

LAr Depth

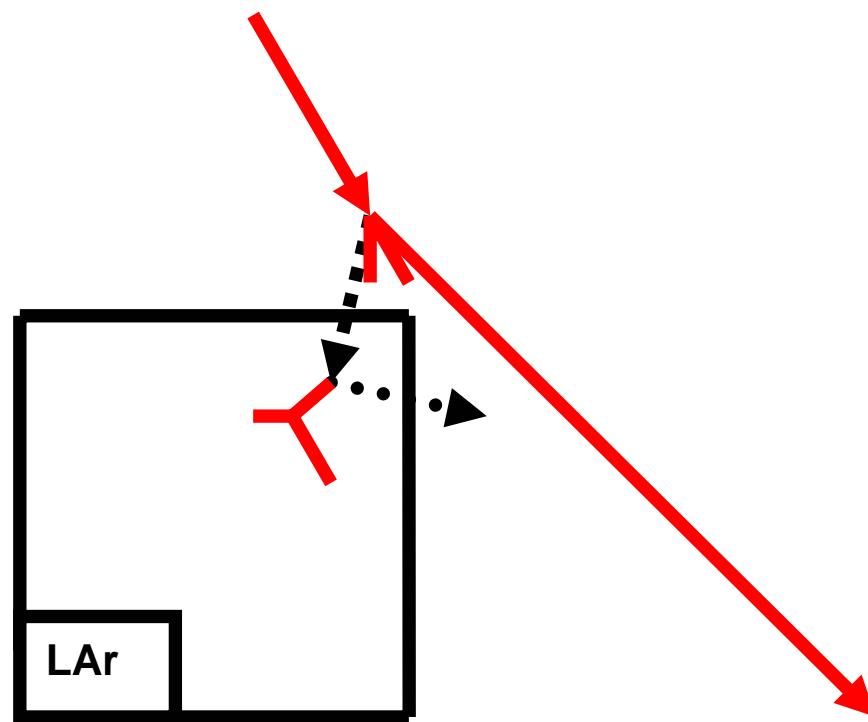
W. Morse - BNL

October 16-17, 2008 BNL

LAr at DUSEL

- Dave Cline LAr talk at April 2008 DUSEL Workshop at Lead, S.D.
- Depth determined by $K_L p \rightarrow K^+ n$ background to $p \rightarrow K^+ \nu$ decay.
- I volunteered to do a “back of the envelope” calculation, ie. not GEANT.
- Another background: K_s regeneration background to $n \rightarrow K_s \nu$ decay.

μ rock $\rightarrow \mu K_L x$; $K_L p \rightarrow K^+ n$ or $K_L \rightarrow K_S$



Depth for LAr Detector

- Assume 100kton cylindrical LAr detector
- June 4, 2008 White Paper
- Miland posted it on the web
- 3.5km w.e. gives 0.1 event background in 3 yr live-time for $p \rightarrow K^+ \nu$.
- A. Bueno et al., hep-ph/0701101(2007).
- 3km w.e. gives 0.1 event background in one yr live-time for $p \rightarrow K^+ \nu$.
- Ray Davis experiment was at 4.5km w.e.

