

Double Beta Decay III: The State of the Field

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National Nuclear Physics Summer School

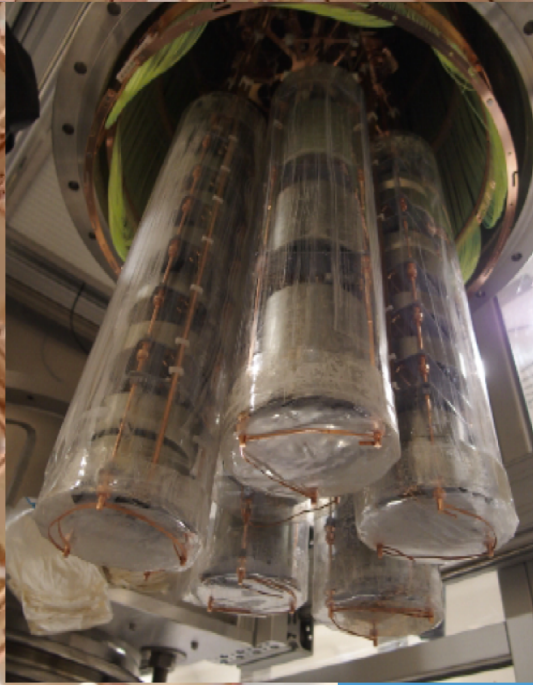
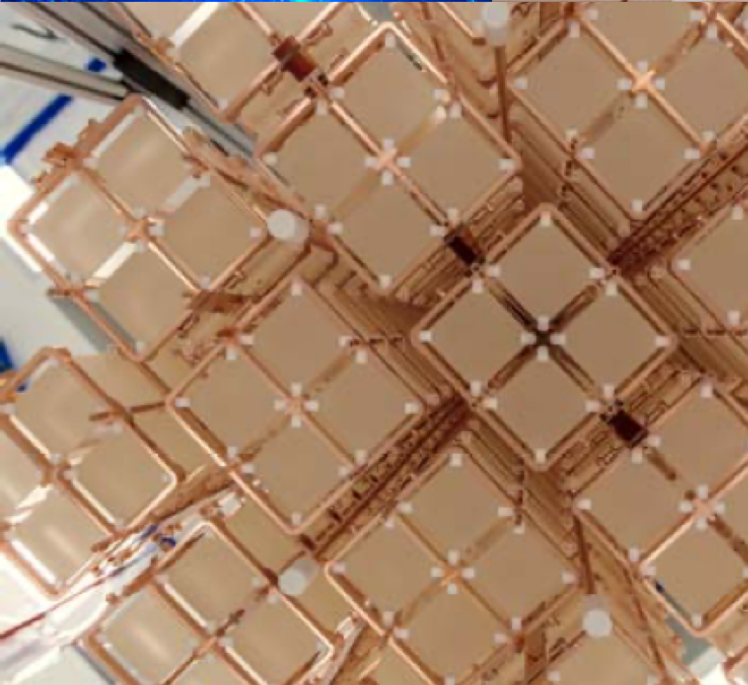
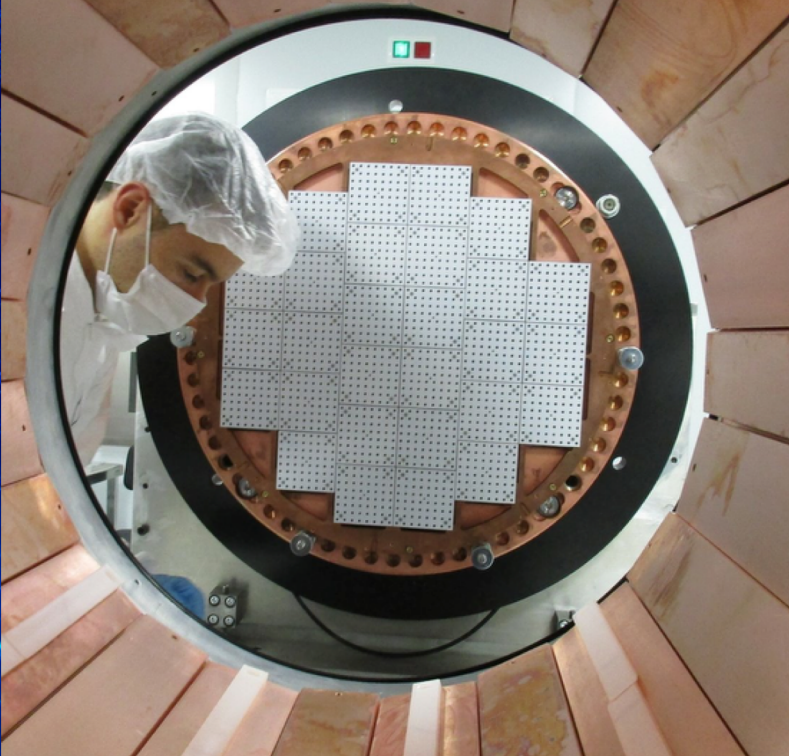
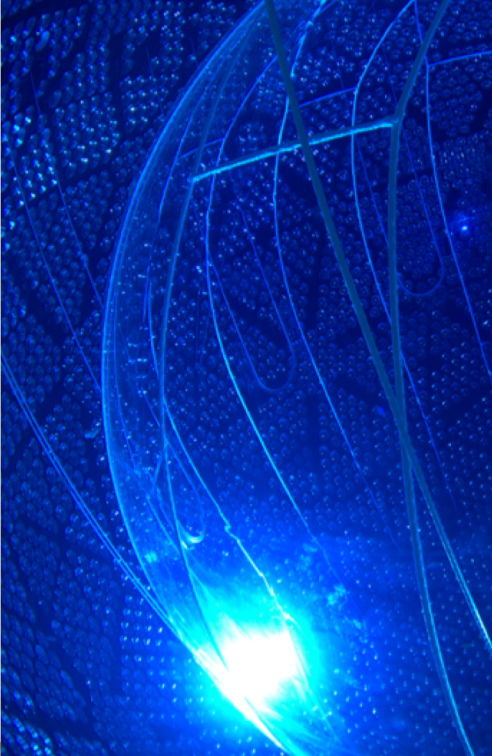
July 15, 2022



THE UNIVERSITY
of NORTH CAROLINA
at CHAPEL HILL

Reminder from Yesterday

- Calculating NMEs for $0\nu\beta\beta$ is challenging:
 - Mean-field methods let you calculate matrix elements for large nuclei, but uncertainties are impossible to quantify
 - Ab-initio methods are advancing quickly, but can't address all $\beta\beta$ decay elements yet
- To discover $0\nu\beta\beta$, we need very large experiments with very low backgrounds
 - Most experiments use a source=detector strategy to maximize efficiency
 - Background rejection based on event topology, location, type of interaction, and other information is helpful



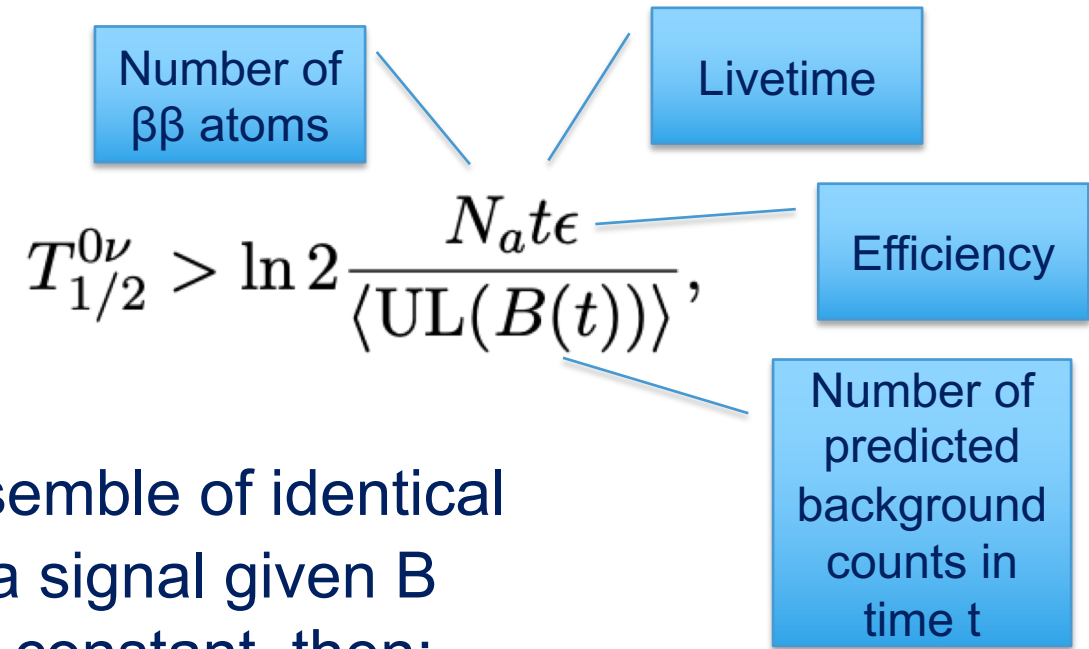
Outline

- Wednesday: Why look for $0\nu\beta\beta$?
- Thursday: How to look for $0\nu\beta\beta$
- Friday: The State of the Field
 - How Measurements Work
 - The $0\nu\beta\beta$ Search Landscape
 - The Experiments:
 - Liquid Scintillator Experiments
 - TPCs: Liquid and Gas
 - Bolometer Experiments
 - Germanium Experiments
 - Tracking Experiments
 - Doing other physics with $0\nu\beta\beta$ experiments
 - What comes after discovery?

How Measurements Work

Measuring and Setting Limits on $T_{1/2}$

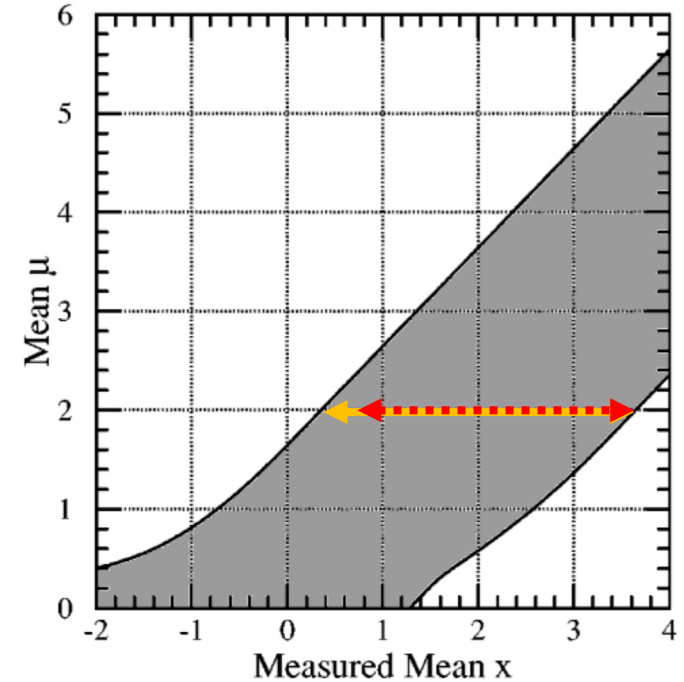
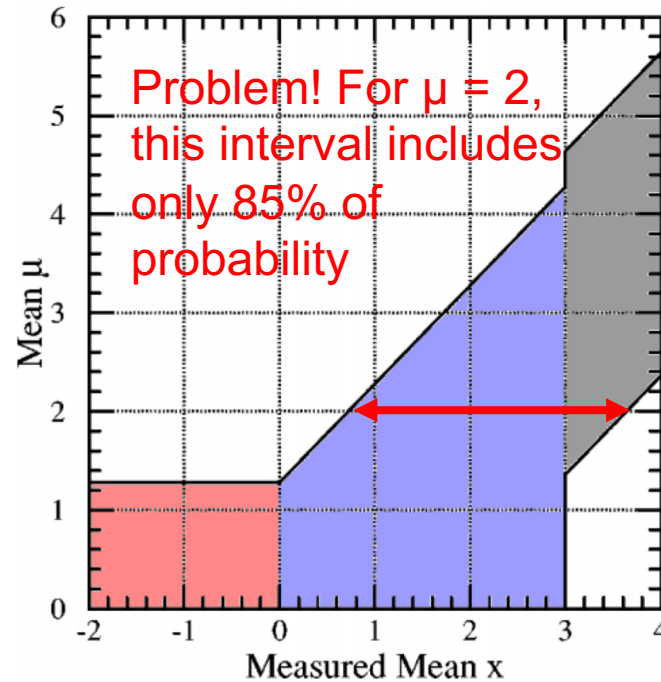
- $0\nu\beta\beta$ experiments measure a half-life
- Before you run an experiment, you can calculate its expected half-life sensitivity:



- $\langle UL(B(t)) \rangle$ is the average upper limit an ensemble of identical experiments would place in the absence of a signal given B background counts. If $B(t) = bt$, where b is a constant, then:
 - If $bt \gg 1$, $\langle UL(B(t)) \rangle = \sqrt{bt}$ (background limited)
 - If $bt \ll 1$, $\langle UL(B(t)) \rangle = \text{constant}$ (background free)
 - In between, you have to use Feldman-Cousins statistics

Feldman-Cousins Statistics

- In classical frequentist statistics, you have to decide **ahead of making a measurement** whether you'll be setting confidence intervals or an upper limit.
- If you “flip-flop,” you will end up with under-coverage.



- Flip-flopping:
If the experiment measures...
- $x < 0$, use a 3σ UL
 - $x < 3\sigma$, use a 3σ UL
 - $x > 3\sigma$, use a CL

- Feldman-Cousins gives a recipe for building confidence intervals that don't lead to under-coverage, and transition smoothly between UL and CL

How to Use Feldman-Cousins

- Look-up table for 90% CL from the Feldman-Cousins paper
- ROOT and python packages for calculating F.C. intervals are also available

$n_0 \backslash b$	0.0	0.5	1.0	1.5	2.0	2.5	3.0	3.5	4.0	5.0
0	0.00, 2.44	0.00, 1.94	0.00, 1.61	0.00, 1.33	0.00, 1.26	0.00, 1.18	0.00, 1.08	0.00, 1.06	0.00, 1.01	0.00, 0.98
1	0.11, 4.36	0.00, 3.86	0.00, 3.36	0.00, 2.91	0.00, 2.53	0.00, 2.19	0.00, 1.88	0.00, 1.59	0.00, 1.39	0.00, 1.22
2	0.53, 5.91	0.03, 5.41	0.00, 4.91	0.00, 4.41	0.00, 3.91	0.00, 3.45	0.00, 3.04	0.00, 2.67	0.00, 2.33	0.00, 1.73
3	1.10, 7.42	0.60, 6.92	0.10, 6.42	0.00, 5.92	0.00, 5.42	0.00, 4.92	0.00, 4.42	0.00, 3.95	0.00, 3.53	0.00, 2.78
4	1.47, 8.60	1.17, 8.10	0.74, 7.60	0.24, 7.10	0.00, 6.60	0.00, 6.10	0.00, 5.60	0.00, 5.10	0.00, 4.60	0.00, 3.60
5	1.84, 9.99	1.53, 9.49	1.25, 8.99	0.93, 8.49	0.43, 7.99	0.00, 7.49	0.00, 6.99	0.00, 6.49	0.00, 5.99	0.00, 4.99
6	2.21, 11.47	1.90, 10.97	1.61, 10.47	1.33, 9.97	1.08, 9.47	0.65, 8.97	0.15, 8.47	0.00, 7.97	0.00, 7.47	0.00, 6.47
7	3.56, 12.53	3.06, 12.03	2.56, 11.53	2.09, 11.03	1.59, 10.53	1.18, 10.03	0.89, 9.53	0.39, 9.03	0.00, 8.53	0.00, 7.53
8	3.96, 13.99	3.46, 13.49	2.96, 12.99	2.51, 12.49	2.14, 11.99	1.81, 11.49	1.51, 10.99	1.06, 10.49	0.66, 9.99	0.00, 8.99
9	4.36, 15.30	3.86, 14.80	3.36, 14.30	2.91, 13.80	2.53, 13.30	2.19, 12.80	1.88, 12.30	1.59, 11.80	1.33, 11.30	0.43, 10.30

The original paper is a classic: [10.1103/PhysRevD.57.3873](https://arxiv.org/abs/10.1103/PhysRevD.57.3873), and some nice slides on the subject can be found at https://www.pas.rochester.edu/~sybenzvi/courses/phy403/2015s/p403_19_intervals.pdf

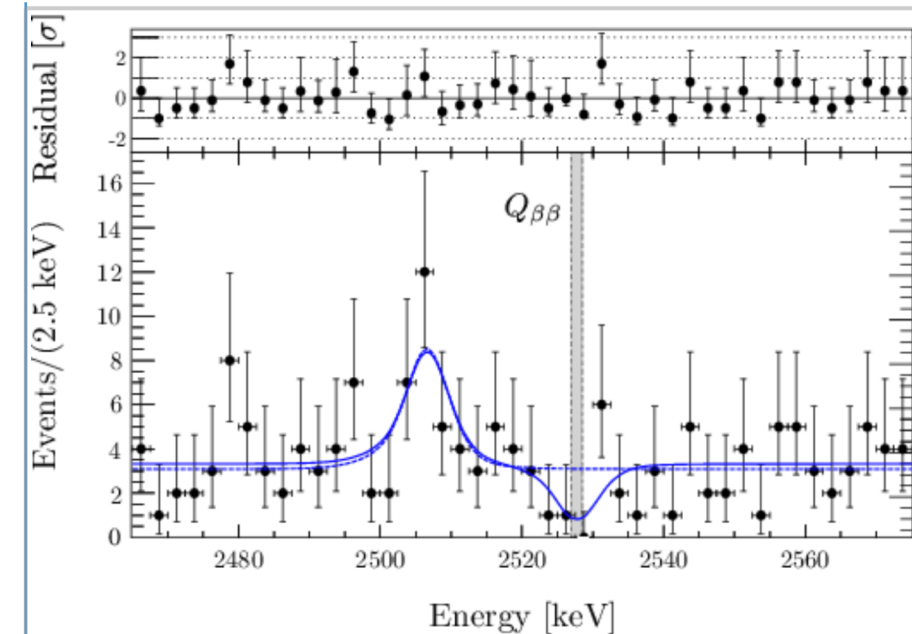
Measuring and Setting Limits on $T_{1/2}$

- After you run a $0\nu\beta\beta$ search...
 - You either see an excess at the Q value, and fit a peak with some rate to it
 - Or you don't see an excess. In that case, you set a lower limit on half-life:

$$T_{1/2}^{0\nu} > \ln(2) \frac{N_a T \epsilon}{S}$$

- S is the upper limit on the signal counts based on the observed data. Again, use F.C. to calculate if it's appropriate.
- Experiments that don't see anything report sensitivity and a limit. If they get lucky and have a downward fluctuation, limit will exceed sensitivity.

CUORE 2018 Result



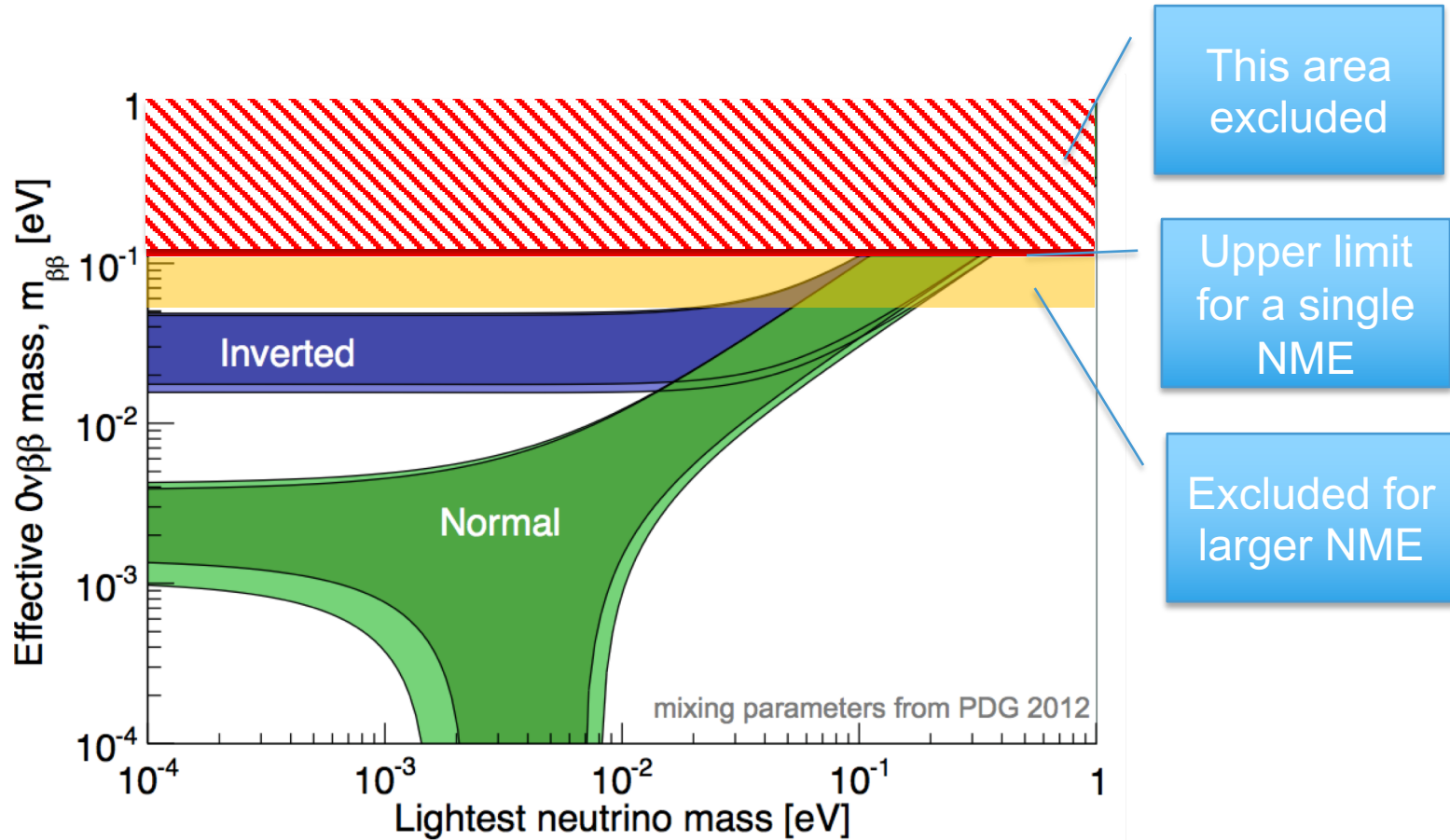
Sensitivity: 7×10^{24} yrs

Limit: $T_{1/2} > 1.3 \times 10^{25}$ yrs

10.1103/PhysRevLett.120.132501

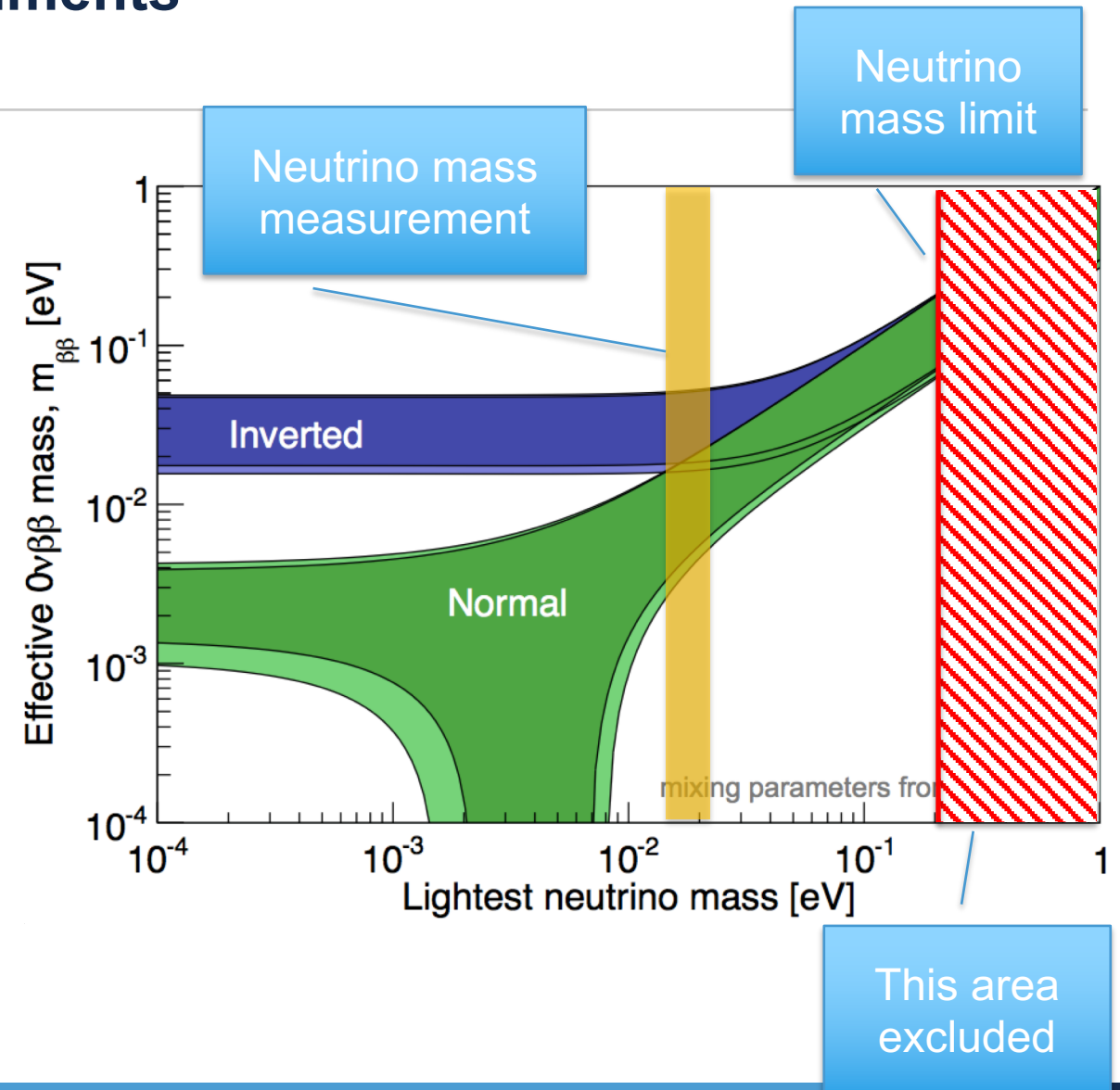
Translating Half-Life to $m_{\beta\beta}$

