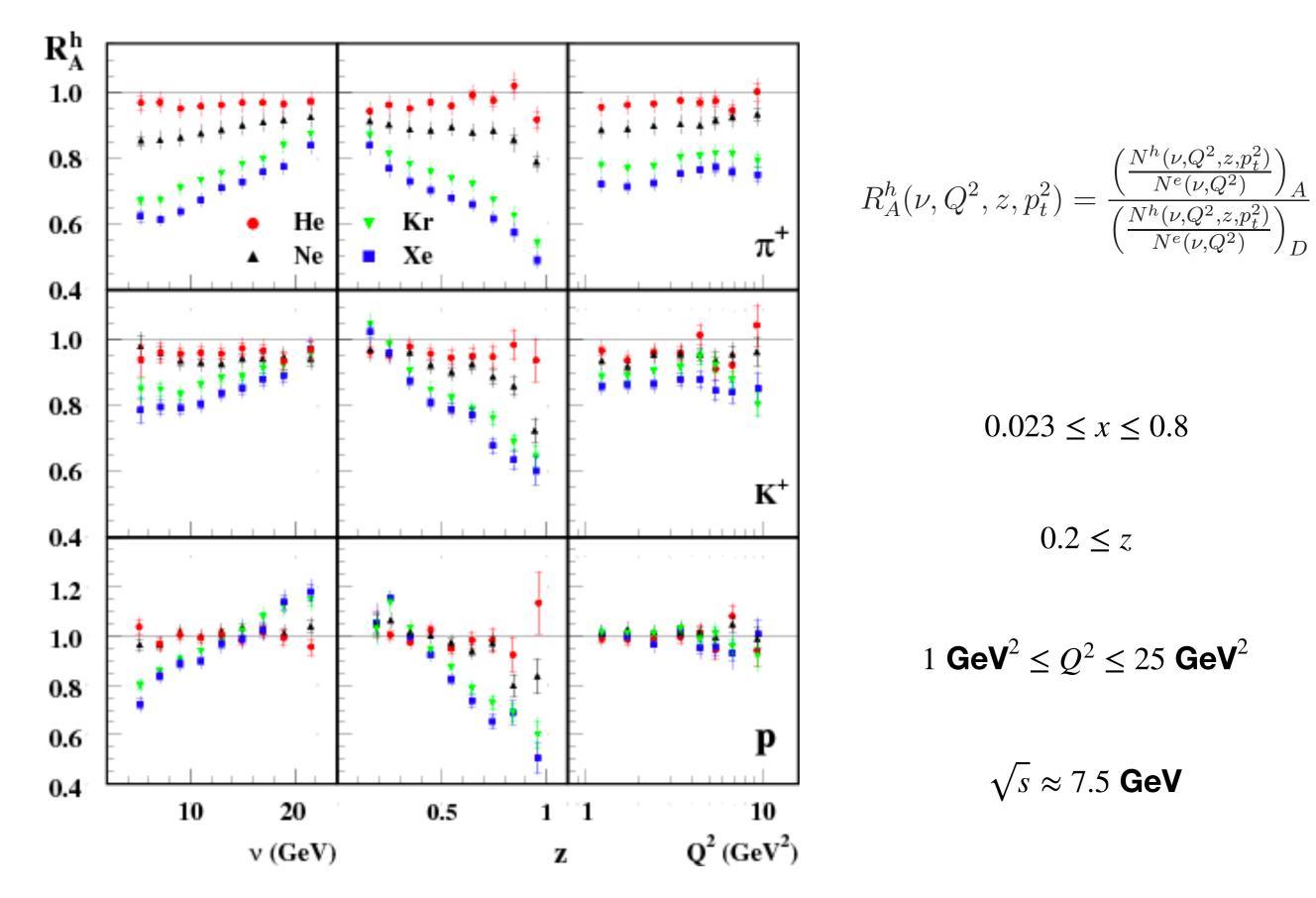
SIDIS in the nuclear medium

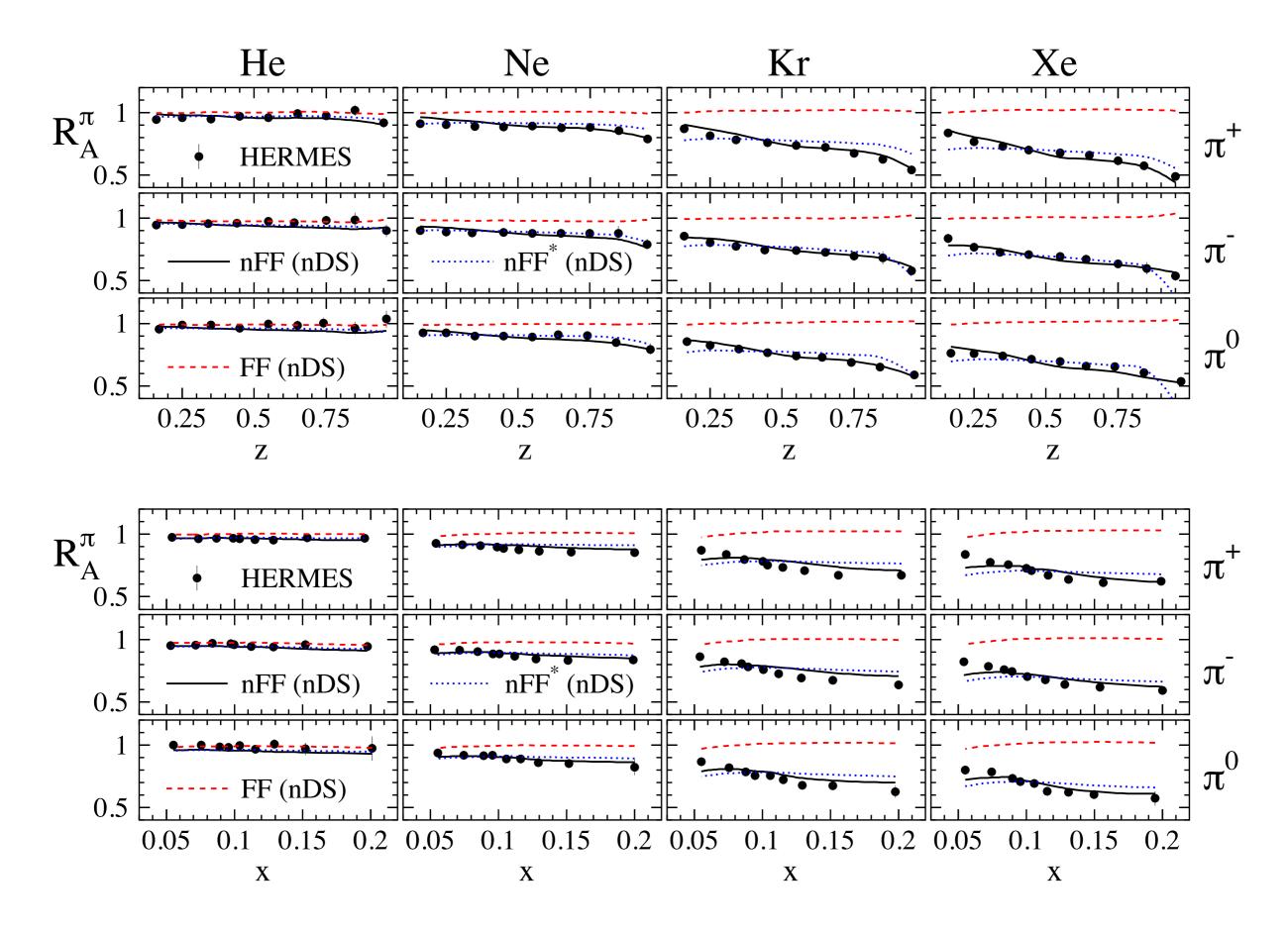
P. Zurita

05/10/20

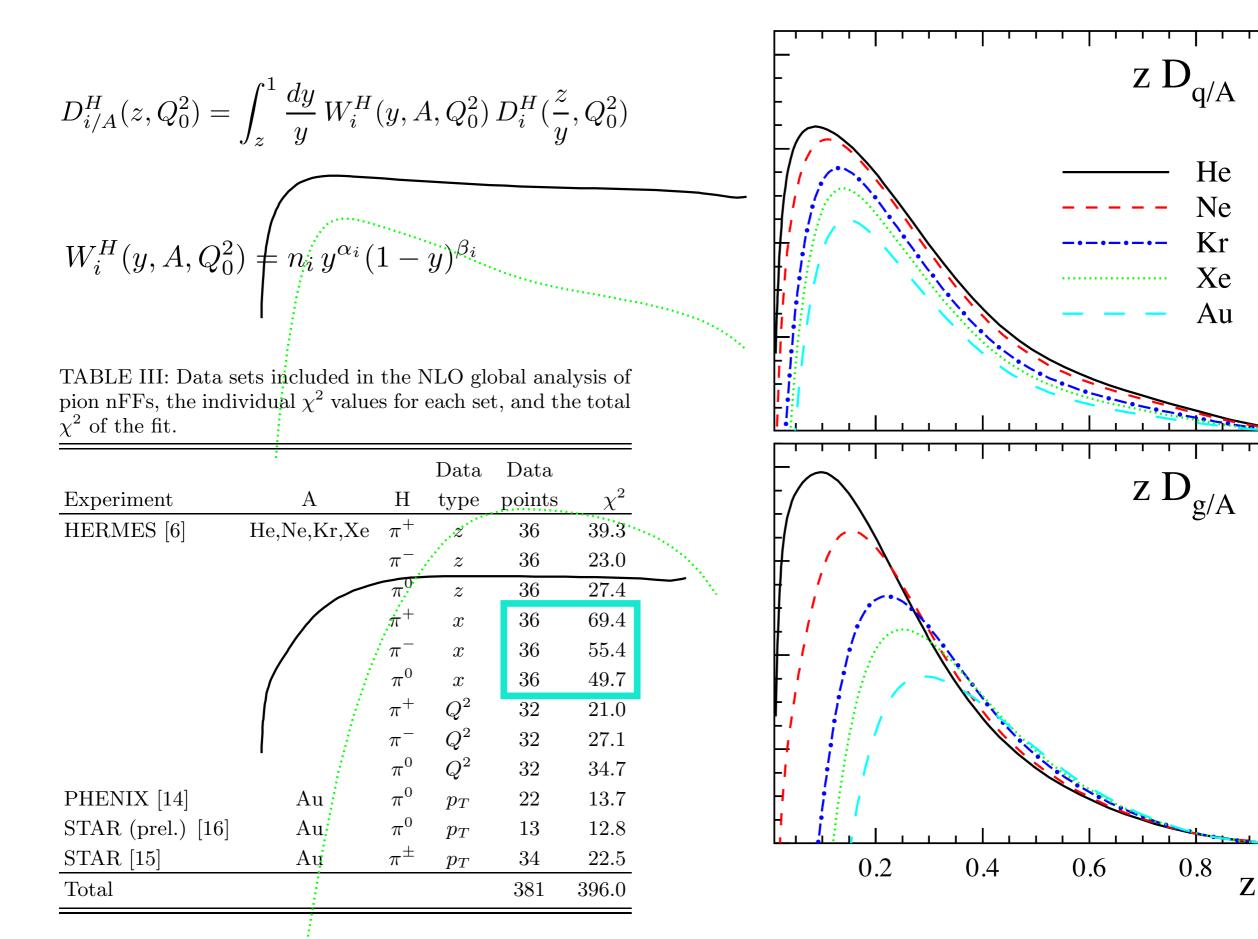
Only one experiment: HERMES



Phys.Rev.D 81 (2010) 054001



