shape and radial structure of nuclei.

Nuclear structure via v_n -ratio

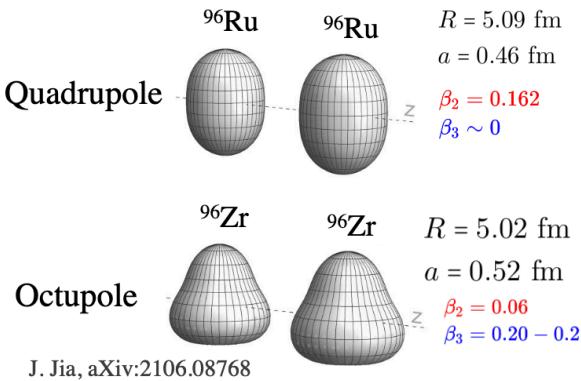


- $\beta_{2\text{Ru}} \sim 0.16$ increase v_2 , no influence on v_3 ratio
- $\beta_{3\text{Zr}} \sim 0.2$ decrease v_2 in mid-central, decrease v_3 ratio
- diffu. $\Delta a_0 = -0.06 \text{ fm}$ increase v_2 mid-central, no influe. on v_3 .
 - Similar study by Haojie et.al.
- Radius $\Delta R_0 = 0.07 \text{ fm}$ only slightly affects v_2 and v_3 ratio.

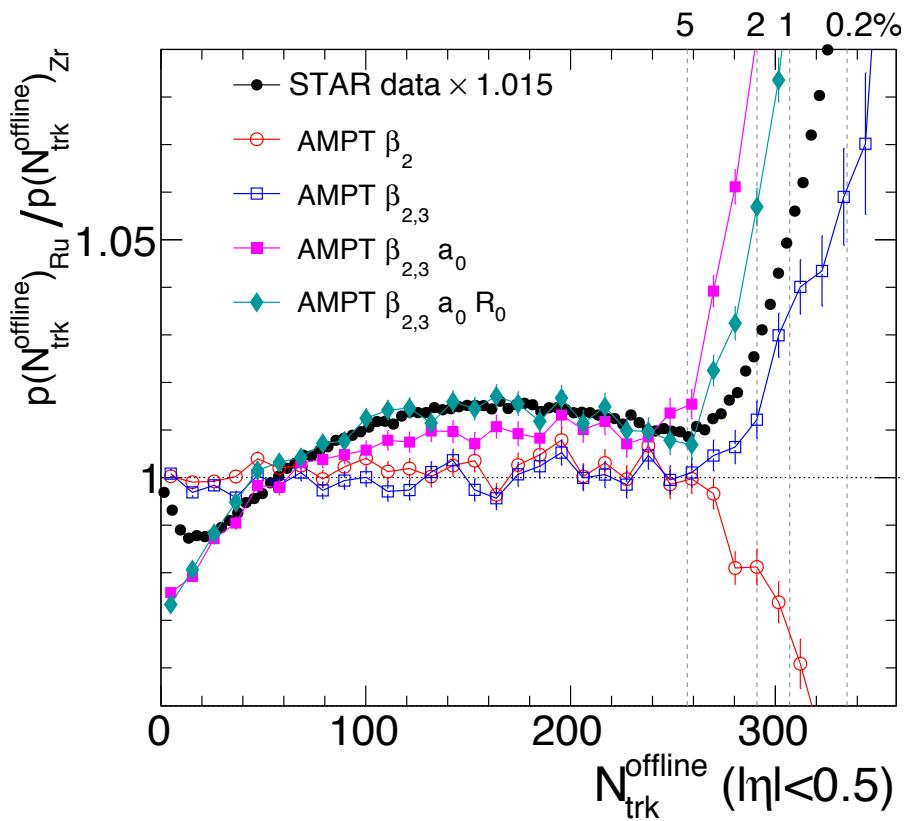


Simultaneously constrain these parameters using different N_{ch} regions

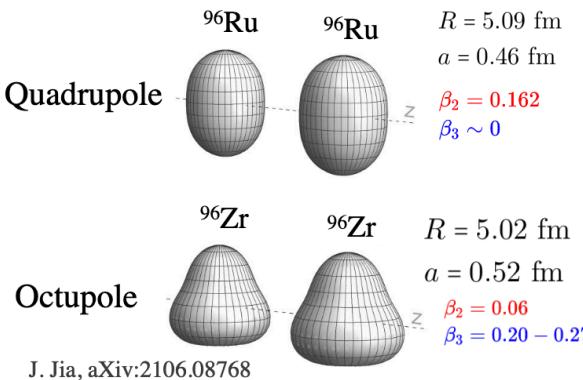
Nuclear structure via p(N_{ch})-ratio



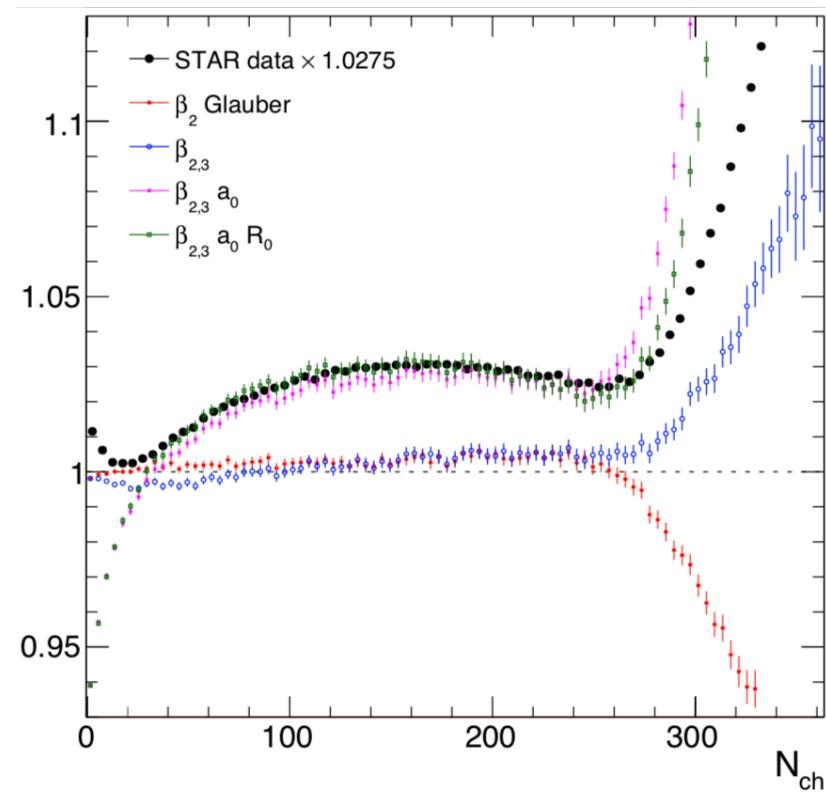
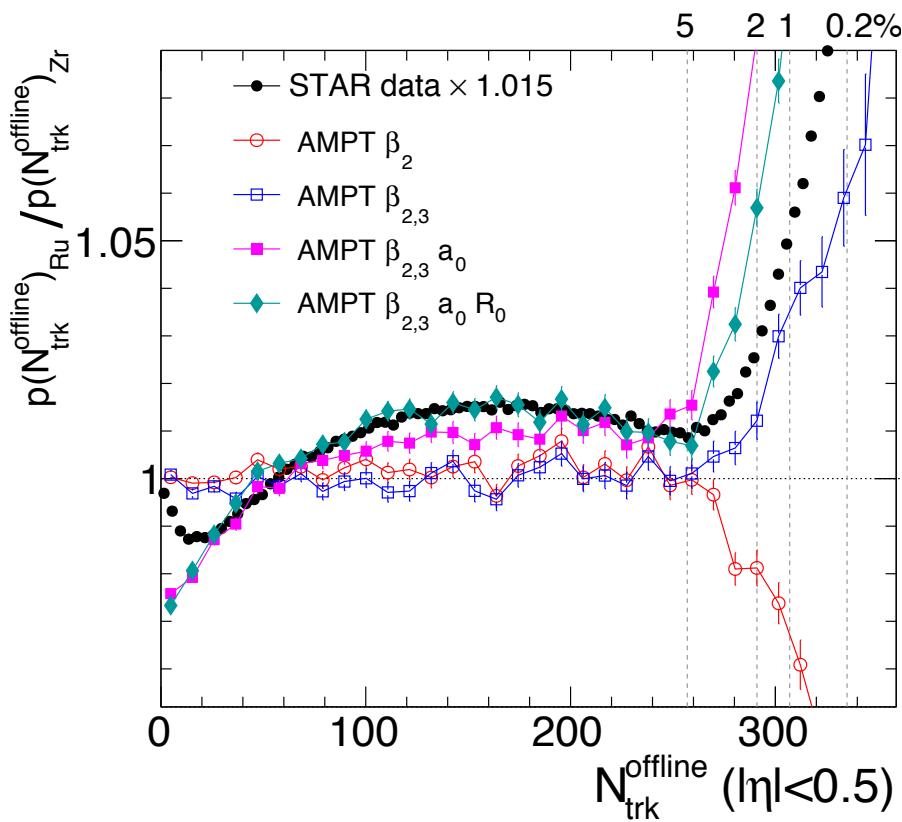
- $\beta_{2\text{Ru}} \sim 0.16$ decrease ratio, increase after considering $\beta_{3\text{Zr}} \sim 0.2$
- The bump structure in non-central region mostly sensitive to differences in surface diffuseness Δa_0 and radius ΔR_0