- Need to use a particular model, the phase space factor and a nuclear matrix element to turn half-life into $m_{\beta\beta}$
- Results are generally reported for the full set of NMEs, so the upper limit in $m_{\beta\beta}$ has a range



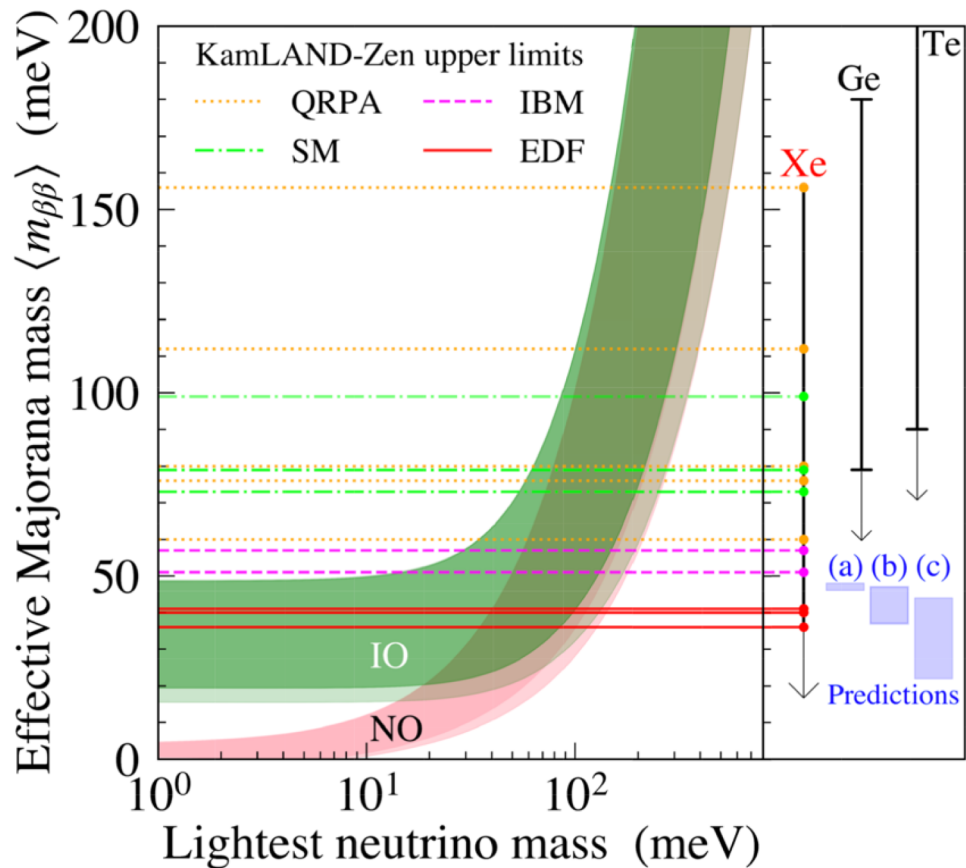
Information from Other Neutrino Experiments

- Light-colored edges are 3σ uncertainty on neutrino mixing and mass splittings
- Measuring hierarchy would tell us which branch we need to look in
- Mass measurement would tell us which vertical band to look in



The $0\nu\beta\beta$ Search Landscape

Current Best Limits on $0\nu\beta\beta$



arXiv: 2203.02139

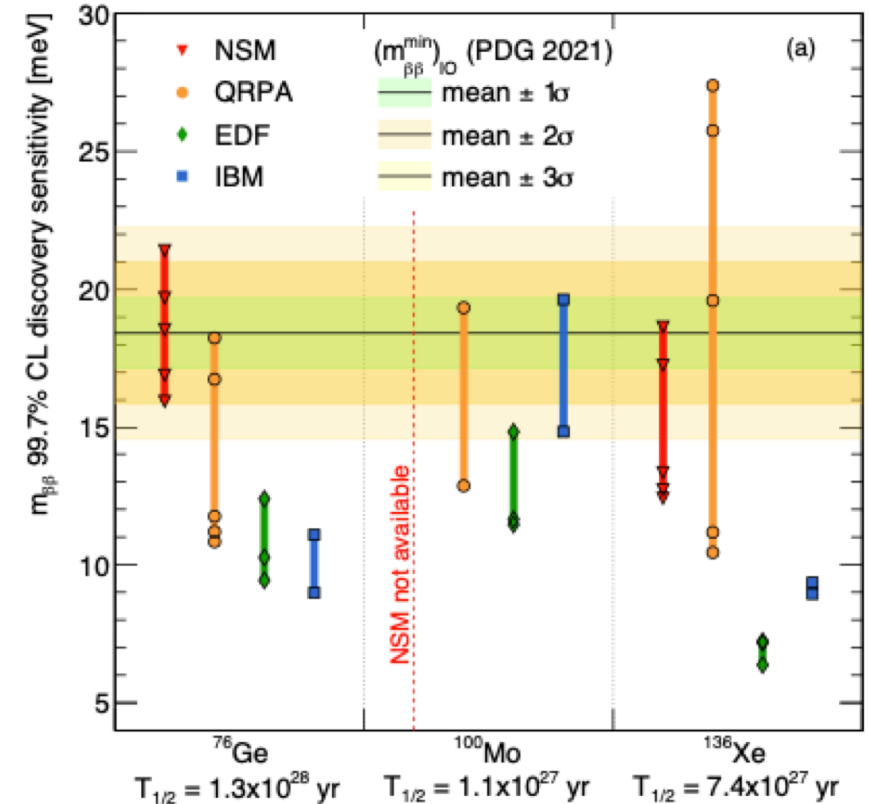
Experiment	Isotope	Exposure [kg yr]	$T_{1/2}^{0\nu}$ [10^{25} yr]	$m_{\beta\beta}$ [meV]
Gerda	^{76}Ge	127.2	18	79-180
Majorana	^{76}Ge	26	8.3	113-269
KamLAND-Zen	^{136}Xe	970	23	36-156
EXO-200	^{136}Xe	234.1	3.5	93-286
CUORE	^{130}Te	1038.4	2.2	90-305

NSAC recommendation: quote a range of $m_{\beta\beta}$ using the largest and smallest available NME from the 4 main calculation methods; $g_A=1.27$; no contribution from the contact term

The Ton-Scale Generation

- The next-generation of experiments seeks to cover the IO region in discovery mode
- That will take $O(1 \text{ ton})$ of isotope
- 3 candidate experiments with US participation, in addition to other efforts: LEGEND, nEXO, and CUPID
- All 3 were evaluated by the DOE in Summer 2021. DOE-NP is seeking international support to pursue all 3 experiments.

Discovery Sensitivity for the “Big 3”



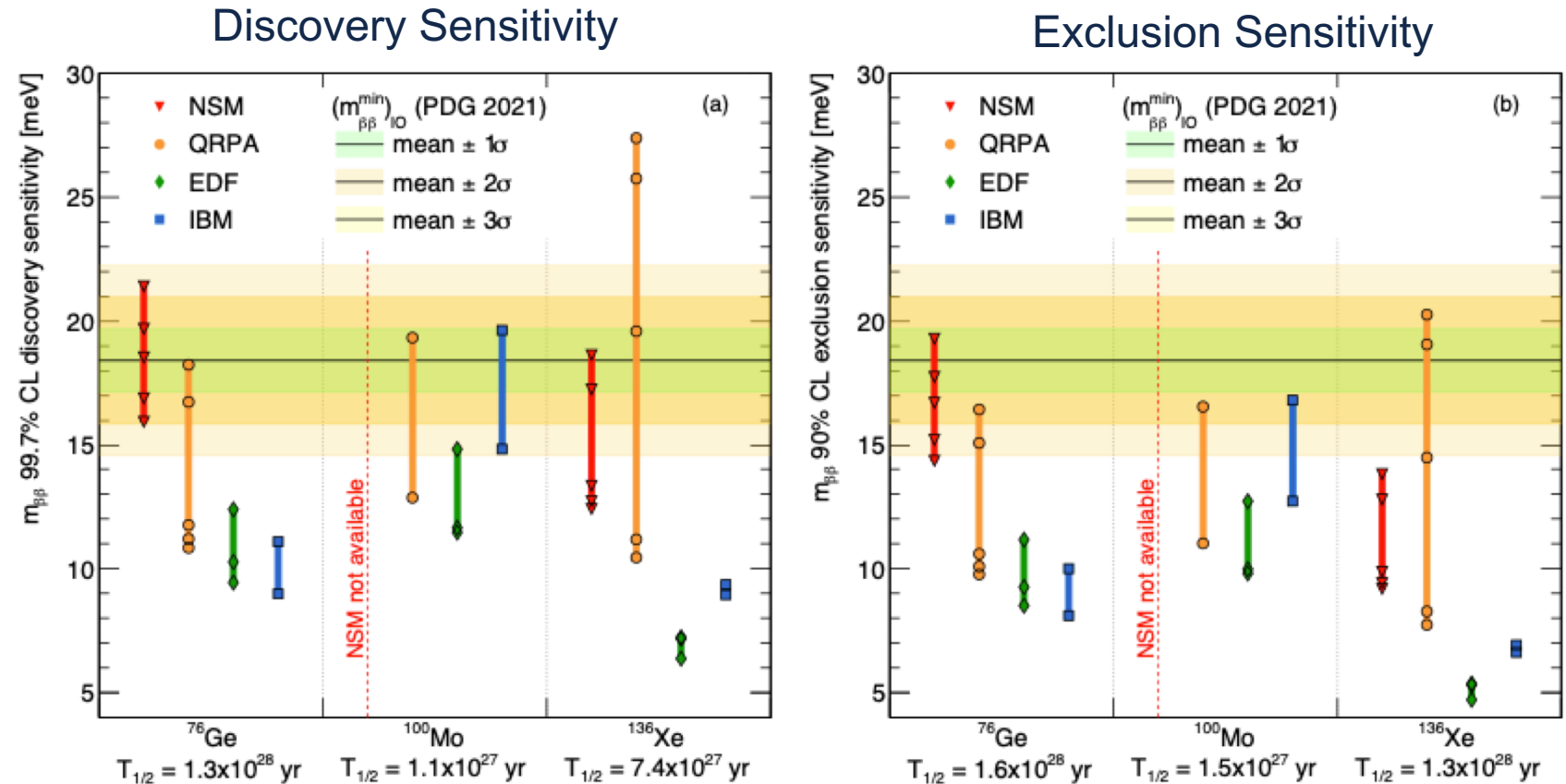
10.1103/PhysRevC.104.L042501

Discovery and Sensitivity

- All 3 experiments cover the IO for some matrix elements, and miss for others
- Larger background = more difference between discovery and exclusion

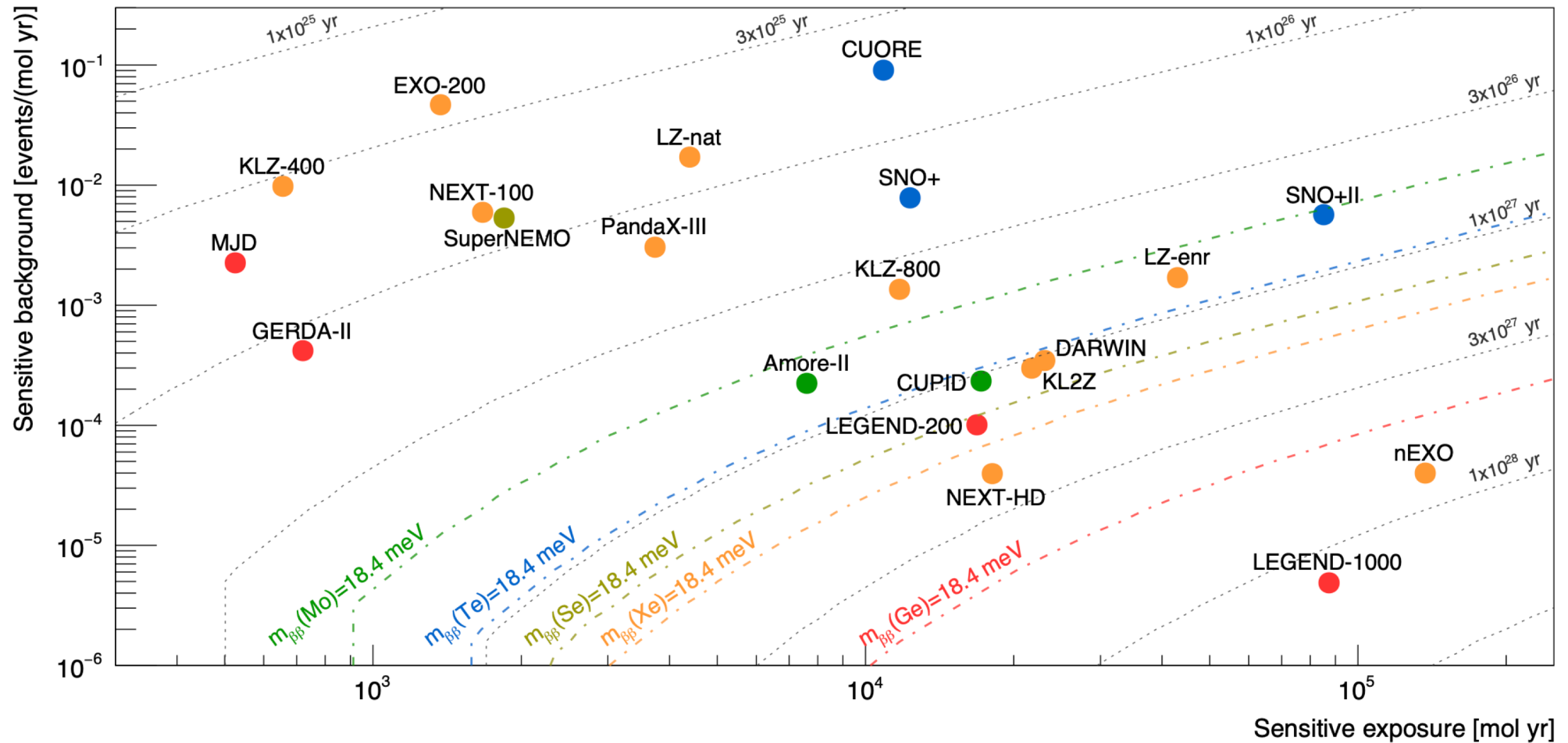
TABLE I. Nuclear matrix elements M for $0\nu\beta\beta$ decay mediated by light neutrinos, calculated with the NSM, QRPA, EDF, and IBM methods. The ranges correspond to the minimum and maximum values obtained with the same many-body method.

	Ref.	^{76}Ge	^{100}Mo	^{136}Xe
NSM	[35]	2.89, 3.07	–	2.28, 2.45
	[36]	3.37, 3.57	–	1.63, 1.76
	[37]	2.66	–	2.39
	All	2.66-3.57	–	1.63-2.45
QRPA	[38]	5.09	–	1.55
	[39]	5.26	3.90	2.91
	[40]	4.85	5.87	2.72
	[41]	3.12, 3.40	–	1.11, 1.18
	[42]	–	–	3.38
All	3.12-5.26	3.90-5.87	1.11-3.38	
EDF	[43]	4.60	5.08	4.20
	[44]	5.55	6.59	4.77
	[45]	6.04	6.48	4.24
	All	4.60-6.04	5.08-6.59	4.20-4.77
IBM	[46] ^a	5.14	3.84	3.25
	[47]	6.34	5.08	3.40
	All	5.14-6.34	3.84-5.08	3.25-3.40



From Agostini et al., PRC 104, L042501 (2021)

Recent and Proposed Experiments



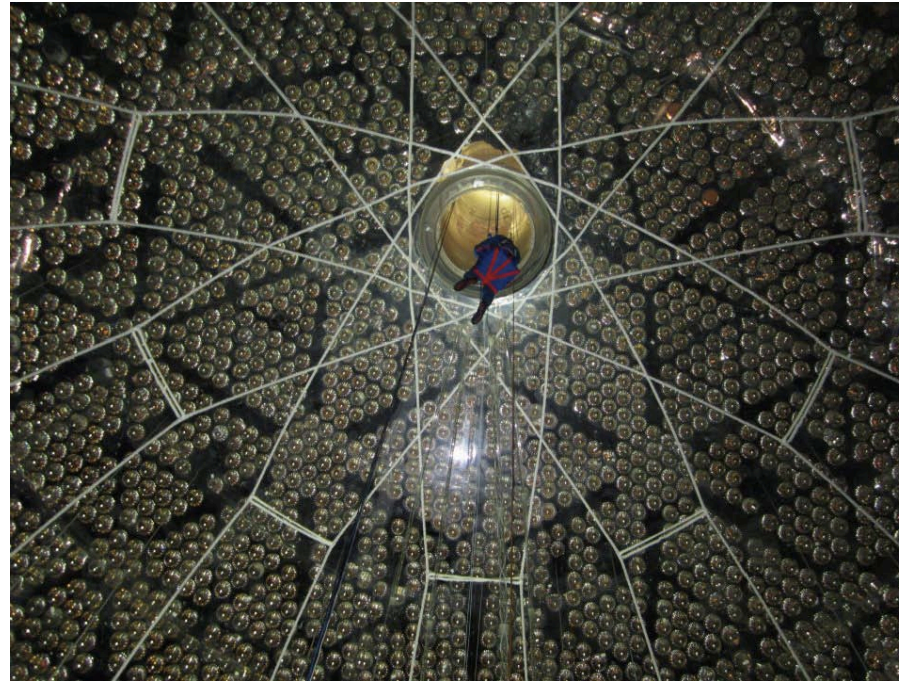
The Experiments

Liquid Scintillators

Liquid Scintillator Detector Concept

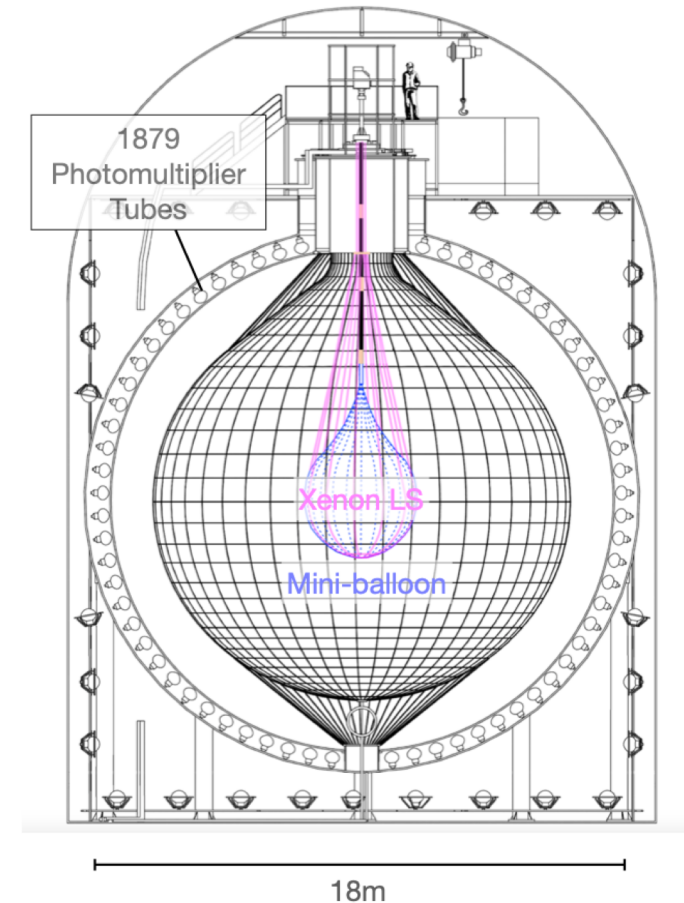
Measure isotropic scintillation light to extract energy

- Self-shielding, fiducialization
- Interior materials can be made extremely pure
- Some event topology and particle ID, with additional future improvements expected
- Measurement with and without isotope is possible
- Other strengths: flexible and scalable design
- Weakness: energy resolution and spatial resolution; large “target” for cosmogenics



KamLAND-Zen

- Adaptation of existing reactor neutrino experiment
- ^{136}Xe concentrated in inner balloon, holds LS doped with 3.13% enriched Xenon by weight
- Relatively shallow site, spallation backgrounds dominate
- KamLAND-Zen 400: 2011-2014
 - Phase I: 320 kg 90% enriched ^{136}Xe
 - Phase II: 380 kg
 - $T_{1/2} > 1.07 \times 10^{26}$ yr, $m_{\beta\beta} < 61\text{--}165$ meV
- KamLAND-Zen 800:
 - Data-taking began in January 2019
 - Scintillator purification campaign
 - Larger, cleaner inner balloon
 - 750 kg enriched ^{136}Xe
- Currently has the largest single-experiment half-life limit: $T_{1/2} > 2.3 \times 10^{26}$ yrs



Backgrounds in KamLAND-Zen

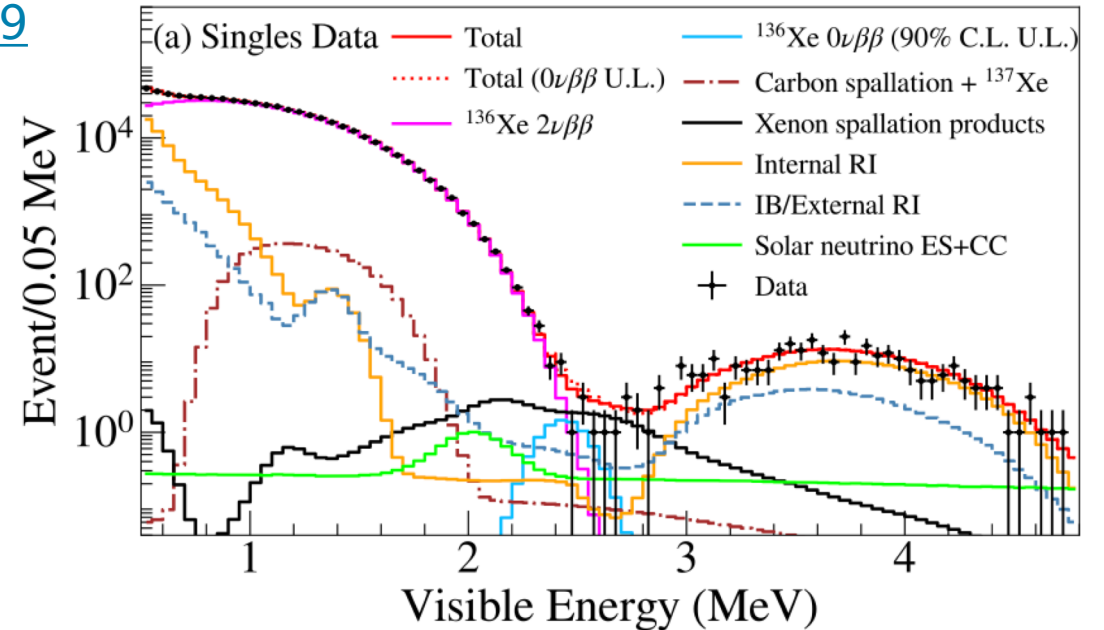
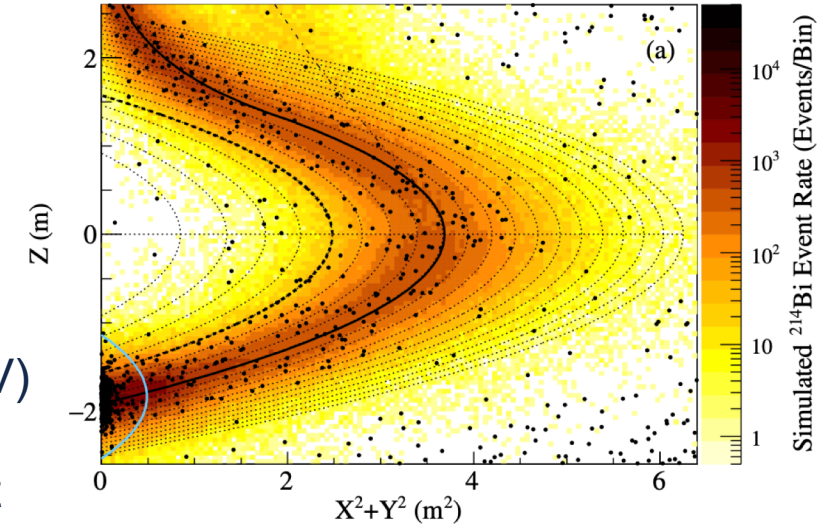
TABLE I: Summary of the estimated and best-fit background contributions for the frequentist and Bayesian analyses in the energy region $2.35 < E < 2.70$ MeV within the 1.57-m-radius spherical volume. In total, 24 events were observed.

