# Higgs' invisible branching fraction at the LHC

#### BNL HEP Seminar April 28, 2016



University of Pittsburgh



Tae Min Hong



#### Thanks



**Announcements** 

Scholars

> 2015-2016 Program

2014-2015 Scholars

#### Institutional

I thank U.S. ATLAS for the funds to be a U.S.A. Scholar at BNL.

#### Individual

I thank Michael and Kétévi for their collaboration on projects.

I thank Philip for talk content, preparation, and discussion.

More thanked in aux. slides.

#### Announcement of 2015-2016 U.S. ATLAS Scholars

Each year, the U.S. ATLAS Analysis Support Centers (ASCs) at Argonne National Laboratory, Brookhaven National Laboratory, and Lawrence Berkeley National Laboratory host distinguished research scholars in a program to foster physics collaboration and enhance U.S. contributions to the ATLAS physics program. The U.S. ATLAS Scholars play an important role in the life of the ASCs as outside guests who bring new ideas and projects. The selected scholars have proposed to use the resources available at an ASC to make major contributions to ATLAS in the areas of physics analysis but also play a leading role in organizing ASC and U.S. ATLAS analysis activities.

The U.S. ATLAS Scholars for 2015-2016 are listed below with their selected Analysis Support Center and their proposed projects.

- Tae Min Hong (University of Pittsburgh) will work at the BNL ASC on searches for Higgs bosons that are produced via vector boson fusion and decay invisibly. In addition, he will work on physics simulation in support of the design of the gFEX board for the Phase 1 upgrade to the Level 1 calorimeter trigger.
- and 5 other scholars listed on the webpage <u>http://www.usatlas.bnl.gov/programoffice/scholars.php</u>





### Warning → measurement



"We apologize... for having <u>no idea</u> what is  $[m_H]$ ... For these reasons we <u>do not want to encourage big experimental searches</u> for the Higgs..."

Nucl. Phys. B 106 (1976) 292



### Is large $H_{125} \rightarrow invisible$ motivated?



Let's review the Higgs width







to indirectly limit invisible

Pittsburgh

Search for this to directly limit invisible

# Hiding in the couplings



#### Measurements



# Hiding in the couplings





#### Summary table



### **Nevent yields**



#### My rough estimate of events in the peak (don't read plots!)



# From $N \rightarrow \mu$





#### From $N \rightarrow \mu \rightarrow \kappa$ Example of VBF *WW*





**K**<sub>V</sub>ector boson

### From $N \rightarrow \mu \rightarrow \kappa \rightarrow B_{inv}$





Summa	ry table	-(µ	$\mu, Z$	$\gamma, t\bar{t}H$
<b>ATLAS</b> 20.3 fb <sup>-1</sup> (8 TeV 4.5-4.7 fb <sup>-1</sup> (7	<sup>/) +</sup> TeV) <u>EPJC 7</u>	6 (20	<u>016) 6,</u>	Fig. 1
YY	Overall ggF VBF WH ZH			
ZZ WW	Overall ggF+ttH VBF+VH Overall ggF VBF VH			
Value			CL	Ndof
$\Delta \mu/\mu = 0.$	127		1σ	1
$B_{inv} < 0.7$	127		68%	1
$B_{inv} < 0.2$	95%	1		
$B_{inv} < 0.8$	508 my e	est.	95%	8

(0.48)

95%

 $B_{inv} < 0.49$ 

You can approx.  $\mu \rightarrow B_{inv}$  by hand

- (1) Write down  $1\sigma$
- <sup>(2)</sup> Rewrite 1 $\sigma$  as 68% limit for 1 d.o.f.
- <sup>③</sup> Rescale 2 as 95% limit for 1 d.o.f.
- ④ Rescale by  $\sqrt{p(\chi^2, dof)} \approx 2$  for 8 d.o.f.
- **5** Compare with ATLAS full fit result

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### From $N \rightarrow \mu \rightarrow \kappa \rightarrow B_{inv}$

#### 8 parameter fit with "invisible"



#### Why do we say "BSM"?

Technically includes "undetected." Examples are  $H \rightarrow soft$  jets. If assume zero, then = invisible.