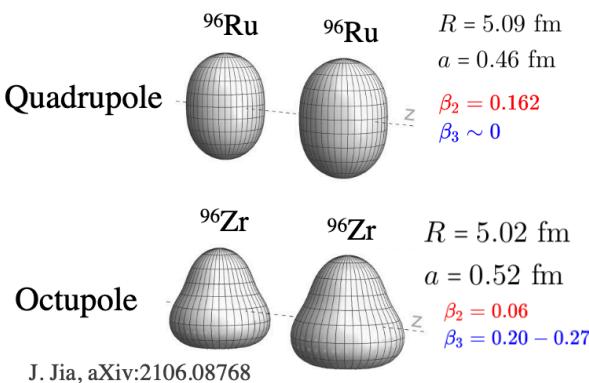
Nuclear structure via p(N_{ch})-ratio



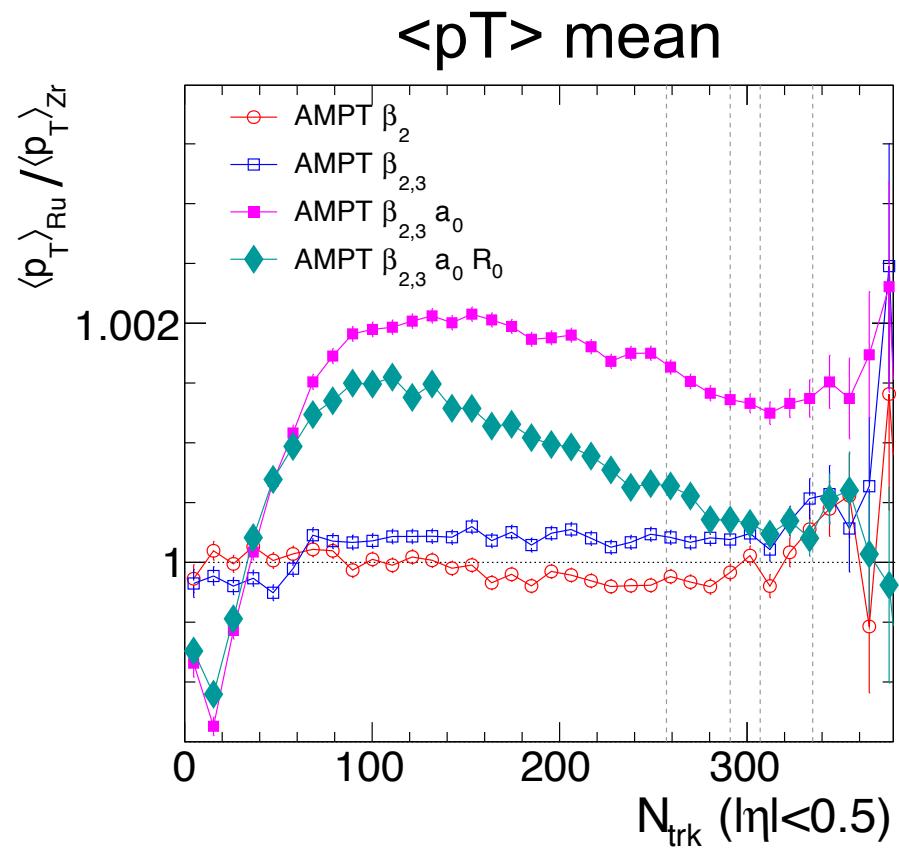
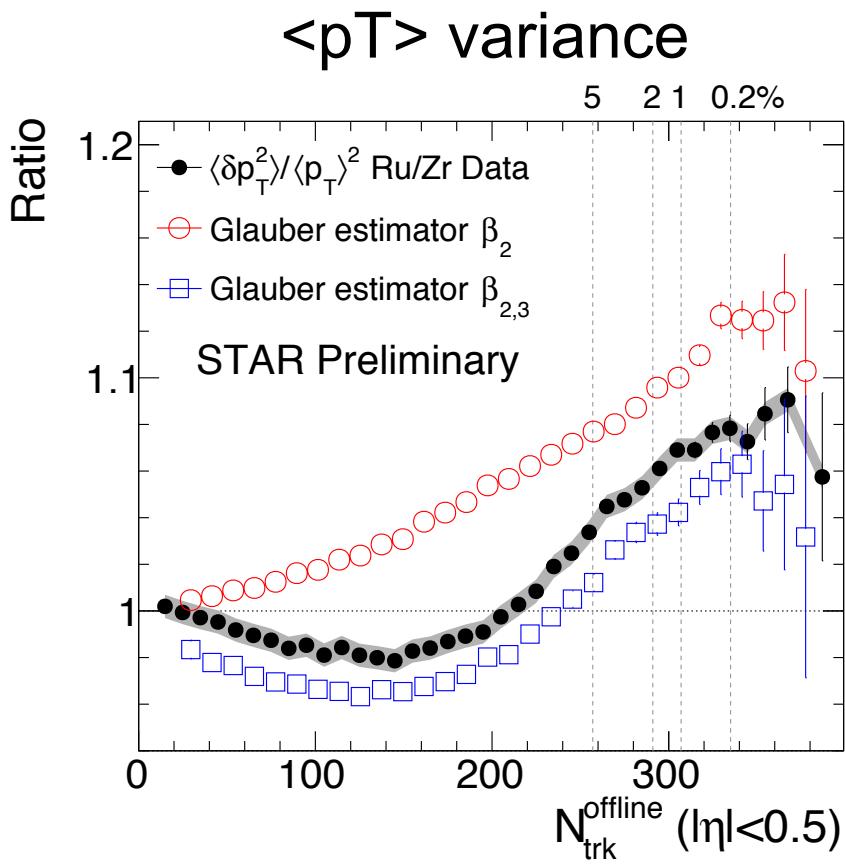
- $\beta_{2\text{Ru}} \sim 0.16$ decrease ratio, increase after considering $\beta_{3\text{Zr}} \sim 0.2$
- The bump structure in non-central region mostly sensitive to differences in surface diffuseness Δa_0 and radius ΔR_0
- All these trends are quantitatively reproduced by Glauber



Structure via radial flow: $\langle p_T \rangle$ & its flucs.



- Glauber model is used by assuming $\frac{\delta[p_T]}{[p_T]} \propto \frac{\delta d_\perp}{d_\perp}$
 - Variance: $\beta_{2\text{Ru}} \sim 0.16$ increase ratio, $\beta_{3\text{Zr}} \sim 0.2$ decrease it
- AMPT has wrong $\langle p_T \rangle$ responses, only for qualitative info
 - R_0 and a plays most important role

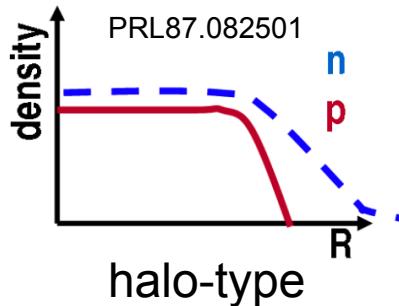
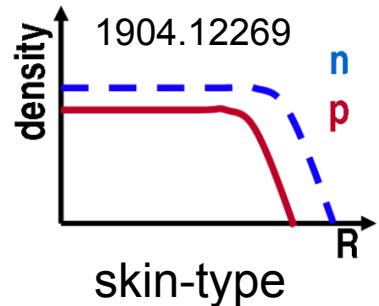


Probing the Neutron Skin

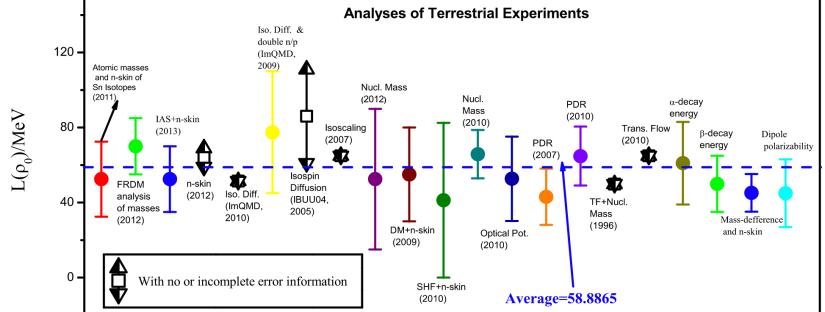
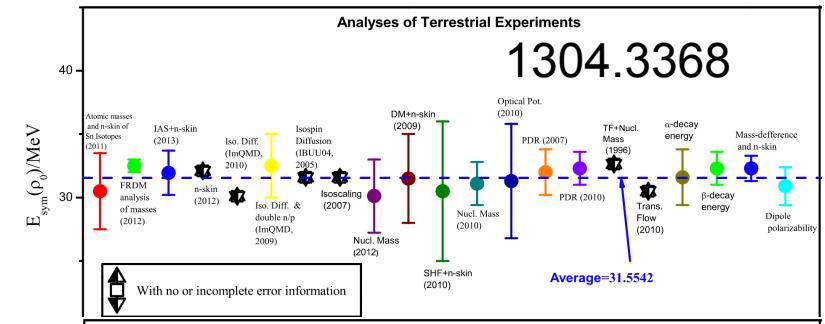
$$\rho = \frac{\rho_0}{1 + e^{-(r-R_0)/a}}$$

Neutron skin:

$$\Delta r_{np} = \langle r_n^2 \rangle^{1/2} - \langle r_p^2 \rangle^{1/2}$$



Constraints from structure and low-energy heavy-ion experiments



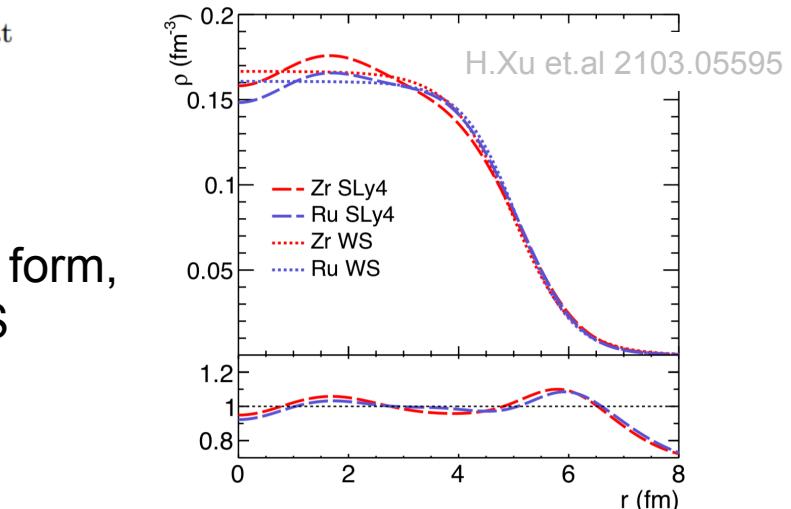
Related to the EOS of symmetry energy, in particular the slope parameter "L"

$$E(\rho, \alpha) \approx E_{\text{SNM}}(\rho) + E_{\text{sym}}(\rho) \cdot \alpha^2 \quad \alpha = (\rho_n - \rho_p)/\rho$$

$$E_{\text{sym}}(\rho) \approx J + Lx + \frac{1}{2}K_{\text{sym}}x^2 \quad x = (\rho - \rho_{\text{sat}})/3\rho_{\text{sat}}$$