Background	Estimated	Best-fit	
		Frequentist	Bayesian
$^{136}\text{Xe } 2\nu\beta\beta$	-	11.98	11.95
Residual radioactivity in Xe-LS			
^{238}U series	0.14 ± 0.04	0.14	0.09
^{232}Th series	-	0.84	0.87
External (Radioactivity in IB)			
^{238}U series	-	3.05	3.46
^{232}Th series	-	0.01	0.01
Neutrino interactions			
^8B solar νe^- ES	1.65 ± 0.04	1.65	1.65
Spallation products			
Long-lived	$7.75 \pm 0.57^\dagger$	12.52	11.80
^{10}C	0.00 ± 0.05	0.00	0.00
^6He	0.20 ± 0.13	0.22	0.21
^{137}Xe	0.33 ± 0.28	0.34	0.34

[†] Estimation based on the spallation MC study. This event rate constraint is not applied to the spectrum fit.

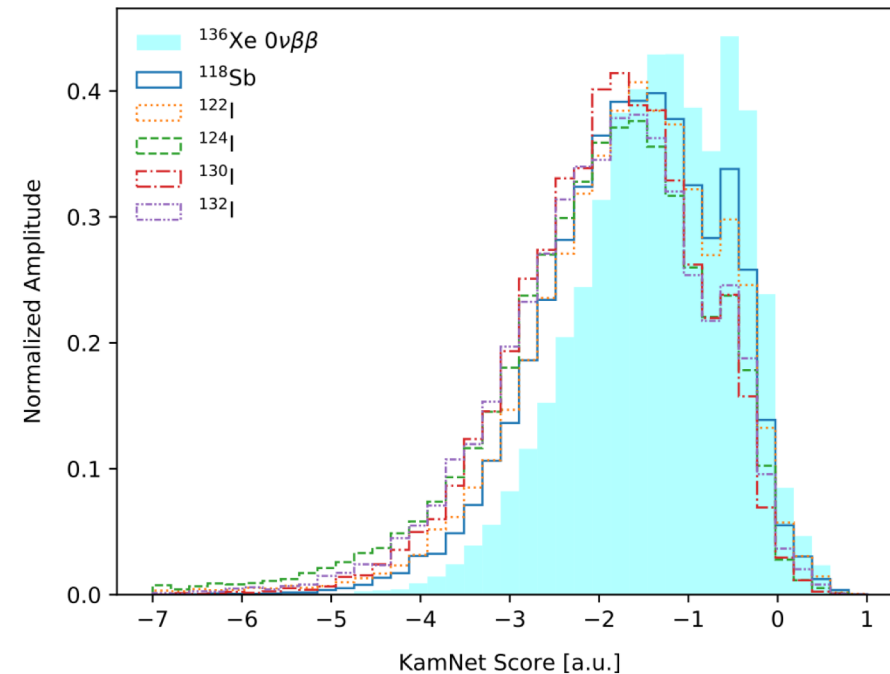
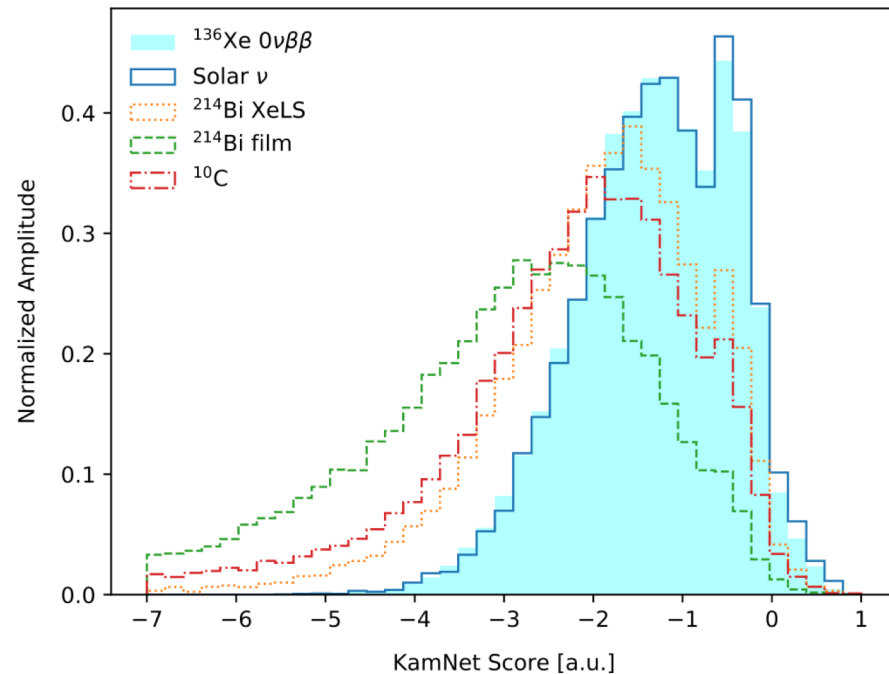
Energy resolution: $6.7\%/\sqrt{E}(\text{MeV})$
 Vertex resolution $13.7 \text{ cm}/\sqrt{E}(\text{MeV})$
 Poor energy resolution means that $2\nu\beta\beta$ is the largest background

[arXiv:2203.02139](https://arxiv.org/abs/2203.02139)



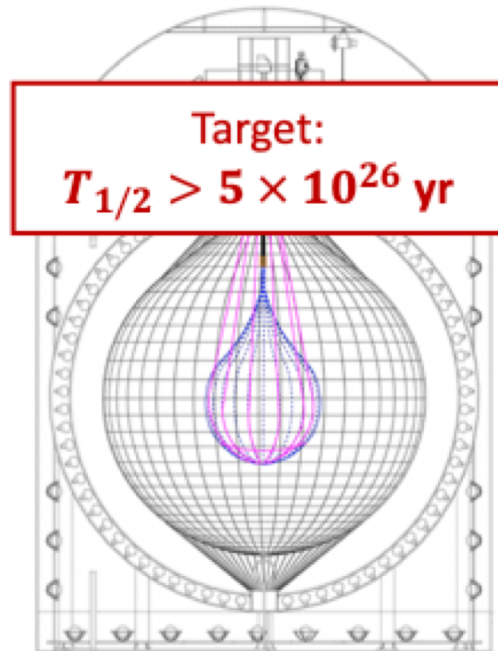
Machine Learning-Based Analysis: KamNET

- Machine learning model using LSTM + spherical CNN can be used to reject balloon and spallation backgrounds
- Complementary to other spallation-tagging methods
- [arXiv:2203.01870](https://arxiv.org/abs/2203.01870)



KamLAND2-Zen

KamLAND-Zen 800



KamLAND-Zen 800

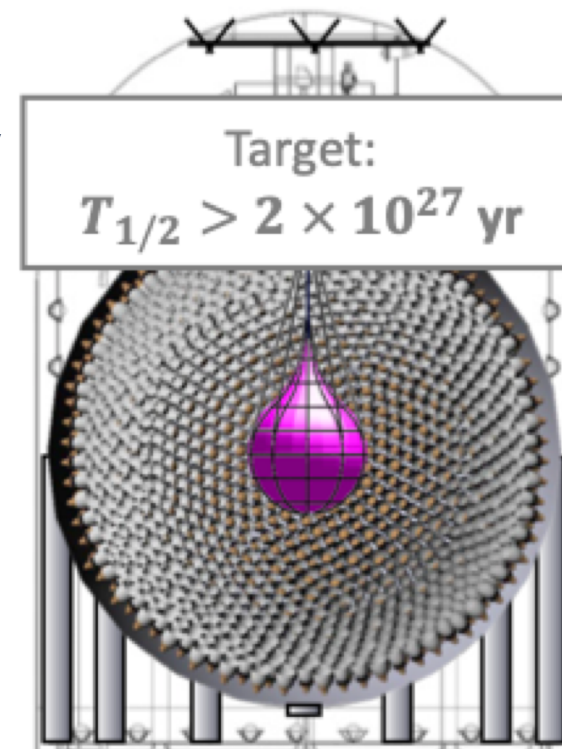
Mini-balloon Radius = 1.90 m

Xenon mass = 745 kg

90%
exclusion
sensitivity



KamLAND2-Zen



Primary goal: improve energy resolution

Secondary goal: record longer-buffer data to improve cosmogenic rejection

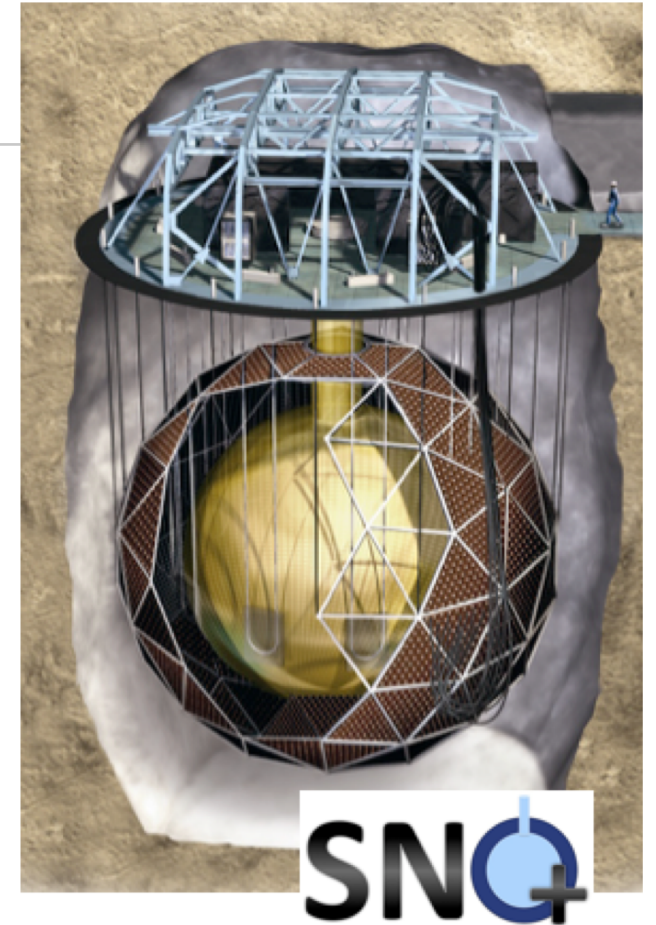
- 1 ton of ^{136}Xe
- New, brighter LAB-based scintillator
- Winston cones and HQE PMTs
- Scintillating balloon film
- Improved front-end electronics

SNO+

- Adaptation of existing solar neutrino experiment
- Switch from heavy water to liquid scintillator
- $^{\text{nat}}\text{Te}$ loaded throughout
- Deeper site, solar ν backgrounds expected to dominate

Status

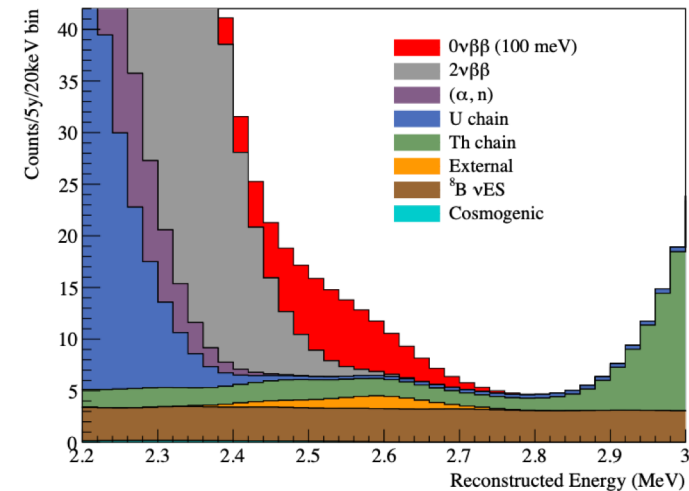
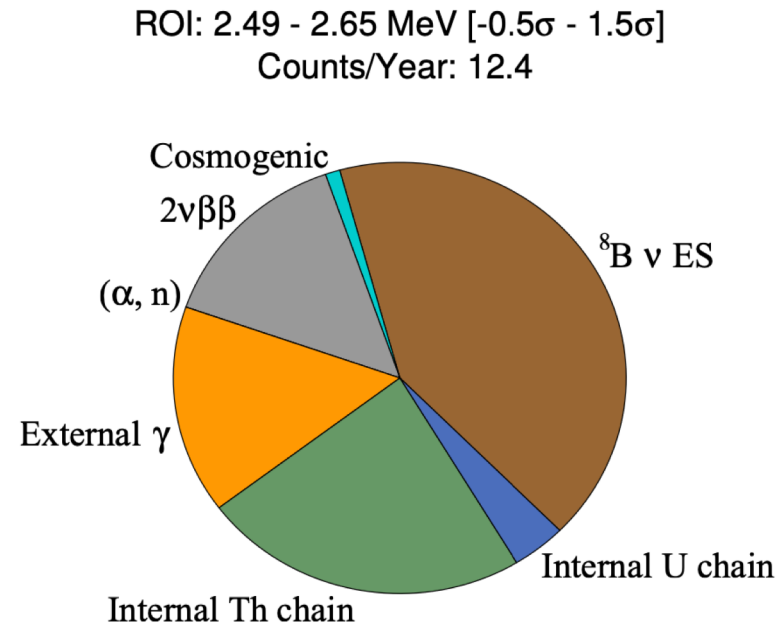
- Ultra-pure water Cherenkov phase completed
 - Solar neutrino and BSM physics results released
- Unloaded liquid scintillator phase completed, scintillator characterized
- Tellurium loading underway @ 0.3%
 - Expected loading: 800 kg of ^{130}Te
- Expected 5 year sensitivity:
 $T_{1/2} > 9 \times 10^{25}$ yrs , $m_{\beta\beta} < 55 - 133$ meV



Initial loading: 0.5%
natural Te by weight

Backgrounds in SNO+

- Expected backgrounds before running
- Since then, U and Th backgrounds have been measured to be smaller than expected!
- Solar neutrino scattering will dominate



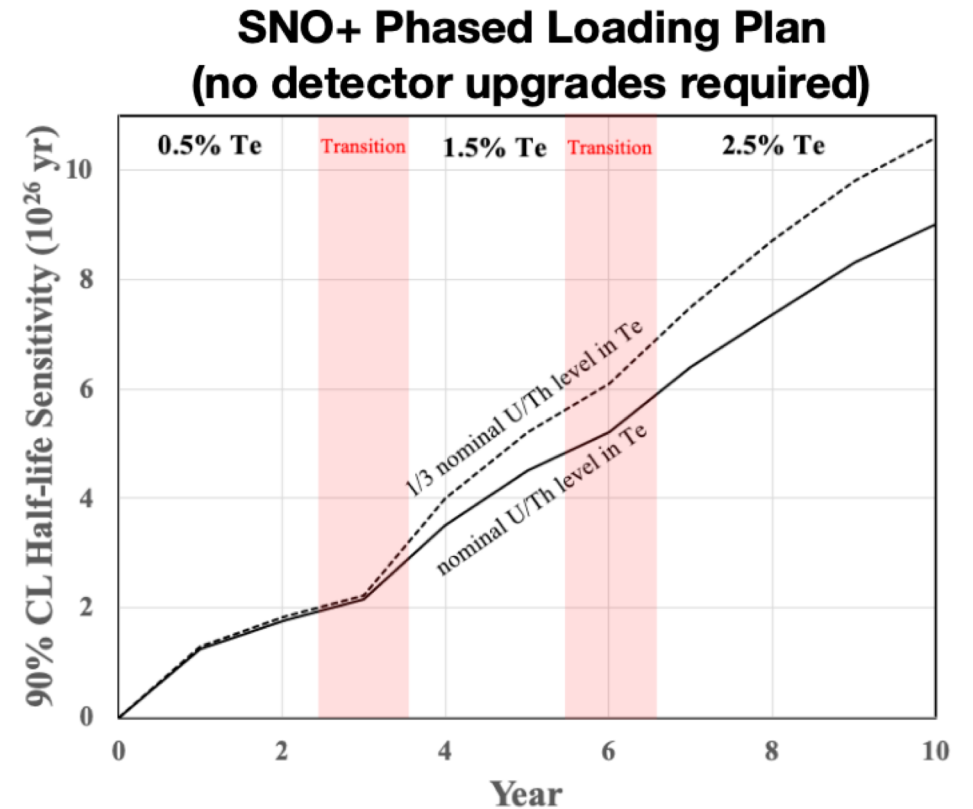
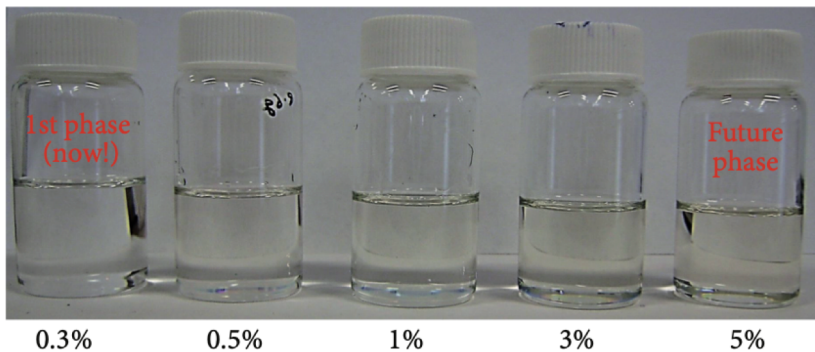
[arXiv:1809.05986](https://arxiv.org/abs/1809.05986)

SNO+ Future Improvements

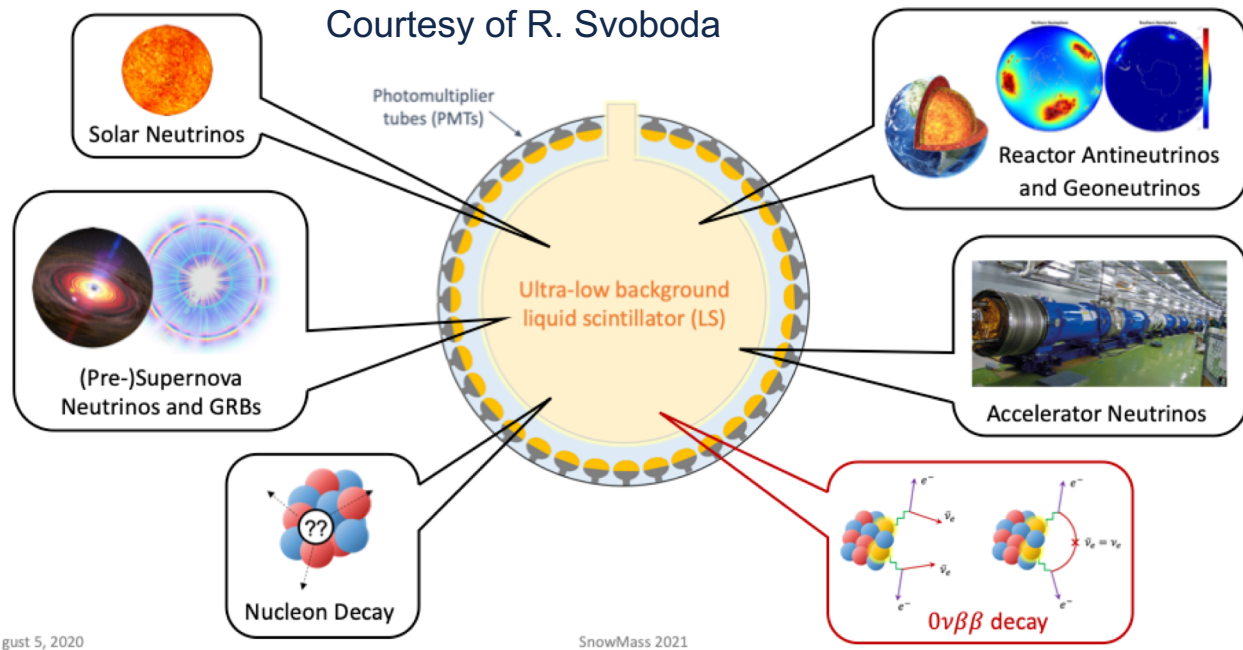
Planning to move to 3% ^{nat}Te loading for future data-taking

No hardware changes needed: just add more Te

Extends exclusion sensitivity to $> \sim 10^{27}$ yrs

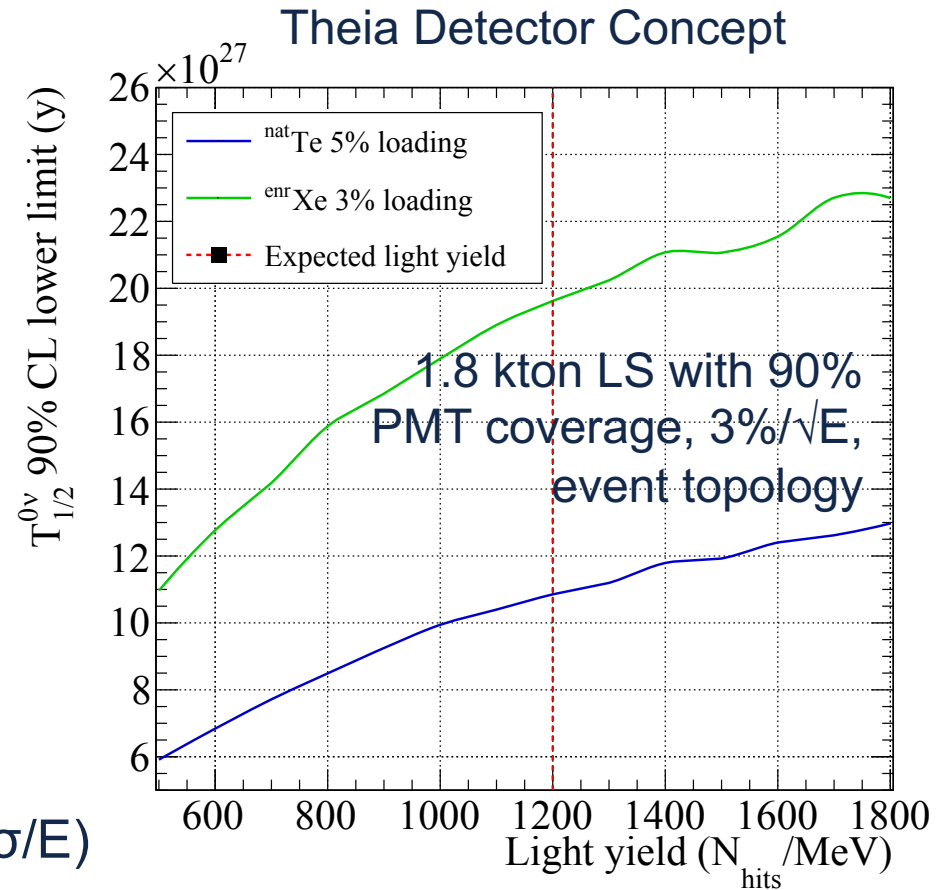


Multi-purpose Liquid Scintillator Experiments



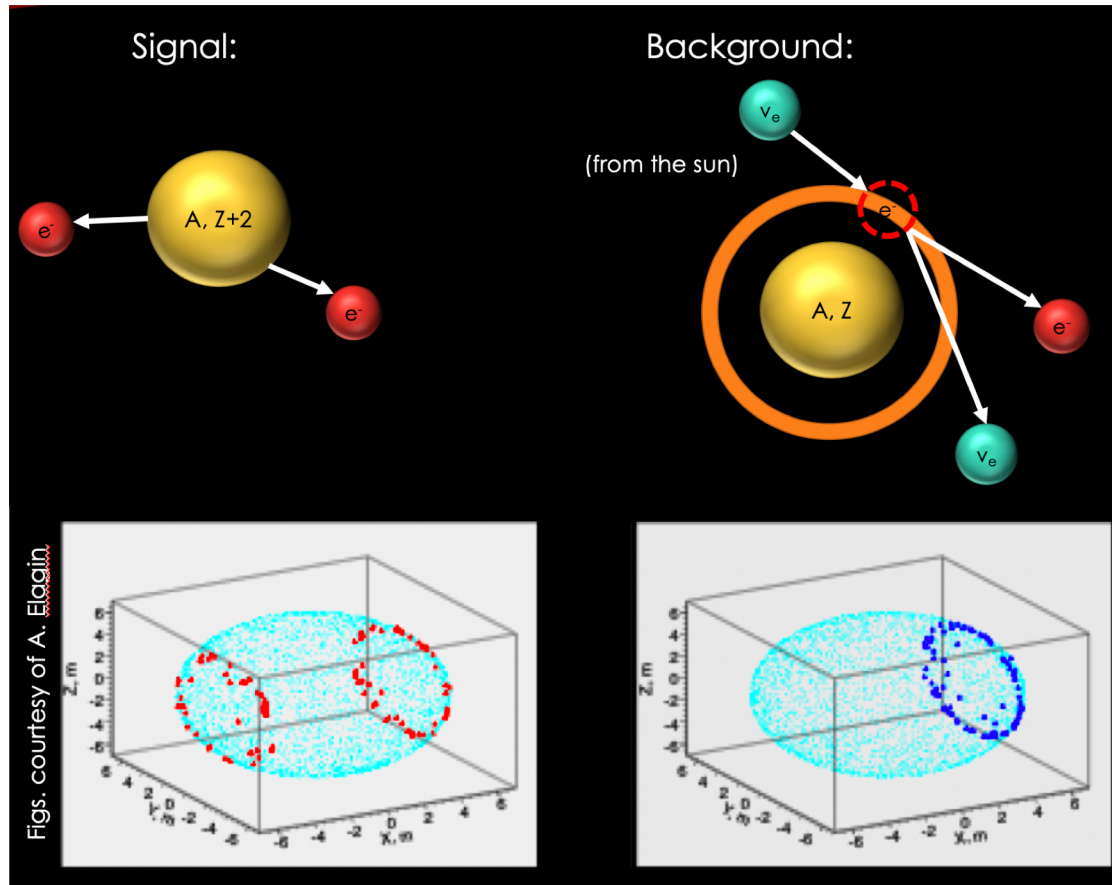
JUNO $0\nu\beta\beta$ Search Proposal:

- 50 tons of ^{136}Xe , expected energy resolution better than 2% (σ/E)
- Exclusion sensitivity: 1.8×10^{28} yr, 5-12 meV
- $0\nu\beta\beta$ upgrade starting in 2030s