10 24. Cosmic rays

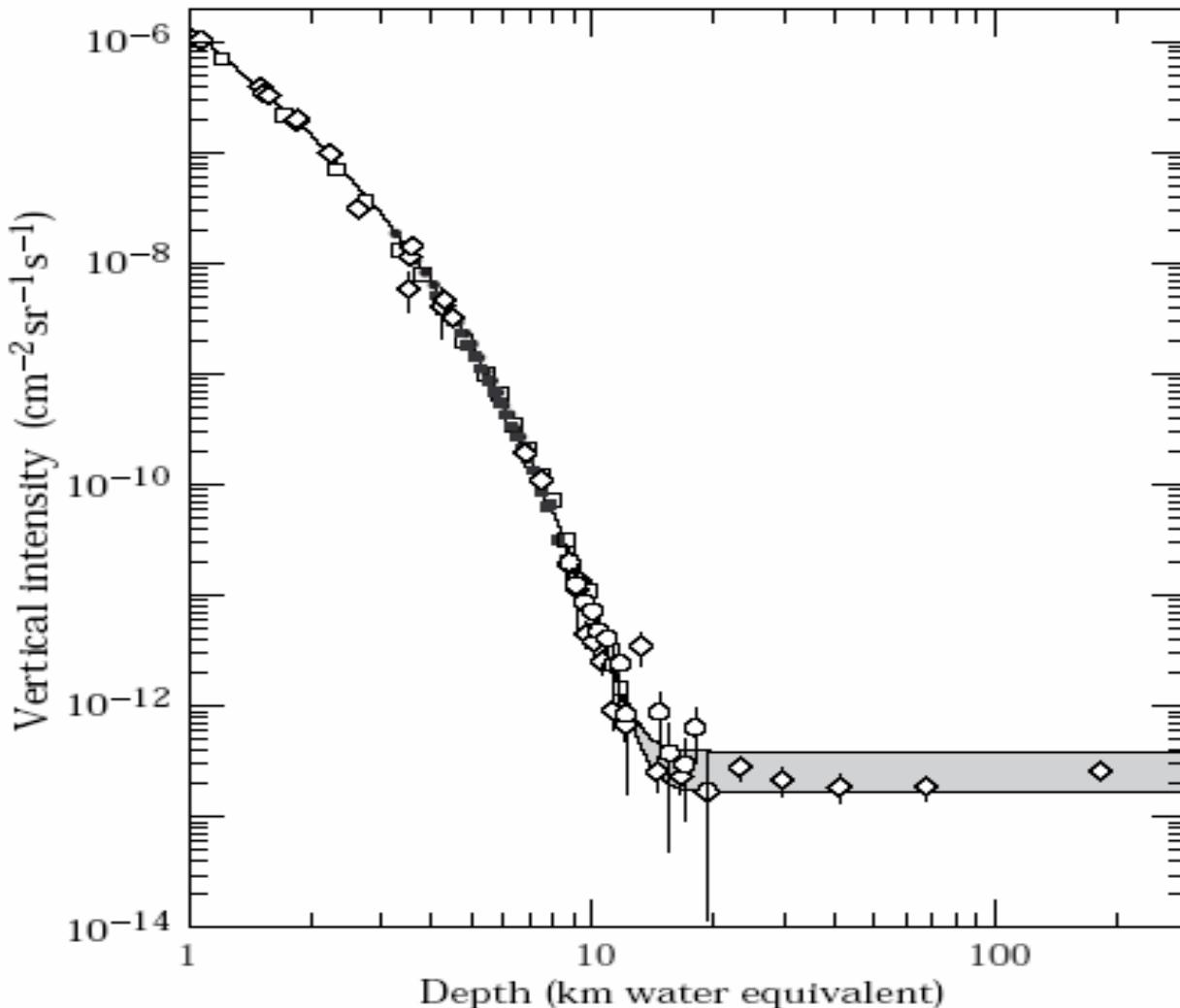