Phys.Rev.D 81 (2010) 054001, Phys.Rev.D 75 (2007) 114010



0.8

0.6

0.4

0.2

0.8

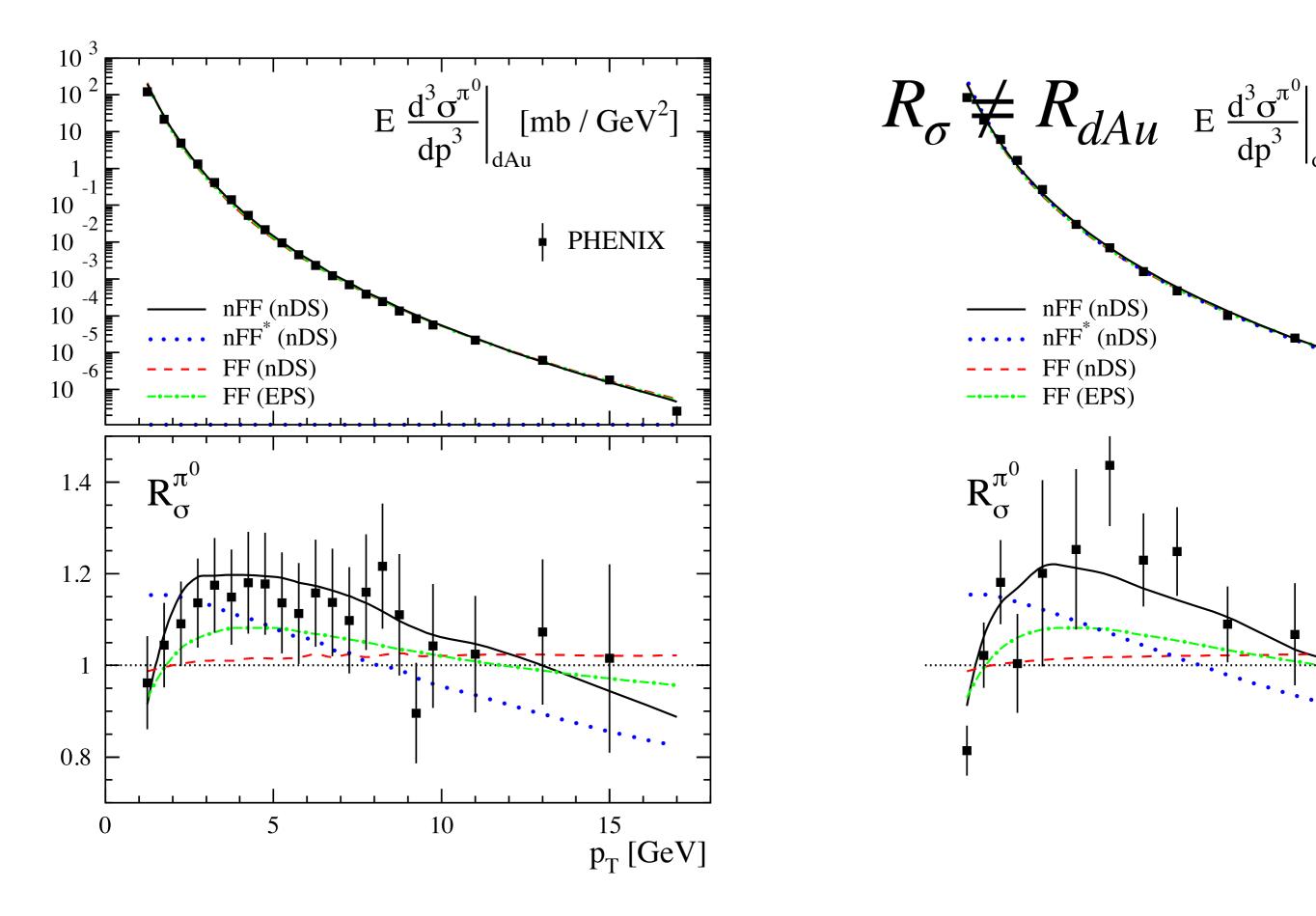
0.6

0.4

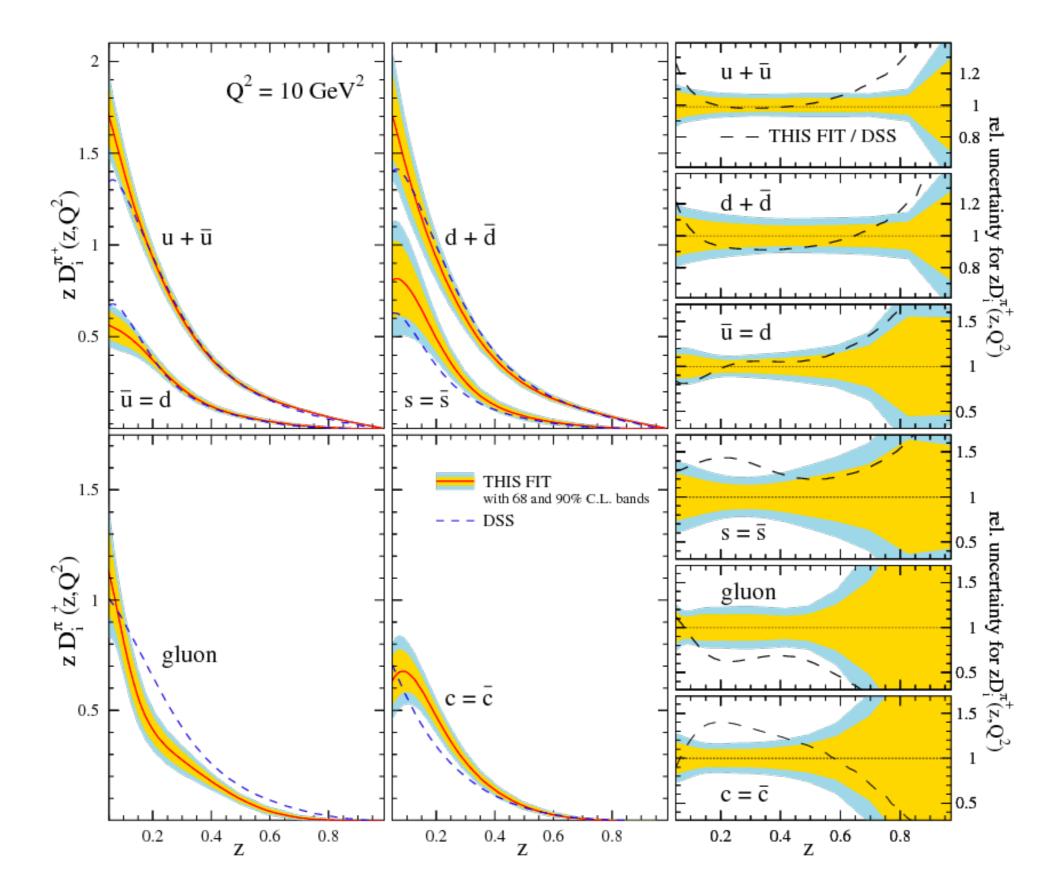
0.2

0

Phys.Rev.D 81 (2010) 054001



DSS2007 is very different from DEHSS2014



Time to update the nFFs: new fit, using xFitter

$$D_i^h(z, Q_0) = N_i x^{\alpha_i} (1-x)^{\beta_i} \left| 1 + \gamma_i (1-x)^{\delta_i} \right| \qquad i = u + \bar{u}, d + \bar{d}, s + \bar{s}, c + \bar{c}, b + \bar{b}, \bar{u}, g$$