Eur. Phys. J. C (2020) 80:416

Solar Neutrino Scattering in Liquid Scintillator Detectors

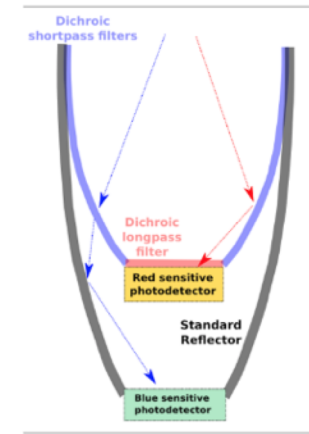


- How to reduce this “irreducible background”?
- If you can measure Cherenkov **and** scintillation light (separately) you can tell them apart
- Timing- and wavelength-based separation have been demonstrated in benchtop test-stands

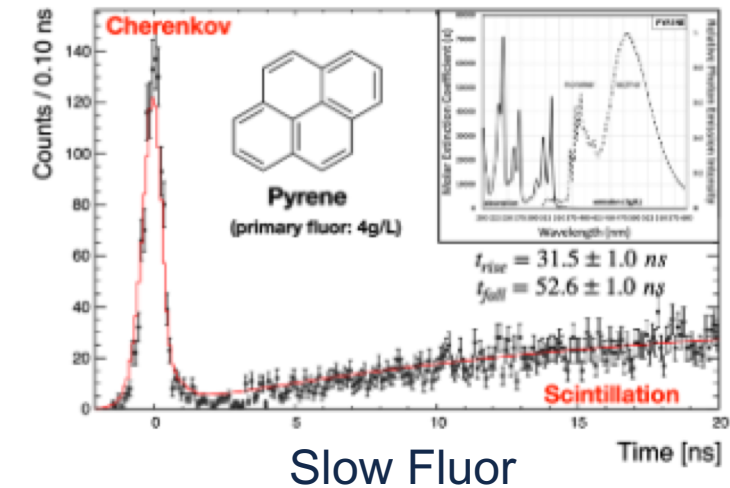
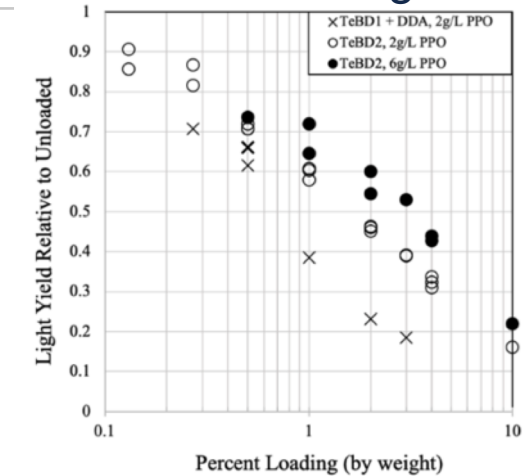
R&D Efforts for Liquid Scintillators

- Hybrid Cherenkov/scintillation detectors:
 - Reduce backgrounds by measuring 2 e- signature
 - Timing-based separation: slower fluors, faster photodetectors
 - Wavelength-based separation: dichroic filters
- New scintillator cocktails and isotopic loading techniques:
 - Water-based Liquid Scintillator: purification and stability, gadolinium loading, pulse shape discrimination
 - Tellurium loading: several % loading demonstrated, increased loading and purification R&D underway
 - Quantum dot-based isotope loading: production scaling, stability, and optical performance studies underway
- Advanced photon sensors and collectors:
 - LAPPDs: ongoing R&D on high-channel-count readout techniques, self-triggering and synchronization, streamlined fabrication
- Advanced simulation and analysis techniques

Dichroicon

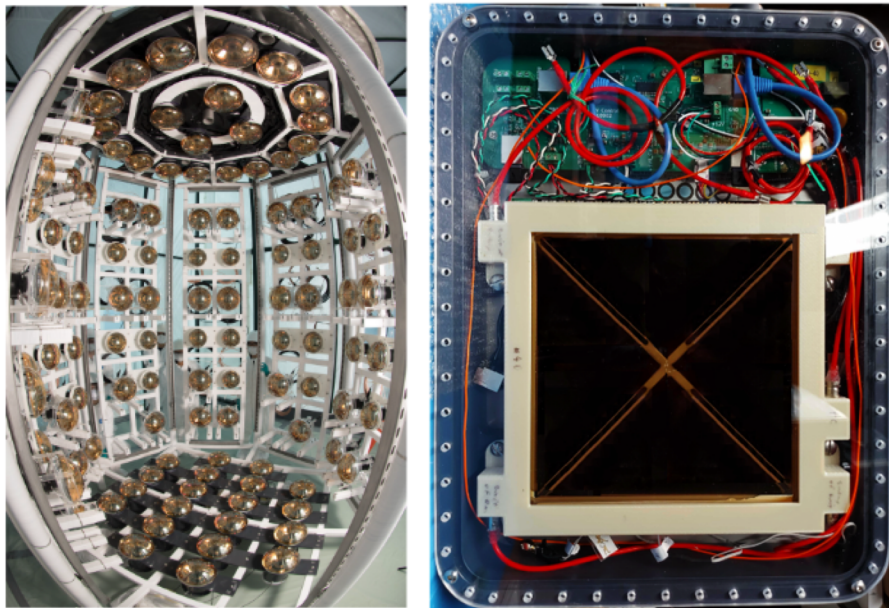


Te Loading

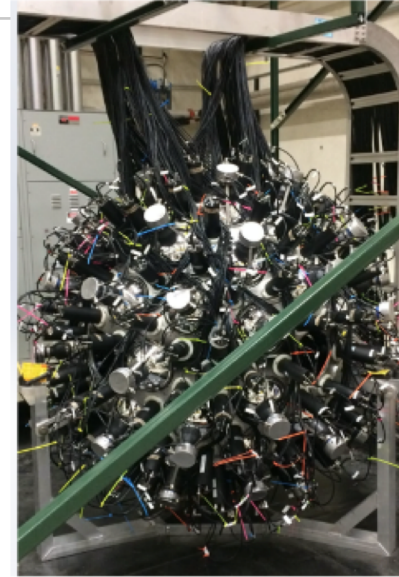


Mid-Scale Test Stands for Future Liquid Scintillator Detectors

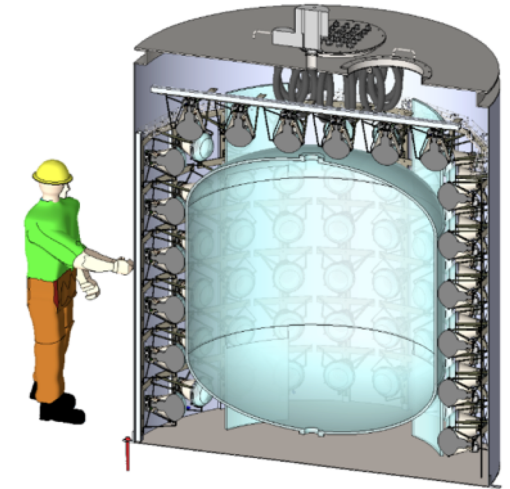
- ANNIE: first large-scale test of LAPPDs, planning for Gd-loaded WbLS



ANNIE detector and LAPPD module



NuDot: 1/2 ton test stand Eos: few-ton WbLS



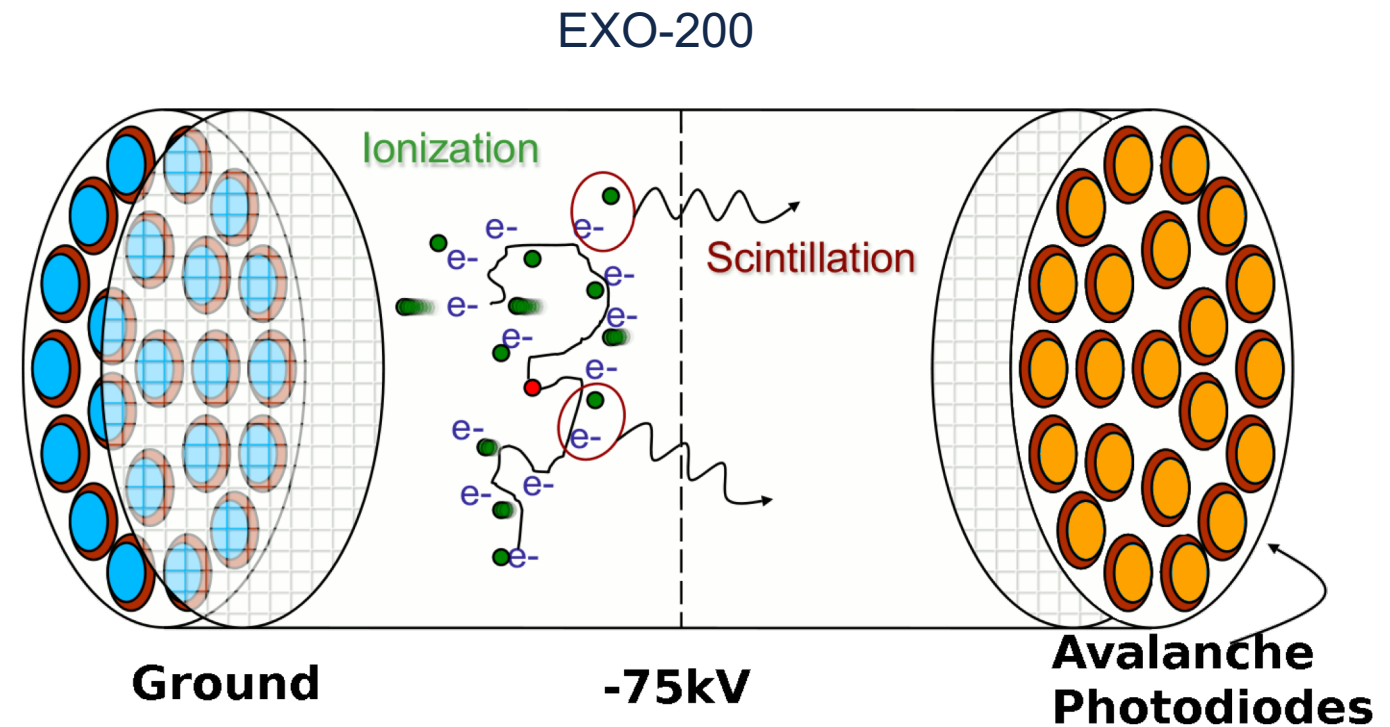
- NuDot: timing-based Cherenkov/scintillation separation and quantum dot loading
- Eos: Cherenkov/scintillation separation and WbLS, validation of microphysics simulations at low energy

The Experiments

TPCs: Liquid and Gas

Liquid Xenon TPC Concept

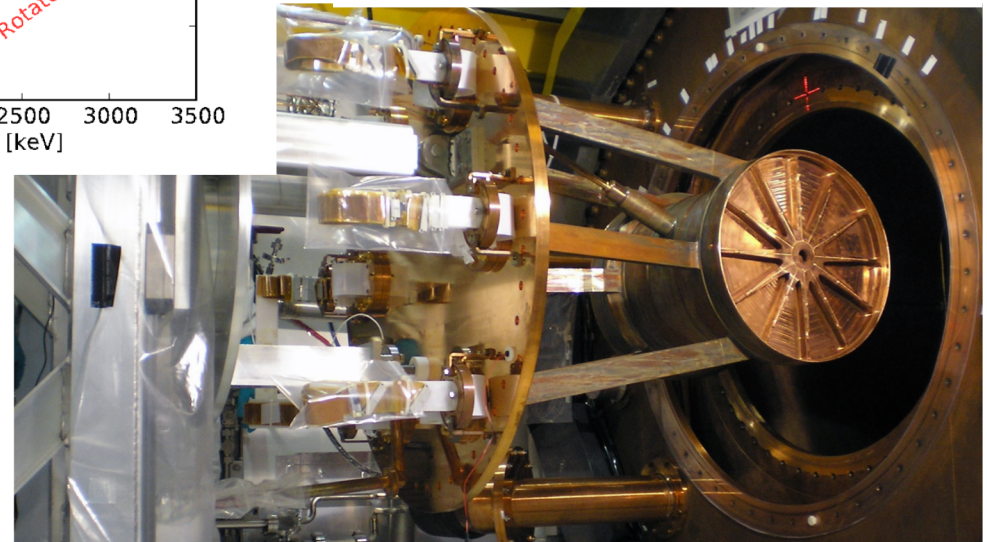
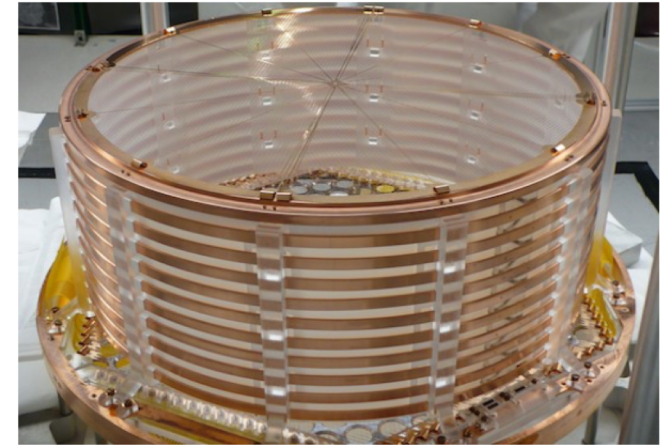
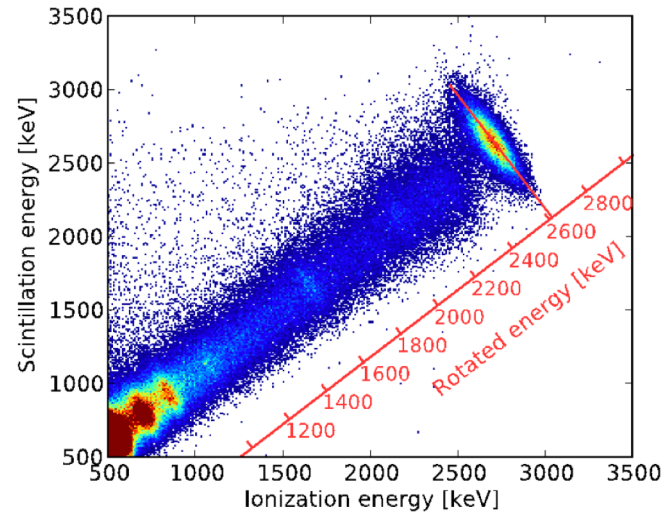
- Single-phase time projection chamber, made out of enriched Xenon
- Read out ionization and scintillation
- Dark matter experiments use dual-phase to amplify the ionization signal – not needed for higher-energy $\beta\beta$
- Take advantage of self-shielding, (non-binary) fiducialization, and event topology information to reduce backgrounds
- Better energy resolution and spatial resolution than LS, while still being monolithic
- Weaknesses: background peak overlapping $Q_{\beta\beta}$, cost/Xe availability



EXO-200

- 200 kg TPC with a center wire grid cathode
- Ran starting in 2011, then stopped in 2014 due to WIPP fire
- Upgrade in 2016: improved electronics led to better energy resolution
- Use anticorrelation between charge and scintillation to improve resolution: $\sigma/E = 1.15\%$

Scintillation vs. ionization, ^{228}Th calibration:

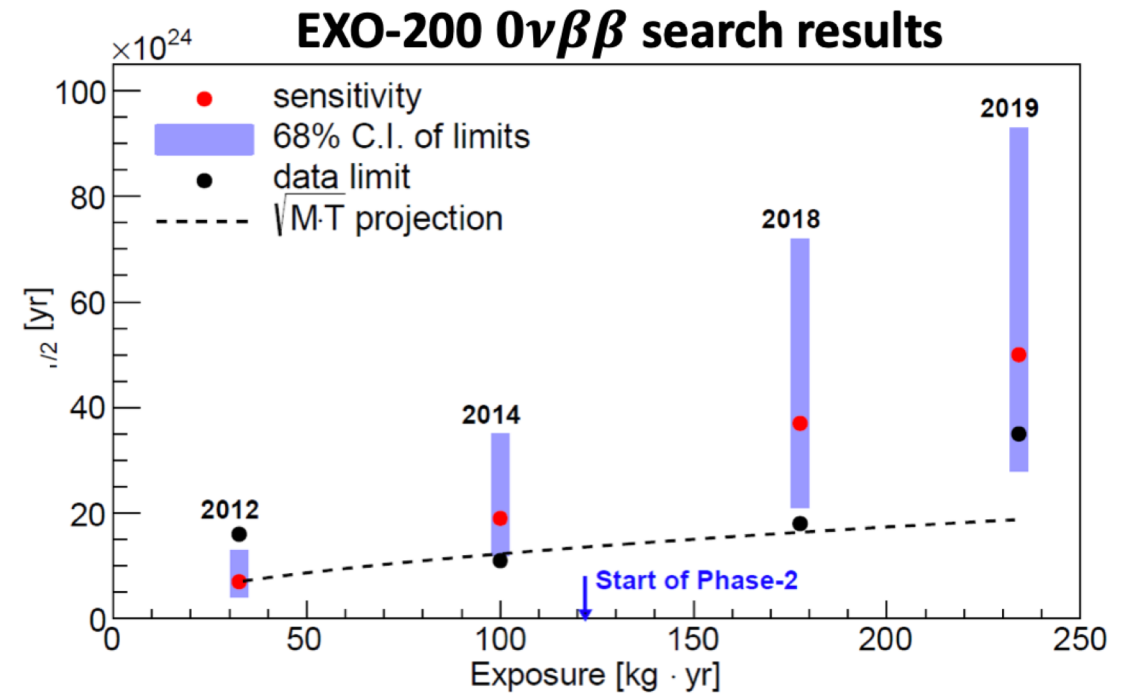
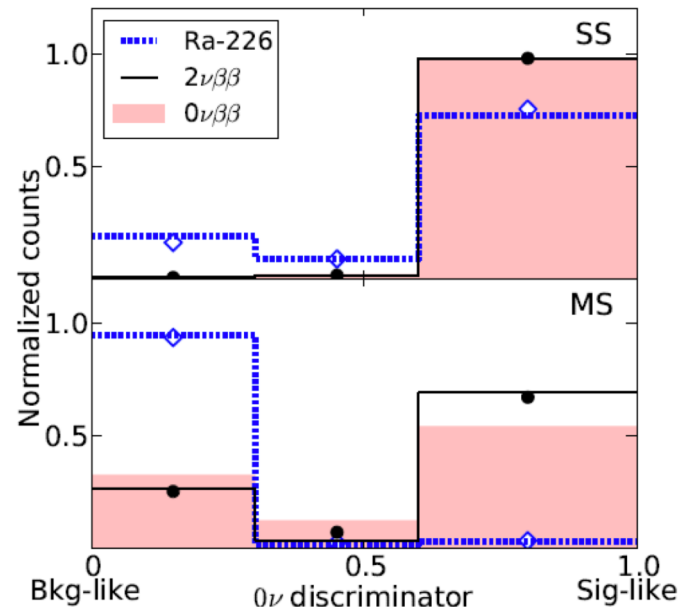


EXO-200 Results

- Detector upgrade and improved analysis techniques have led to linear sensitivity growth over time

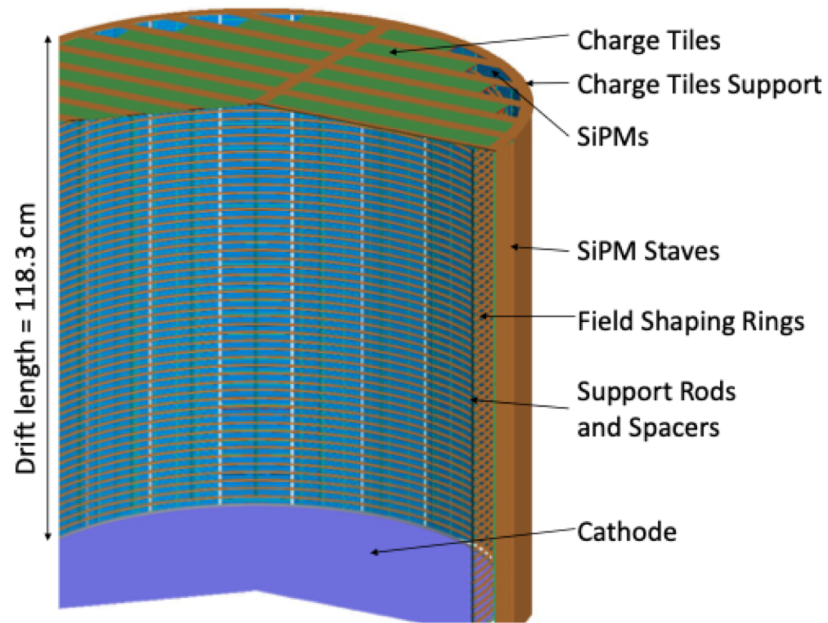
$$T_{1/2} > 3.5 \times 10^{25} \text{ yrs}$$

- Deep Neural Network, stand-off distance, and cluster size used to reduce backgrounds



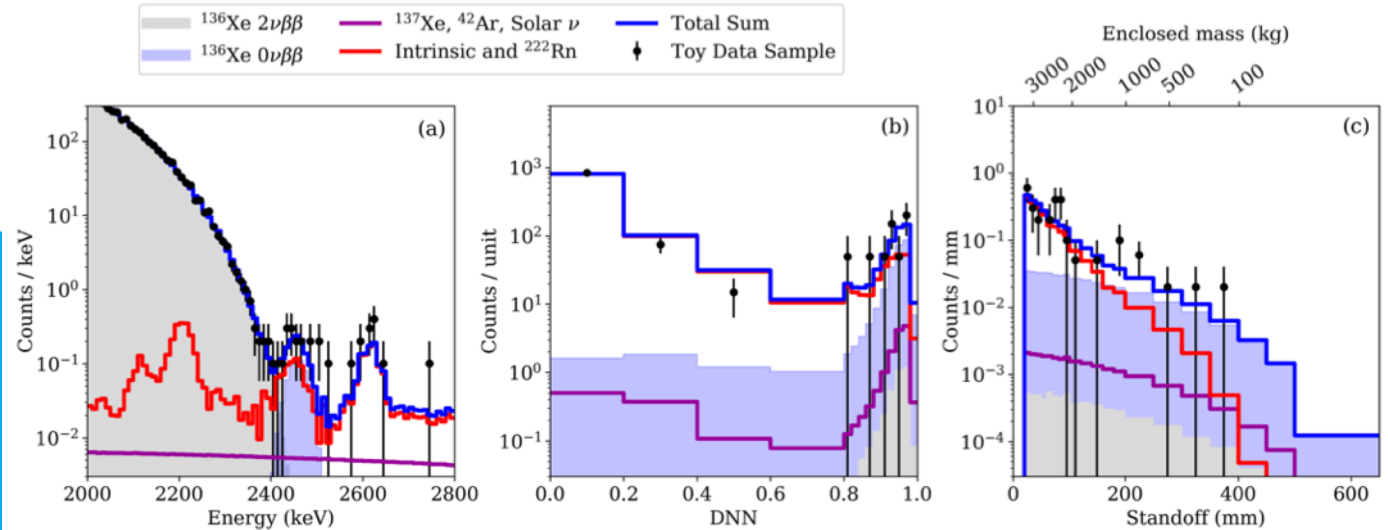
- 2012: *Phys.Rev.Lett.* 109 (2012) 032505
- 2014: *Nature* 510 (2014) 229-234
- 2018: *Phys. Rev. Lett.* 120, 072701 (2018)
- 2019: *arXiv* 1906.02723

Proposed Next-Generation: nEXO



- Large single-phase LXe TPC, building on EXO-200 experience
- Switches to charge tiles for ionization readout, SiPMs for light readout

- 5000 kg of ^{enr}Xe
- Enriched to 90% ^{136}Xe
- Energy res. (σ_E/E): 0.8%
- Discovery sensitivity: $T_{1/2} \sim 7.4 \times 10^{27}$ yrs
- $m_{\beta\beta}$ discovery sensitivity: 5-27 meV



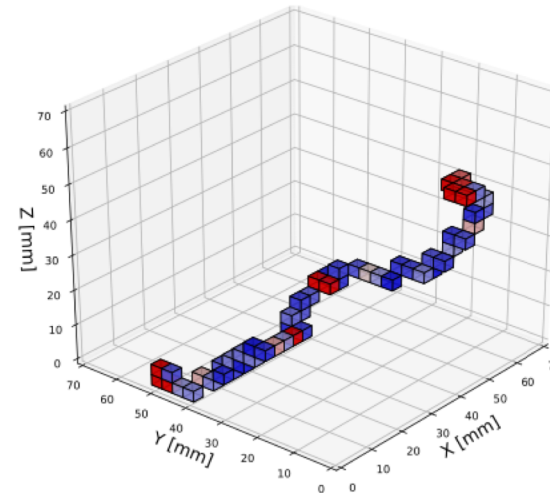
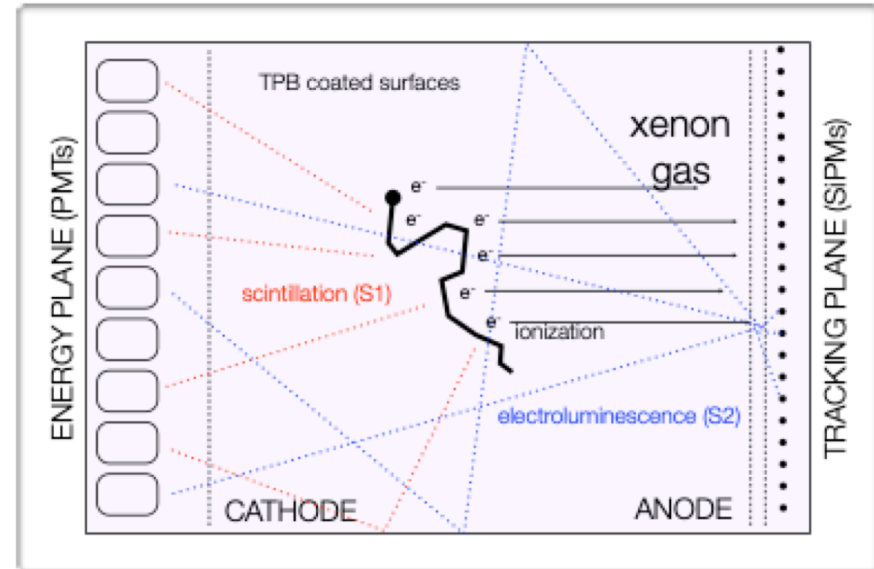
J. Phys. G: Nucl. Part. Phys. 49, 015104 (2022)

High-Pressure Gas TPC Concept

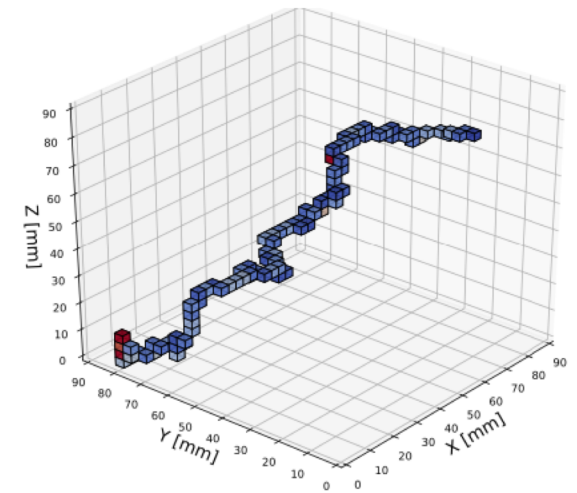
High-pressure gas Xenon time projection chamber:

- Energy resolution is intrinsically better in gas
- Event topology tracking information, fiducialization, and particle ID
- Can actually see β tracks!