Fundamental importance in nuclear and astro-physics

Distribution has more sophisticated form, but typically parameterized with WS



Hydro-response to Neutron skin

Sensitive to L parameter via hydro response: $\langle p_{\perp} \rangle \sim d_{\perp} \equiv \sqrt{N_{\text{part}} / S_{\perp}}$

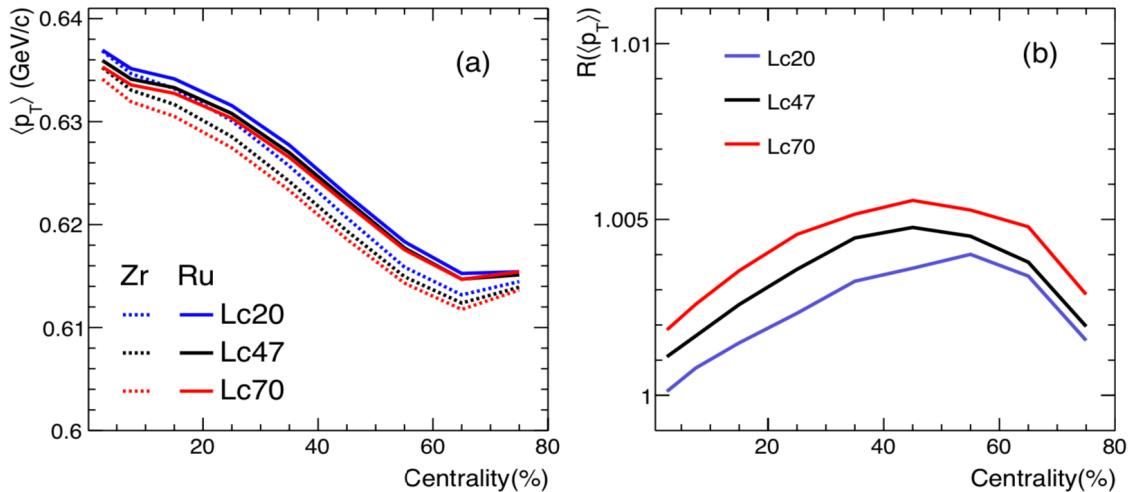
H. Xu et.al.2111.14812

$$E(\rho, \alpha) \approx E_{\text{SNM}}(\rho) + E_{\text{sym}}(\rho) \cdot \alpha^2$$

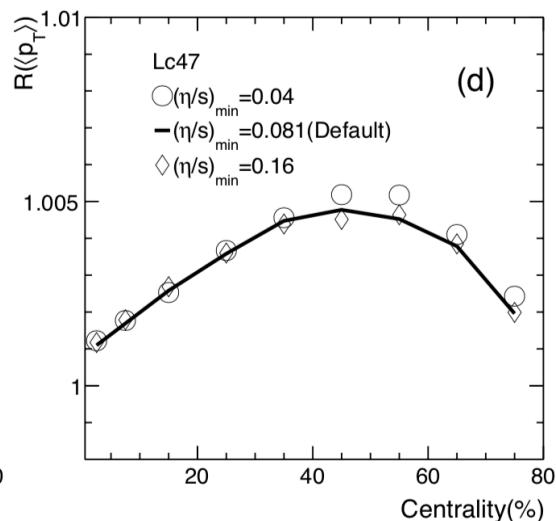
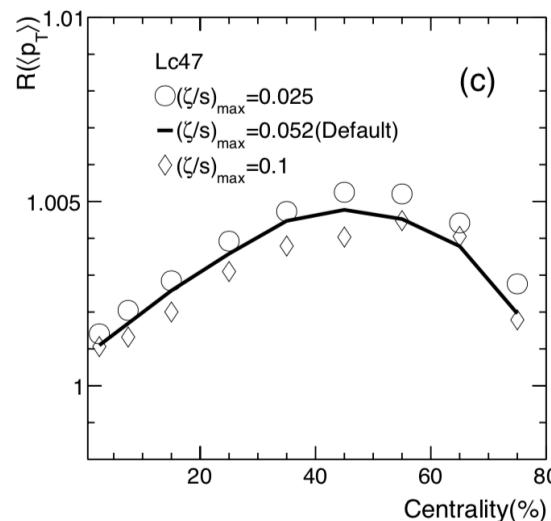
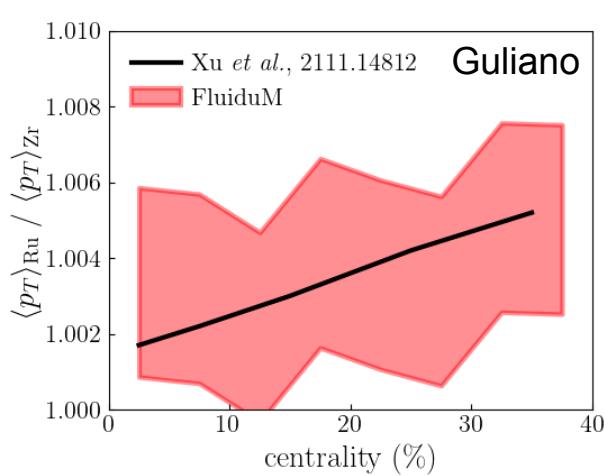
$$E_{\text{sym}}(\rho) \approx J + Lx + \frac{1}{2}K_{\text{sym}}x^2$$

$$\alpha = (\rho_n - \rho_p)/\rho$$

$$x = (\rho - \rho_{\text{sat}})/3\rho_{\text{sat}}$$



Insensitive to final state effects, direct probe of the initial state:



Skin by peripheral isobaric collision

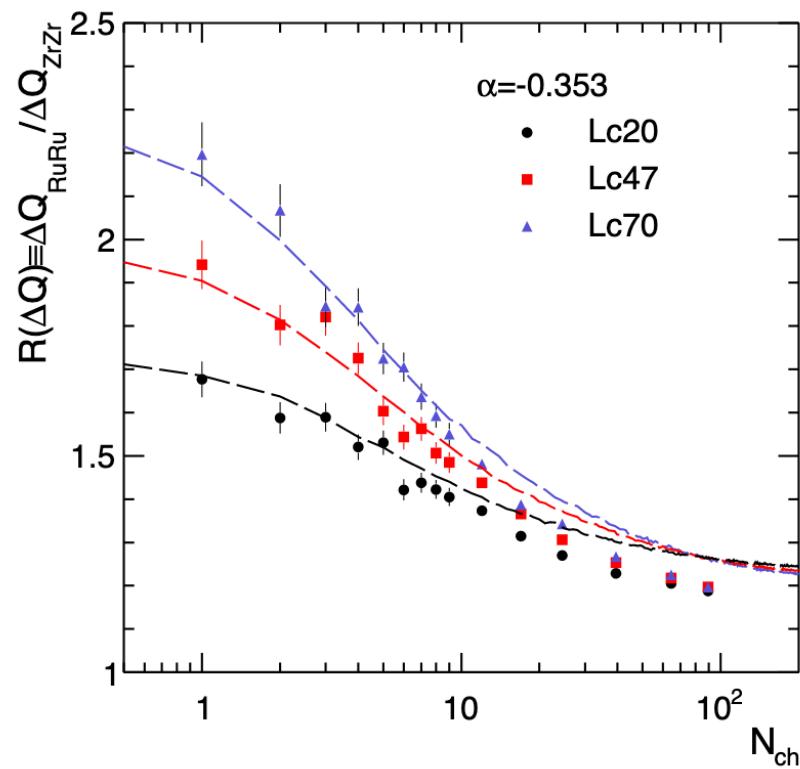
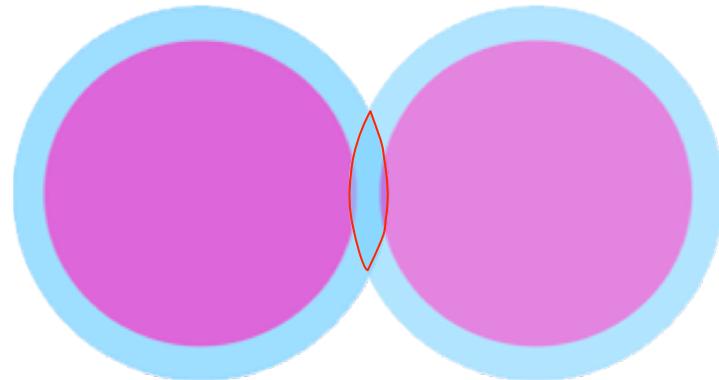
Enhanced skin contribution in the collisions reduces net charge

H. Xu et.al. 2105.04052

$$R_{\Delta Q} \equiv \frac{\Delta Q_{\text{RuRu}}}{\Delta Q_{\text{ZrZr}}} = \frac{q_{\text{RuRu}} + \alpha/(1 - \alpha)}{q_{\text{ZrZr}} + \alpha/(1 - \alpha)}$$

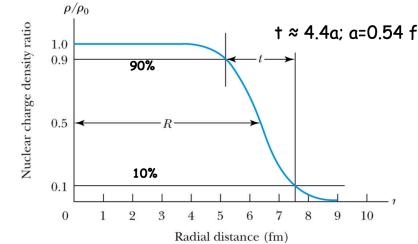
$\alpha \simeq -0.352$ is net-Q ratio in nn vs pp collisions

Large neutron skin of ^{96}Zr
is sensitive to L parameter



A direct link to Neutron Skin

Using relation for WS: $R^2 \equiv \langle r^2 \rangle \approx \left(\frac{3}{5} R_0^2 + \frac{7}{5} \pi^2 a^2 \right) / \left(1 + \frac{5}{4\pi^2} \sum_n \beta_n^2 \right)$



Neutron skin expressed by R and a parameters for nucleons and protons:

$$\Delta r_{np} \approx \frac{R^2 - R_p^2}{R(\delta + 1)} \approx \frac{3(R_0^2 - R_{0,p}^2) + 7\pi^2(a^2 - a_p^2)}{\sqrt{15}R_0 \sqrt{1 + \frac{7\pi^2}{3} \frac{a^2}{R_0^2}} (1 + \delta + \frac{5}{8\pi^2} \sum_n \beta_n^2)} \quad \delta = (N - Z)/A$$