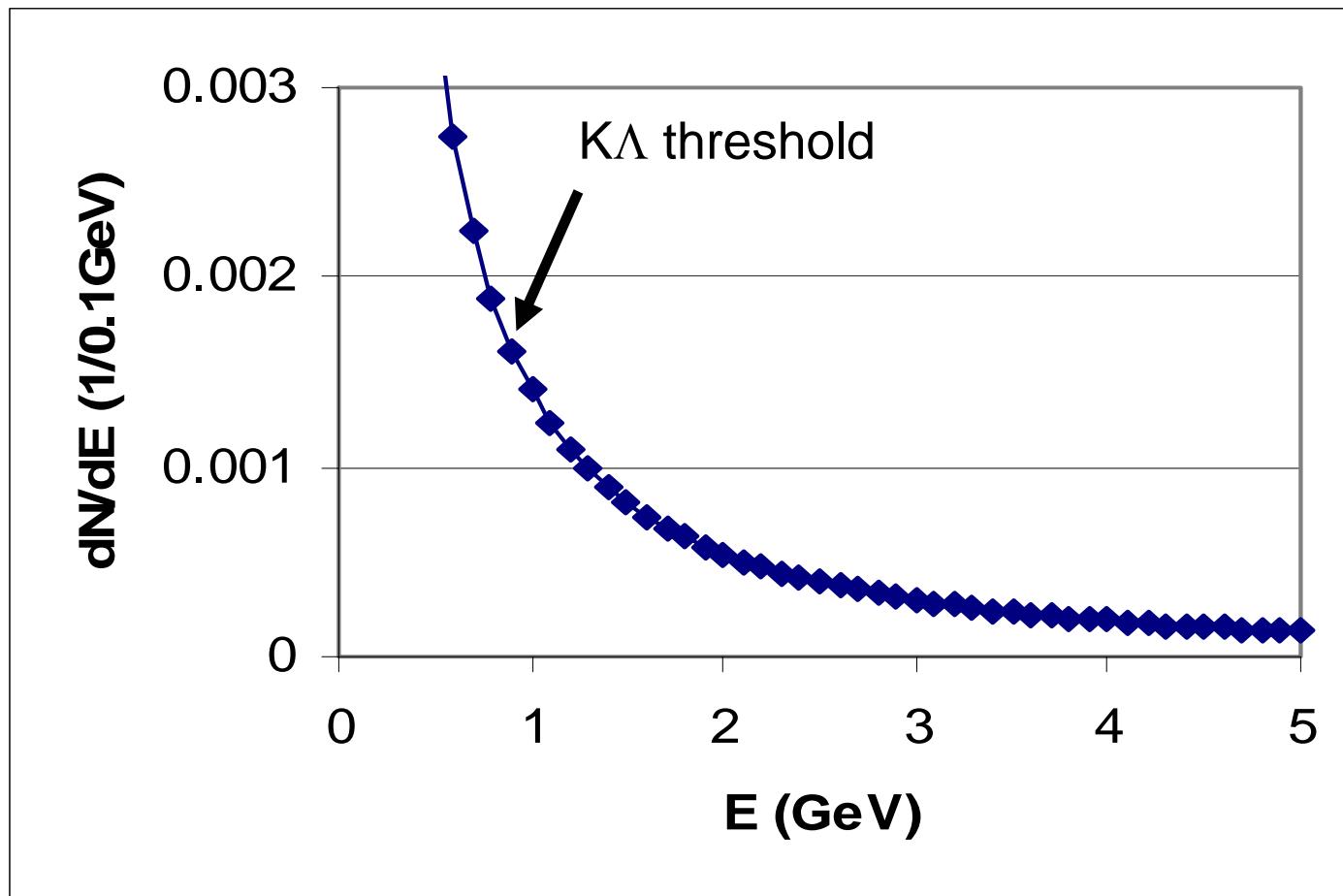