$$D_i^h(z, Q_0) \to D_i^h(z, Q_0, A)$$
 $Q_0 = 1 \text{ GeV}, m_c, m_b$

$$N_i \to N_i \Big[1 + N_{1,i} (1 - A^{N_{2,i}}) \Big]$$

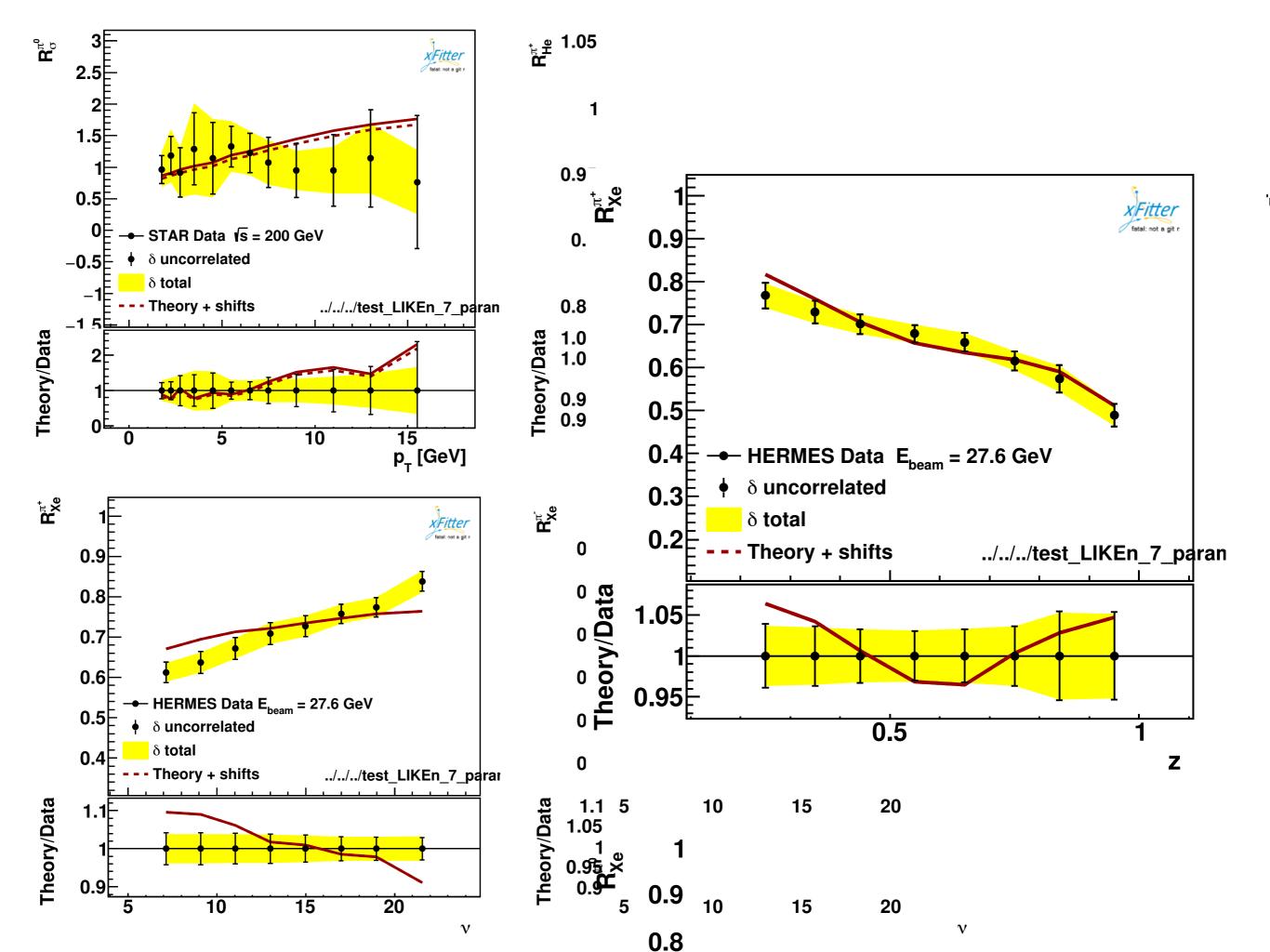
$$p_i \rightarrow p_i + p_{1,i}(1 - A^{p_{2,i}})$$
 $i = \beta, \gamma, \delta$

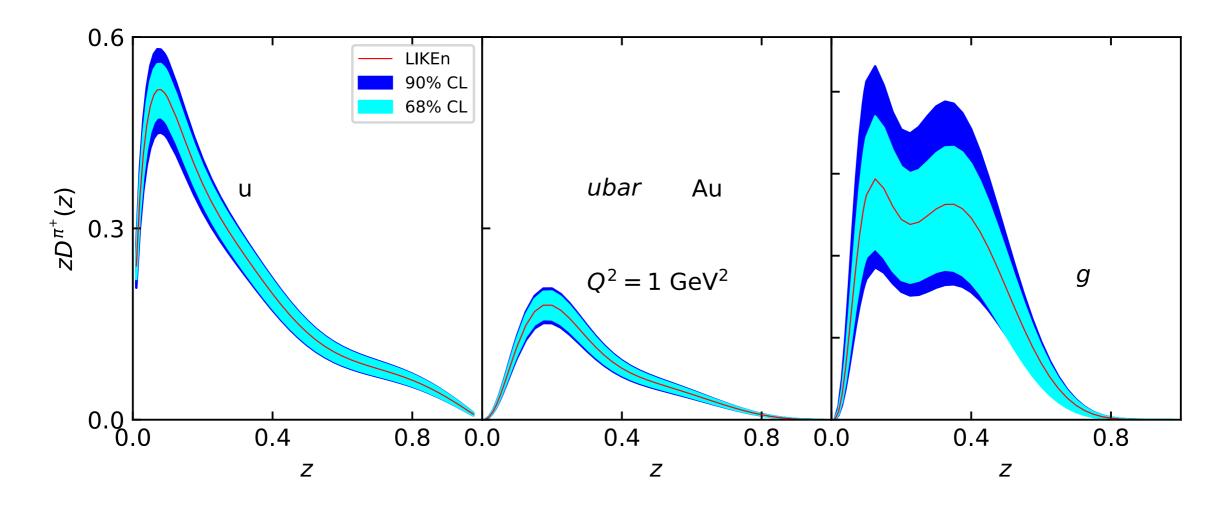
N _{g,1}	N _{g,2}	β _{g,1}	β _{g,2}	γg,1	γg,2	δ _{g,1}	δ _{g,2}	Nq,1	N _{q,2}	β _{q,1}	β _{q,2}	γ q,1	γ q,2	δ _{q,1}	δ _{q,2}
free	free	free	N _{g,2}	free	N _{g,2}	free	N _{g,2}	free	free	$\beta_{g,1}$	N _{q,2}	γg,1	N _{q,2}	δ _{g,1}	N _{q,2}

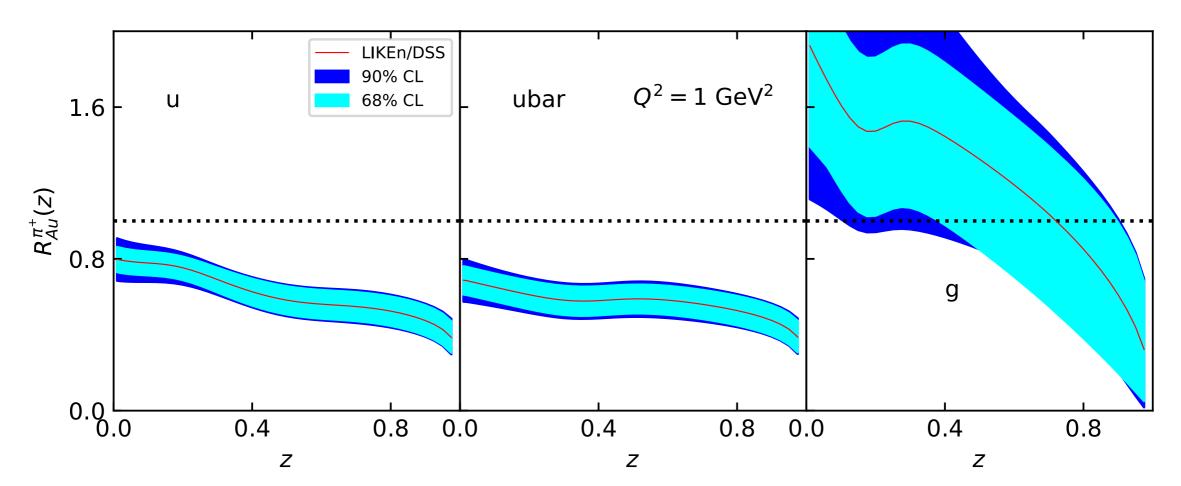
Experiment	Н	Dependence	N° points	X ²	diff	
	π+		36	30.89	-8.41	
	π-	z	36	14.64	-8.36	
	π ⁰		36	28.03	+0.63	
	π+		36	47.76	-21.64	
HERMES	π-	x	36	46.57	-8.83	
	π ⁰		36	35.84	-13.86	
	π+		32	6.89	-14.11	
	π-	Q ²	32	9.17	-17.93	
	π ⁰		32	16.09	-18.61	
	π ⁰		13	5.77	-7.03	
STAR	π+, π-	рт	30	46.14	+11.44	