Other TPC gases could use this technique: SeF_6 in R&D



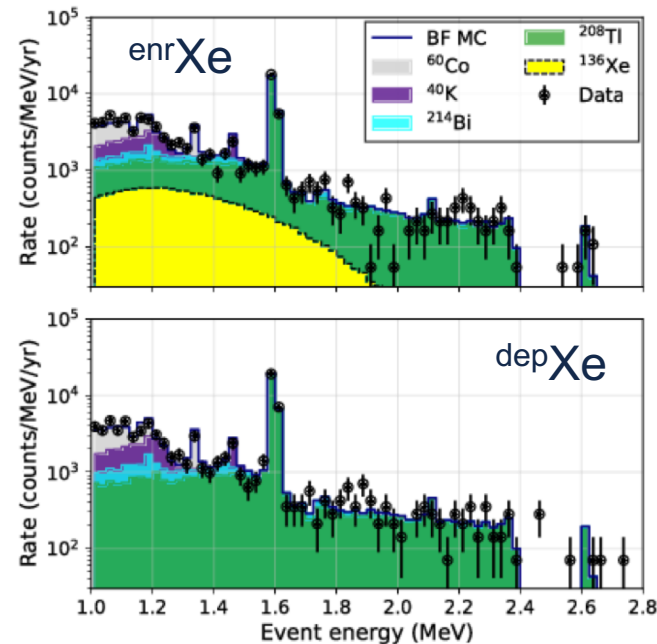
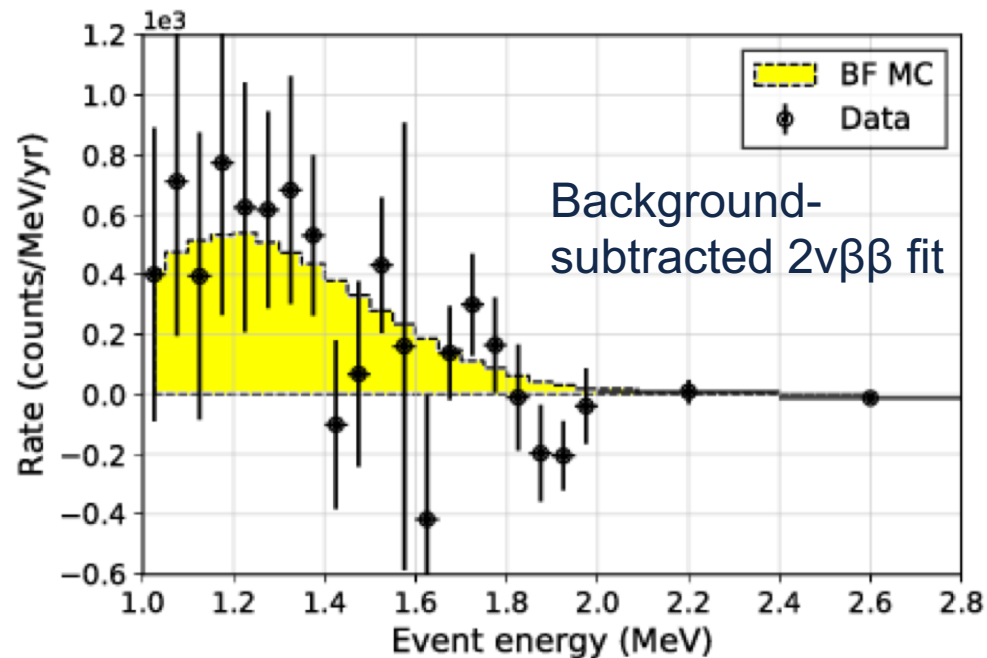
$\beta\beta$ Signal



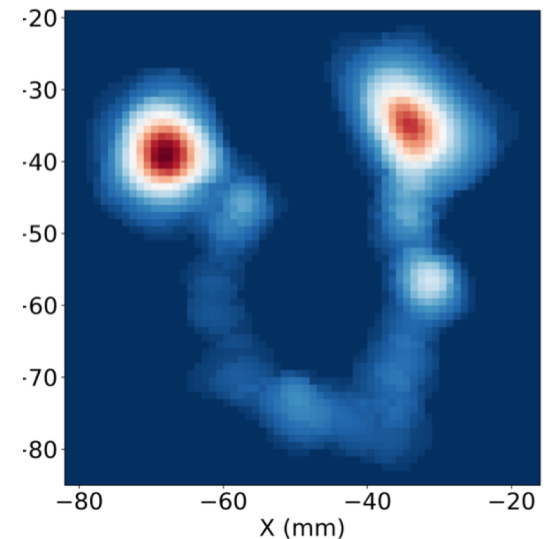
e- Track Background

NEXT-White

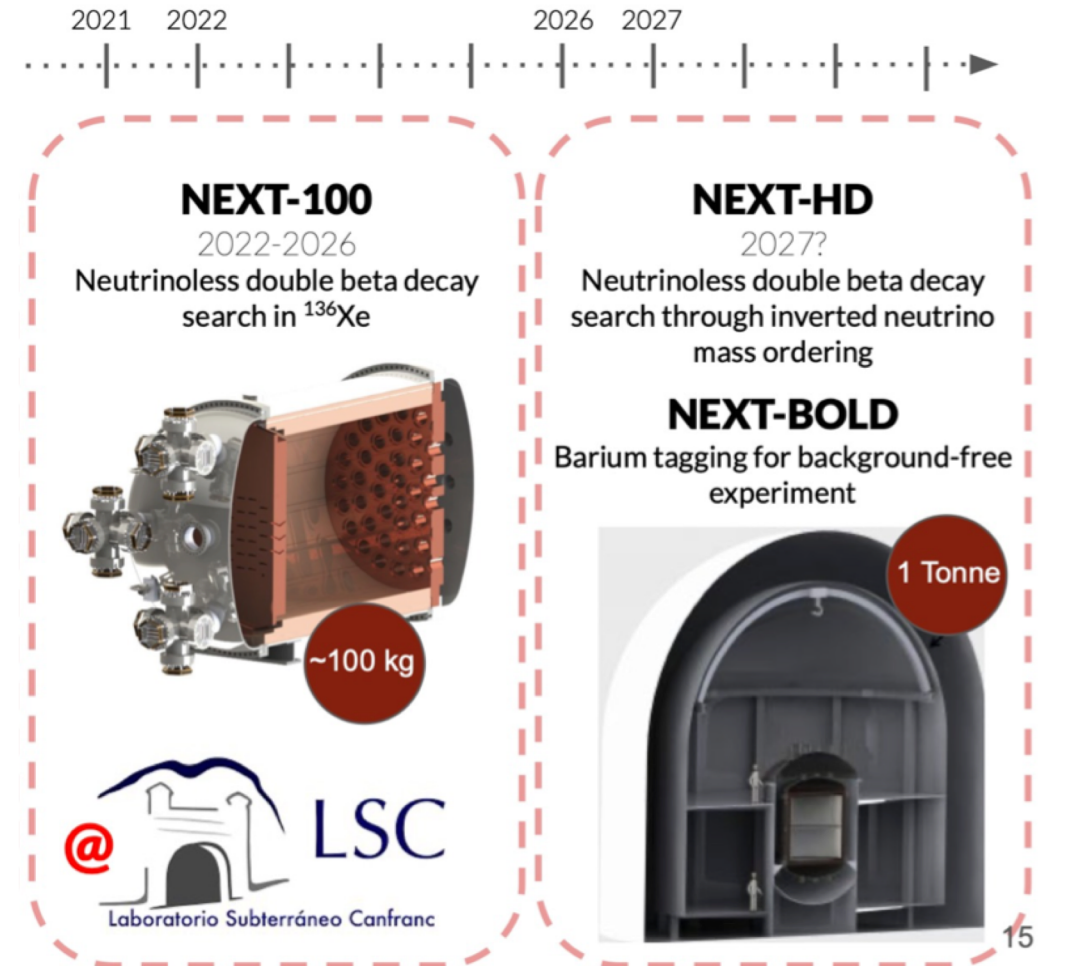
- 5kg demonstrator TPC, ran 2016-2021
- 10 bar pressure, energy resolution: 0.91% σ/E
- Right now, one issue is relatively low efficiency, larger detector will improve that



[arXiv:2201.10907](https://arxiv.org/abs/2201.10907)



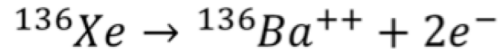
Future Directions for NEXT



- NEXT-100**
- 100 kg at 15 bar
 - Construction started in 2021
 - Expected sensitivity: 6×10^{25} yrs

- NEXT-HD**
- 1230 kg of $^{\text{enr}}\text{Xe}$
 - 1109 kg of ^{136}Xe
 - Energy res. (FWHM/E): 0.5%
 - BI $< 4 \times 10^{-6}$ cnts/(keV kg yr)
 - Discovery sensitivity: $T_{1/2} \sim 2.7 \times 10^{27}$ yrs
 - $m_{\beta\beta}$ discovery sensitivity: 8-45 meV

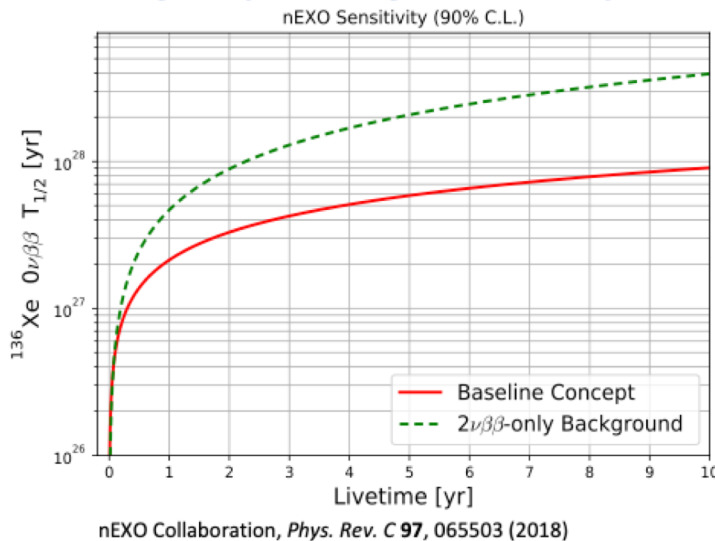
Barium Tagging



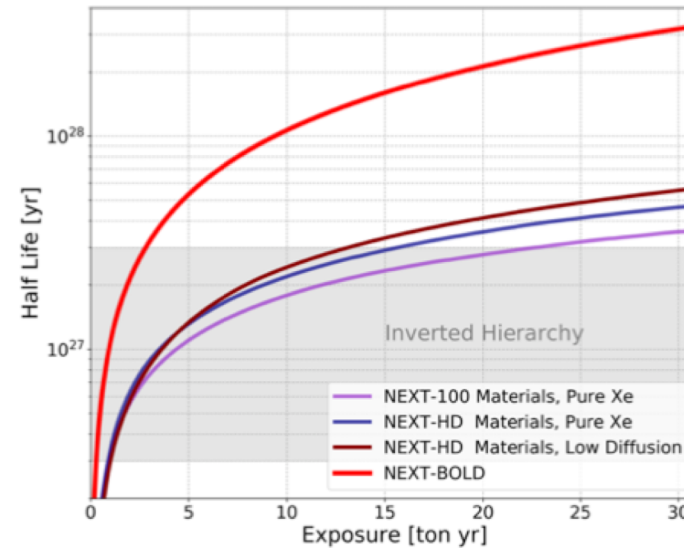
“Tagging” Ba daughter has potential to eliminate all but $2\nu\beta\beta$ backgrounds

M. Moe, Phys. Rev. C 44, R931 (1991)

In nEXO, eliminating other backgrounds could give up to 4x higher sensitivity



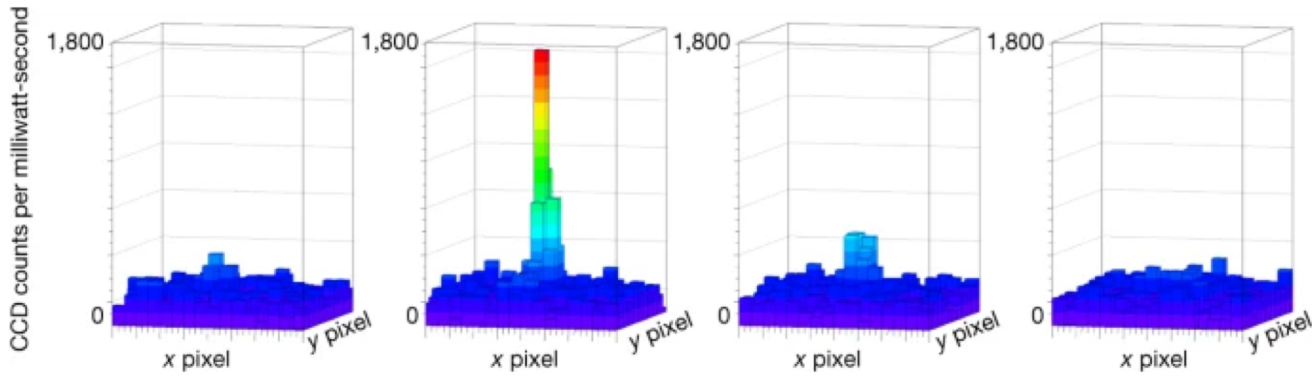
In NEXT, higher efficiency with Ba tagging and eliminating other backgrounds could provide up to a factor of 6 higher sensitivity



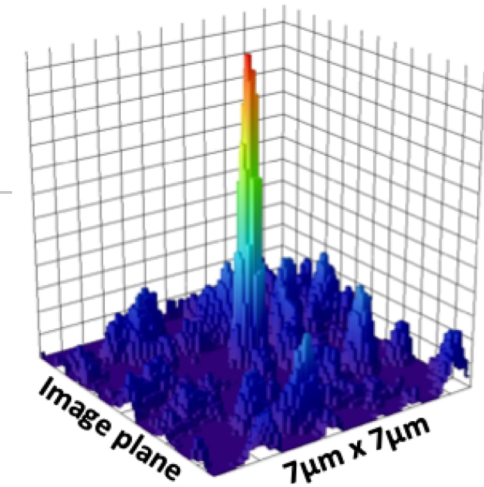
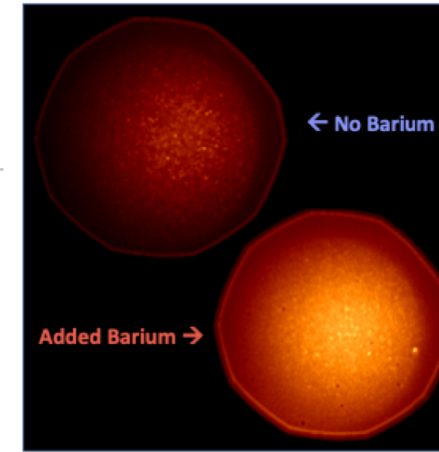
Materials courtesy of the NEXT and nEXO Collaborations, from B. Fairbank

- Considered a possible upgrade path for the tonne-scale TPC experiments
- Could extend sensitivity (further) into the normal ordering region!

R&D for Barium Tagging



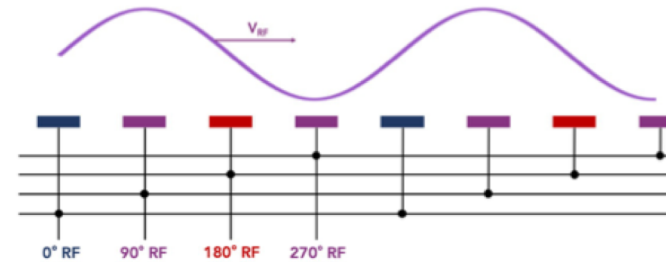
Laser-based ID in solid Xe for nEXO, *Nature* 569, 203-207 (2019)



Fluorescent molecule-based ID for NEXT, *ACS Sens.* 2021, 6, 1, 192–202 (2021)



Cryoprobe-based extraction for nEXO



RF carpet-based transport for NEXT, *arXiv:2111.11091* (2021)

- Feasible single-ion sensing techniques have been demonstrated in GXe and LXe
- Next steps: Barium capture, transport, and sensing in more-realistic detector environments

Other TPC R&D

For Xe, isotope acquisition is a challenge:

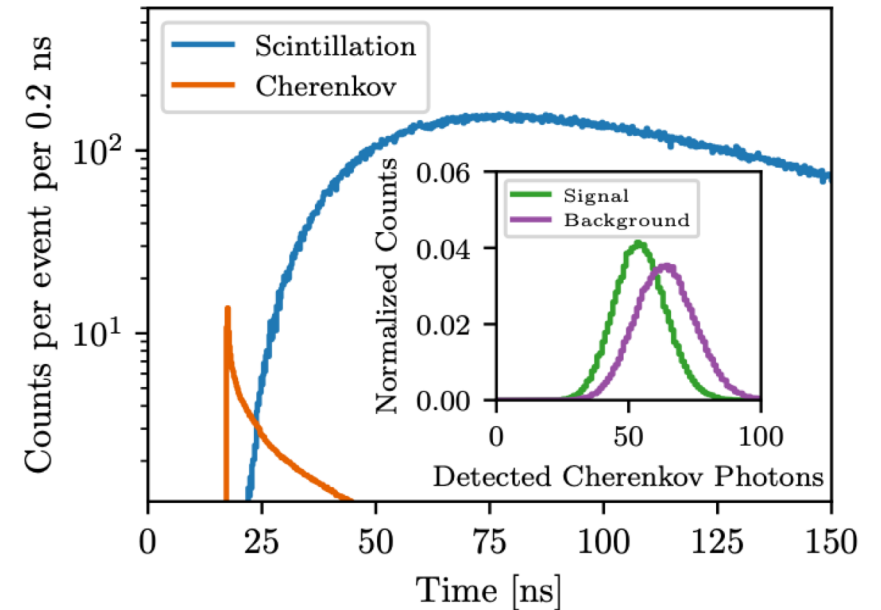
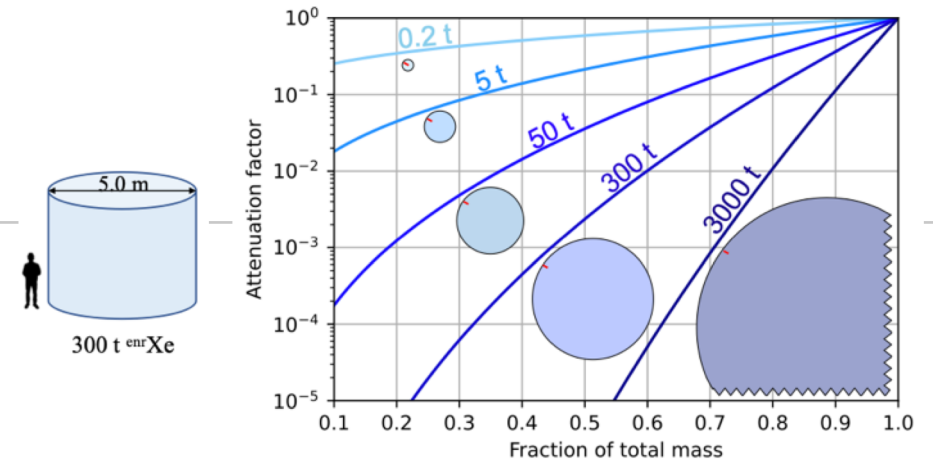
- Currently depends on liquid oxygen production for steel industry, limiting supply
- R&D on alternative extraction methods: Xe-adsorbing materials, could be implemented at CO2 capture plants

If acquisition can be resolved, kiloton-scale GXe and LXe TPCs should be feasible:

- R&D: increasing light detection efficiency, Cherenkov light-based background reduction
- Projected sensitivity $\sim 10^{30}$ yrs

Other ideas:

- DUNE and DarkNoon: Xe-doped LAr; R&D on energy resolution, gas mixture handling, and Cherenkov/scintillation response
- SeF₆ TPCs: R&D on ion readout techniques



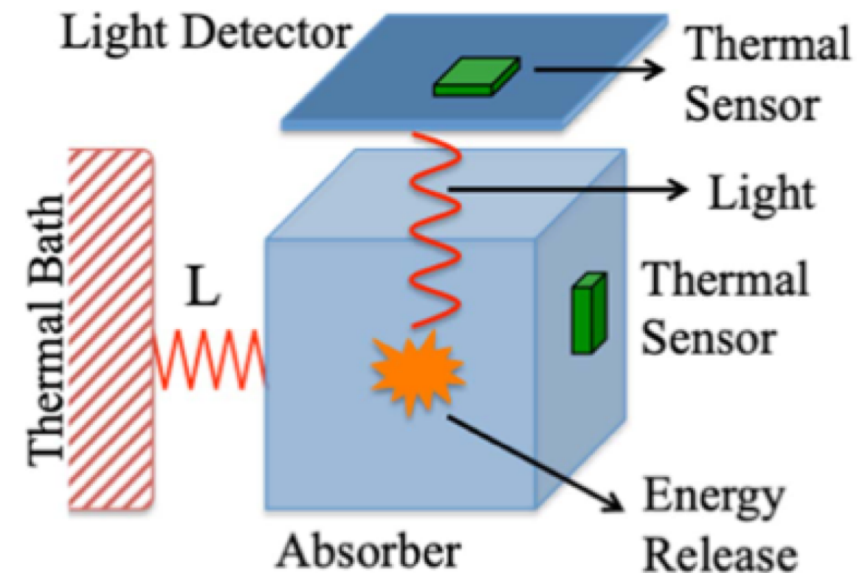
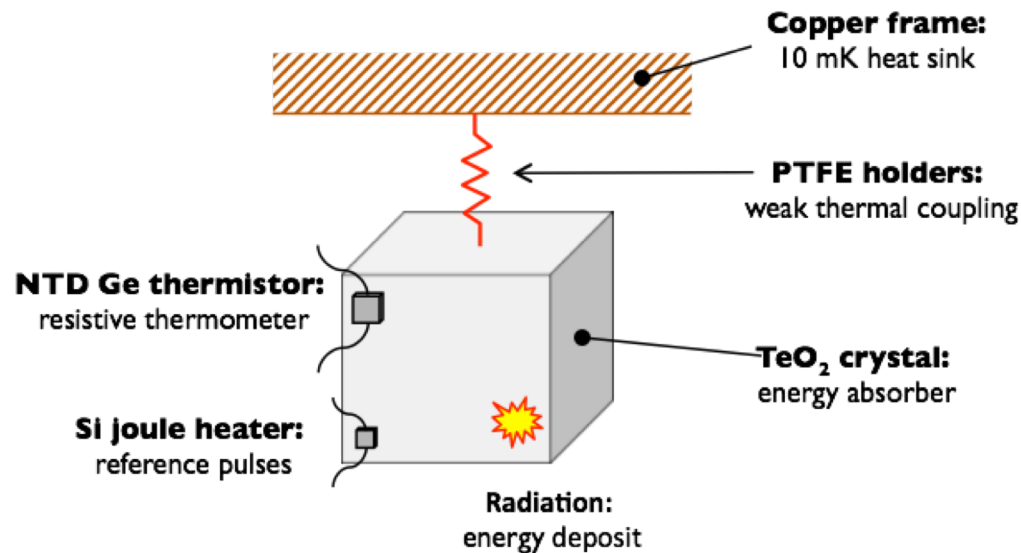
Physical Review D 104 (11): 112007 (2021)