Figure 24.5: Vertical muon intensity vs depth (1 km.w.e. = 10^5 g cm $^{-2}$ of standard rock). The experimental data are from: \diamond : the compilations of Crouch [45], \square : Baksan [46], \circ : LVD [47], \bullet : MACRO [48], \blacksquare : Frejus [49]. The shaded area at large depths represents neutrino-induced muons of energy above 2 GeV. The upper line is for horizontal neutrino-induced muons, the lower one for vertically upward muons.

$$dN/dE \approx 2\alpha/\pi E (\ln 1.4 E_\mu/E - 1/2)$$



$E_{\mu} > 1 \text{ TeV}$ (2.5km w.e.)

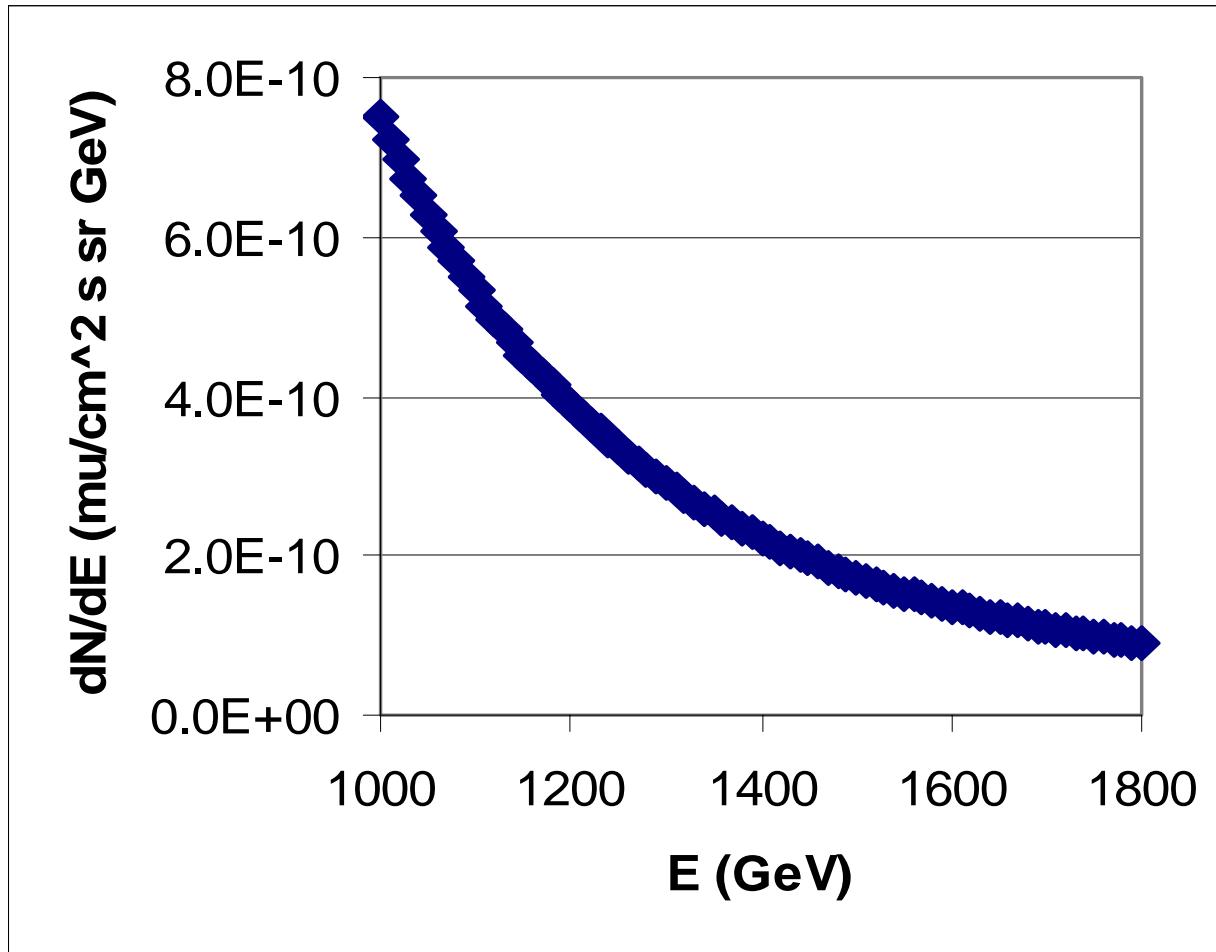
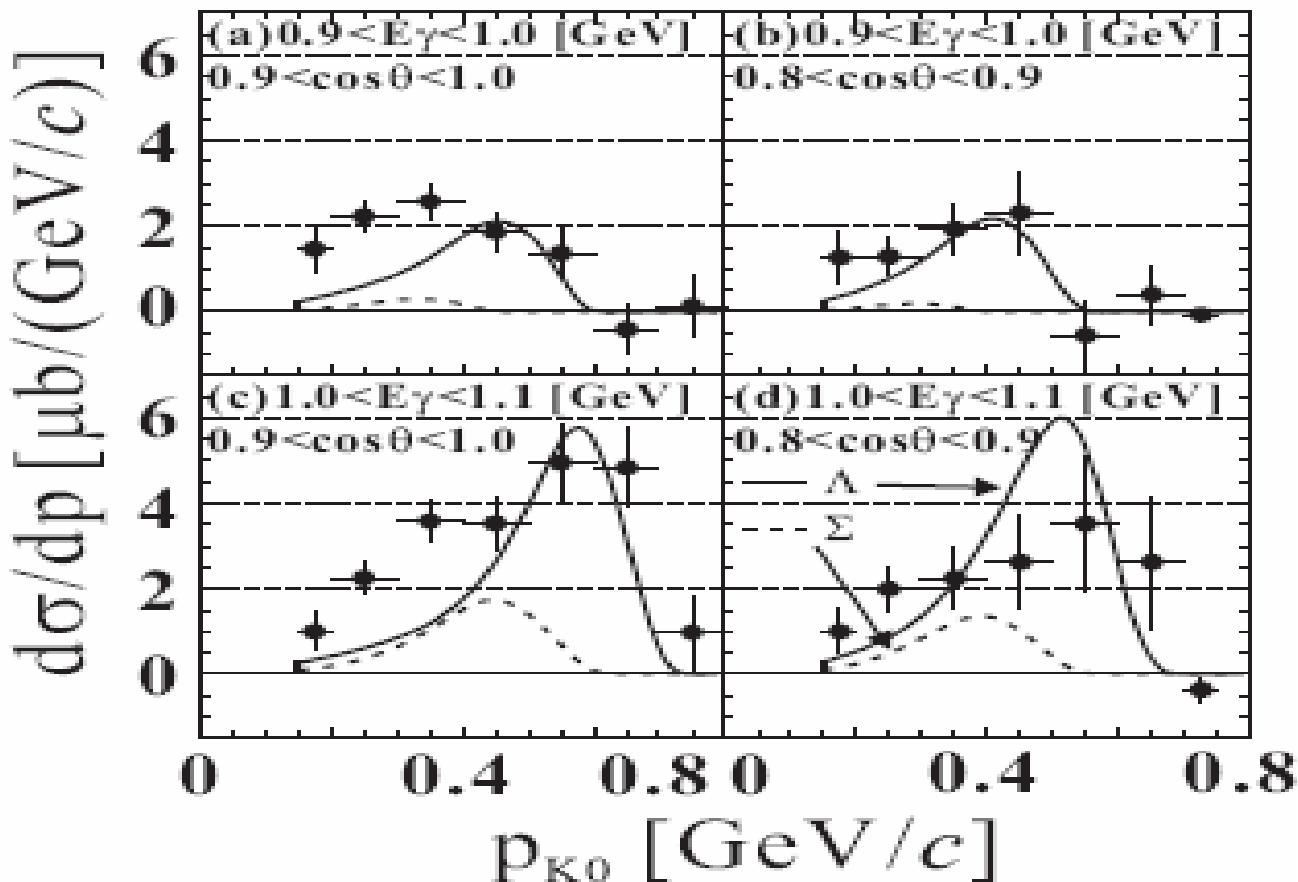
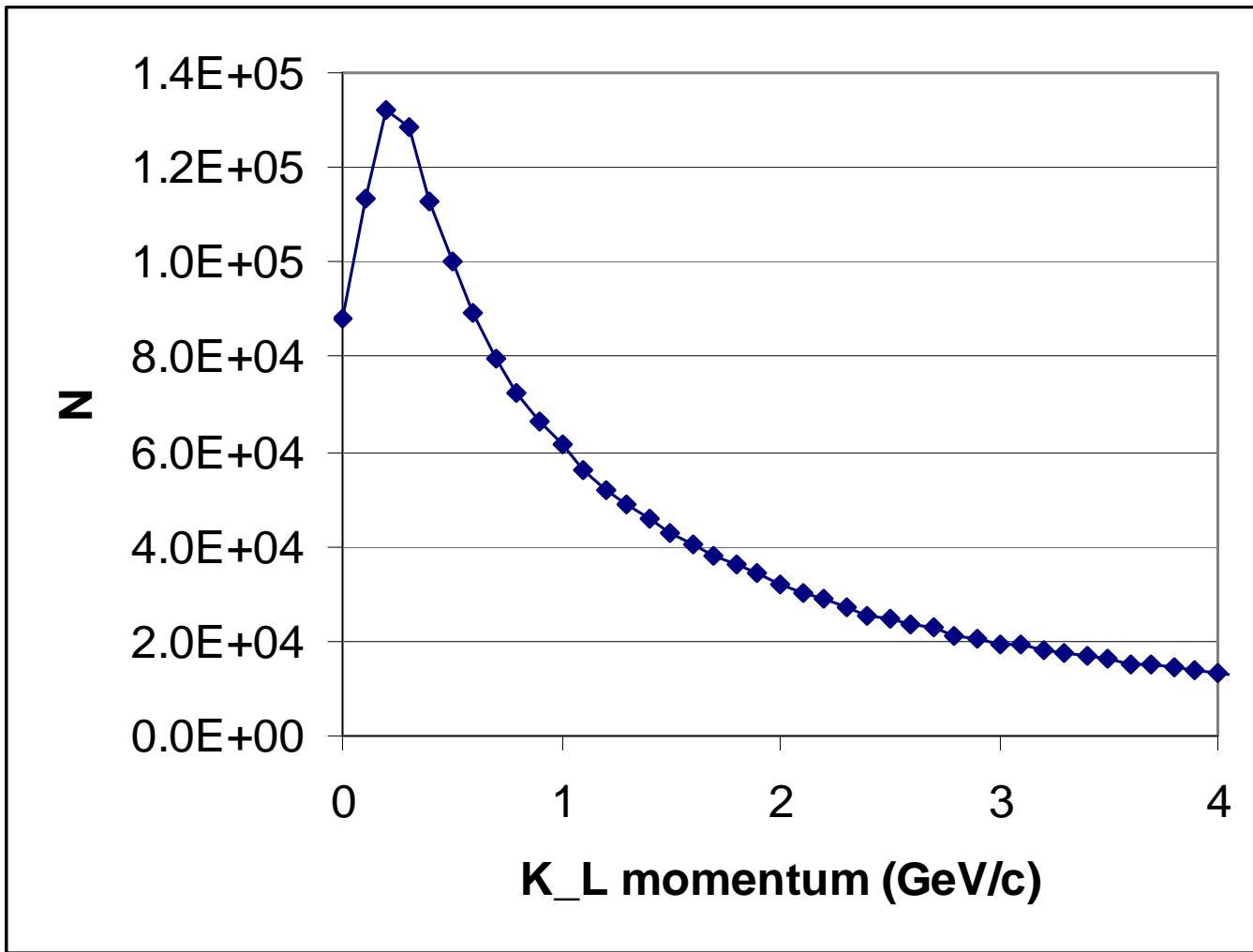