not a fit

Older STAR neutral pion data were preliminary PHENIX data not available

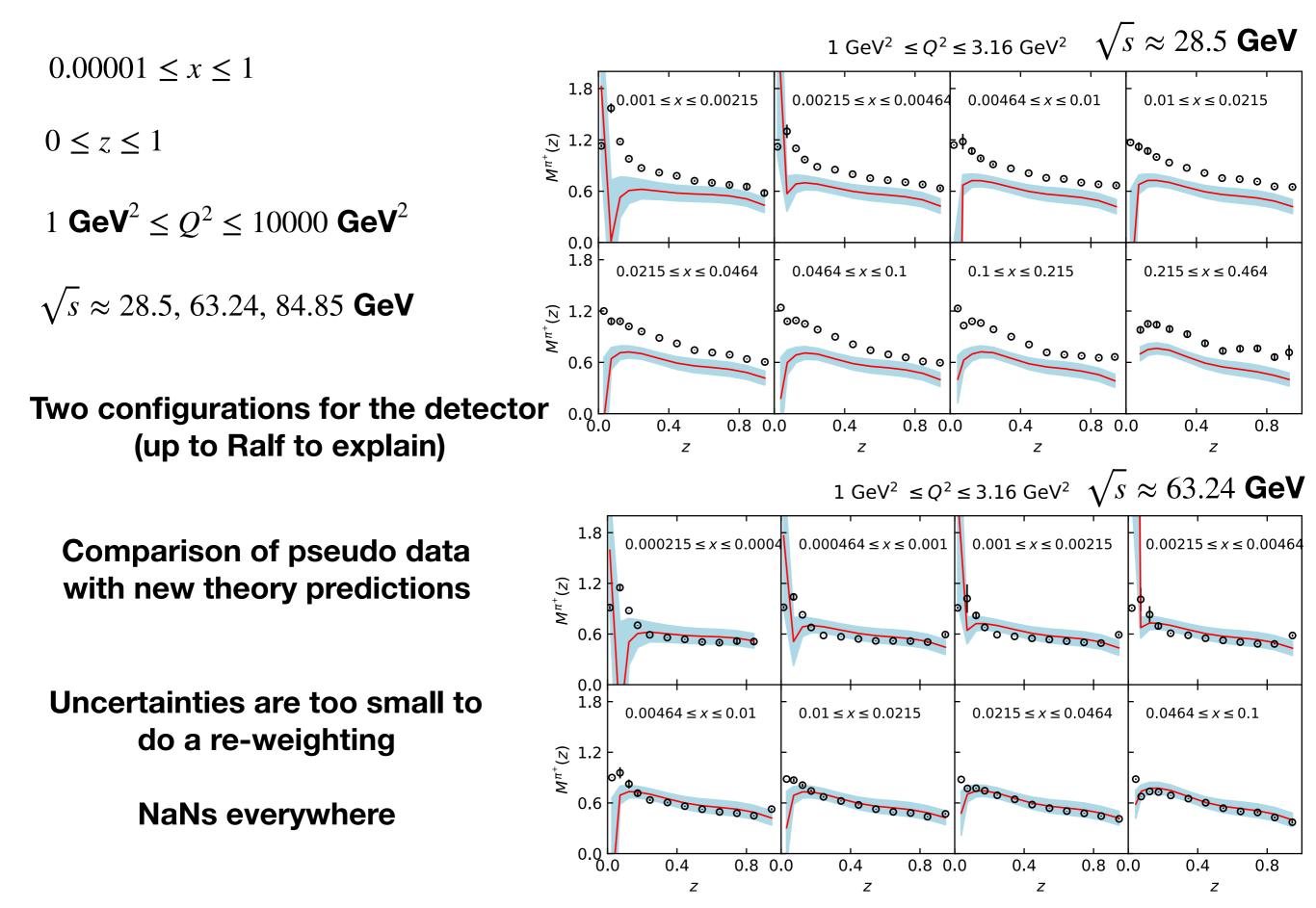






NOW: EIC

I gave the OLD grid for Au to Ralf. He sent back (tons) of pseudo data



I tried creating pseudo data using the old set of nFFs: didn't work

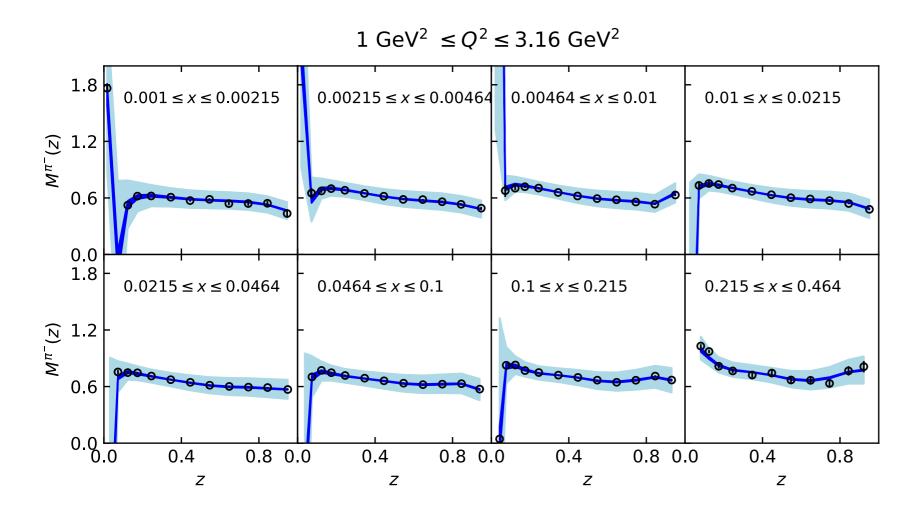
I tried creating pseudo data using the new set of nFFs

$$M_i \to M_i + \delta_i r_i$$
 $\delta_i = \sqrt{\Delta_{syst,i}^2 + \Delta_{stat,i}^2}$