The Experiments

Bolometer Experiments

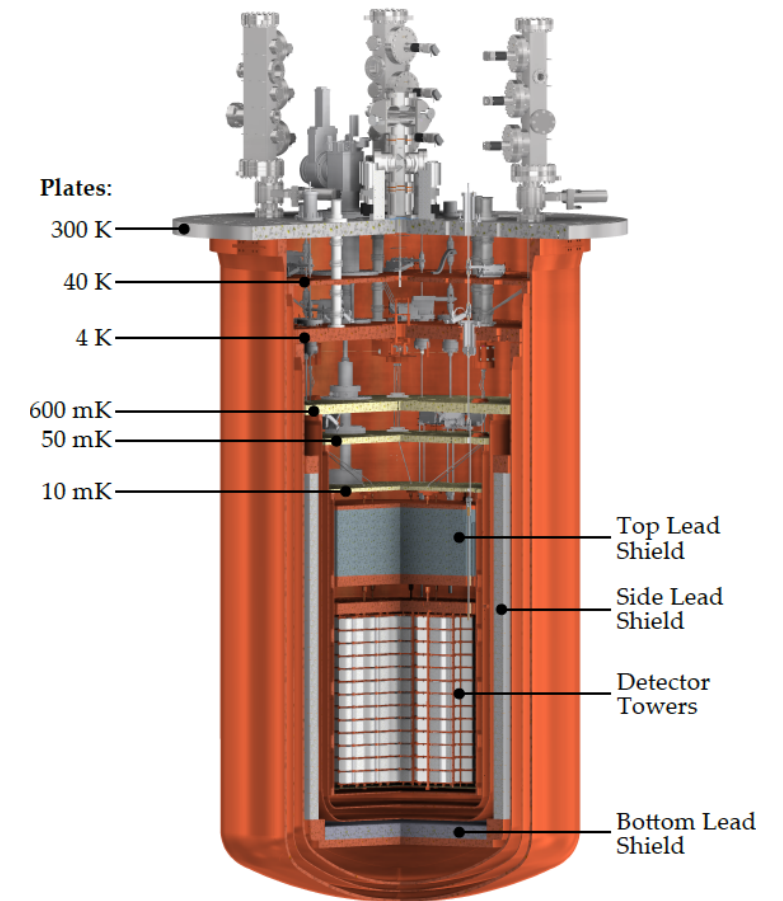
Bolometer Detector Concept

- Keep crystals at ~ 15 mK temperature in dilution fridge
- Interactions create phonons, read out with temperature sensor
- Some experiments use scintillating crystals and add a light detector: lets you distinguish α from β/γ

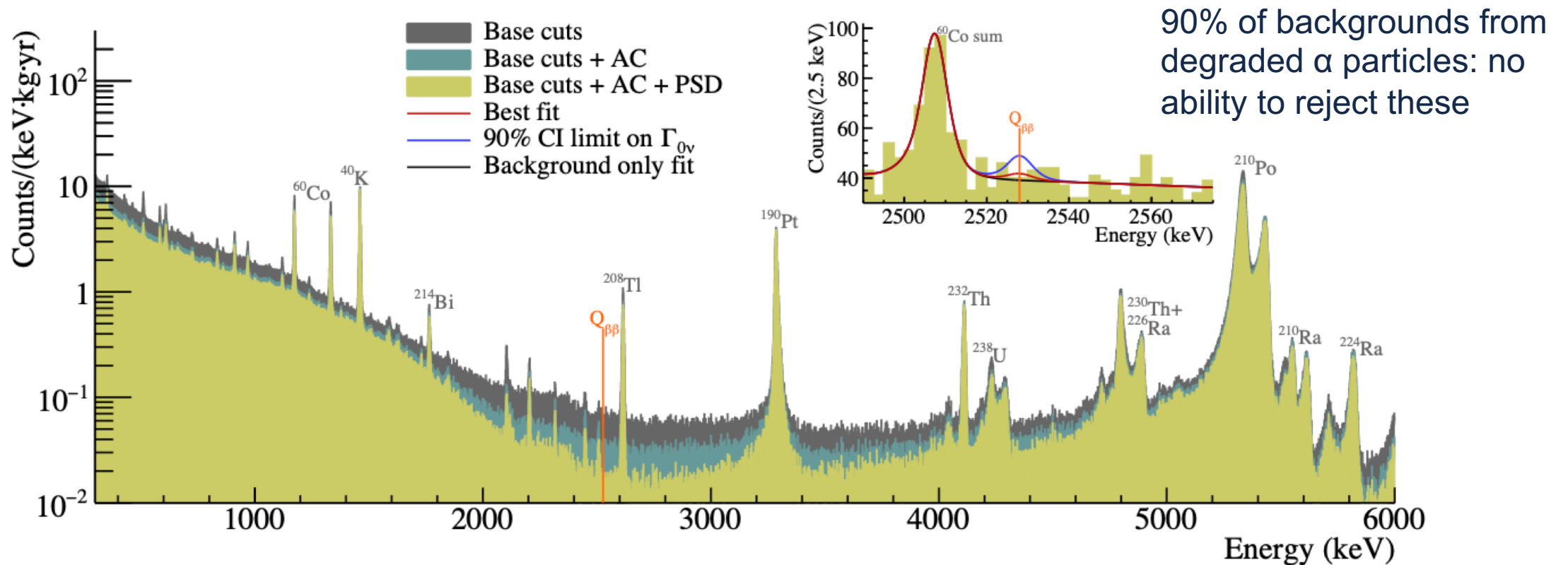


CUORE

- 988 natural-abundance TeO_2 crystals operated as bolometers
- 742 kg of detectors, 206 kg of ^{130}Te
- Energy resolution is 7.4 - 8.3 keV FWHM at $Q_{\beta\beta}$
- Taking data since 2017
- Will continue to take data until CUPID begins



CUORE Results



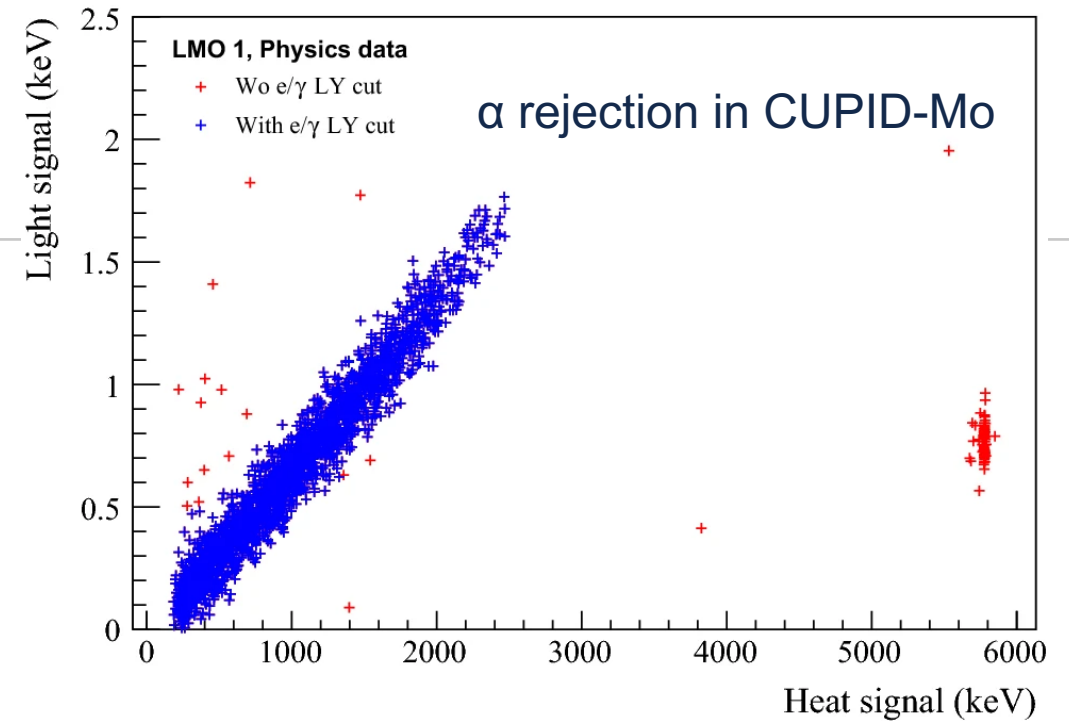
2021 Results: $T_{1/2} > 2.2 \times 10^{25}$ yr, $m_{\beta\beta} > 90 - 305$ meV

DOI: 10.1038/s41586-022-04497-4

CUPID

- Tonne-scale bolometer approach demonstrated in CUORE
- ^{100}Mo Q-value: 3.03 MeV
- Scintillating bolometer technique demonstrated in CUPID-Mo and other experiments, allows for α rejection
- Switch from CUORE crystals to scintillating bolometers with light readout in existing infrastructure
- Other options for crystal/isotope: ZnSe (candidate ^{82}Se), CdWO_4 (candidate ^{116}Cd), and TeO_2 (candidate ^{130}Te)

Material provided by CUORE, CUPID, CUPID-Mo, and CUPID-0 Collaborations

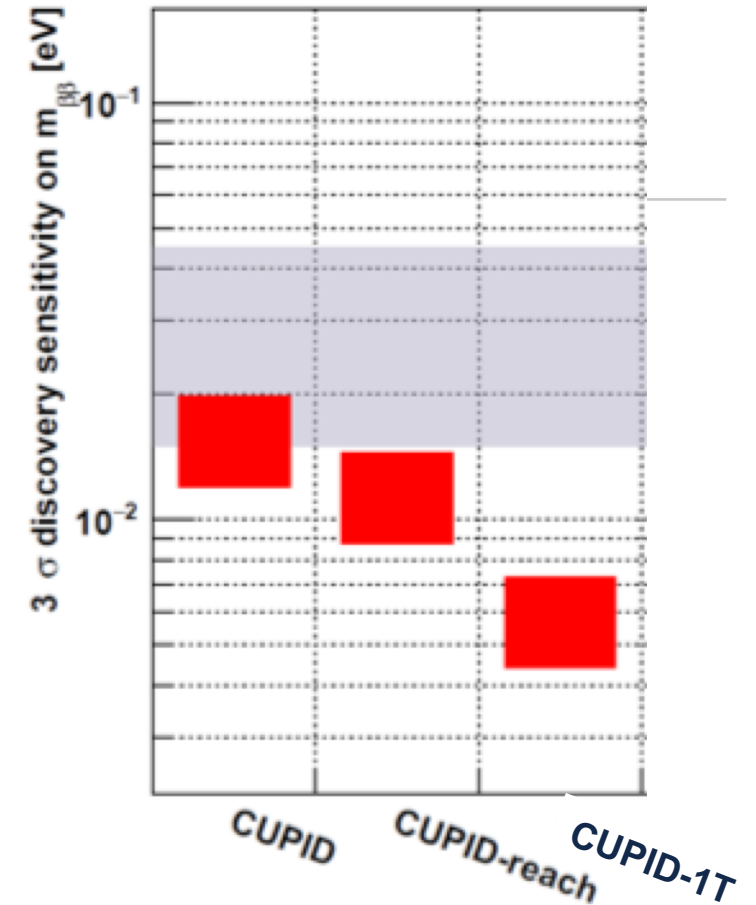


- Crystal: $\text{Li}_2^{100}\text{MoO}_4$
- Enrichment > 95% \rightarrow 253 kg of ^{100}Mo
- Energy res. (FWHM): 5 keV
- BI < 10^{-4} cnts/(keV kg yr)
- Discovery sensitivity: $T_{1/2} \sim 1.1 \times 10^{27}$ yrs
- $m_{\beta\beta}$ discovery sensitivity: 12-20 meV

Beyond Ton-Scale in Bolometers

- R&D Areas: high-speed superconducting sensors, multiplexed readout technologies, active γ veto, CMOS and ASIC instrumentation for quantum sensors, superconducting crystal coatings for improved PSD
- Could adopt a diffuse staging technique, with sites around the world

Parameter	CUPID Baseline	CUPID-reach	CUPID-1T
Crystal	$\text{Li}_2^{100}\text{MoO}_4$	$\text{Li}_2^{100}\text{MoO}_4$	$\text{Li}_2^{100}\text{MoO}_4$
Detector mass (kg)	472	472	1871
^{100}Mo mass (kg)	253	253	1000
Energy resolution FWHM (keV)	5	5	5
Background index (counts/(keV·kg·yr))	10^{-4}	2×10^{-5}	5×10^{-6}
Containment efficiency	79%	79%	79%
Selection efficiency	90%	90%	90%
Livetime (years)	10	10	10
Half-life exclusion sensitivity (90% C.L.)	1.5×10^{27} y	2.3×10^{27} y	9.2×10^{27} y
Half-life discovery sensitivity (3σ)	1.1×10^{27} y	2×10^{27} y	8×10^{27} y
$m_{\beta\beta}$ exclusion sensitivity (90% C.L.)	10–17 meV	8.2–14 meV	4.1–6.8 MeV
$m_{\beta\beta}$ discovery sensitivity (3σ)	12–20 meV	8.8–15 meV	4.4–7.3 meV



“Toward Sensitivity to the Neutrino Normal Hierarchy with Quantum Calorimetry,” D. Speller, Y. Kolomensky, L. Winslow, Snowmass LOI

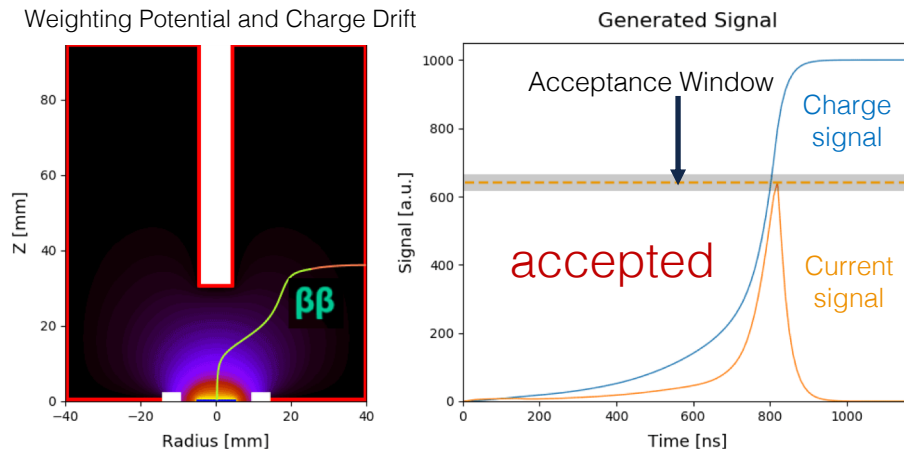
The Experiments

Germanium Experiments

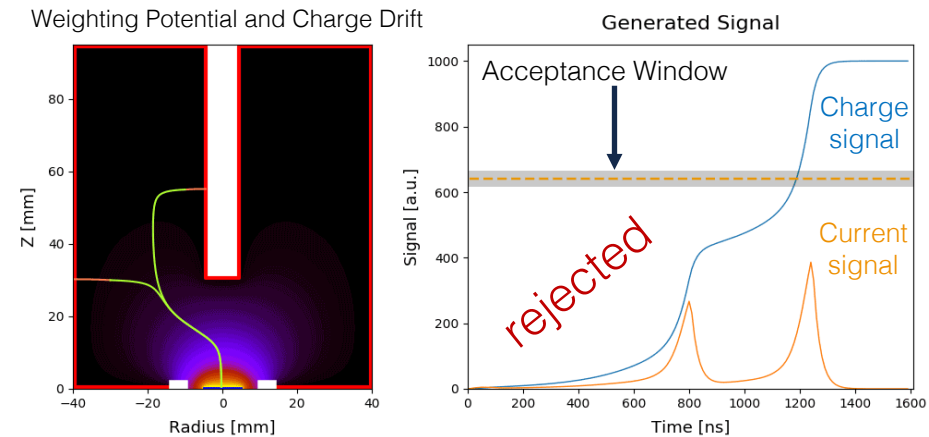
Germanium Detector Concept

- Single-crystal semiconductor diode made from enriched Ge
- Read out ionization signal– pulse shape can be used to distinguish multi-site and surface events from $\beta\beta$ events
- Excellent energy resolution: $\sim 0.05\% \sigma/E$

$0\nu\beta\beta$ signal candidate (single-site)



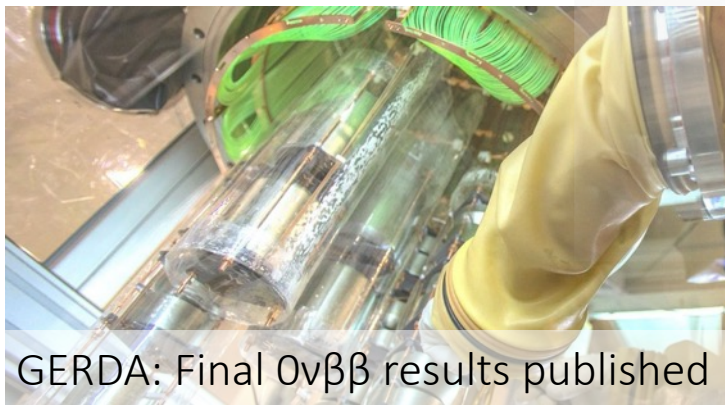
γ -background (multi-site)



Ge-Based $0\nu\beta\beta$ Program

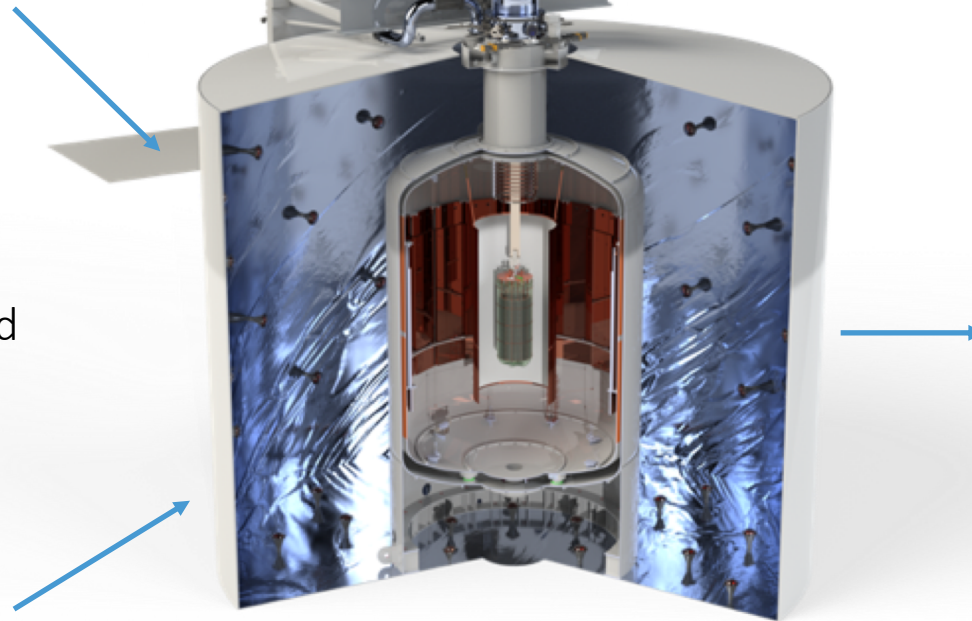


MJD: Final results recently announced

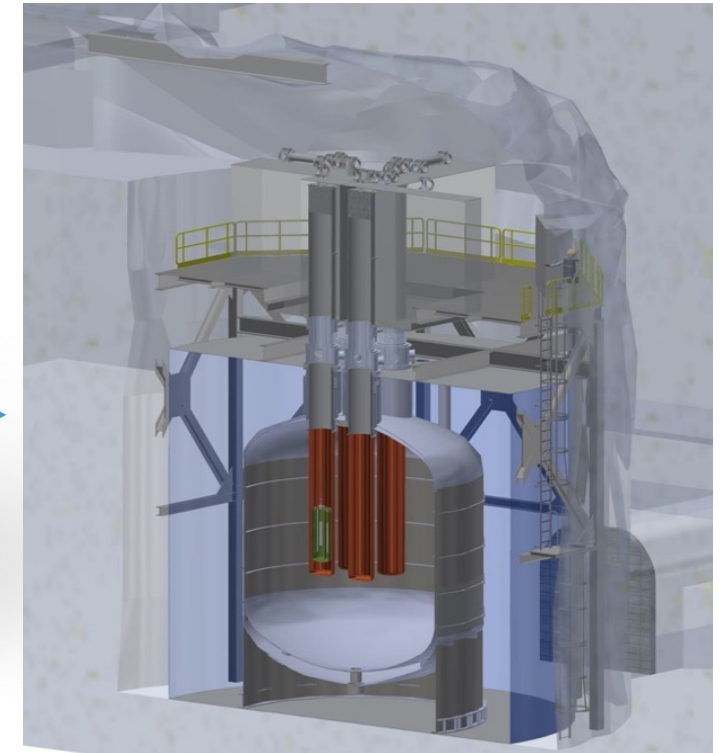


GERDA: Final $0\nu\beta\beta$ results published

[PRL 125, 252502 \(2020\)](#)



LEGEND-200: Now in commissioning

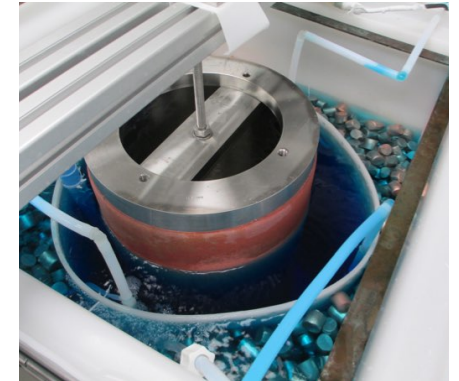
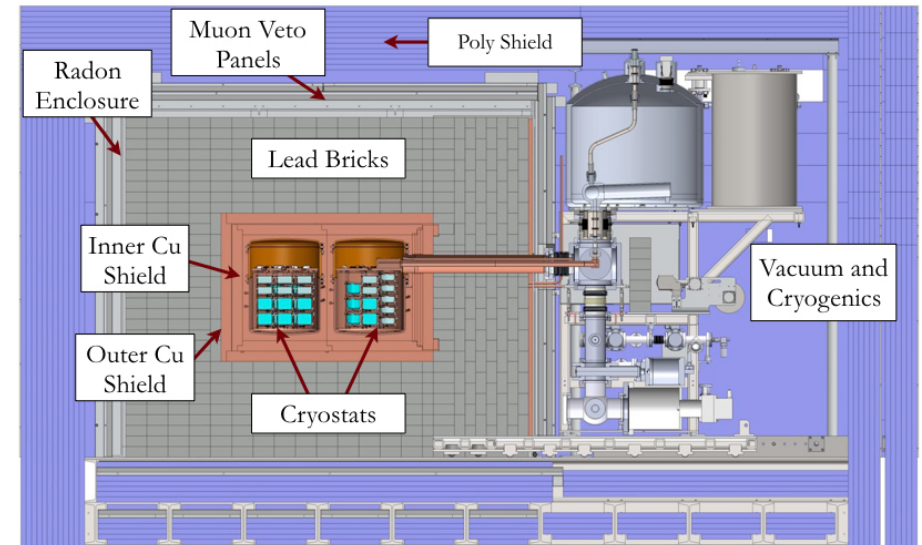


LEGEND-1000: Conceptual design development ongoing

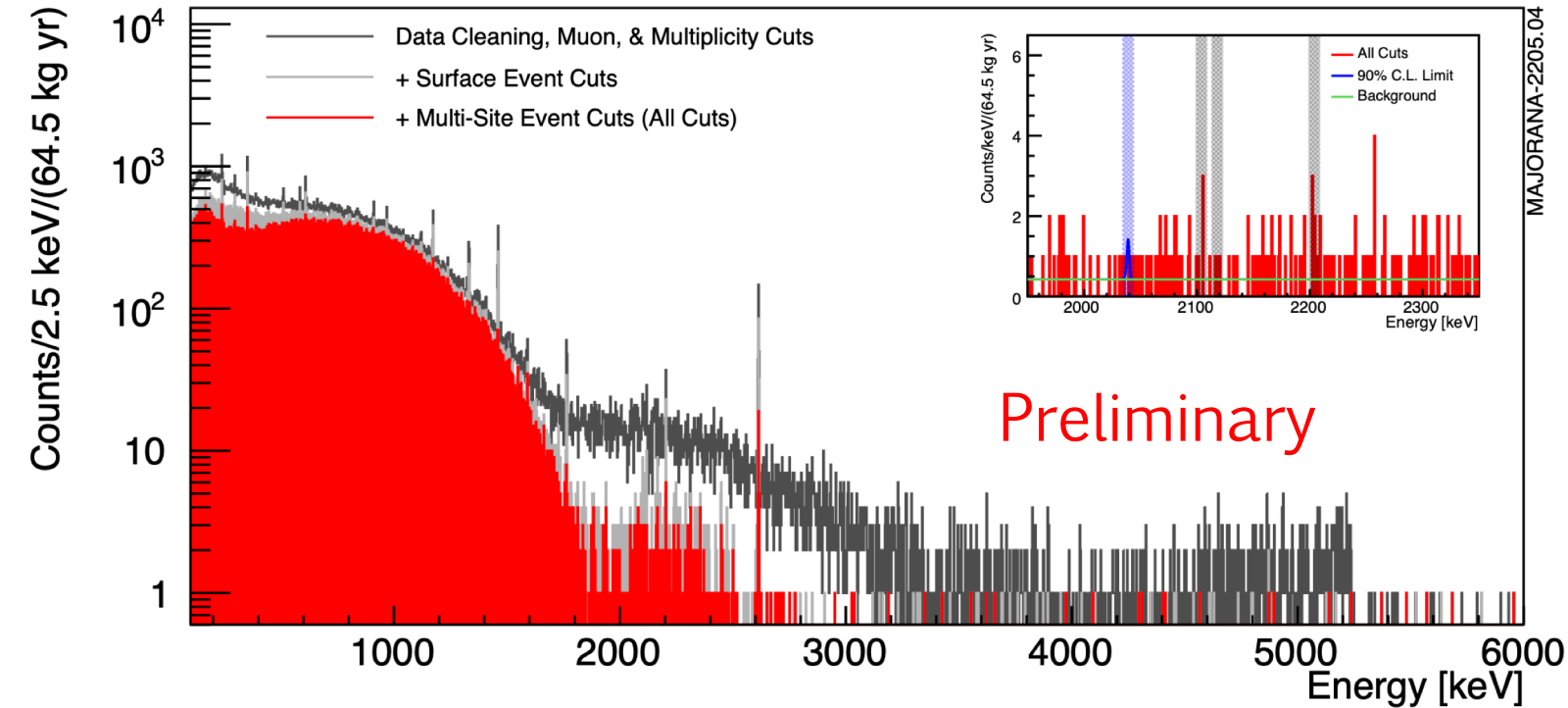
[arXiv: 2107.11462](#)