Using couplings alone, indirectly limit  $B_{BSM} < 0.34$  ATLAS+CMS (0.35)

### Closer look at $\mu_{ggF}$ v. $\mu_{VBF}$



WW and yy are similar size circles in  $\mu$ 



W strongest input, establishes VBF production

#### Closer look at KVector boson V. KFermion



#### "The most precise determination of $\kappa_V$ and $\kappa_F$ is obtained from **WW**."



"At first look, the LHC is unlikely to ever get to 6% sensitivity." PRD 86 (2012) 095001, received 18 July 2012

### ATLAS v. CMS



#### Compare the expected contour. Here I overlay them by hand.



Difference in the upper bound in  $\kappa_{\underline{Fermion}}$  is due to VBF WW.

#### Physics of VBF $H \rightarrow WW^* \rightarrow ev\mu v$



LHC is a vector boson collider



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### **VBF** topology





#### $\mathsf{VBF} \to H \to WW^* \to ev\mu v$





#### **Transverse mass**



#### You can see with your eyes a broad peak for signal



Difficulty estimating  $t\bar{t}$ , leading systematic

## Variables for VBF $H \rightarrow WW^* \rightarrow ev\mu v$ Hong



Three groups of variables

#### Higgs decay

- *m*<sub>T</sub> ≈ *m*<sub>H</sub>
- $m_{\ell\ell}$  small
- $\Delta \phi_{\ell\ell}$  small

#### VBF configuration

- *m<sub>jj</sub>* large
- Δy<sub>jj</sub> large
- centrality of  $\ell\ell$

Top quark

- Σ**p**<sub>T</sub>
- $\Sigma m_{\ell j}$  of lep-jet



We tried **O**(1k) variable combinations & matrix element methods

### **Multivariate analysis**



BDT can be thought of as an *S*/*B* grouping of cut-and-count bins



#### Achieved 2-to-1 ratio of S-to-B

# Difficulty estimating $t\bar{t} \rightarrow WbW\bar{b}$



Can't reject b jets where there is no tracker





Only  $N_{b-tag} = 1$  faithful, extrapolation good to 30%

#### ATLAS v. CMS



Comparison of a sensitive eµ bin in Run 1

Not the fit distributions, but representative samples.



Small statistics, room for improvement in Run 2

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### The future

Back-of-the-envelope using the most sensitive  $e\mu$  bin with 25 fb<sup>-1</sup>

#### All contributions scale up by a factor of 2-3 for Run 2



#### Have $3\sigma$ in Run 1, will likely approach $5\sigma$ in Run 2



#### **VBF Higgs established**









### **VBF** history



 $\overline{V}$ 

#### The crux of all these differences is on estimating $Z \rightarrow v\bar{v}$





"If (CMS normalized Z using W similar to ATLAS) it would be ... 33%."

### Physics of VBF *H* → *invisible*



Production established by WW\*



Not as many handles, background est. crucial

# Why not ggF?

The problem is invisible

Disappears

- Can only measure recoil
- Decay products conserve momentum

### There is nothing to measure





# Wait, why not ggF + <u>1 jet</u>?



The problem is the  $Z \rightarrow v\bar{v}$ 



Not worth the trouble for ggF

### **VBF** analysis





$Z_{vv}$ estimates	Method	Who	Pro	Con	Precision
$z \sim \sqrt{\frac{v}{\bar{v}}}$	<b>N</b> <sub>MC</sub>	-	WYSIWYG	Jet energy, QCD scale	± 50%
$z \sim \ell_{\ell}$	Nzee	CMS Run 1 µµ only (?)	$Z_{\ell\ell} = Z_{vv}$	Low stats (~20 evts)	± 40%
$w \sim \sqrt{\frac{\ell}{\bar{v}}}$	N <sub>MC</sub> • R <sub>Wev</sub>	ATLAS R1, CMS Run 2	Large stats (~600 evts)	$W_{\ell v} \neq Z_{vv}$	± 10%

Reduce even more in Run 2? Maybe  $\gamma_{\ell\ell} \sim Z_{\ell\ell}$ 

#### **Dark matter interpretation**





Once you have a collider, every problem starts to look like a particle. look for invisible dark matter

#### **Dark matter interpretation**





# Many direct detection results

LHC measurement









Direct detection exclude to  $m_{\chi} \sim 5$  GeV of  $\sigma \sim$  ab-zb range