 r_i random number

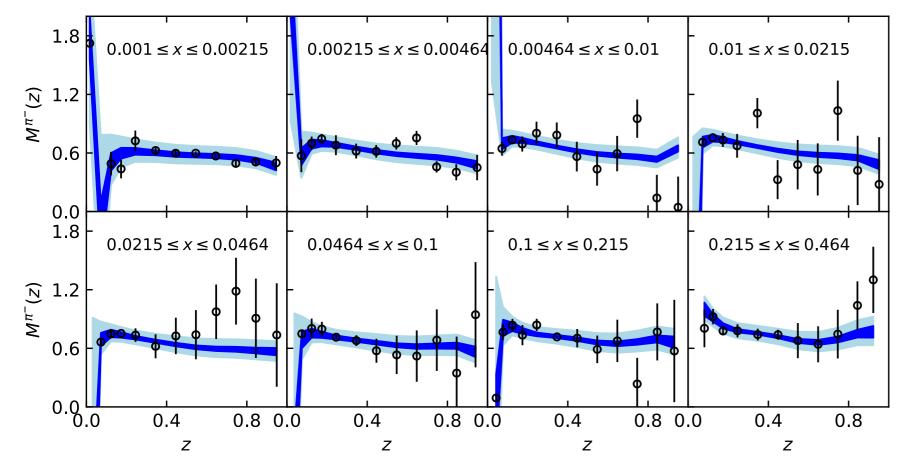
With this I did the re-weighting, considering:

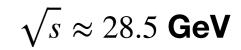
- For some kinematic bins the prediction becomes negative, so those are NOT included in the weights (only in the comparison). Idem with z < 0.2
- The estimated uncertainties are so small that most of the time one obtains one remaining replica (out of 1000)
- Best case scenario 10% of the replicas remain