The difference between two isobar can be expressed as

$$\Delta(\Delta r_{np}) = \Delta r_{np,1} - \Delta r_{np,2} \approx \frac{\Delta Y - \frac{7\pi^2}{3} \frac{\bar{a}^2}{R_0^2} \left(\frac{\Delta Y}{2} + \bar{Y} \left(\frac{\Delta a}{\bar{a}} - \frac{\Delta R_0}{R_0} \right) \right)}{\sqrt{15}\bar{R}_0 (1 + \bar{\delta} + \frac{5}{8\pi^2} \sum_n \bar{\beta}_n^2)}$$

where $Y \equiv 3(R_0^2 - R_{0,p}^2) + 7\pi^2(a^2 - a_p^2)$ $\Delta x = x_1 - x_2$ $\bar{x} = (x_1 + x_2)/2$

Can obtain skin diff. from ΔR_0 Δa for nucleons and known ΔR_0 Δa for protons

Example: 2103.05595

	⁹⁶ Ru		⁹⁶ Zr	
	R	a	R	a
p	5.060	0.493	4.915	0.521
n	5.075	0.505	5.015	0.574
p+n	5.067	0.500	4.965	0.556

Direct calc.: $\Delta(\Delta r_{np}) = 0.0296 \text{ fm} - 0.1606 \text{ fm} = -0.1310 \text{ fm}$

Formula: $\Delta(\Delta r_{np}) = -0.1319 \text{ fm}$

How to do system (isobar) scan

arXiv:2106.08768

$$\langle \epsilon_2^2 \rangle = a' + b' \beta_2^2$$

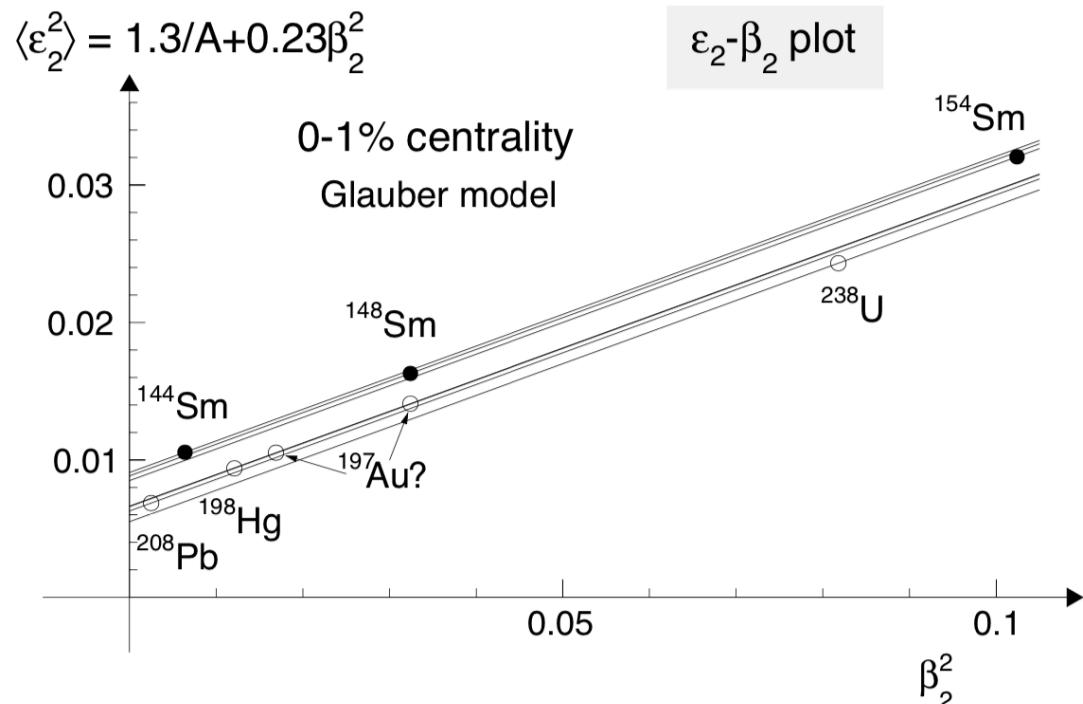
$$\langle v_2^2 \rangle = a + b \beta_2^2$$

In central collisions

$$a' = \langle \epsilon_2^2 \rangle_{|\beta_2=0} \propto 1/A$$

$$a = \langle v_2^2 \rangle_{|\beta_2=0} \propto 1/A$$

b' , b are \sim independent of system



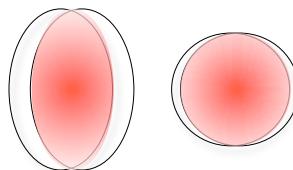
Systems with similar A fall on the same curve.

Fix a and b with two isobar systems with known β_n ,
then make predictions for third one

Similar approach also works for R_0 and diffusivity.

Application in $^{197}\text{Au}+^{197}\text{Au}$ vs $^{238}\text{U}+^{238}\text{U}$

body+body tip+tip



$\text{U}+\text{U}$

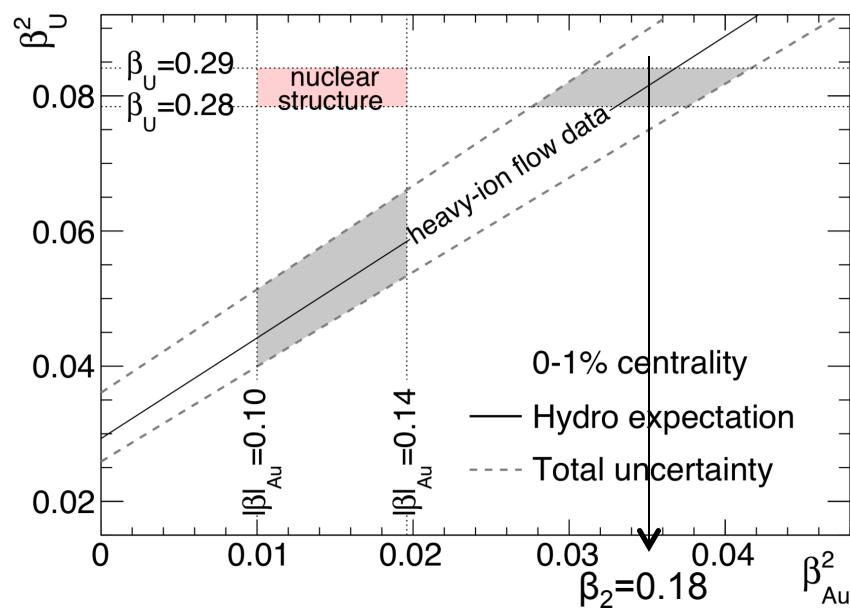
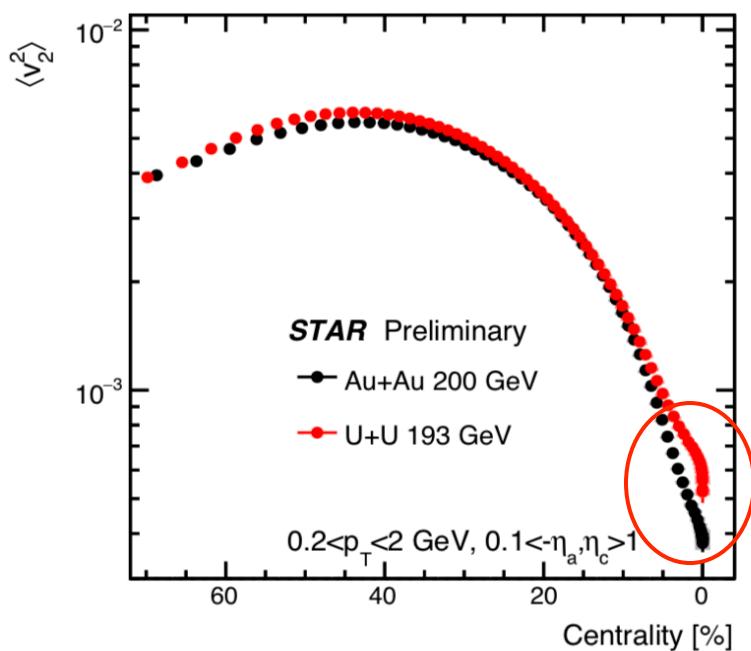
Collisions at $\sqrt{s_{\text{NN}}} = 193\text{-}200 \text{ GeV}$ See:arXiv:2105.01638

$$v_{2,\text{Au}}^2 = a_{\text{Au}} + b\beta_2^2$$

$$v_{2,\text{U}}^2 = a_{\text{U}} + b\beta_2^2$$

Need to correct for slightly different size: $a \propto 1/A$, $r_a = \frac{a_{\text{Au}}}{a_{\text{U}}} = \frac{238}{197} = 1.21$

A simple relation for $\beta_{2\text{U}}$ and $\beta_{2\text{Au}}$: $\beta_{\text{U}}^2 = \frac{r_{v_2^2} r_a - 1}{b/a_{\text{U}}} + r_{v_2^2} \beta_{\text{Au}}^2$ $r_{v_2^2} = \frac{v_{2,\text{U}}^2}{v_{2,\text{Au}}^2}$



Suggests $|\beta_2|_{\text{Au}} \sim 0.18 \pm 0.02$, larger than NS model of 0.13 ± 0.02

Note: ^{197}Au is an odd-mass nucleus, β_2 not measured.

Higher-order correlations

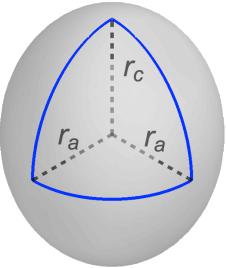
A wide range of possibilities, focus only one topic...