### LHC overlay





#### <u>Reinterpret</u> VBF → *invisible*?





Mono-jet search for DM



If DM is fermiophobic, mono-jet not sensitive

on  $\sigma_{any}$ , e.g.,

- Fermiophobic scalars coupling to electroweak bosons, 1604.07975 (yesterday)
- Electrowk SUSY sector with compressed sleptons, PRD 87 (2013) 035029, 91 (2015) 055025

#### Tell us your favorite model!

### The future is now



#### Collisions started this week, ATLAS collected O(1) pb<sup>-1</sup>



#### Hope to have $\mathcal{O}(10)$ fb<sup>-1</sup> of data this calendar year

### The future







### My future







# **Auxiliary material**

### **More thanks**



#### I thank Elliot for his clear slides on VBF invisible at a PITT workshop. https://indico.cern.ch/event/460471/contributions/1132574/ attachments/1199947/1745600/HInvis\_PittPacc.pdf

I thank George for the SUSY chat. It turns out I knew about the compressed scenario for my job talk & forgot it until our chat!

I thank Ben for slides feedback. Especially on the details of VBF *invisible*.

I thank Alex for the talk rehearsal. Especially on the prospects for VBF  $H \rightarrow WW^*$ .

I thank Josh for discussions. Especially on going from  $\Gamma_{inv}$  to  $\sigma_{DM}$ .



E. Lipeles U Pennsylvania



G. Redlinger



B. Carlson U Pittsburgh



A. Tuna Harvard U



J. Kunkle U Maryland

### **Abstract for this seminar**



https://indico.bnl.gov/conferenceDisplay.py?confld=1765

Does the Higgs have a large invisible branching fraction? Two approaches are presented. The first is an indirect constraint of the invisible branching fraction using precision Higgs couplings measurements. The second is a direct search of invisible decays. In particular, I will discuss in detail two of ATLAS's results: the  $H \rightarrow WW$  in VBF, which is one of the strongest inputs for the couplings and the evidence for VBF Higgs production, and the  $H \rightarrow invisible$  in VBF, which gives the strongest direct limit. Comparisons with CMS's results are made.

### No idea on Higgs mass



Nucl. Phys. B 106 (1976) 292

"We should perhaps finish with an apology and a caution. We apologize to the experimentalists for having no idea what is the mass of the Higgs boson, unlike the case with charm, and for not being sure of its couplings to other particles, except that they are probably all very small. For these reasons we do not want to encourage big experimental searches for the Higgs boson, but we feel that people performing experiments vulnerable to the Higgs boson should know how it may turn up."

#### **Guidelines for the next 2 slides**





#### **CMS overlay**



#### CMS-HIG-13-023



#### **ATLAS overlay**

ATLAS-HIGG-2013-13





### **VBF** *WW*\* comparison



PRD 92 (2015) 012006 v. JHEP 01 (2014) 096

#### ATLAS

	Summary		Composition of $N_{\rm bkg}$								
			$N_{ m sig}$	gnal		Nt	top	$N_{ m r}$	nisid		
Channel	$N_{\rm obs}$	$N_{ m bkg}$	$N_{\rm ggF}$	$N_{\rm VBF}$	$N_{WW}$	$N_t$	$N_{t\bar{t}}$	$N_{Wj}$	$N_{jj}$	$N_{VV}$	N <sub>DY</sub>
(a) 8 TeV data san	nple										
$n_i \ge 2$ , VBF	130	$99\pm9$	$7.7\pm2.6$	$21\pm3$	$11 \pm 3.5$	$5.5\pm0.7$	$29\pm5$	$4.7\pm1.4$	$2.8 \pm 1.0$	$4.4\pm0.9$	$38\pm7$
$e\mu$ bin 1	37	$36\pm4$	$3.3\pm1.2$	$4.9\pm0.5$	$5.0\pm1.5$	$3.0\pm0.6$	$15.6 \pm 2.6$	$3.2\pm1.0$	$2.3\pm0.8$	$2.3\pm0.7$	$3.6\pm1.5$
$e\mu$ bin 2	14	$6.5\pm1.3$	$1.4\pm0.5$	$4.9\pm0.5$	$1.7\pm0.7$	$0.3\pm0.4$	$2.0 \pm 1.0$	$0.4\pm0.1$	$0.3\pm0.1$	$0.7\pm0.2$	$0.6\pm0.2$
$e\mu$ bin 3	6	$1.2 \pm 0.3$	$0.4 \pm 0.3$	3.8 ± 0.7	$0.3 \pm 0.1$	$0.1 \pm 0.0$	$0.3 \pm 0.1$	-	-	$0.1 \pm 0.0$	$0.2\pm0.1$
		1		1 I							