Photo-production of Neutral Kaons on C. Nucl-ex/067022(2006)



K_L from muons



K_L interaction cross-section

Particle	K^-	K_s	K_L	K^+
$c\tau$	3.71 m	2.7 cm	15.3 m	3.71 m
state	$\underline{u} \ s$	$d \ \underline{s} - \underline{d} \ s$	$d \ \underline{s} + \underline{d} \ s$	$u \ \underline{s}$

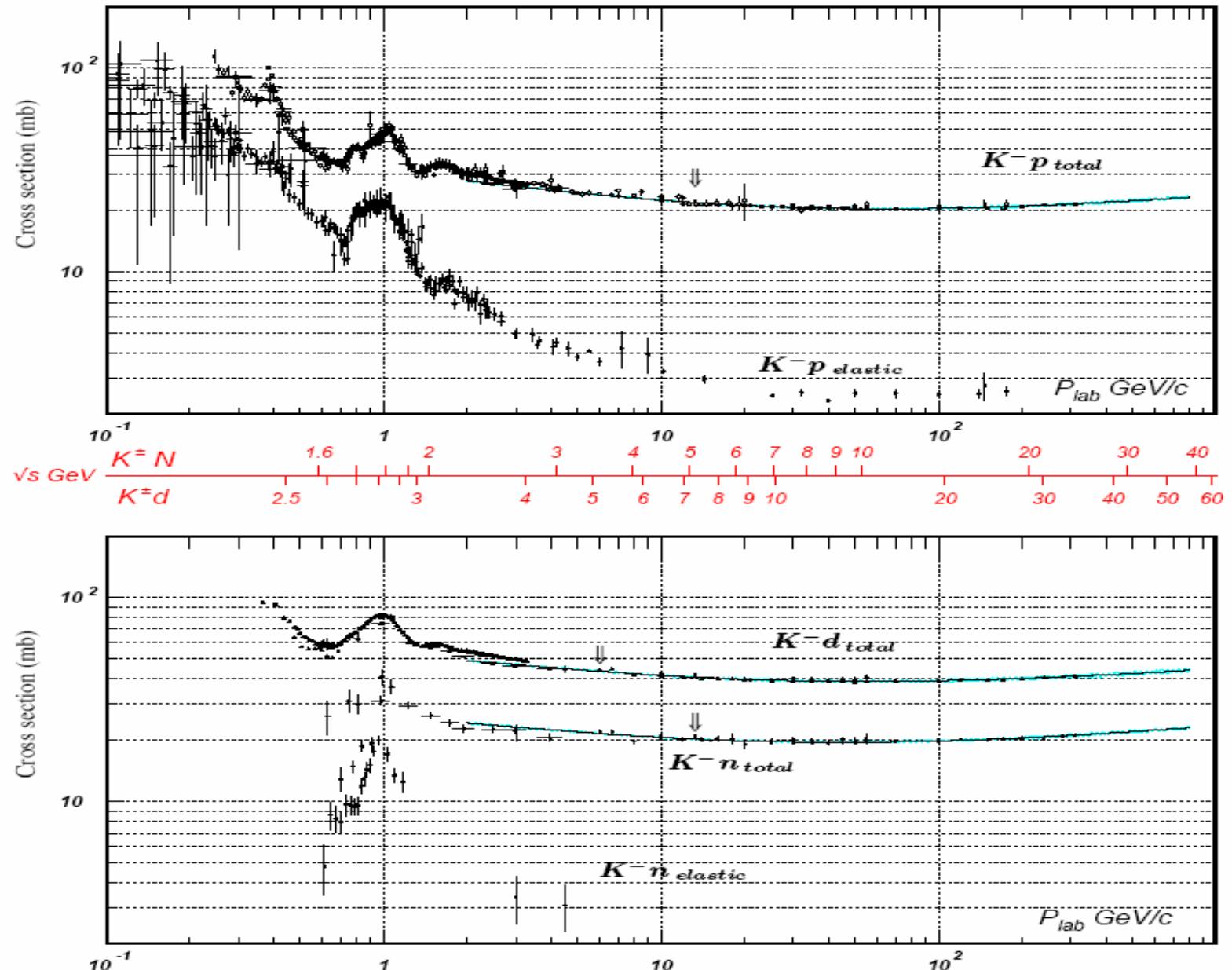


Figure 40.14: Total and elastic cross sections for $K^- p$ and $K^- d$ (total