Triaxiality γ :

$$R(\theta, \phi) = R_0 \left(1 + \beta_2 [\cos \gamma Y_{2,0} + \sin \gamma Y_{2,2}] \right)$$

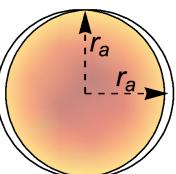
Prolate

$$\beta_2 = 0.25, \cos(3\gamma) = 1$$



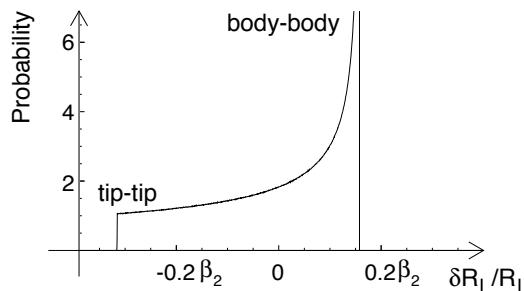
tip+tip

body+body



$$\epsilon_2 \downarrow, R_\perp \downarrow$$

$$\text{cov}(\epsilon_2^2, R_\perp) > 0$$



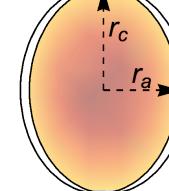
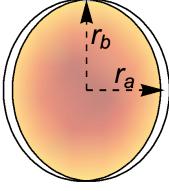
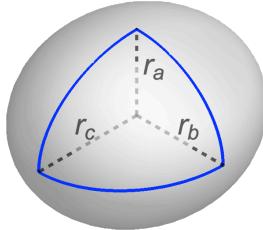
$$\langle (\delta R_\perp)^3 \rangle < 0$$

$$\text{cov}(v_2^2, p_T) < 0 \quad [p_T] \sim 1/R_\perp = d_\perp$$

$$\langle (\delta p_T)^3 \rangle > 0$$

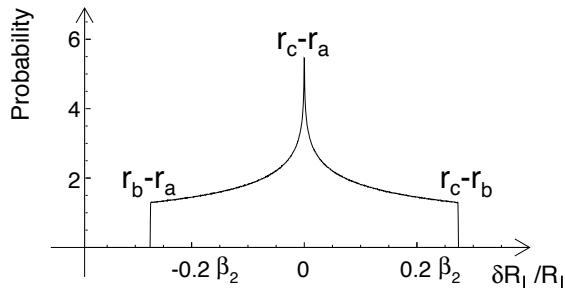
Triaxial

$$\beta_2 = 0.25, \cos(3\gamma) = 0$$



ϵ_2, R_\perp no linear correlation

$$\text{cov}(\epsilon_2^2, R_\perp) = 0$$



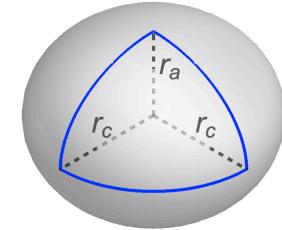
$$\langle (\delta R_\perp)^3 \rangle = 0$$

$$\text{cov}(v_2^2, p_T) = 0$$

$$\langle (\delta p_T)^3 \rangle = 0$$

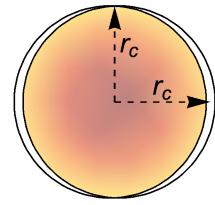
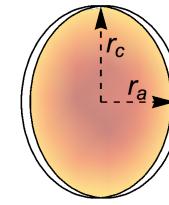
Oblate

$$\beta_2 = 0.25, \cos(3\gamma) = -1$$



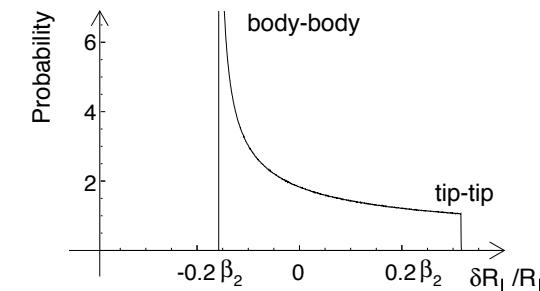
body+body

tip+tip



$$\epsilon_2 \uparrow, R_\perp \downarrow$$

$$\text{cov}(\epsilon_2^2, R_\perp) < 0$$



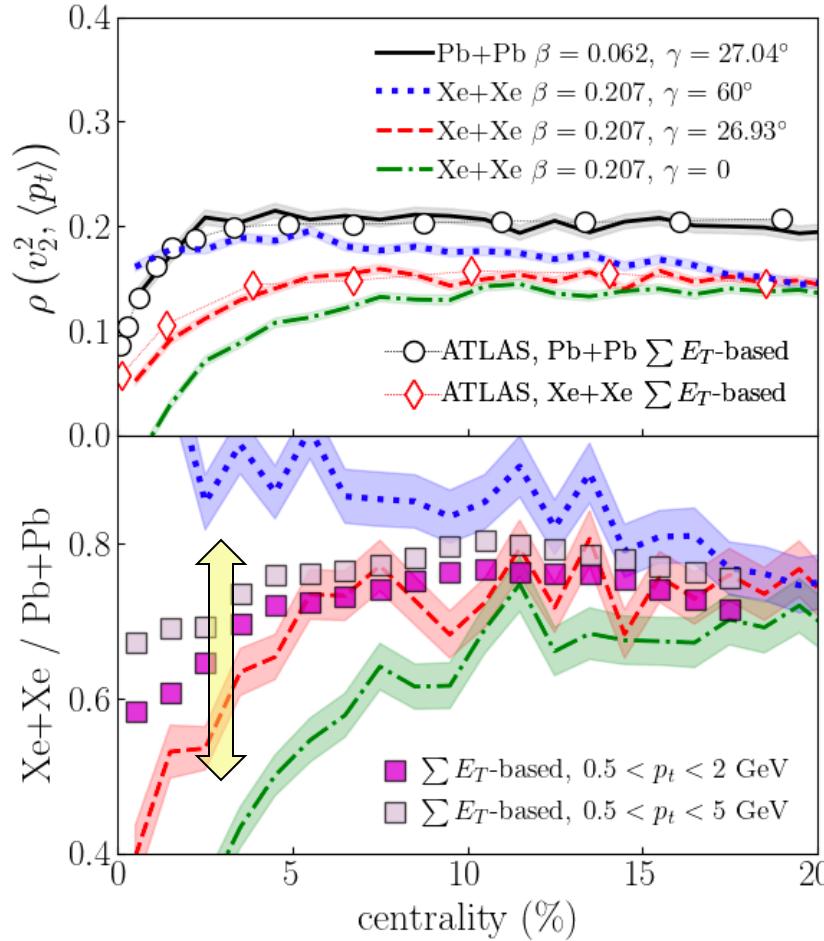
$$\langle (\delta R_\perp)^3 \rangle > 0$$

$$\text{cov}(v_2^2, p_T) > 0$$

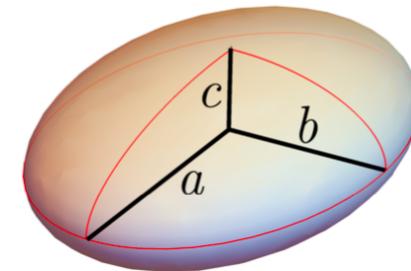
$$\langle (\delta p_T)^3 \rangle < 0$$

v_2^2 - p_T correlation at LHC

B Bally, M Bender, G Giacalone, V Somà 2108.09578



$$R(\theta, \phi) = R_0 \left(1 + \beta_2 [\cos \gamma Y_{2,0} + \sin \gamma Y_{2,2}] \right)$$



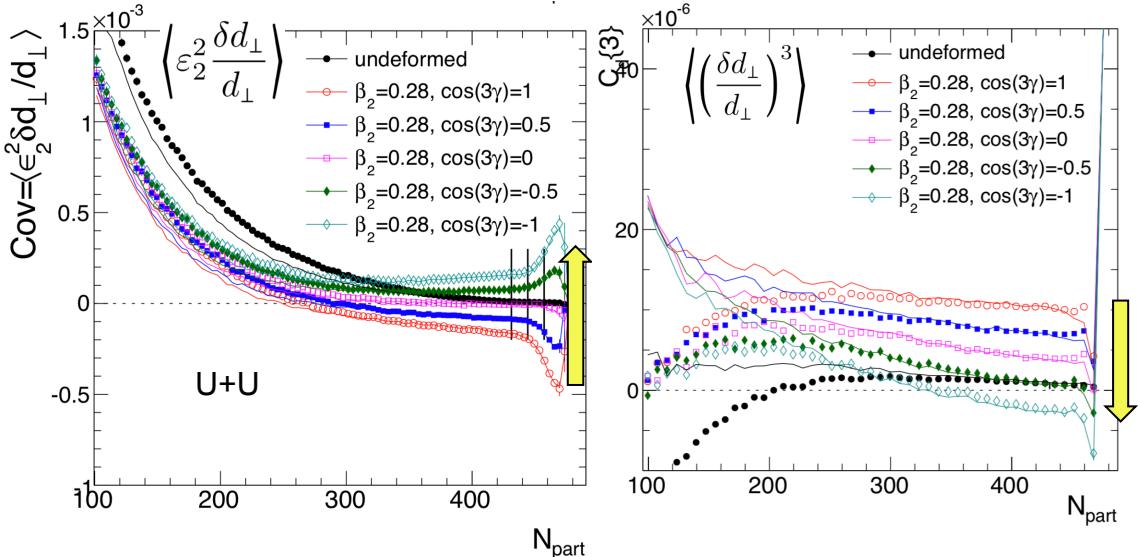
- Clear sensitivity to the triaxiality of ^{129}Xe .