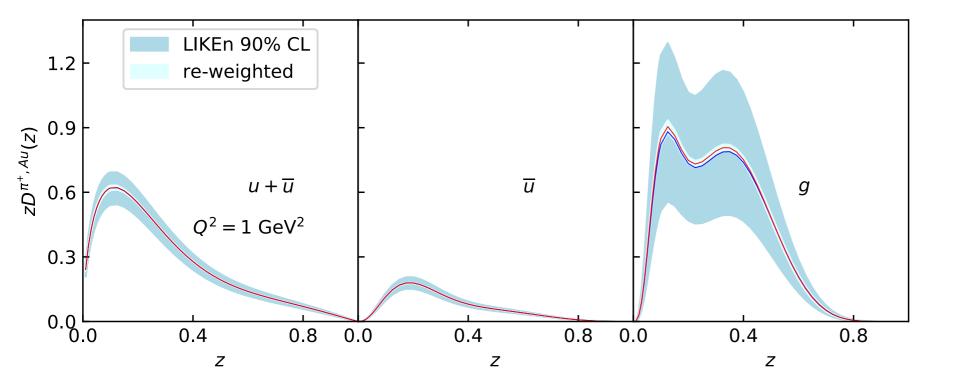


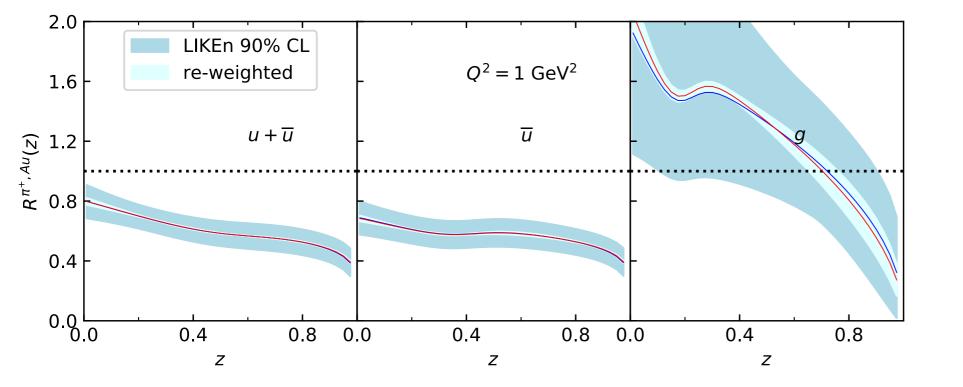
 $\sqrt{s} \approx 28.5 \text{ GeV}$

 $1 \text{ GeV}^2 \le Q^2 \le 3.16 \text{ GeV}^2$

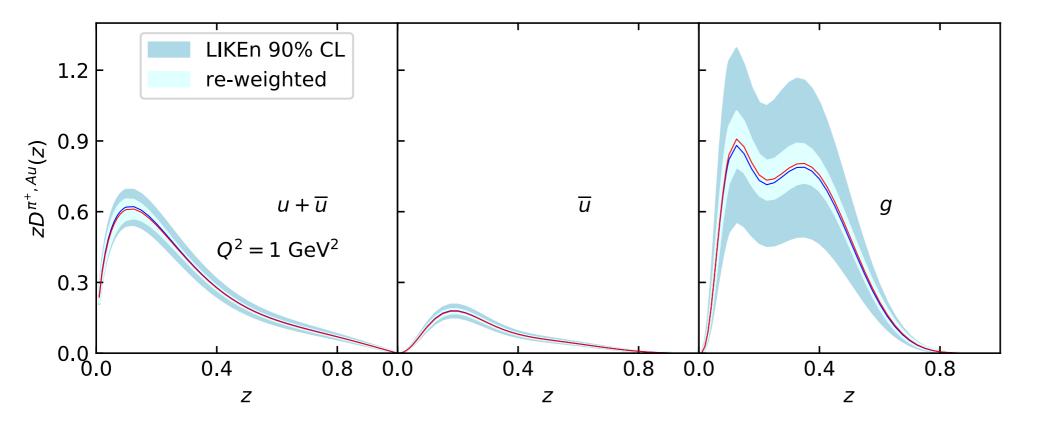


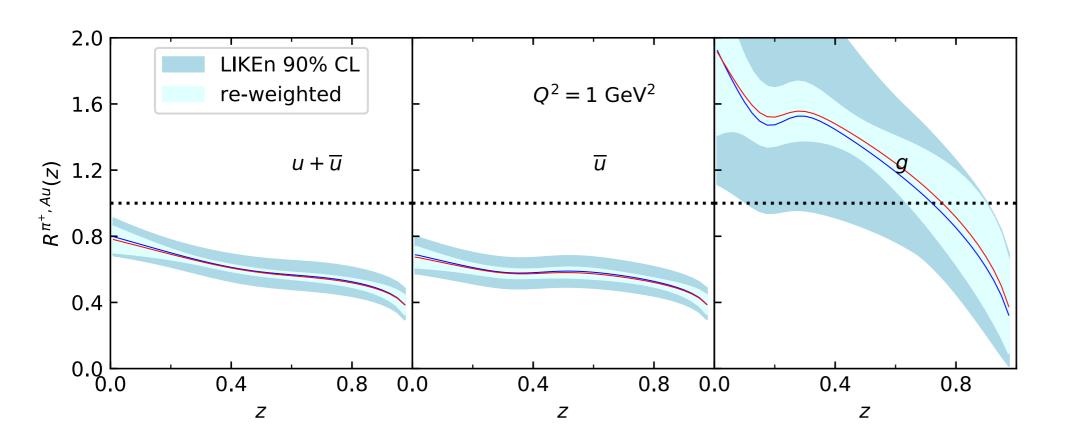


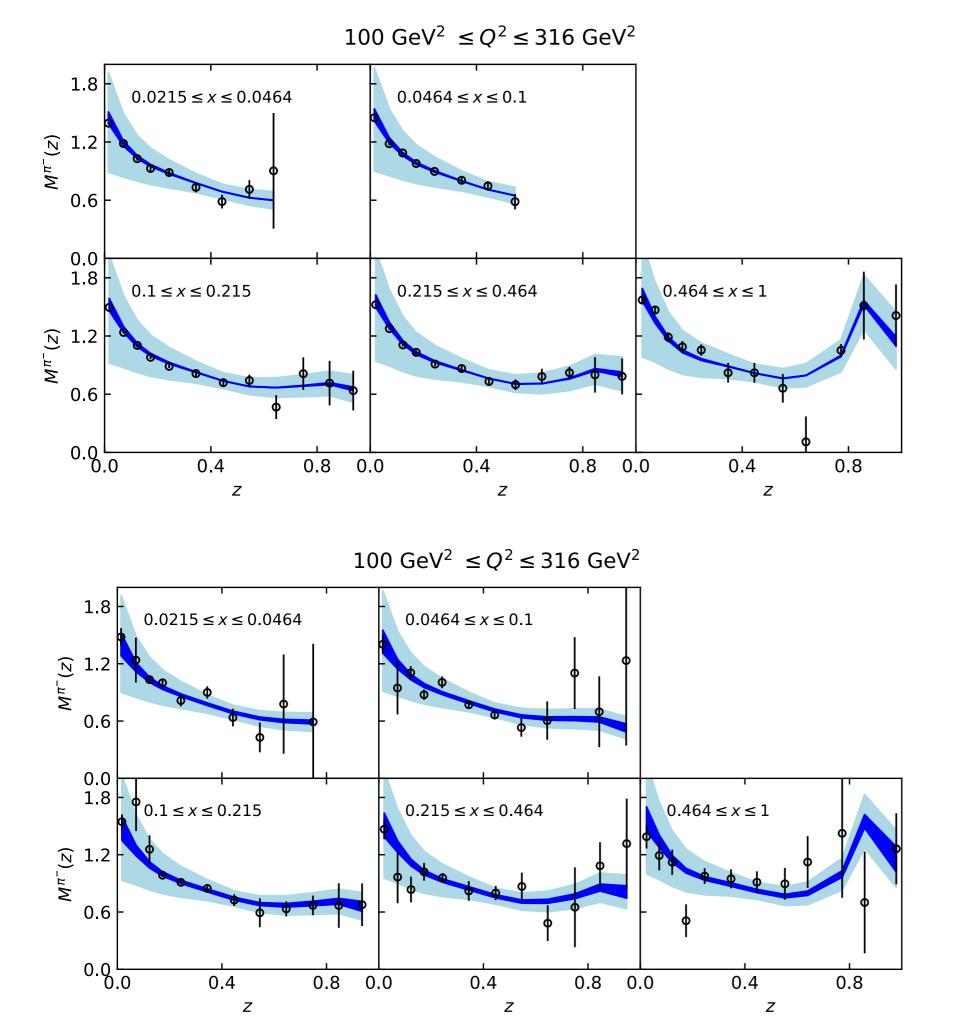




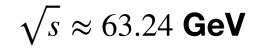