only), and $K^- n$ collisions as a function of laboratory beam momentum and total center-of-mass energy. Corresponding computer-readable data files may be found at <http://pdg.lbl.gov/current/xsect/>. (Courtesy of the COMPAS Group, IHEP, Protvino, August 2005)

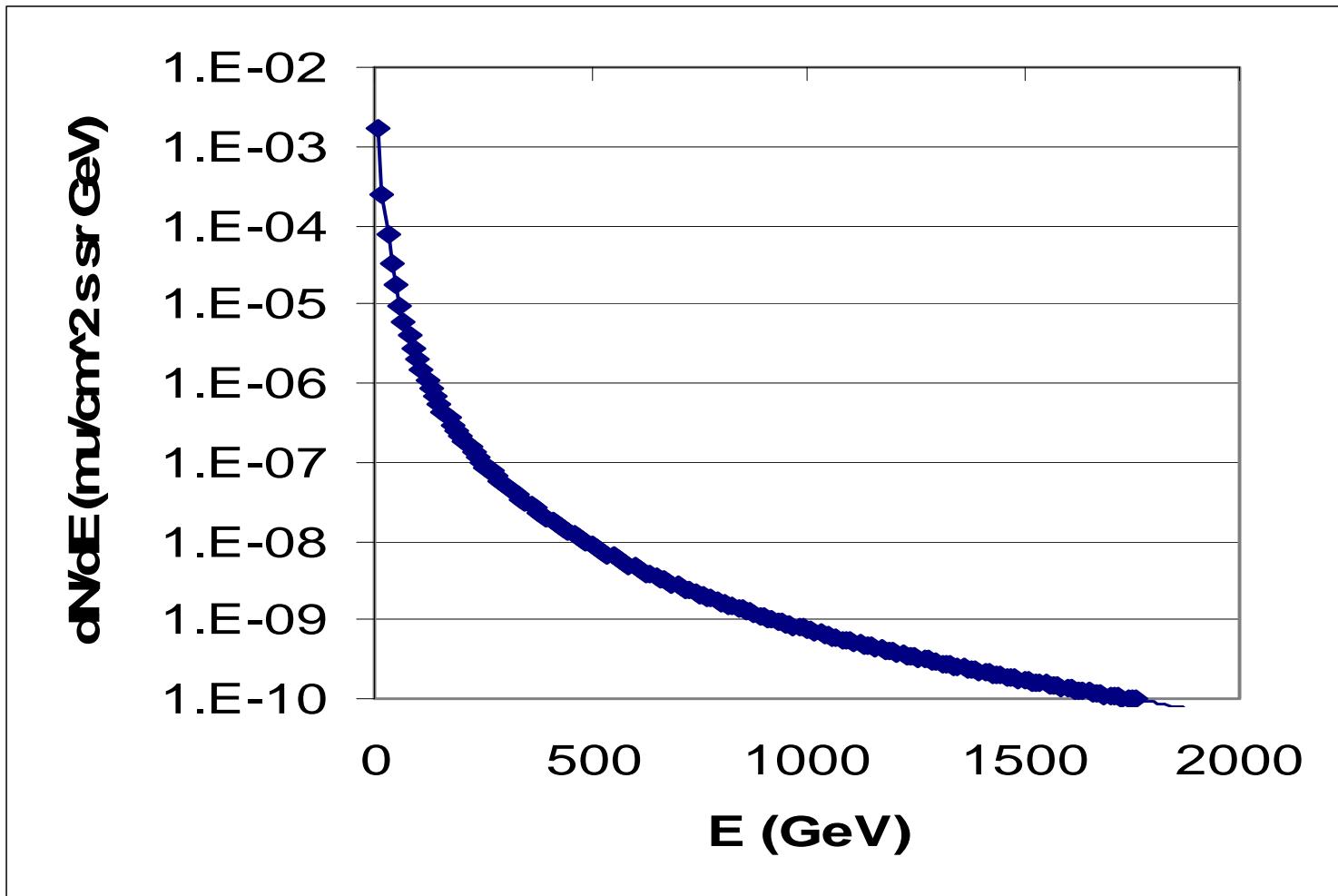
K_L Ar regeneration/ interaction L

- I get 1.1m from the last slide.
- A. Bueno et al. used 0.9m.
- Pretty good agreement.