Influence of triaxiality on initial state

Skewness super sensitive ☺

Described by

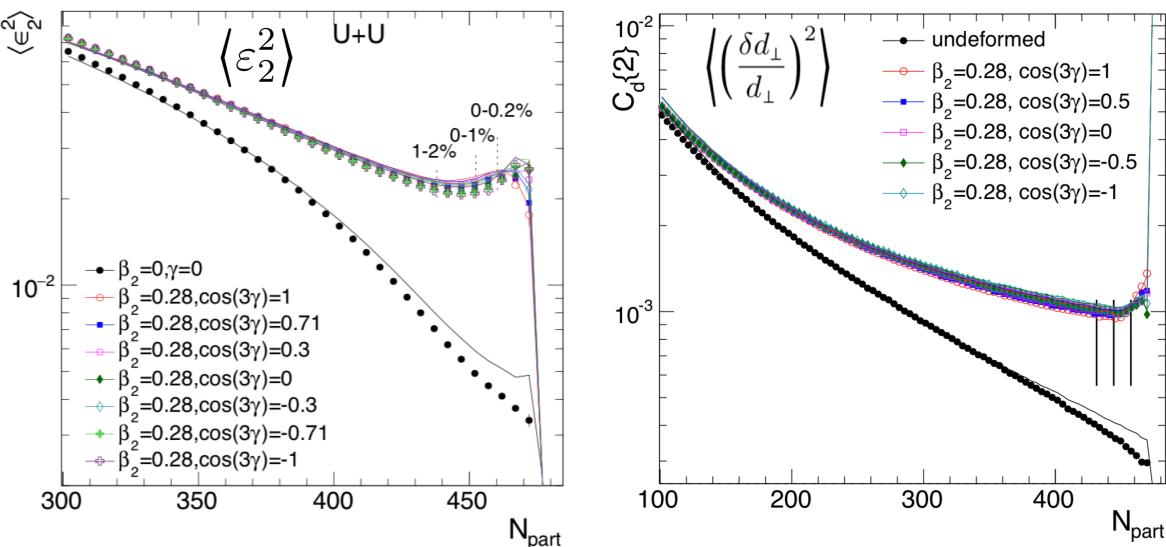
$$a' + (b' + c' \cos(3\gamma)) \beta_2^3$$



variances insensitive to γ

Only a function of β_2 in central

$$a' + b' \beta_2^2$$



Use variance to constrain β_2 , use skewness to constrain γ

(β_2, γ) diagram in heavy-ion collisions

The (β_2, γ) dependence in 0-1% U+U Glauber model can be approximated by:

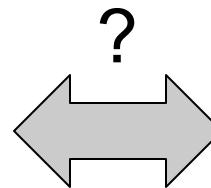
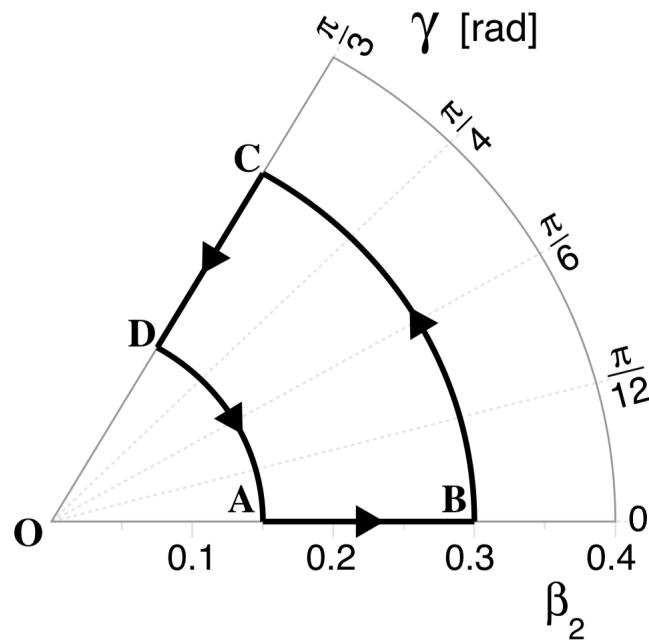
$$\langle \varepsilon_2^2 \rangle \approx [0.02 + \beta_2^2] \times 0.235$$

$$\langle (\delta d_\perp/d_\perp)^2 \rangle \approx [0.035 + \beta_2^2] \times 0.0093$$

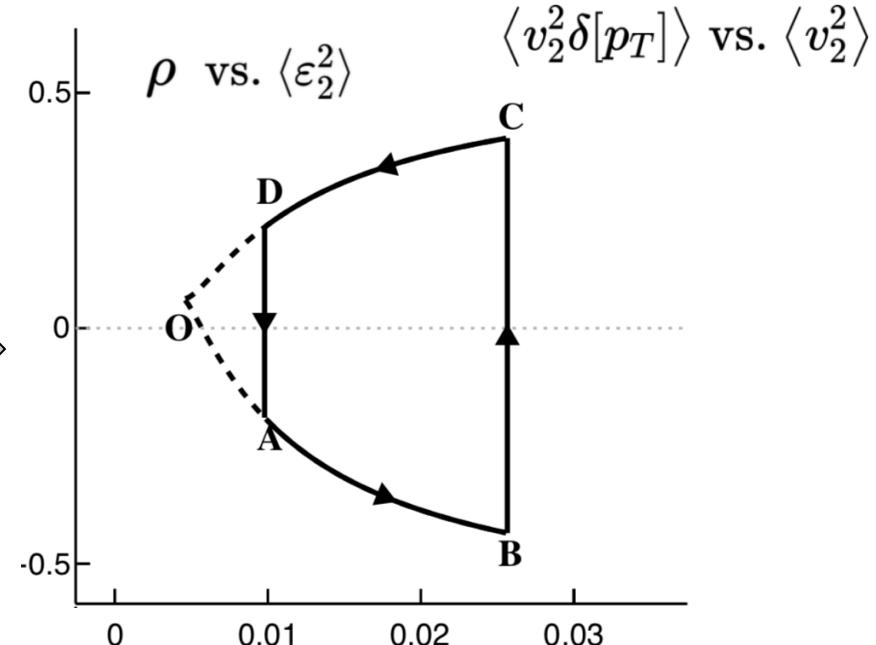
$$\langle \varepsilon_2^2 \delta d_\perp/d_\perp \rangle \approx [0.0005 - (0.07 + 1.36 \cos(3\gamma)) \beta_2^3] \times 10^{-2}$$

$$\rho = \frac{\langle \varepsilon_2^2 \delta d_\perp \rangle}{\langle \varepsilon_2^2 \rangle \sqrt{\langle (\delta d_\perp)^2 \rangle}}$$

Map from (β_2, γ) plane to HI observables

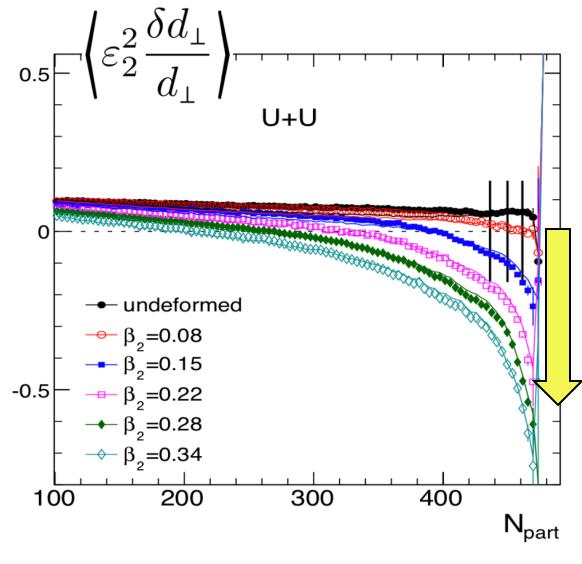
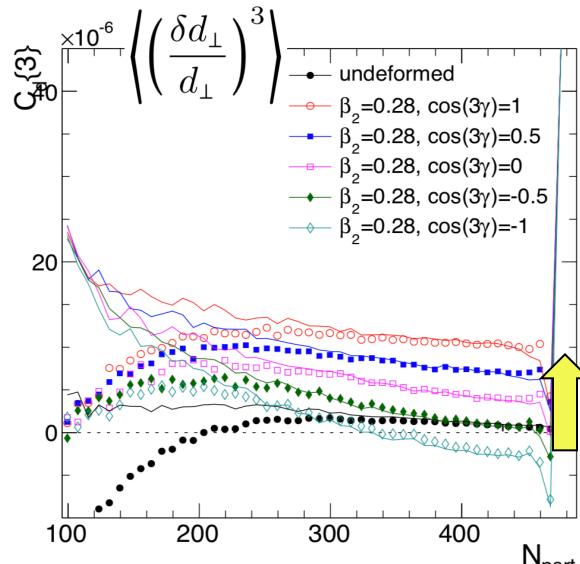
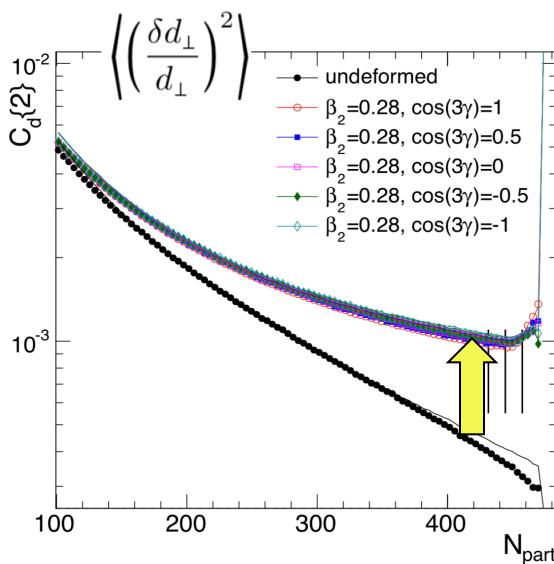
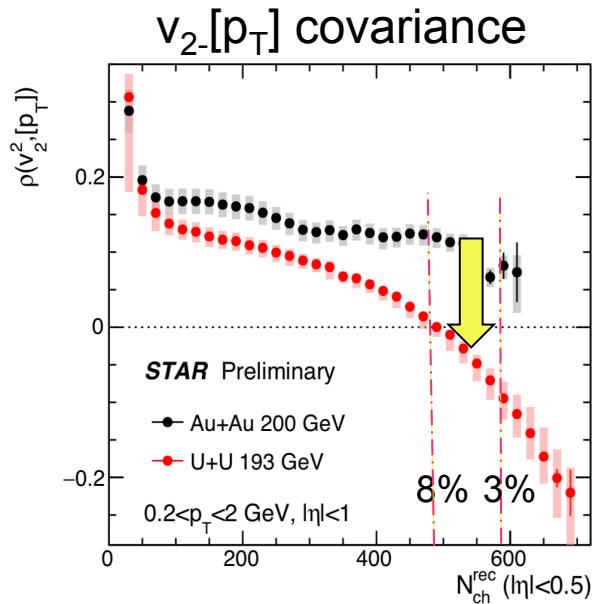
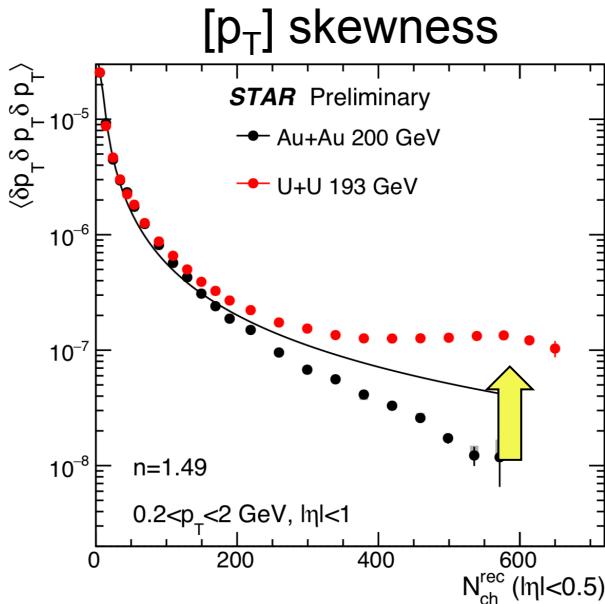
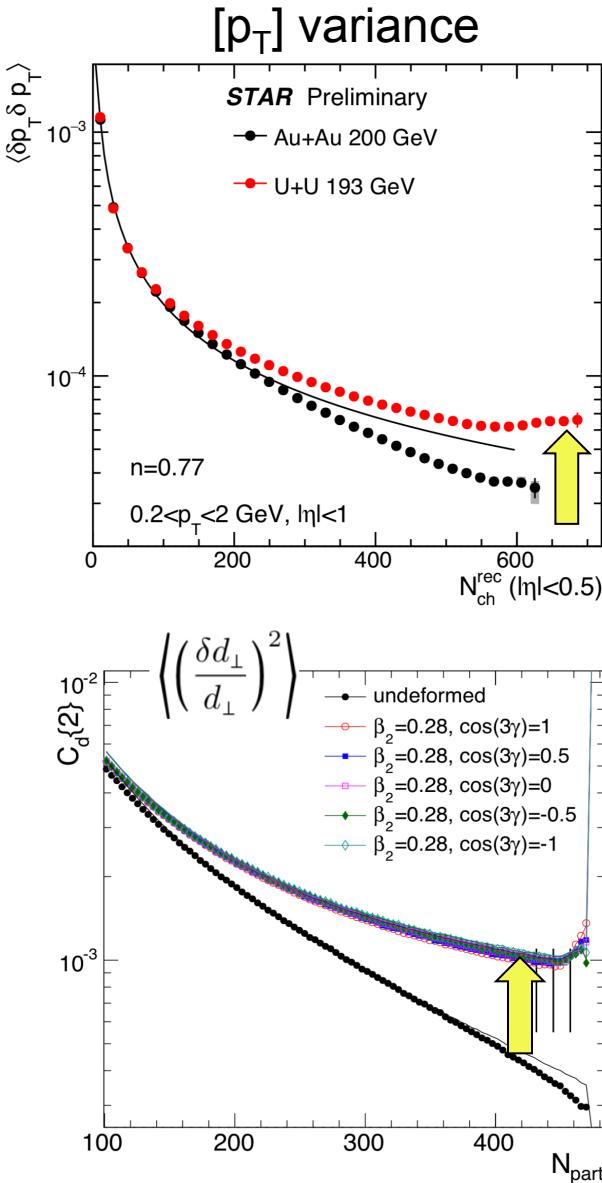