#### CMS

$m_{\rm H} \; [{\rm GeV}]$	ggH	VBF+VH	Data	All bkg.	WW	$VZ + W\gamma^{(*)} + Z/\gamma^* \rightarrow \ell\ell$	$t\bar{t}+tW$	W + jets
		$8\mathrm{TeV}$	$e^{\mu}$ e final	al state, 2-jets	s category, VB	SF tag		
120	$0.43 \pm 0.18$	$2.06\pm0.28$	2	$3.34\pm0.55$	$0.75\pm0.22$	$0.36\pm0.12$	$1.75\pm0.42$	$0.48 \pm 0.26$
125	$0.89 \pm 0.35$	$3.41\pm0.47$	2	$4.38\pm0.81$	$0.86 \pm 0.24$	$0.49\pm0.14$	$2.67\pm0.73$	$0.36\pm0.22$
		1 I		1				

#### **VBF** *WW\** comparison

PRD 92 (2015) 012006 v. JHEP 01 (2014) 096

#### CMS



#### ATLAS

	Signal significance	Expected	Observed uncertainty	Observed central value	_
Sample	Exp. Obs. Bar graph of $Z_0$ $Z_0$ observed $Z_0$	Tot. err. $+$ $-$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{c c} \mu_{\rm obs} & \mu_{\rm obs} \pm {\rm stat.} \ ({\rm thick}) \\ & \pm {\rm total} \ ({\rm thin}) \end{array}$	_
	3.70     4.08       2.89     3.07       2.36     3.12       1.43     0.71	$\begin{array}{c} 0.35 \ 0.30 \\ 0.41 \ 0.36 \\ 0.49 \ 0.44 \\ 0.74 \ 0.70 \end{array}$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	-
$n_j = 1$ $e\mu \text{ category}$ $ee/\mu\mu \text{ category}$	2.60 2.49 2.56 2.83 1.02 0.21	$\begin{array}{cccc} 0.51 & 0.41 \\ 0.51 & 0.42 \\ 1.12 & 0.98 \end{array}$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	_
$n_j \ge 2$ , ggF, $e\mu$	1.21 1.44	$0.96 \ 0.83$	0.91 0.84 0.70 0.68 0.70 0.49	1.20 —	
$n_j \ge 2$ , VBF-enr. $e\mu$ category $ee/\mu\mu$ category	3.38     3.84       3.01     3.02       1.58     2.96	$\begin{array}{c cccc} 0.42 & 0.36 \\ 0.48 & 0.40 \\ 0.84 & 0.67 \end{array}$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	← ±40%
All $n_j$ , all signal ggF as signal VBF as signal	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccc} 0.23 & 0.20 \\ 0.30 & 0.24 \\ 0.50 & 0.43 \end{array}$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	-

+100\_80%



#### **Higgs-DM models**



(9)

#### CMS Run-1 paper on VBF and ZH, EPJC 74 (2014) 2980

#### **9** Dark matter interactions

We now interpret the experimental upper limit on  $\mathcal{B}(H \rightarrow inv)$ , under the assumption of SM production cross section, in the context of a Higgs-portal model of DM interactions [7–9]. In these models, a hidden sector can provide viable stable DM particles with direct renormalizable couplings to the Higgs sector of the SM. In direct detection experiments, the elastic interaction between DM and nuclei exchanged through the Higgs boson results in nuclear recoil which can be reinterpreted in terms of DM mass,  $M_{\chi}$ , and DM-nucleon cross section. If the DM candidate has a mass below  $m_{\rm H}/2$ , the invisible Higgs boson decay width,  $\Gamma_{\rm inv}$ , can be directly translated to the spin-independent DM-nucleon elastic cross section, as follows for scalar (S), vector (V), and fermionic (f) DM, respectively [8]:

$$\sigma_{\rm S-N}^{\rm SI} = \frac{4\Gamma_{\rm inv}}{m_{\rm H}^3 v^2 \beta} \frac{m_{\rm N}^4 f_{\rm N}^2}{(M_{\chi} + m_{\rm N})^2},\tag{8}$$
$$\sigma_{\rm V-N}^{\rm SI} = \frac{16\Gamma_{\rm inv} M_{\chi}^4}{m_{\rm H}^3 v^2 \beta (m_{\rm H}^4 - 4M_{\chi}^2 m_{\rm H}^2 + 12M_{\chi}^4)} \frac{m_{\rm N}^4 f_{\rm N}^2}{(M_{\chi} + m_{N})^2},$$

$$\sigma_{\rm f-N}^{\rm SI} = \frac{8\Gamma_{\rm inv}M_{\chi}^2}{m_{\rm H}^5 v^2 \beta^3} \frac{m_{\rm N}^4 f_{\rm N}^2}{(M_{\chi} + m_{\rm N})^2}.$$
(10)

Here,  $m_{\rm N}$  represents the nucleon mass, taken as the average of proton and neutron masses, 0.939 GeV, while  $\sqrt{2}v$  is the Higgs vacuum expectation value of 246 GeV, and  $\beta = \sqrt{1 - 4M_{\chi}^2/m_{\rm H}^2}$ . The dimensionless quantity  $f_{\rm N}$  [8] parameterizes the Higgs-nucleon coupling; we take the central values of  $f_{\rm N} = 0.326$  from a lattice calculation [69], while we use results from the MILC Collaboration [70] for the minimum (0.260) and maximum (0.629) values. We convert the invisible branching fraction to the invisible width using  $\mathcal{B}(\rm H \rightarrow inv) = \Gamma_{\rm inv}/(\Gamma_{\rm SM} + \Gamma_{\rm inv})$ , where  $\Gamma_{\rm SM} = 4.07$  MeV.

### **VBF invisible results CMS Run 1**



CMS PAS HIG-14-038

#### 7 Limits on the cross section of invisibly decaying Higgs bosons

Upper limits on the Higgs boson production cross section times  $\mathcal{B}(H \rightarrow inv)$  are placed at 95% C.L. using an asymptotic CLs method [20, 38, 39], following the standard LHC Higgs combination technique [40, 41]. Systematic uncertainties are treated as nuisance parameters in a frequentist paradigm, as described in [41], and all correlations between processes are taken into account.

Using this procedure and assuming SM Higgs boson production cross sections and acceptances, the observed (expected) 95% C.L. limit on  $\mathcal{B}(H \to inv)$  of a SM 125 GeV Higgs boson is 57% (40%). The 95% C.L. limit on  $\mathcal{B}(H \to inv)$  and the 95% C.L. limit on the cross section times  $\mathcal{B}(H \to inv)$ , both assuming SM Higgs boson acceptances are shown as a function of Higgs boson mass in Fig. 7. As can be seen from Table 3 the dominant systematic uncertainty in the analysis is that from the limited numbers of data events in some control regions, in particular the Z control region. If the Z control region statistical uncertainty were to be reduced to the level of that from the  $W \to \mu v$  control region the expected 95% C.L. limit on the cross section times  $\mathcal{B}(H \to inv)$  for a SM 125 GeV Higgs boson would be reduced to 33%.

The result is also combined with that obtained by CMS in searches in the channel where the Higgs boson is produced in association with a Z which was reported in [8]. The procedure for this combination is also described in [8]. The 95% C.L. observed (expected) limit on  $\mathcal{B}(H \rightarrow inv)$  after combination is 47% (35%) for a SM 125 GeV Higgs boson.