The MAJORANA DEMONSTRATOR

- ~30kg of enriched Ge detectors in a compact graded shield
- Innovations:
 - make underground electroformed copper (“the cleanest copper in the world”)
 - Low-mass low-noise low-background electronics enable the best energy resolution of any $0\nu\beta\beta$ experiment and a low energy threshold



The MAJORANA DEMONSTRATOR Results



Final enriched detector active exposure:

$$64.5 \pm 0.9 \text{ kg yrs}$$

Background Index at 2039 keV in lowest background config:

$$15.7 \pm 1.4 \text{ cts}/(\text{FWHM t yr})$$

Background Index:

$$(6.2 \pm 0.6) \times 10^{-3} \text{ cts}/(\text{keV kg yr})$$

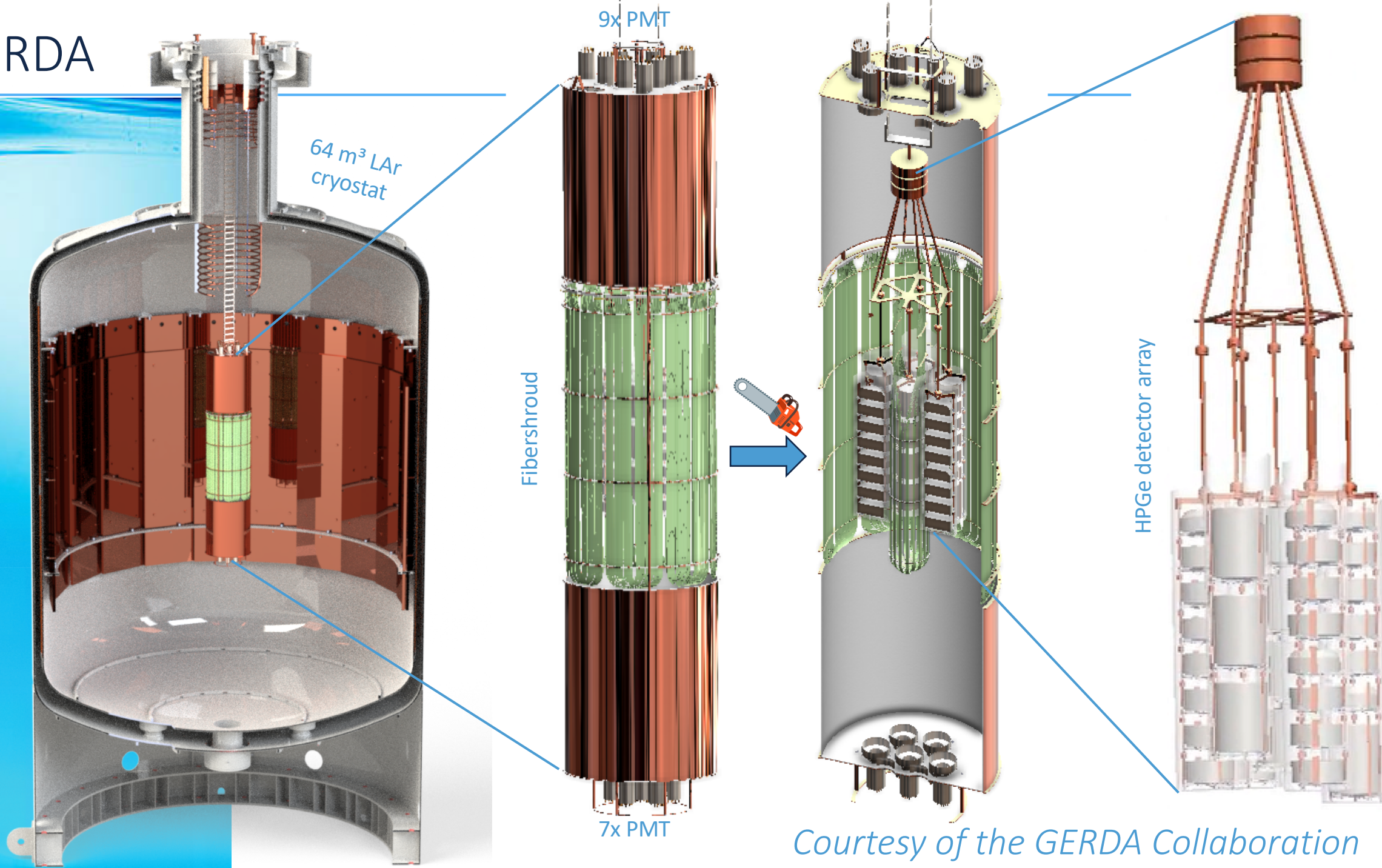
Energy resolution (FWHM) : 2.5 keV

Median $T_{1/2}$ Sensitivity: $8.1 \times 10^{25} \text{ yr}$ (90% C.I.)

65 kg-yr Exposure Limit: $T_{1/2} > 8.3 \times 10^{25} \text{ yr}$ (90% C.I.)

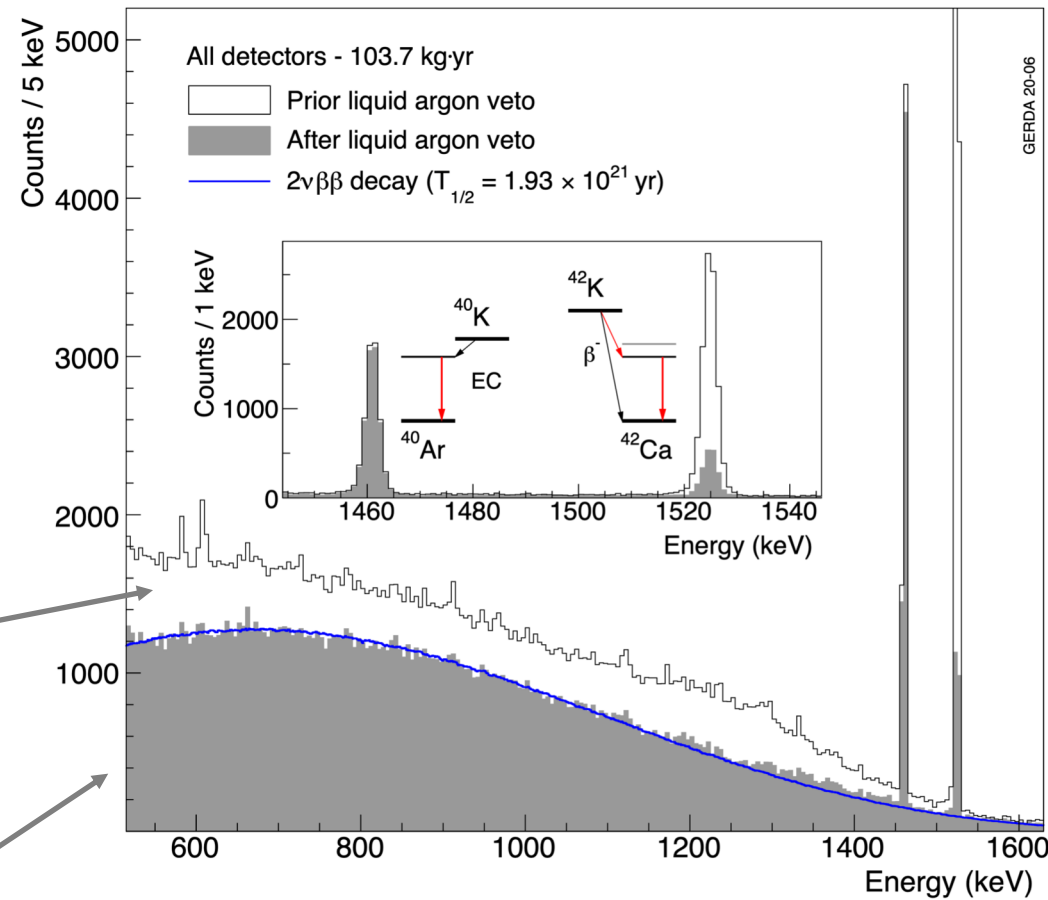
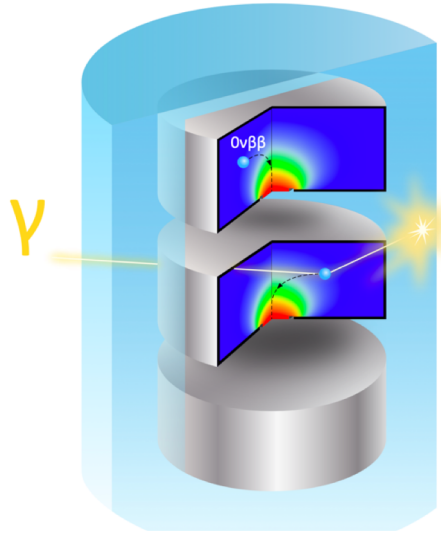
GERDA

590 m³ instrumented water tank



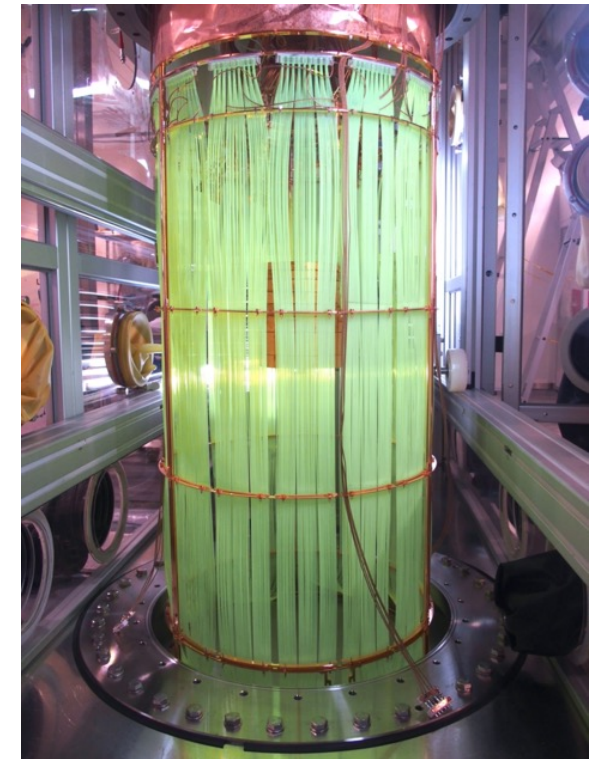
Courtesy of the GERDA Collaboration

LAr Active Veto



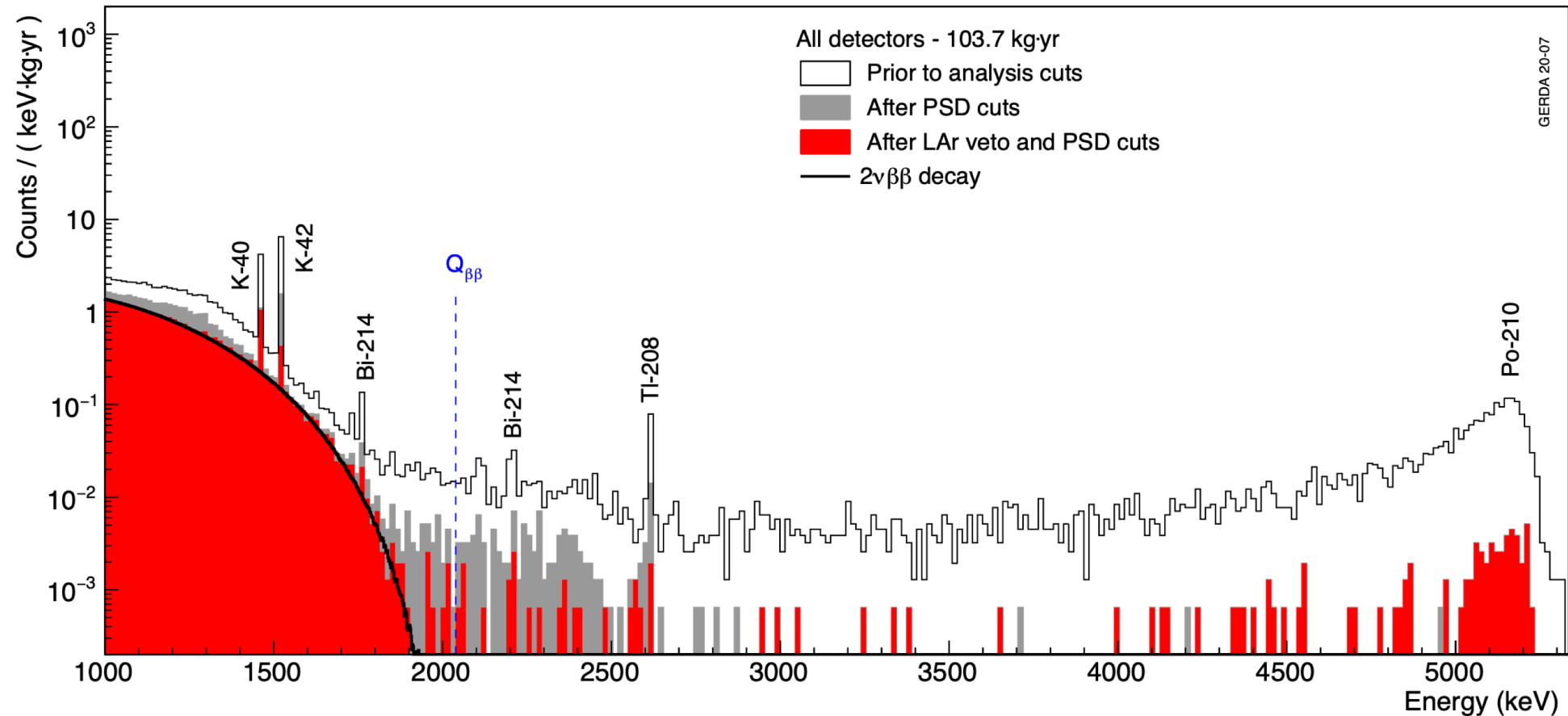
Compton events with energy deposition in the LAr

Pure $2\nu\beta\beta$ spectrum after LAr signal suppression



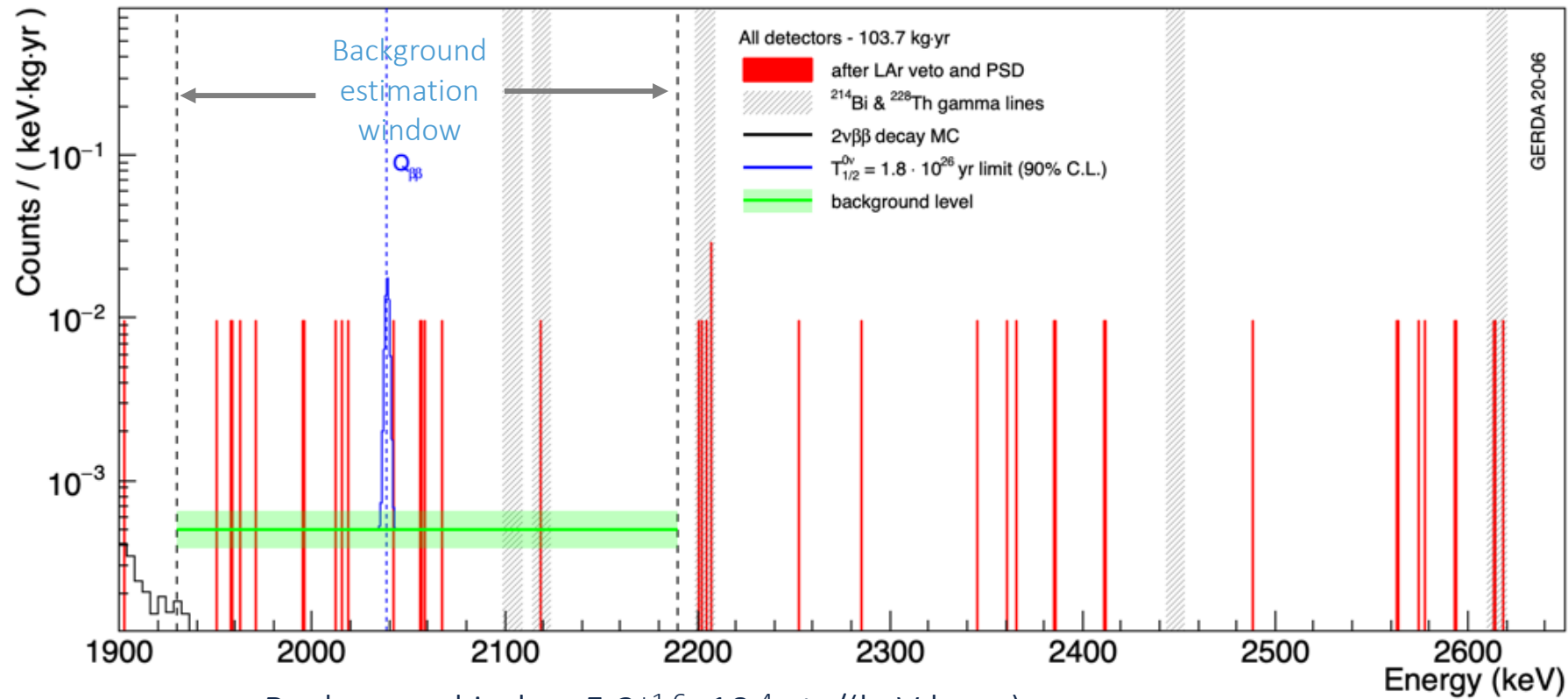
Courtesy of the GERDA Collaboration

GERDA Final-Exposure Spectrum



Courtesy of the GERDA Collaboration

GERDA Results



PRL 125, 252502 (2020)

Background index: $5.2^{+1.6} \cdot 10^{-4}$ cts/(keV kg yr),

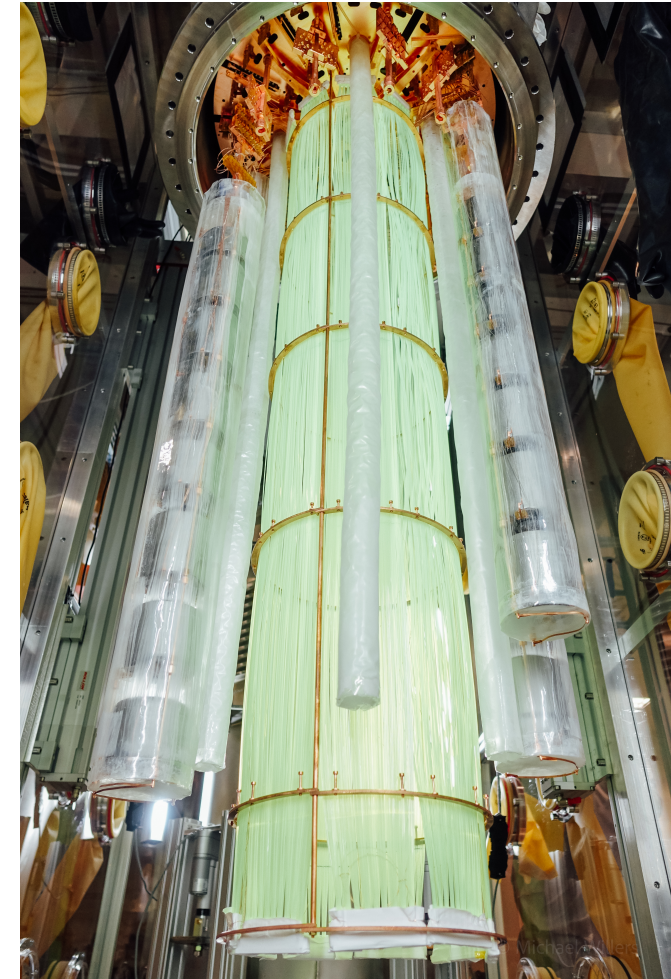
Energy resolution ~ 3 keV (FWHM)

Frequentist: $N^{0\nu} = 0$ best fit, $T_{1/2} > 1.8 \cdot 10^{26}$ yr at 90% C.L.

*Courtesy of the
GERDA Collaboration*

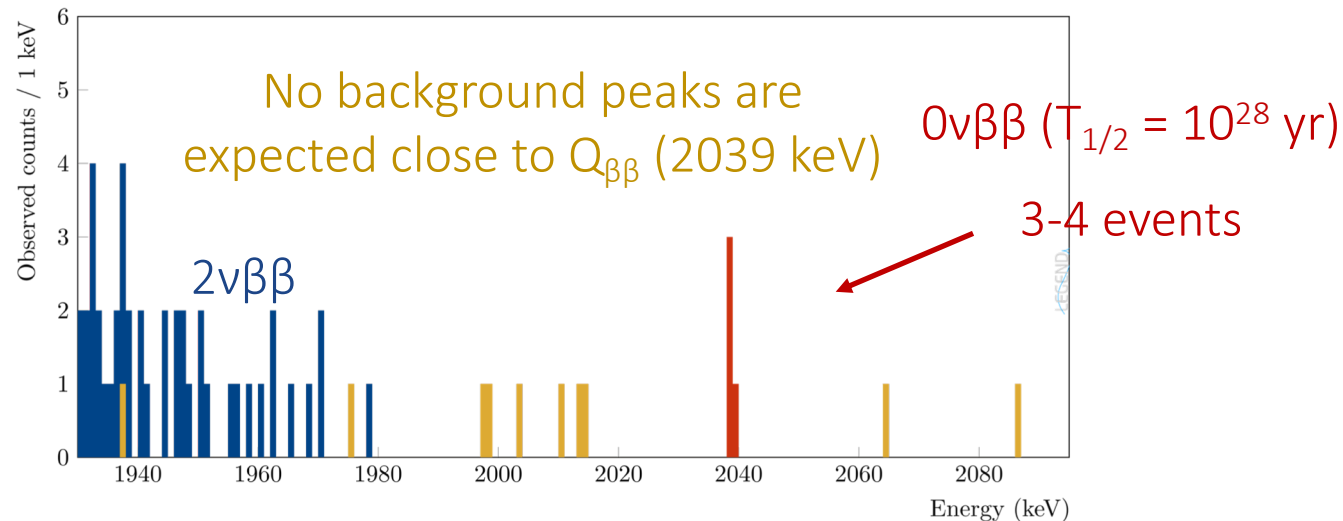
LEGEND-200

- 200 kg of HPGe ICPC detectors enriched to 91% ^{76}Ge , operated in active LAr shield
- Upgrade of GERDA infrastructure
- Uses new, larger inverted-coaxial point-contact detectors
- >2.5x background reduction relative to GERDA expected
- Commissioning underway, full detector deployment expected this year
- Final expected sensitivity:
 $T_{1/2} > 1 \times 10^{27}$ yrs



LEGEND-1000

- New cryostat either at LNGS or SNOLAB; 4 independent payloads with depleted UAr in inner volumes
- 1000 kg of enriched Ge detectors: fabricate 870 kg of new detectors; use 130 kg from LEGEND-200; recycle 50 kg of small detectors
- Multi-dimensional analysis nearly eliminates γ , surface α and β backgrounds
- LEGEND goal: BI $< 1 \times 10^{-5}$ cnts/keV kg yr, Discovery up to 1.3×10^{28} yrs



Projected spectrum: even a signal at the bottom of the inverted ordering will be visible to the eye

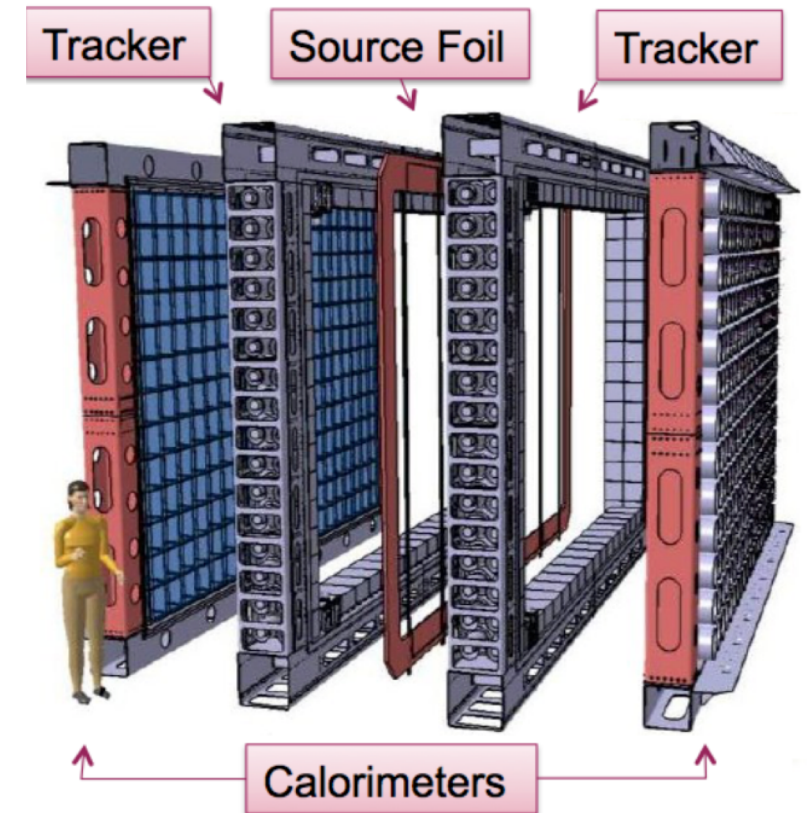
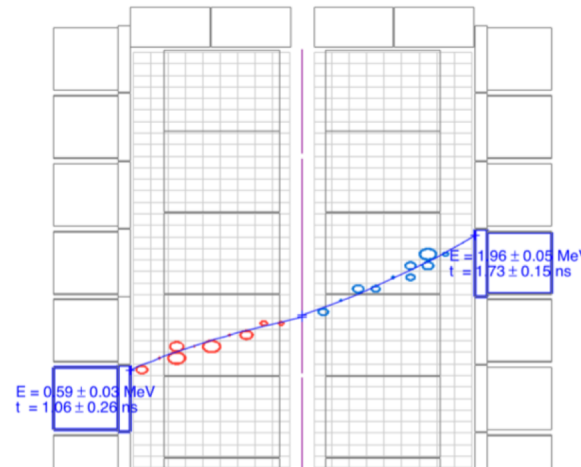
Multi-dimensional analysis can also be conducted