How about



Collision system scan to map out this trajectory: calib. coefficients with species with known β, γ , then predict for species of interest.

Contrast Glauber model with STAR data

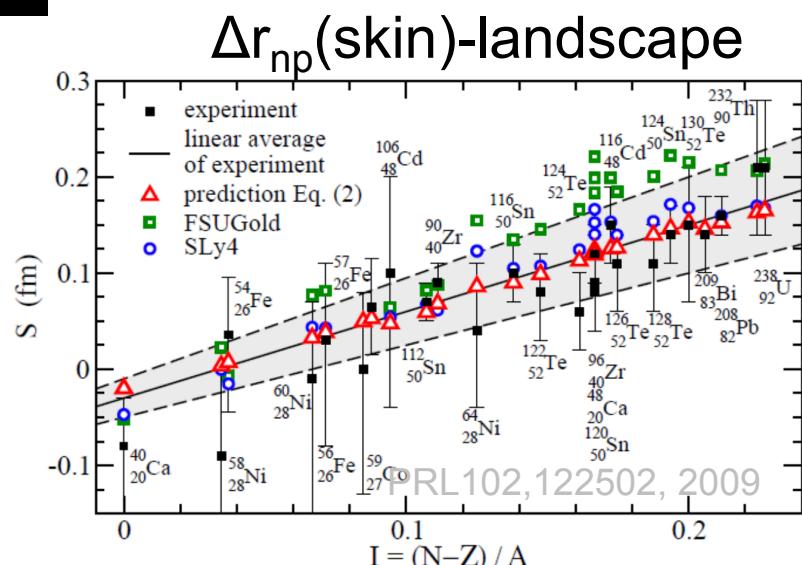
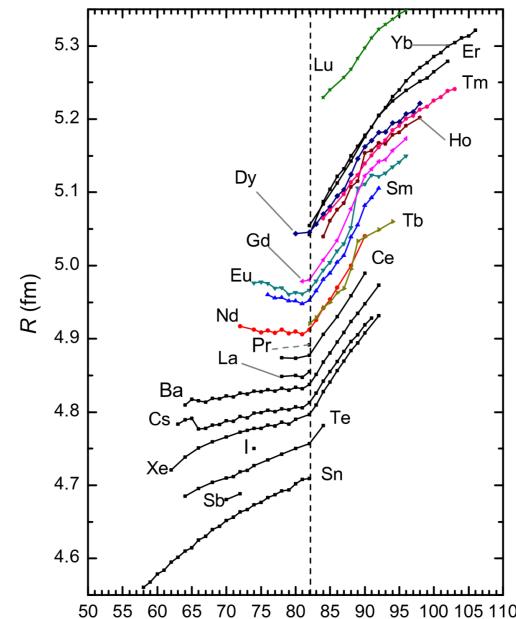
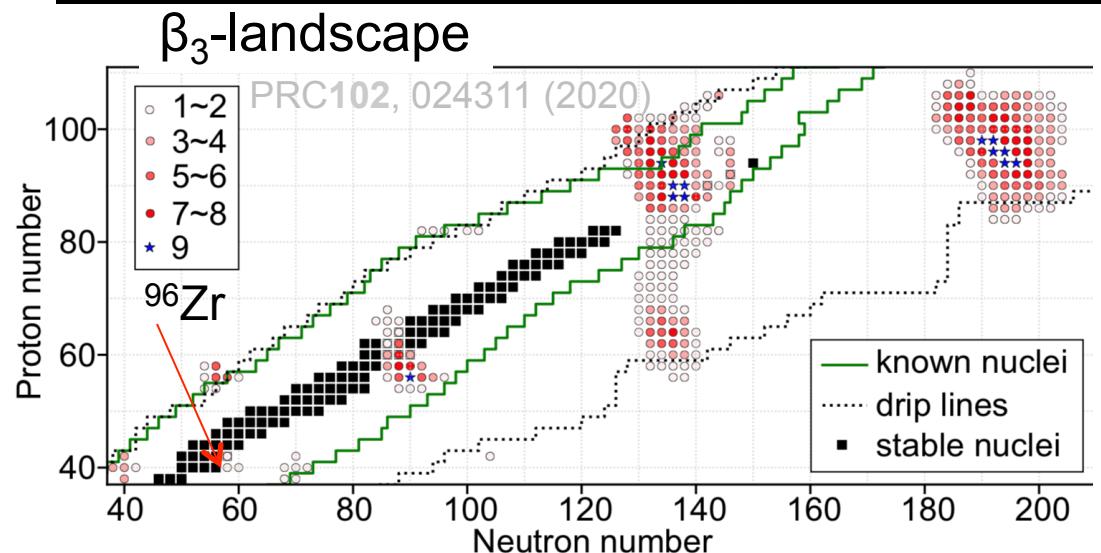
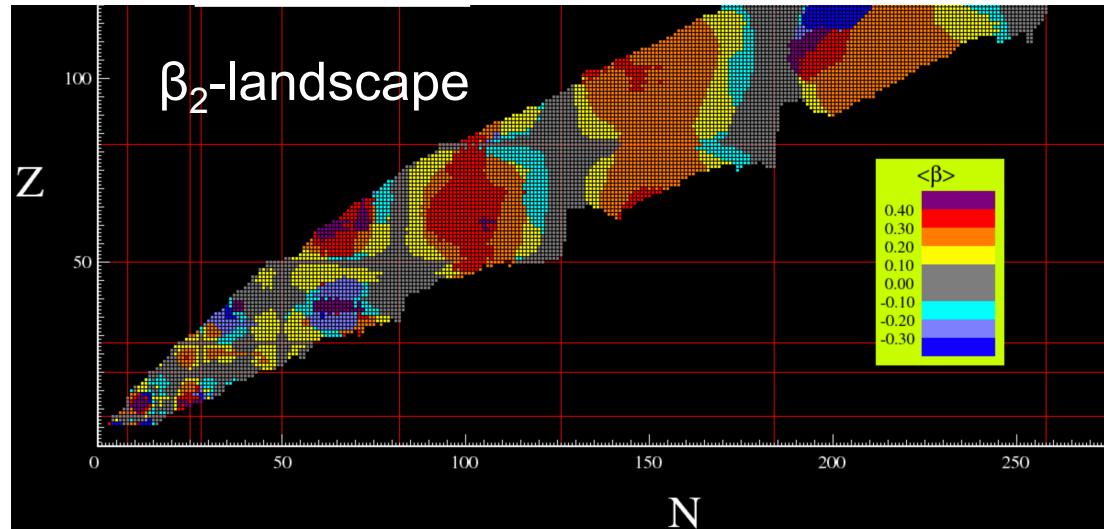


Require high-stat. hydro simulation to quantify the response!

nuclear structure shape/size landscapes

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A lot of possibilities, need to identify most impactful ones!