Conclusions

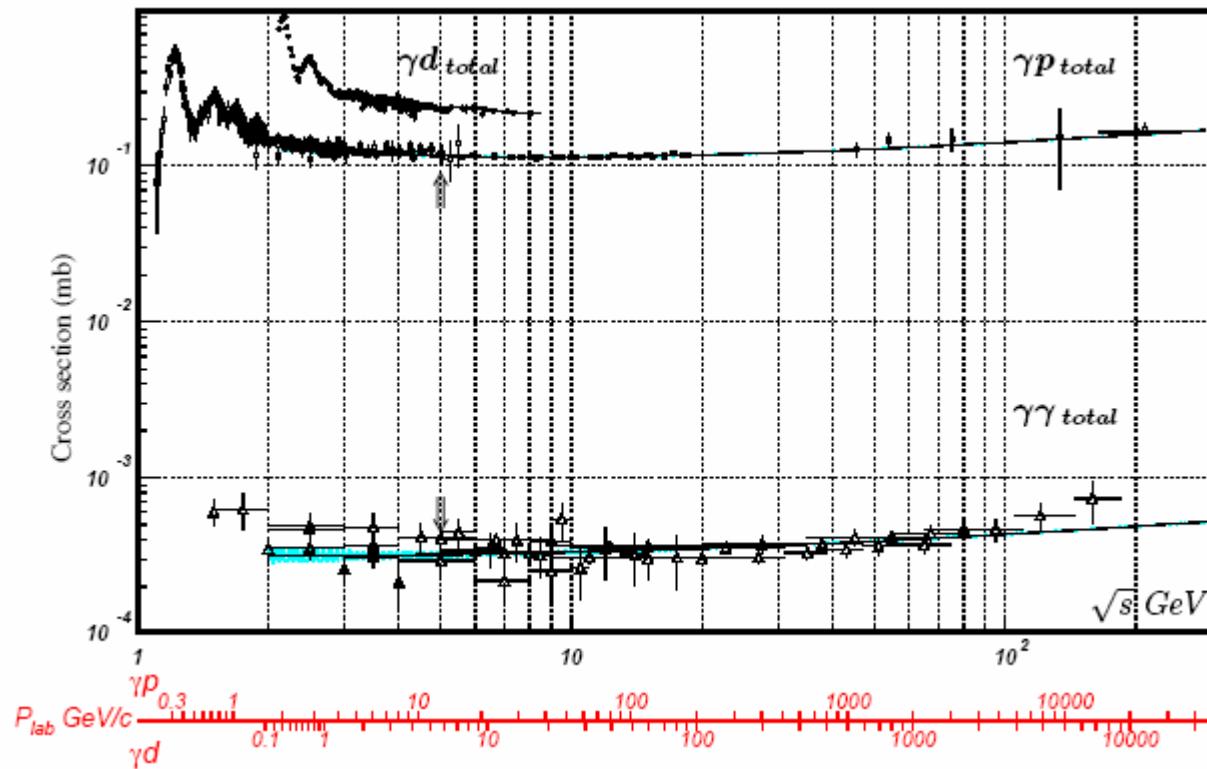
- 3.5km w.e. depth gives:
- 0.1 evt background from $K_L p \rightarrow K^+ n$ in 3 yr live-time for $p \rightarrow K^+ \nu$.
- 0.4 evt background from $K_L \rightarrow K_S$ in 3 yr live-time for $n \rightarrow K_s \nu$.
- A. Bueno et al. also studied sacrificing fiducial volume to reduce background.
- I disagree with this philosophy,
- Ray Davis experiment 4.5km w.e.

Extras



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γp total hadronic cross sections



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A. Bueno et al.

- Section III E. An active very large area cosmic muon veto.
- 100% efficiency counters of area $10K\ m^2$.

A. Bueno et al. – no muon veto

Depth	Fiducial Mass/Cut
0.5km w.e.	66 kton/ 6.6m
1km w.e.	71 kton/ 5.45m
3km w.e.	90 kton/ 1.8m