### VBF invisible syst. for ATLAS Run 1



JHEP 01 (2016) 172

Uncertainty	VBF	ggF	Z  or  W	$Z_{\rm SR}/W_{\rm CR}$ or $W_{\rm SR}/W_{\rm CR}$
Int opprav scale	16	43	17 - 33	3–5
Jet energy scale	9	12	0 - 11	1-4
Ist operation	Negligible	Negligible	Negligible	Negligible
Jet energy resolution	3.1	3.2	0.2 - 7.6	0.5 – 5.8
Luminosity	2.8	2.8	2.8	Irrelevant
OCD seels	0.2	7 9	5-36	7.8–12
QUD scale	0.2	1.0	7.5 - 21	1-2
DDE	2.3	75	3 - 5	1.9
I DF	2.8	1.0	0.1 – 2.6	1-2
Parton shower	4.4	41	9–10	5
Veto on third jet	4.4	29	Negligible	Negligible
Higgs boson $p_{\rm T}$	Negligible	9.7	Irrelevant	Irrelevant
MC statistics	2	46	2.3 - 6.4	3 3-6 6
INIC STATISTICS	0.6	13	0.8 – 4.5	0.0-0.0

Table 7. Detector and theory uncertainties (%) after all SR or CR selections. For each source of uncertainty, where relevant, the first and second rows correspond to the uncertainties in SR1 and SR2 respectively. The ranges of uncertainties in the Z or W column correspond to uncertainties in the Z+jets and W+jets MC yields in the SR or CR. The search uses the uncertainties in the ratios of SR to CR yields shown in the last column.

#### **Fermiophobic scalars**



https://arxiv.org/abs/1604.07975 (27 April 2016)

Let us consider a simple UV-complete model to illustrate these points and further investigate the nature of the couplings. We will consider a scalar mediator coupled to dark matter as  $\phi \chi^2$ . This scalar mediator can obtain couplings to SM states via a Higgs portal mixing with the Higgs. Due to this mixing it inherits all of the SM Higgs couplings, suppressed by a factor  $\sin \theta$ , where  $\theta$  is the mixing angle. This model thus has couplings to quarks, leptons, and vector bosons

$$\mathcal{L} = \sin\theta \phi \left(\frac{m_q}{v}\overline{q}q + \frac{m_l}{v}\overline{l}l + 2\left(\frac{M_W^2}{v}W^{+\mu}W^{-}_{\mu} + \frac{M_Z^2}{2v}Z^{\mu}Z_{\mu}\right)\right) \quad . \tag{3.4}$$

First of all, this demonstrates that in UV-complete models realising the  $\lambda_q \phi \bar{q} q$  interaction, the interaction of Eq. (3.2) also typically arises. Second, the results of [71, 87, 88] demonstrate that when a mediator has these couplings the strongest collider bounds will arise from VBF production of the DM, shown in Fig. 2. Since the monojet bounds arise from the mediator couplings to quarks, and the VBF bounds from the mediator couplings to vectors, it is clear that it may be possible to overlook the strongest probes of DM for scalar mediators at the LHC if one only considers the  $\lambda_q \phi \bar{q} q$  interaction for scalar mediators.

#### If $\phi$ is fermiophobic, mono-jet isn't the best (VBF may be).

### **Jet energy uncertainties**



ATLAS performance



① Jet energy measurement and its systematic uncertainty in proton-proton collisions at  $\sqrt{s} = 7$  TeV with the ATLAS detector, EPJC 75 (2015) 17.

② Jet calibration and systematic uncertainties for jets reconstructed in the ATLAS detector at √s=13 TeV, ATL-PHYS-PUB-2015-015 (24 July 2015),

#### LUX latest results this week



PRL 116 (2016) 161301, 161302



FIG. 3. Upper limits on the spin-independent elastic WIMPnucleon cross section at 90% C.L. Observed limit in black, with the 1- and 2- $\sigma$  ranges of background-only trials shaded green and yellow. Also shown are limits from the first LUX analysis [6] (gray), SuperCDMS [40] (green), CDMSlite [41] (light blue), XENON100 [42] (red), DarkSide-50 [43] (orange), and PandaX [44] (purple). The expected spectrum of coherent neutrinonucleus scattering by <sup>8</sup>B solar neutrinos can be fit by a WIMP model as in [45], plotted here as a black dot.



FIG. 1. LUX upper limits on the WIMP-neutron (top) and proton (bottom) elastic SD cross sections at 90% C.L. The observed limit is shown in black with the  $\pm 1\sigma$  ( $\pm 2\sigma$ ) band from simulated background-only trials in green (yellow). Also shown