The Experiments

Tracking Experiments

Tracking Detector Concept

- Source \neq Detector!
- Thin foils of $\beta\beta$ decay materials sandwiched between tracking planes and surrounded by calorimeters
- Magnetic field for charge ID
- Allows full track reconstruction of $\beta\beta$ decays and energy measurement
- Can use any isotope
- Foils have to be just 0.3mm thick so β s can escape

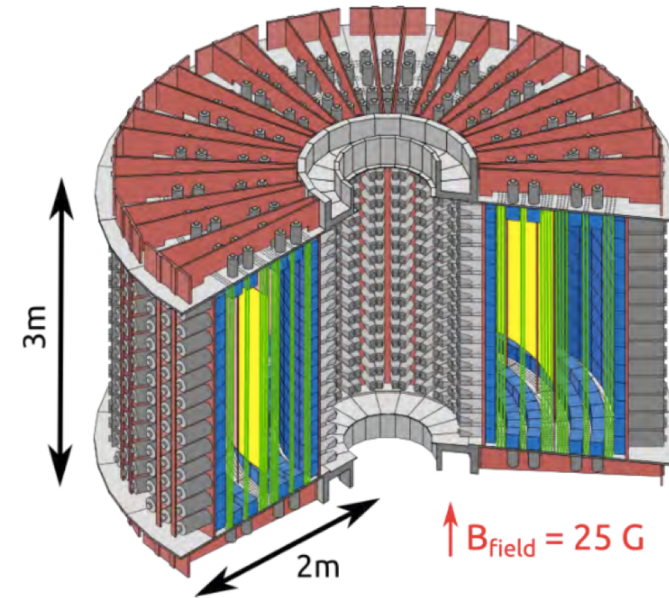
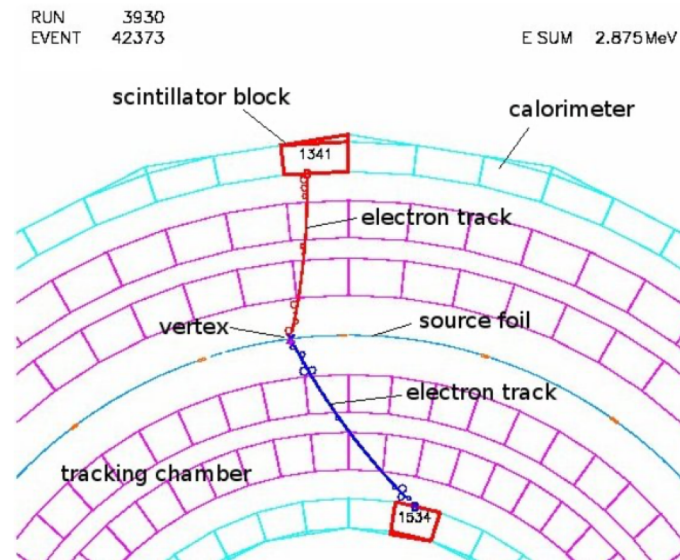
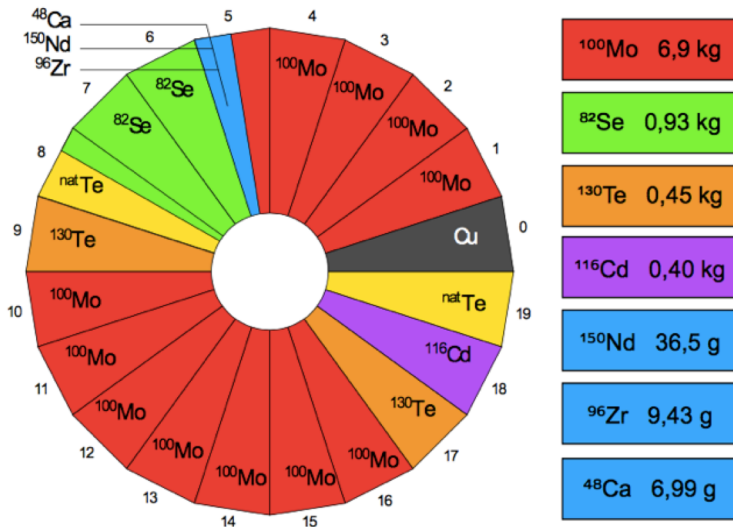


Images courtesy of the NEMO Collaboration

NEMO-3

- Ran from 2003-2011 with 10 kg total mass
- Measured many isotopes simultaneously (including some first measurements)

NEMO-3 "camembert" (source top view)



sources
60 mg/cm² foils
10 kg of $\beta\beta$ isotopes

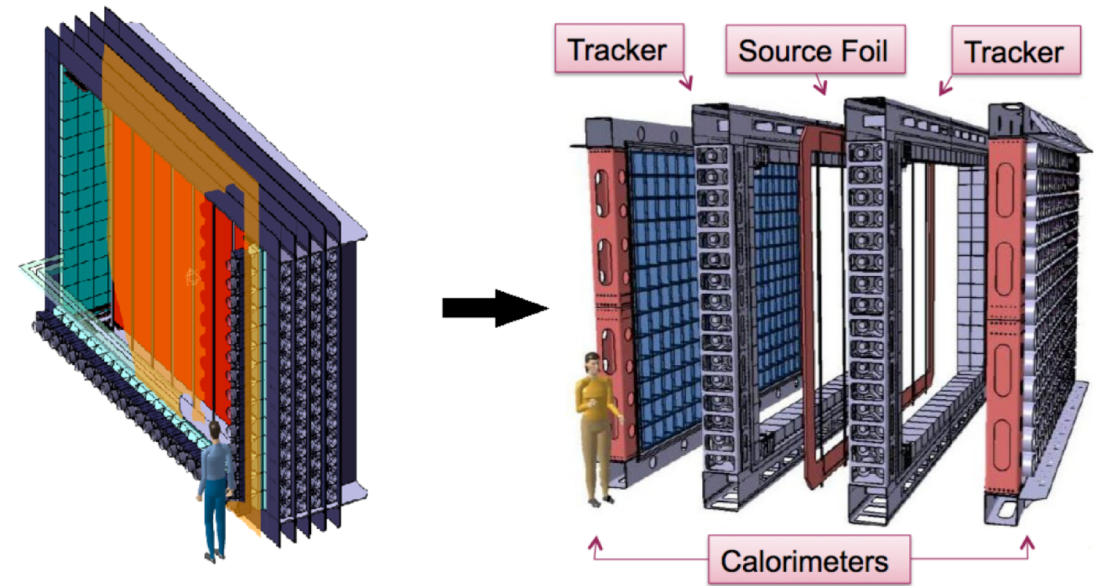
tracker
6180 Geiger cells
vertex resolution :
 $\sigma_{xy} \sim 3 \text{ mm}$ $\sigma_z \sim 10 \text{ mm}$

calorimeter
1940 optical modules :
polystyren scintillators
+ 3" and 5" PMTs
 $\text{FWHM}_E \sim 15\% / \sqrt{E_{\text{MeV}}}$
 $\sigma_t \sim 250 \text{ ps}$

Images courtesy of the
NEMO Collaboration

Super-NEMO and Science Goals

- Using ^{82}Se foils for initial run
- 1 module to start, up to 20 after that
- Low mass, so not much sensitivity ($\sim 10^{24}$ yrs)
- Tracking allows you to look for new physics in $2\nu\beta\beta$ decays
- R&D for future Majorana mass mechanism investigations?



Doing Other Physics
with $0\nu\beta\beta$ Experiments

Not Just $0\nu\beta\beta$...

- Once you've built a large low-background detector, there's a lot of other physics you can do with it!
- $\beta\beta$ -related searches
- Standard Model rare-event searches
- Dark matter searches
- Other BSM physics

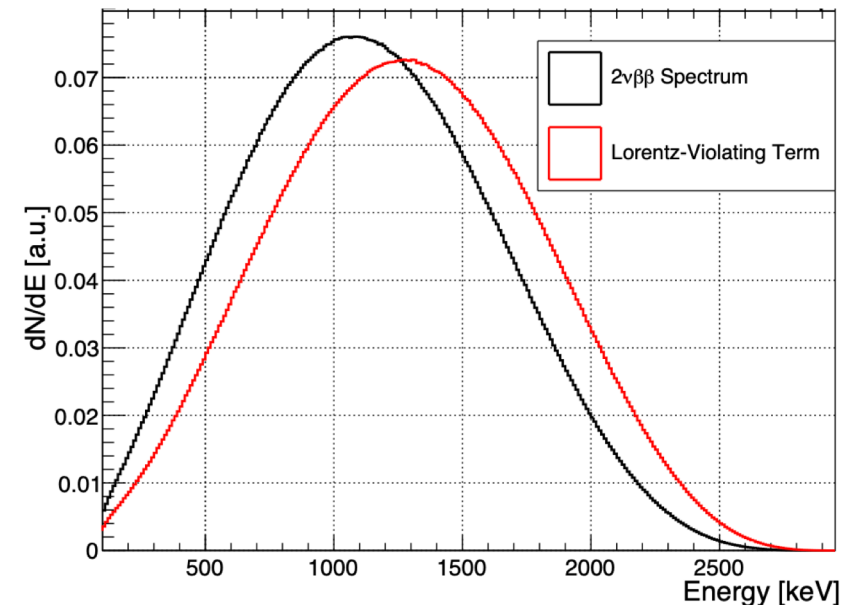
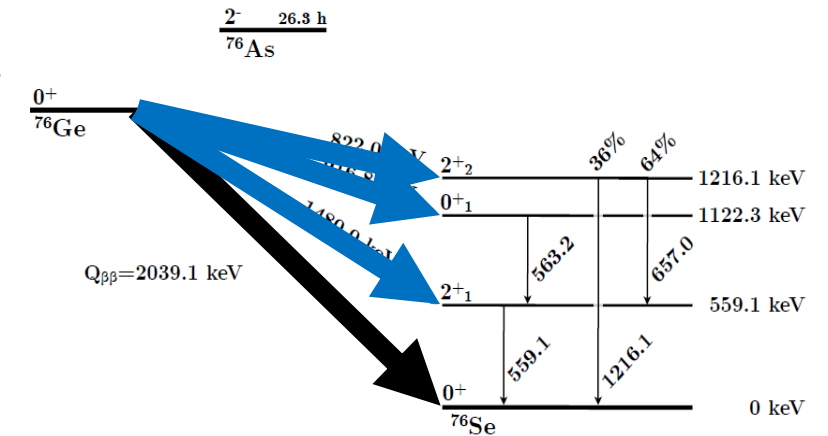
Other $\beta\beta$ -related Physics

Excited-state decays:

- $2\nu\beta\beta$ and $0\nu\beta\beta$ decays to excited states can also occur
- $\beta\beta$ decay, followed by γ emission from de-excitation
- The $2\nu\beta\beta$ decay to the 0_1^+ state has been measured in ^{150}Nd and ^{100}Mo . Other first measurements may be around the corner!
- $0\nu\beta\beta$ excited state decay measurements could help constrain the Majorana mass mechanism some day

Lorentz violation:

- Violation in $2\nu\beta\beta$ decay would change the spectral shape



DOI: 10.1103/PhysRevD.100.092002

Standard Model Rare-Event Searches

Addressing uncertainties in expected rates of SM processes:

- Studies of in-situ cosmogenics for future low-background experiments

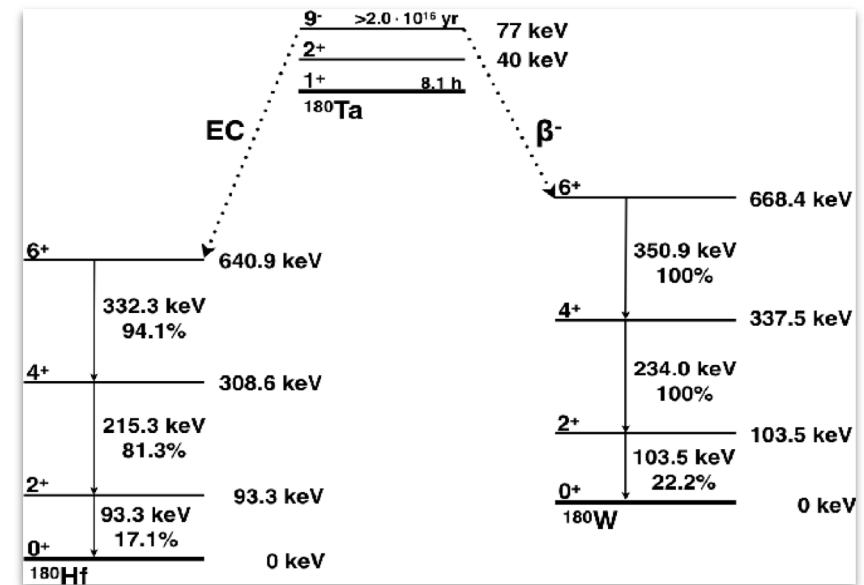
- (α, n) reactions from calibration sources

Rare-event searches with added sources:

- The Majorana Demonstrator is now searching for the decay of ^{180m}Ta : “nature’s longest-lived excited state”

- Expected sensitivity: 10^{19} yrs

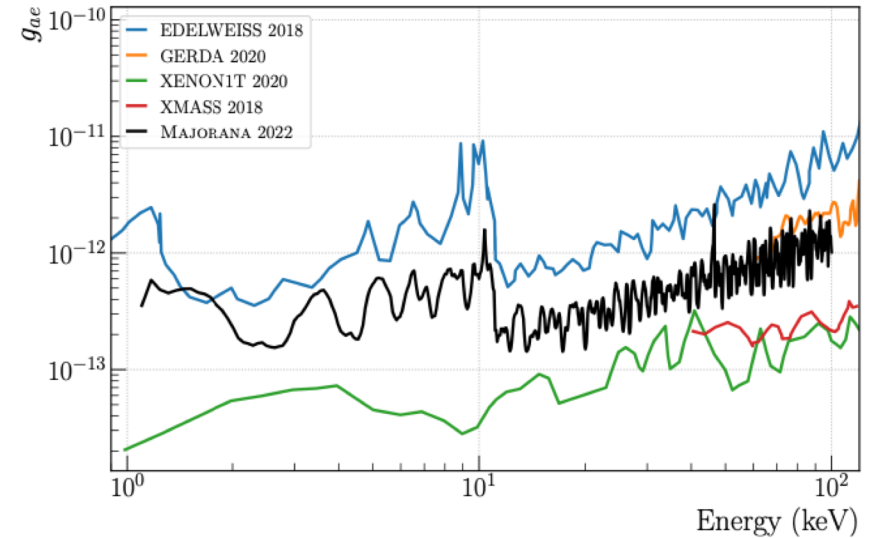
Added Ta foils



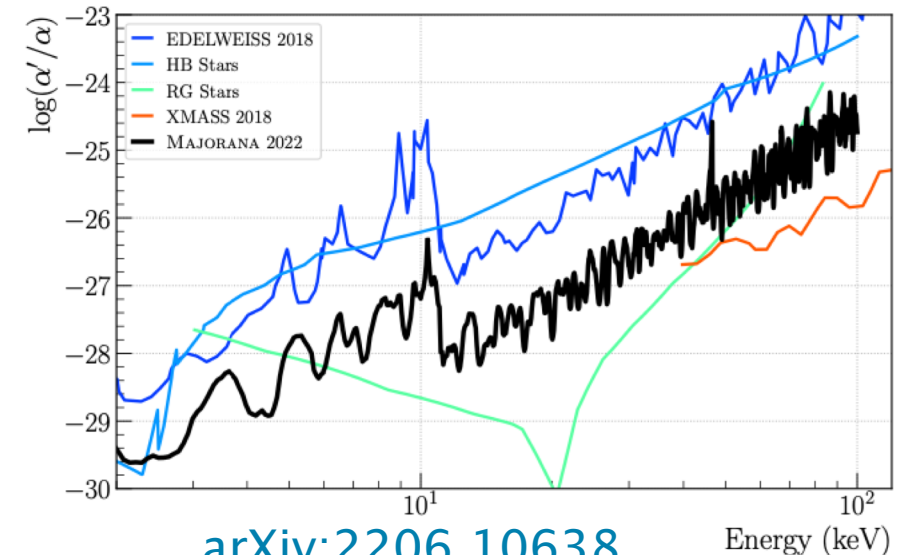
Dark Matter Searches

- Many searches enabled by low energy threshold and low backgrounds:
 - Pseudoscalar dark matter: axion-like particles
 - Vector dark matter: dark photons
 - Fermionic dark matter
 - Sterile neutrino
 - Primakoff solar axion
 - 14.4-keV solar axion
- Dark matter searches at higher energy:
 - keV-scale dark matter
 - Bosonic super-WIMPs
- $0\nu\beta\beta$ experiments can be competitive with dedicated dark matter experiments!

Pseudoscalar Dark Matter



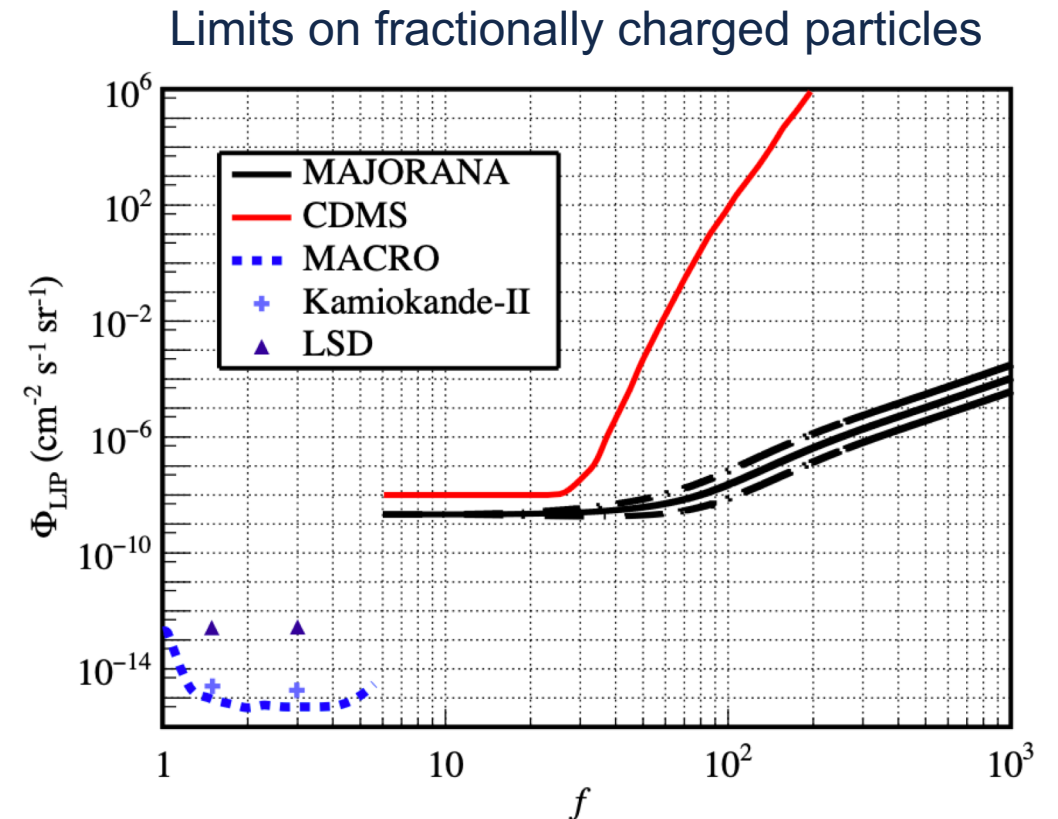
Vector Boson Dark Matter



[arXiv:2206.10638](https://arxiv.org/abs/2206.10638)

Other BSM Searches

- Tests of fundamental symmetries and conservations:
 - Baryon number violation
 - Pauli exclusion principle violation
- Other exotic physics:
 - Quantum wavefunction collapse
 - Lightly-ionizing (fractionally charged) particles



DOI: 10.1103/PhysRevLett.120.211804

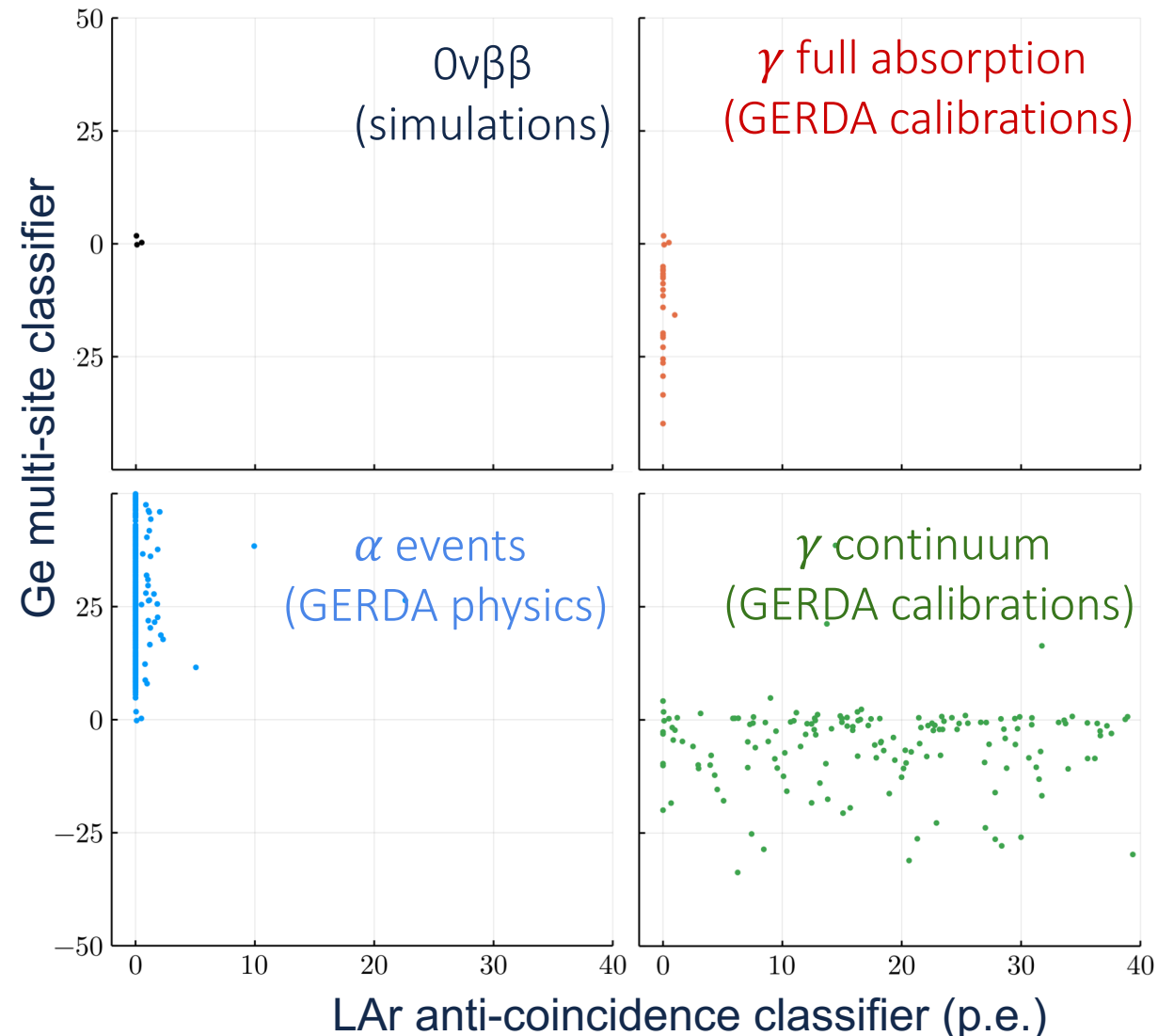
What comes after
discovery?

Confirming a Signal

- Make sure your signal behaves correctly as a function of your analysis parameters
- Depending on the experiment, you may be able to do an “on/off” measurement.
- Or, if your backgrounds are low enough, run for longer and increase your signal!
- Add more mass, if you can!
- If the signal is confirmed...

Example from LEGEND:

Three counts in $0\nu\beta\beta$ region: What is their origin?



Hand out the Nobel Prizes!

Where do we go from there?

- If we discover $0\nu\beta\beta$, we'll want to measure it in more isotopes: ratios between isotopes could provide insight into mechanism
- Precision measurements, along with measurements of other neutrino properties, would let us figure out the Majorana phases: many theories of flavor predict specific phases
- Try to measure excited-state decays to get more insight into mechanism
- Measure outgoing electron energies and momenta: some Majorana mass mechanism models make specific predictions
 - Build HyperNEMO for this?

Conclusion

- I hope you've enjoyed learning something about $0\nu\beta\beta$!
- This is a rich field with a lot of ongoing activity, and big things coming soon!
- We're entering a regime where discovery could come at any time, giving us vital insight into the nature of neutrino mass and potentially, the origin of the matter/anti-matter asymmetry.