Thoughts on scan strategy

- Even-even nuclei (0^+ ground state) can be compared with Nuclear Structure exp.
 - We collided several odd-mass ones ☹

A list of large systems
from RHIC and LHC

	β_2	β_3	β_4		β_2	β_3	β_4
^{238}U	0.286 [9]	0.078 [10]	0.094 [10]	^{208}Pb	0.06 [9]	0.04 [11]	?
^{197}Au	-(0.13-0.16) [12, 13]	?	-0.03 [12]	^{129}Xe	0.16 [12]	?	?
^{96}Ru	0.16 [14]	?	?	^{96}Zr	0.06 [14]	0.20-0.27	0.06 [12]

2102.08158

A	isobars	A	isobars	A	isobars
36	Ar, S	106	Pd, Cd	148	Nd, Sm
40	Ca, Ar	108	Pd, Cd	150	Nd, Sm
46	Ca, Ti	110	Pd, Cd	152	Sm, Gd
48	Ca, Ti	112	Cd, Sn	154	Sm, Gd
50	Ti, V, Cr	113	Cd, In	156	Gd, Dy
54	Cr, Fe	114	Cd, Sn	158	Gd, Dy
64	Ni, Zn	115	In, Sn	160	Gd, Dy
70	Zn, Ge	116	Cd, Sn	162	Dy, Er
74	Ge, Se	120	Sn, Te	164	Dy, Er
76	Ge, Se	122	Sn, Te	168	Er, Yb
78	Se, Kr	123	Sb, Te	170	Er, Yb
80	Se, Kr	124	Sn, Te, Xe	174	Yb, Hf
84	Kr, Sr, Mo	126	Te, Xe	176	Yb, Lu, Hf
86	Kr, Sr	128	Te, Xe	180	Hf, W
87	Rb, Sr	130	Te, Xe, Ba	184	W, Os
92	Zr, Nb, Mo	132	Xe, Ba	186	W, Os
94	Zr, Mo	134	Xe, Ba	187	Re, Os
96	Zr, Mo, Ru	136	Xe, Ba, Ce	190	Os, Pt
98	Mo, Ru	138	Ba, La, Ce	192	Os, Pt
100	Mo, Ru	142	Ce, Nd	198	Pt, Hg
102	Ru, Pd	144	Nd, Sm	204	Hg, Pb
104	Ru, Pd	146	Nd, Sm		

- A partial list of isobar pairs
 - Mapping the nuclear structure within each isobar pair.
 - Control system-size dependence hydro-response by stepping between different isobar pairs
 - How to optimize the choices?

Future Opportunities

- STAR proposed to explore such opportunities in the next few years
 - Part1: calibrate systematics with two species around ^{197}Au : ^{208}Pb & ^{198}Hg 5 days each opportunistically
 - ^{208}Pb $\sqrt{s}=0.2$ RHIC vs 5 TeV @LHC: Precision on IS and pre-equilibrium dynamics
 - ^{208}Pb $\sqrt{s}=0.2$ vs ^{197}Au $\sqrt{s}=0.2$ TeV: Quantify effects of Au deformation
 - ^{198}Hg $\sqrt{s}=0.2$ TeV: with known β_2 cross-check the consistency of $\beta_{2\text{Au}}$, γ in ^{197}Au .
 - Part2: explore more exotic regions for triaxial and octupole deformations
 - Scan a isotopic chain: ^{144}Sm ($\beta_2=0.08$), ^{148}Sm ($\beta_2=0.14$, triaxial), ^{154}Sm ($\beta_2=0.34$) potentially for 2026 if running
 - Compare a pair with equal mass: ^{154}Sm ($\beta_2 = 0.34$) and ^{154}Gd ($\beta_2 = 0.31$) potentially for 2026 if running

But priority is sPHENIX, very limited possibility at RHIC.

- Maybe at LHC beyond 2030 (RUN5)? See CERN yellow report

Table 4: Parameters and performance for a range of light nuclei with a moderately optimistic value of the scaling parameter $p = 1.5$ in (5).

	$^{16}\text{O}^{8+}$	$^{40}\text{Ar}^{18+}$	$^{40}\text{Ca}^{20+}$	$^{78}\text{Kr}^{36+}$	$^{129}\text{Xe}^{54+}$	$^{208}\text{Pb}^{82+}$
γ	3760.	3390.	3760.	3470.	3150.	2960.
$\sqrt{s_{\text{NN}}}/\text{TeV}$	7.	6.3	7.	6.46	5.86	5.52

- Other heavy ion facilities such as NICA? $\sqrt{s_{\text{NN}}}$ up to 11 GeV
 - Measure the same isobar pair at different energies.

Manifestation of nuclear structure are \sqrt{s} and rapidity dependent!

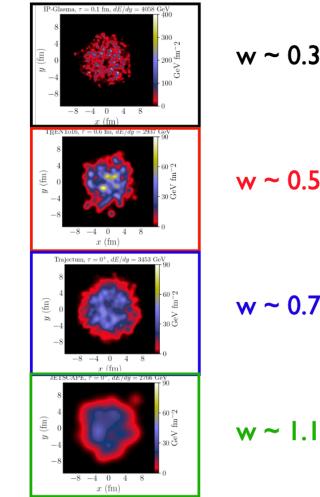
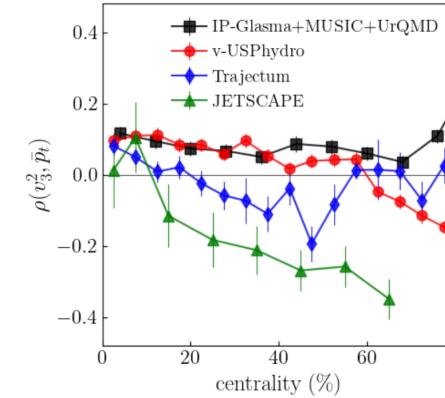
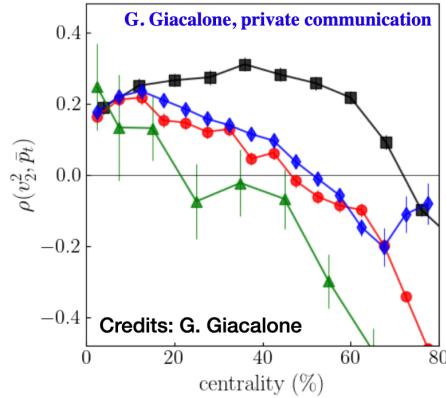
Energy dependence of nucleon size

❖ Sensitive to the nucleon width parameter (size of nucleon)

- IP-Glasma ~ 0.3 ; v-USPhydro ~ 0.5 ; Trajectum ~ 0.7 ; JETSCAPE (T_{RENT}) ~ 1.1
- $w(\text{IP-Glasma}) < w(\text{v-USPhydro}) < w(\text{Trajectum}) < w(\text{JETSCAPE})$

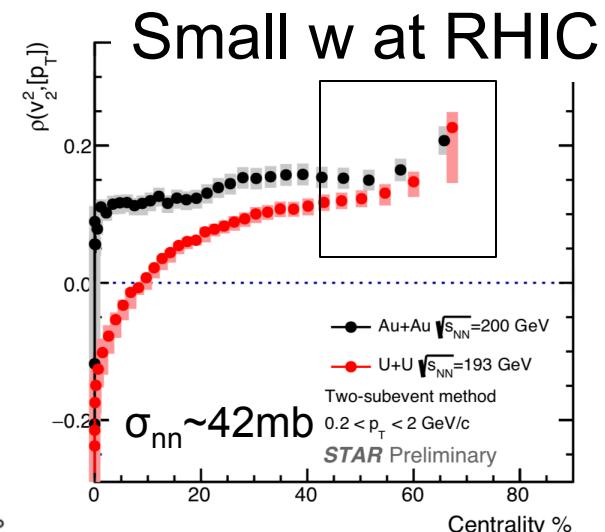
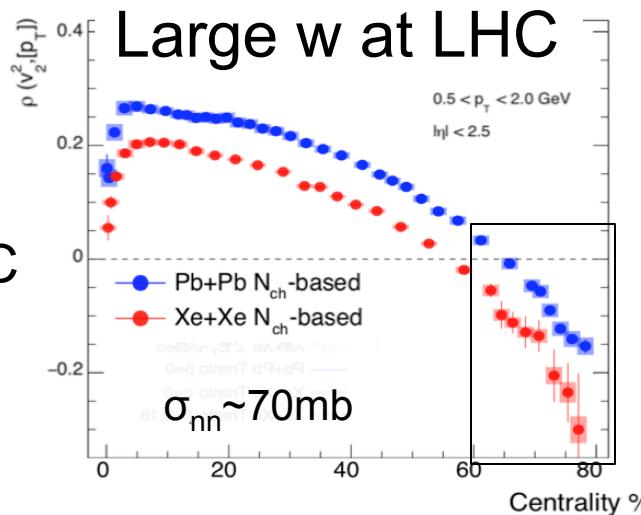
Slide from You Zhou

$v_n^2 - p_T$ has strong sensitivity to nucleon width



Nucleon size have strong $\sqrt{s_{nn}}$ dependence

Very different peripheral behavior at RHIC vs LHC



Questions

- How are nuclear shape and radial profile inferred from hydrodynamic response related to properties measured in nuclear structure experiments?
- How does the uncertainty brought by nuclear structure impact the initial state of heavy-ion collisions and extraction of QGP transport properties?
- What is the energy and longitudinal dependence of nuclear structure?

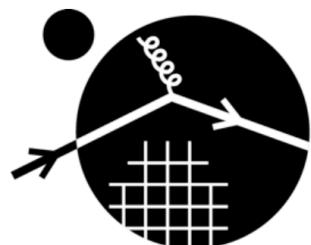
RIKEN BNL Research Center

Physics Opportunities from the RHIC Isobar Run

This workshop will be held virtually.

January 25–28, 2022

<https://www.bnl.gov/porir2022/index.php>



INSTITUTE for
NUCLEAR THEORY

**“Intersection of nuclear structure and
high-energy nuclear collisions”**

January 23 – February 24, 2023

Giuliano Giacalone, Jiangyong Jia, Dean Lee,
Matthew Luzum, Jaki Noronha-Hostler , Fuqiang Wang