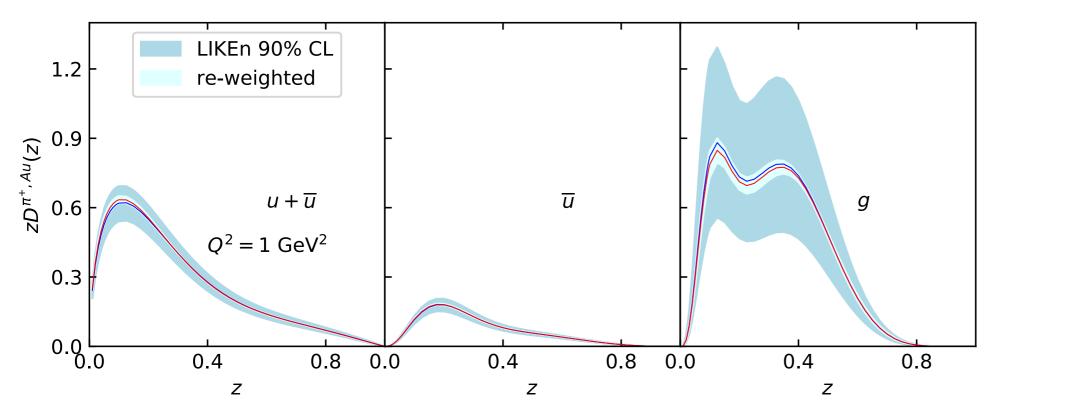


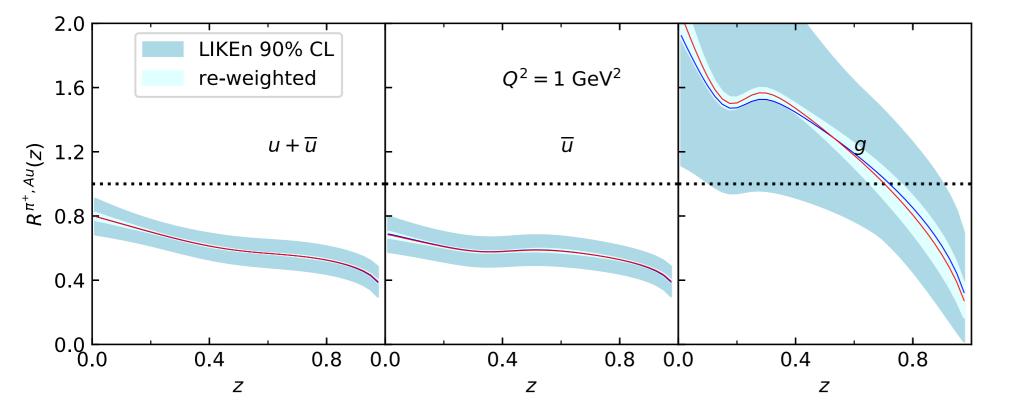




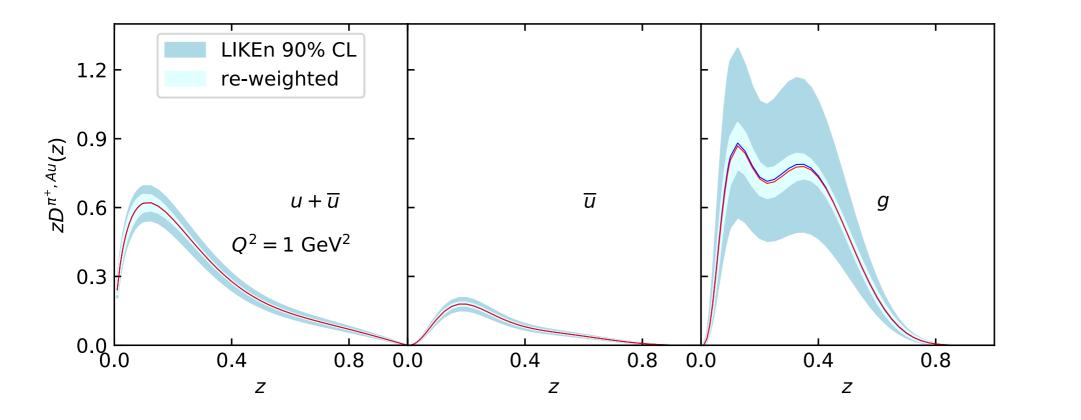
 $\sqrt{s} \approx 63.24 \text{ GeV}$

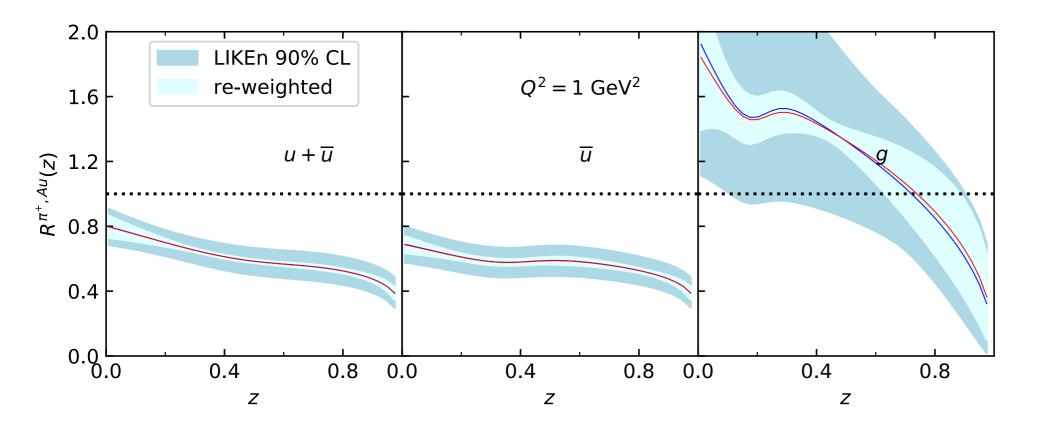




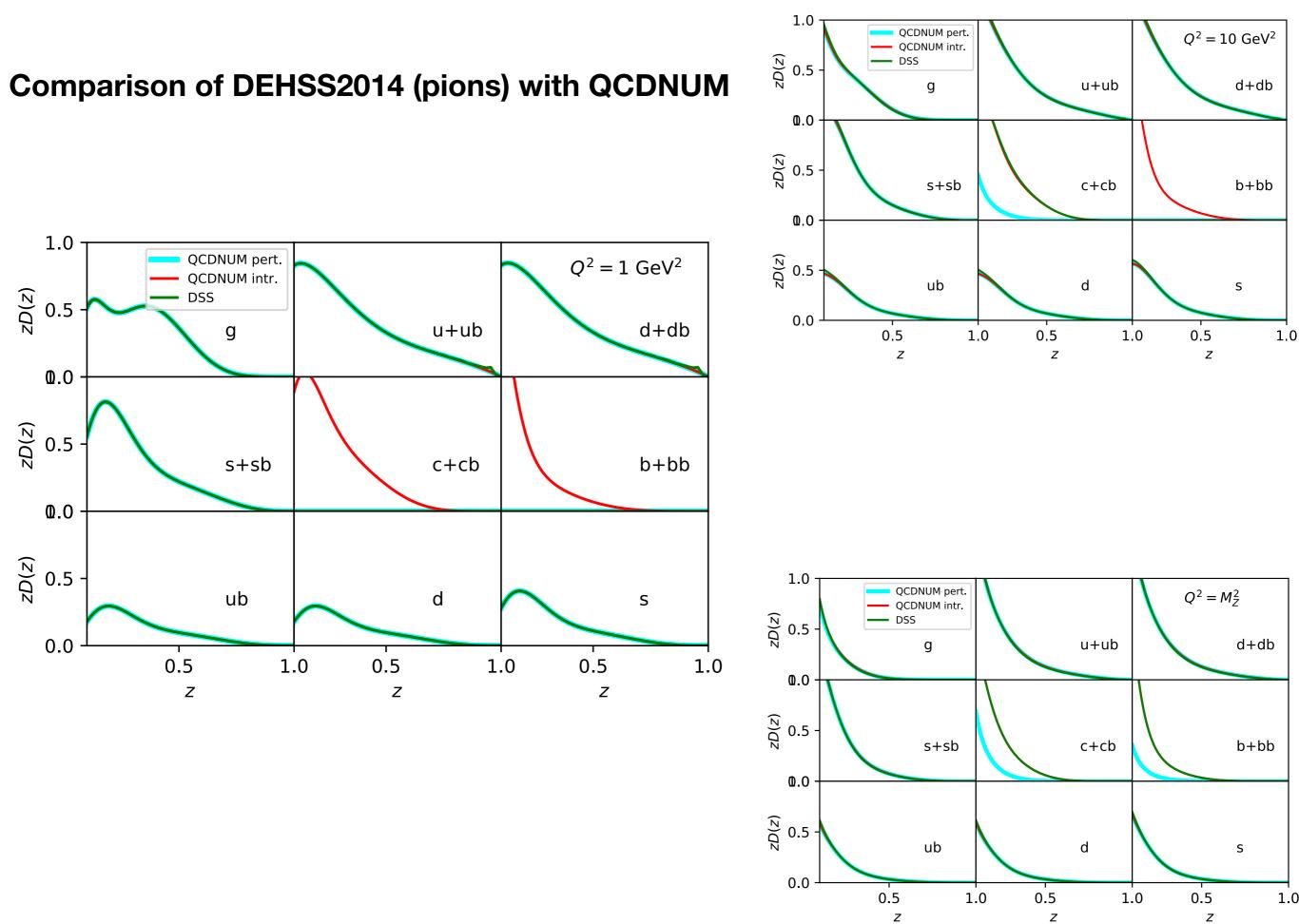


$\sqrt{s} \approx 63.24 \text{ GeV}$





BACK UP



Comparison of DEHSS2017 (kaons) with QCDNUM

